

1 Estimating carbon dioxide residence time scales through
2 noble gas and stable isotope diffusion profiles

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12 **ABSTRACT**

13 The study of natural carbon dioxide reservoirs provides fundamental insight into
14 processes involved in carbon capture and storage. However, the calculations of process
15 rates such as dissolution of CO₂ into formation water remain uncertain due to indirectly
16 determined ages of the CO₂ influx. The proposed ages for the Bravo Dome gas field in
17 New Mexico, USA, vary from 56 ka to 1.5 Ma. Here we demonstrate that residence times
18 can be estimated from simple modeling of noble gas and stable isotope diffusion profiles
19 from the gas-water contact through the gas column. The Bravo Dome gas field shows a
20 gradient in noble gas concentrations and isotopic ratios from east to west across the 70-
21 km-wide field. A mantle-like end member with a ³He/⁴He (R/R_A) ratio of up to 4.7 is
22 found in the west in contrast to a groundwater end member with high concentrations of

23 air- and crustal-derived noble gases in the east. The air- and crustal-derived noble gases
24 decrease gradually toward the west. Stable isotope compositions (C and O) also vary
25 across the field. Diffusion modeling of He, Ne, Ar, Kr, Xe, and $\delta^{13}\text{C}$ and data yield
26 residence times for the CO_2 between 14.1 ± 0.2 ka and $16.9 +1.1/-0.5$ ka. This is far less
27 than the previous estimates of 1.2–1.5 Ma based on apatite (U-Th)/He
28 thermochronology, leading to a dissolution rate of **[[Precede these values with “~”**
29 **symbol, given that they are rounded from the values reported in**
30 **Conclusions?]]**30,000 +12,000/-11,000 t/a to **36,000** $\pm 12,000$ t/a, implying that 32% of
31 the total emplaced CO_2 dissolved. This new method can be applied to a wide variety of
32 gas fields with variation in the concentration of groundwater-derived noble gases and
33 allow a better assessment of the time scale of other diffusive fluid-fluid interactions.

34 INTRODUCTION

35 Natural carbon dioxide reservoirs serve as natural analogues for geological CO_2
36 storage and hence are studied extensively to understand processes that act over geological
37 time scales (Bickle et al., 2013). In order to quantify the rates of interactions between
38 CO_2 , subsurface fluids, and the host rock, it is crucial to know the residence time of the
39 CO_2 in the reservoir. Previous estimates of CO_2 reservoir residence times include dating
40 of surface intrusive igneous rocks from the **Colorado Plateau** (Armstrong, 1969), dating
41 of close-by effusive igneous rocks (Condit, 1995; Stroud, 1997), and apatite (U-Th)/He
42 thermochronology (Sathaye et al., 2014). These methods have been used to infer
43 residence times for the Bravo Dome (**New Mexico, USA**) CO_2 reservoir ranging between
44 56 ka (Stroud, 1997) and 1.5 Ma (Sathaye et al., 2014). Petrographic studies on drill cores
45 suggest the carbon dioxide filling of Bravo Dome in the mid-Cenozoic (Hartig et al.,

46 2011), but rely exclusively on relative time constraints of diagenetic mineral growths
47 with poorly developed textures. Here we present a model for estimating directly the CO₂
48 residence time applied to the Bravo Dome gas field. The model relies on spatial sampling
49 of gases across the field and uses the diffusion of groundwater-derived noble gases
50 entering the gas column at the gas-water contact, with known diffusion constants, to
51 estimate residence ages.

52 BRAVO DOME FIELD

53 Bravo Dome is a well-studied sandstone-hosted CO₂ reservoir (Baines and
54 Worden, 2004; Cassidy, 2006; Dubacq et al., 2012; Gilfillan et al., 2008, 2009; Hartig et
55 al., 2011; Holland and Ballentine, 2006; Kessler et al., 2001; Sathaye et al., 2014). The
56 gas is primarily of magmatic origin (Staudacher, 1987) and has ³He/⁴He (R/R_A) ratios up
57 to 4.7 and CO₂/³He ratios in the **mid-oceanic ridge basalt (MORB)** range (Ballentine et
58 al., 2001; Cassidy, 2006; Gilfillan et al., 2008; Holland and Ballentine, 2006). The gas
59 field comprises a thin laterally extensive reservoir with an east-west extent of >70 km
60 (Fig. 1) and supercritical reservoir conditions (Broadhead, 1993). The western portion of
61 the gas field is not in contact with formation water because it overlies dry impermeable
62 basement (Cassidy, 2006), and we assume that the most western fault on the map in
63 Figure 1 represents the east-west gas-water contact (Appendix DR4 in the GSA Data
64 Repository¹). On the east side of this fault, the water table is within the sandstone
65 formation and the CO₂ overlies water-saturated rock, with lateral juxtaposition of CO₂
66 and formation water also occurring laterally to the extreme east (Fig. 1). Oxygen isotope
67 data of CO₂ gas confirm the western fault as a gas-water contact because the gas samples
68 east of the fault have δ¹⁸O of 29‰ **[[relative to?]] standard mean ocean water**

