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- Two-dimensional numerical simulation of *O* to *Z*-mode wave conversion during ionospheric modification
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# Two-dimensional numerical simulation of **O**-mode to **Z**-mode conversion in the ionosphere

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**Abstract** Experiments in the illumination of the *F* region of the ionosphere via radio frequency waves polarized in the ordinary mode (*O*-mode) have revealed that the magnitude of artificial heating-induced effects depends strongly on the inclination angle of the pump beam, with a greater modification to the plasma observed when the heating beam is directed close to or along the magnetic zenith direction. Numerical simulations performed using a recently developed finite-difference time-domain (FDTD) code are used to investigate the contribution of the *O*-mode to *Z*-mode conversion process to this effect. The aspect angle dependence and angular size of the radio window for which conversion of an *O*-mode pump wave to the *Z*-mode occurs is simulated for a variety of plasma density profiles including 2-D linear gradients representative of large-scale plasma depletions, density-depleted plasma ducts, and periodic field-aligned irregularities. The angular shape of the conversion window is found to be strongly influenced by the background plasma profile. If the *Z*-mode wave is reflected, it can propagate back toward the *O*-mode reflection region leading to resonant enhancement of the electric field in this region. Simulation results presented in this paper demonstrate that this process can make a significant contribution to the magnitude of electron density depletion and temperature enhancement around the resonance height and contributes to a strong dependence of the magnitude of plasma perturbation with the direction of the pump wave.

#### 1. Introduction

The use of high-power radio frequency (RF) transmitters to perturb the Earth's ionosphere gives a unique opportunity to study the complex and often nonlinear mechanisms that underpin the interaction between an electromagnetic (EM) wave and ionospheric plasma. The ionosphere can be treated as a vast plasma laboratory, in which experiments can be performed without the limited spatial scale size and chamber edge effects one may encounter while performing plasma experiments in a laboratory. Experiments in the illumination of the *F* region of the ionosphere via RF waves polarized in the ordinary mode (*O*-mode) have resulted in the observation of a wide range of nonlinear thermal phenomena associated with the excitation of plasma waves and instabilities [*Robinson*, 1989; *Rietveld et al.*, 2003]. Many of these processes can induce significant enhancements of the electron temperature around the region in which the EM wave most strongly interacts with the plasma [*Gordon and Carlson*, 1974; *Meltz et al.*, 1974; *Djuth et al.*, 1987; *Stocker et al.*, 1992; *Honary et al.*, 1993]. Substantial changes to the electron density have also been commonly observed on a variety of different spatial scales, from large-scale density depletions spanning several kilometers [*Gurevich et al.*, 2002] to small-scale irregularities aligned with the geomagnetic field and having transverse scale sizes ranging from a few meters down to a few centimeters, as detected by rockets and VHF-UHF backscattering [*Kelley et al.*, 1995].

The main processes driving the enhancement of electron temperature in ionospheric modification experiments involve the excitement of large-amplitude plasma waves below the *O*-mode refection height where the frequency of the incident pump wave is close to the local upper hybrid (UH) wave frequency. In this region, provided that the pump frequency is not close to an electron gyroharmonic [*Rao and Kaup*, 1990; *Stubbe et al.*, 1994], electrostatic (ES) UH waves propagating in a direction perpendicular to the geomagnetic field are readily excited [*Wong et al.*, 1981; *Dysthe et al.*, 1982; *Antani et al.*, 1996]. ES waves of this nature induce the formation of striations in the electron density profile, elongated along the magnetic field direction and with a comparatively small-transverse scale size. The region of interaction is initially narrow but broadens as the striations grow with time. In the stationary state, the vertical scale of this region is several kilometers [*Borisov and Robinson*, 2003]. Explosive heating of electron gas local to the striations then proceeds via the nonlinear resonance instability [*Gurevich et al.*, 1995, 2001; *Gondarenko et al.*, 2005], as incident pump wave electric field

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becomes trapped in the density depletions and amplified through efficient excitation of further UH modes. Collisional damping of the trapped waves is highly efficient within the striations and leads to an increase in the local plasma temperature and consequently to an increase in the electron density depletion within the irregularities. As the striations deepen, the efficiency of trapping increases, and leads to a rapid, nonlinear increase in the electron temperature around the UH resonance height.

The formation of small-scale density irregularities reduces the average electron plasma density around the UH resonance region and leads to a focusing of the pump wave and thus to a further enhancement