69 (SMOW)[[Should this be “Vienna standard mean ocean water (V-SMOW) as
70 indicated in fig. 2A?]] on average, which is the value predicted from CO₂-H₂O
71 fractionation (Brenninkmeijer et al., 1983) and complete equilibration with formation
72 water of $\delta^{18}\text{O} = -11\text{‰}$ at 33 °C (Cassidy, 2006) (Fig. 2A).

73 MODEL

74 Data and Corrections

75 The gas samples were collected in two different sampling campaigns (*b* and *c*) of
76 the Bravo Dome gas field (Gilfillan et al., 2008, 2009; Holland and Ballentine, 2006).
77 Noble gas data of sample set *b* are from previous studies (Gilfillan et al., 2008, 2009;
78 Holland and Ballentine, 2006), whereas sample set *c* has been measured following
79 previous methods (Holland and Ballentine, 2006) (Appendix DR1). The oxygen isotopic
80 and carbon isotopic analysis of sample set *b* was performed at the University of East
81 Anglia (Norwich, UK) using a MIRA (multi-isotope ratio analyzer) mass spectrometer
82 (Dennis, 2014) with procedures following Royle (2015) and precision for standard
83 replicates of <0.01‰. The carbon isotopes of sample set *c* were measured at Lancaster
84 University (UK) using an Isoprime100 spectrometer with a multiflow inlet. The data
85 were corrected to PDB (Peedee belemnite[[Or should it be “Vienna Peedee belemnite
86 (V-PDB)”, as indicated in Fig. 2B? (all instances of “PDB” in text)]) using
87 international standards LSVEC (lithium carbonate; $\delta^{13}\text{C} -46.6\text{‰}$), NBS 18 (calcite; $\delta^{13}\text{C}$
88 -5.014‰) and CO1 (marble; $\delta^{13}\text{C} +2.492\text{‰}$). Precision for standard replicates ($n = 5$, 1
89 standard deviation) was <0.1‰ for carbon isotopes.

90 The data for the model comprise the noble gas measurements and their sampling
91 distance to the gas-water contact. The distance from the 26 selected wells to the gas-

92 water contact has been measured horizontally for the wells west of the main fault and
93 vertically for the wells east of it, because these are the shortest distances to the gas-water
94 contact (Fig. 1). The perforation interval depths from which the CO₂ is produced have
95 been considered for the wells with vertical distances to the gas-water contact.

96 At the gas-water contact, the air- and crustal-derived noble gases, which
97 accumulated in the formation water, degas into the CO₂ because of the higher
98 concentration and lower solubility in the formation water than in the CO₂ phase. In
99 addition to this increase in noble gas concentration in the CO₂, noble gas concentrations
100 in the gas phase are also enhanced by contemporaneous CO₂ dissolution into the
101 formation water. This solvent loss effect can be corrected based on ³He concentrations
102 (Appendix DR3) (Dubacq et al., 2012).

103 **Parameters and Error Calculation**

104 Following the noble gas concentration correction, the data can then be compared
105 to a diffusion profile based on Fick's second law of diffusion:

$$106 \quad \left[\frac{\partial C}{\partial t} = D_{\text{eff}} \frac{\partial^2 C}{\partial x^2} \right], \quad (1)$$

107 where C is concentration, t is time, D_{eff} is the effective diffusion coefficient, and x is the
108 **[[diffusion?]]** distance. The solution of the equation involves two boundary conditions
109 which are set to be a constant noble gas concentration in the CO₂ at the gas-water contact
110 (C_{start}) and a magmatic-derived noble gas concentration (C_{end}) at the western end of the
111 gas field:

$$112 \quad [C(x, t) = C_{\text{start}} - (C_{\text{start}} - C_{\text{end}}) * \text{erf}\left(\frac{x}{2\sqrt{t*D_{\text{eff}}}}\right)] . \quad (2)$$

113 The effective diffusion coefficient compensates for the indirect travel paths of the
114 particles and differs from the common diffusion coefficient by taking into account the
115 porosity and tortuosity:

$$116 \quad [D_{\text{eff}} = D \frac{\phi}{\tau}], \quad (3)$$

117 where D is the diffusion coefficient, ϕ is porosity, and τ is tortuosity. The diffusion
118 coefficients have been calculated after the method of He and Yu (1998) because no
119 experimental data exist for diffusion of noble gases in supercritical CO_2 (Appendix DR2).
120 Poling et al. (2001) tested this approach for the vitamin K_1 in supercritical CO_2 and found
121 a low error of 0.4% compared to the empirically determined diffusion coefficient. The
122 calculated diffusion coefficients for the noble gases in CO_2 are approximately one order
123 of magnitude higher compared to those in water. The reported porosity of the Tubb
124 Formation in New Mexico is 20% on average (Johnson, 1983; Cassidy, 2006), and the
125 tortuosity has been calculated to be 2.4, which is a typical value for sandstones (Cassidy,
126 2006; Epstein, 1989; Matthews and Spearing, 1992).

127 Solutions for the diffusion profiles are calculated using a Monte Carlo simulation
128 in Matlab where both boundary conditions and the diffusion time multiplied by the
129 diffusion coefficient are fitted 1000× to the data, taking into account the error on
130 measurements through minimizing χ^2 :

$$131 \quad \chi^2 = \sum_i^n \frac{(x_i - x_m)^2}{\frac{1}{2}x_{(\text{error range})}^2} + \frac{[y_i - y(x_m)]^2}{\sigma_y^2}, \quad (4)$$

132 where x_m and $y(x_m)$ are model values for the distance and the concentration, x_i and y_i are
133 the distance and concentration of the data points, $x_{(\text{error range})}$ is the error range of the
134 distance, and σ_y is the standard error of the concentration measurements (Appendix

135 DR2). The data points in the simulation are picked randomly within the normal
136 distribution of their standard error for the concentration and within the error range of their
137 distance to the gas-water contact. A second Monte Carlo simulation is set up to evaluate
138 the error of the diffusion coefficient on the diffusion time. The most probable diffusion
139 time, the diffusion coefficient, C_{start} , and C_{end} are reported with asymmetric error brackets
140 defining the 95% quantile (Fig. 2; Appendix DR2). In the simulation for xenon, about
141 one-fifth of the C_{end} concentrations resulted in negative values which were excluded
142 because negative concentrations are meaningless. The model simplifies the field and does
143 not account for lithology heterogeneities, residual water, compartmentalization of the
144 field, lateral heterogeneities, and temporal variation in the noble gas concentration in the
145 formation water which would have the tendency to increase the diffusion time.

146 The fitted carbon isotope C_{start} composition of -7.9‰ (Appendix DR2) fits well
147 with a CO_2 composition in equilibrium with formation water at Bravo Dome that is in
148 equilibrium with dolomite. Dolomite with a $\delta^{13}\text{C}$ of $\sim 2\text{‰}$ PDB (Hoefs, 2008) in
149 equilibrium with the formation water at Bravo Dome (Kessler et al., 2001; Hartig et al.,
150 2011) results in a CO_2 groundwater end member of -8‰ based on a $\text{CO}_2(\text{g})$ -
151 dolomite[[**Explain what “(g)” means**]] fractionation of -10‰ (Sheppard and Schwarz,
152 1970; Clark and Fritz, 1997) at 33 °C (Cassidy, 2006). The fitted C_{end} composition of the
153 carbon isotope is -4.1‰ which is in the range of magmatic-derived CO_2 (Jenden et al.,
154 1993).

155 AGE ESTIMATION

156 The model generates noble gas diffusion profiles from the gas-water contact in the
157 east toward the western end of the field that are consistent with the observed well gas

158 concentrations and place age constraints for **[[the beginning of?]]** the process at between
159 14.1 ± 0.2 ka and $16.9 +1.1/-0.5$ ka (Figs. 2C–2H; Appendices DR2 and DR4). The age
160 difference between the noble gases could arise from the uncertainty of the diffusion
161 coefficients because the noble gas diffusion coefficients are calculated and not
162 empirically measured. The diffusion ages are strongly influenced by the wells close to the
163 gas-water contact because these define the curvature and hence the age of the diffusion
164 profile.

165 The diffusion of the groundwater-derived noble gases through the CO₂ most
166 likely started when the CO₂ first contacted the formation water. The gas-water contact
167 probably moved down dip from the inferred entry point (see Sathaye et al., 2014, their
168 figure 4b) throughout the filling of the CO₂ but this does not affect the diffusion of the
169 noble gases in the CO₂, assuming a slug-like displacement of the formation water, which
170 is reasonable given the likely slow rates of filling. Today the gas-water contact is >30 km
171 away from the inferred entry point of the CO₂ (Sathaye et al., 2014). Therefore, later CO₂
172 pulses, as proposed in previous studies (Baines and Worden, 2004; Sathaye et al., 2014),
173 would not alter the noble gas diffusion profile at the gas-water contact. The water table in
174 the Bravo Dome field lies horizontally (Cassidy, 2006), and the anhydrite seal likely
175 prevents groundwater recharge, which suggests a stagnant aquifer (Akhbari and Hesse,
176 2017). The estimated CO₂ residence times are younger than the youngest **[[Residence
177 time is not an age, and therefore cannot be younger than something – please
178 reword]]** dated effusive rock in the area. Deep igneous activity, however, could have
179 continued after the last lava flow at 56 ka (Stroud, 1997).

180 Additionally, the change in carbon and oxygen isotopes has been investigated;
181 this can give insight into the source of the CO₂ and extent of gas-water exchange and CO₂
182 dissolution. The δ¹³C varies from −3.2‰ PDB in the west to −5.8‰ PDB in the east.
183 This isotopic difference originates from the presence of two sources of carbon. There is a
184 magmatic carbon input in the west, which is isotopically distinct from the formation
185 water-derived carbon in the east. Similarly to the noble gases, diffusion controls the
186 mixing between these two end members (Fig. 2B). Diffusion modeling of the δ¹³C profile
187 indicates a development period of 15.0 +0.2/−0.5 ka (Fig. 2B). This **age** **[[period?]]**
188 agrees well with the age **[[Can't directly compare a "period" to an "age" – reword]]**
189 estimates derived from diffusion modeling of the noble gas concentration profiles.

190 However, the evolution of oxygen isotopes is more complex. The δ¹⁸O
191 **[[Correct?]]** varies from a minimum of 18.6‰ in the west to 29‰ in the east. The wells
192 in the west have lighter isotopic composition but differ by **more than** 9‰–12‰ from an
193 expected mantle-like CO₂ composition of 6‰–9‰ (Eiler, 2001). During the
194 displacement of the formation water by the invading CO₂, an irreducible water saturation
195 of typically 5%–22% in sandstones (Krevor et al., 2012) will be trapped in the CO₂-filled
196 pores. The pores in the Tubb Formation thus contain an irreducible water saturation with
197 a sufficient mass of oxygen such that O-isotopic exchange can modify the O-isotopic
198 composition of across the entire CO₂-filled volume. A residual water content correction
199 has been made to the O-isotope ratios based on the mass balance equation:

$$200 \quad \delta^{18}\text{O}^f_{\text{CO}_2} = \delta^{18}\text{O}^i_{\text{CO}_2}(1 - X^0_{\text{H}_2\text{O}}) + X^0_{\text{H}_2\text{O}}(\delta^{18}\text{O}_{\text{H}_2\text{O}} + \varepsilon), \quad (5)$$

201 **[[Johnson et al., 2011]]** **[[Johnson et al., 2011 is not in the reference list.]]** where ε is the
202 CO₂(g)-H₂O fractionation factor at 33 °C (Brenninkmeijer et al., 1983), δ¹⁸Oⁱ_{CO₂} and

203 $\delta^{18}\text{O}_{\text{CO}_2}^f$ refer to initial and final composition of the CO_2 , respectively, and $X_{\text{H}_2\text{O}}^0$ is the
204 fraction of oxygen in the system from water. This amendment decreases the measured
205 values only very little because the residual water saturation is as low as 5–10 vol%
206 (Sathaye et al., 2014). Due to this complexity, the spatial trends in O-isotopic
207 composition are not a useful constraint on the emplacement age.

208 DISSOLUTION RATE

209 Combining the range of residence times of 14.1 ± 0.2 ka and $16.9 +1.1/-0.5$ ka
210 with the gas mass, which has been subject to dissolution, enables the estimation of the
211 dissolution rate. The dissolved gas can simply be estimated by subtracting the present
212 mass of gas in the reservoir from the mass of gas that was in place originally. The present
213 gas volume in the reservoir has previously been estimated to be 1.3 ± 0.6 Gt CO_2
214 (Sathaye et al., 2014). The original gas mass can be calculated from the difference
215 between the present gas and the deviation from the highest $\text{CO}_2/{}^3\text{He}$ ratio (7.4×10^9)
216 (Appendix DR1), representing dissolution of CO_2 through apparent increase in ${}^3\text{He}$. The
217 dissolved gas mass has been calculated in the past (Sathaye et al., 2014), however, the
218 highest $\text{CO}_2/{}^3\text{He}$ ratio used was 5.35×10^9 (Cassidy, 2006; Gilfillan et al., 2008, 2009).
219 Therefore, the dissolved mass is **one-third** larger than previously estimated (Sathaye et
220 al., 2014) and equates to 506 ± 166 Mt CO_2 . This amount is 32% of the total estimated
221 emplacement of 1.6 ± 0.67 Gt CO_2 (Sathaye et al., 2014). The resulting dissolution rate
222 ranges from $23,000 +1200/-1100$ t/a to $36,000 \pm 12,000$ t/a. The gas-water contact covers
223 an estimated area of 623 km^2 (Sathaye et al., 2014) and enables estimating the flux that
224 ranges from $48 +19/-17$ g/(m^2a) to 58 ± 19.7 g/(m^2a). This flux is $>500\times$ larger compared
225 to previous estimates [0.1 g/(m^2a)] (Sathaye et al., 2014) but still almost $400\times$ smaller

226 than the estimates for the permeable Sleipner CO₂ storage site in the North Sea [20
227 kg/(m²a)] (Neufeld et al., 2010).

228 CONCLUSIONS

229 This diffusion model has, for the first time, been applied to estimation of the
230 residence time of CO₂ in Bravo Dome. This is key for quantifying the rate of CO₂
231 dissolution occurring in carbon dioxide reservoirs. Combining the range of residence
232 times of 14.1 ± 0.2 ka and $16.9 +1.1/-0.5$ ka with the gas volume that has been subject to
233 dissolution, we could estimate the CO₂ dissolution rate of $29,900 +11,800/-10,700$ t/a to
234 $35,900 \pm 12,300$ t/a. Furthermore, the model can be tested on other natural CO₂ and
235 enhanced oil recovery (EOR) reservoirs to investigate gas-water interactions and quantify
236 gas dissolution rates into the formation waters that are ubiquitously found in subsurface
237 gas reservoirs.

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345 **FIGURE CAPTIONS**

346 Figure 1. **A:** Map of Bravo Dome gas field (New Mexico, USA), showing top of CO₂-
347 hosting Tubb Formation (black contour lines, in m above sea level [a.s.l.]), gas-water
348 contact (short dashed line), the ³He/⁴He ratio (long dashed line), and 26 sampling wells
349 (circles). **B:** West-east cross-section highlighting gas-water contact within Tubb
350 Formation on eastern side of westernmost fault, as opposed to west side, where gas
351 reservoir is not in contact with formation water. Sampling wells are projected with their
352 perforation depth (circles).

353 **[[In the figure, label panels A and B with uppercase letters; change all instances of**
354 **“gas/water” to “gas-water”. In panel A, insert hyphen in “east-west” on the gas-**
355 **water fault contact; in symbol explanation, insert space between “Fm” and “[m**
356 **a.s.l.]”, between “water” and “contact”, and between “10” and “km” on the scale**
357 **bar. In panel B, sense of slip on the faults doesn’t seem to match that shown in panel**
358 **A (barbs typically on upthrown side); change “Tubb Sandstone” to “Tubb**
359 **Formation” for consistency; include horizontal scale bar (or reference the one in**
360 **panel A if the same)]]**

361

362 Figure 2. A–H: CO₂ dissolution–corrected noble gas concentrations and measured stable
363 isotopes plotted against distance to gas-water contact. Groundwater-derived noble gases
364 diffusing from gas-water contact through gas column are fitted by Monte Carlo
365 simulation to Fick’s second law of diffusion equation (C–H). This reveals residence time
366 of CO₂. Additionally, Ne diffusion profile is drawn for 1.2 Ma (D). Carbon isotopes show
367 diffusive mixing between magmatic carbon source and formation water carbon source
368 that is in equilibrium with dolomite (B). Time estimated using $\delta^{13}\text{C}$ isotopes is within
369 range of ages estimated from noble gases. Oxygen isotopes show full equilibration with
370 formation water in eastern part of field close to gas-water contact (A). Concentration and
371 distance errors are smaller than symbols (circles). Symbol fills correspond to well
372 locations in Figure 1. $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ values are relative to: V-SMOW—Vienna standard
373 mean ocean water; V-PDB—Vienna Peedee belemnite. cm³STP—standard cubic
374 centimeter.

375 **[[In the figure, change “gas/water” to “gas-water”]]**

376

377 ¹GSA Data Repository item 2017, Appendices DR1–DR4, is available online at
378 <http://www.geosociety.org/datarepository/2017/> or on request from
379 editing@geosociety.org.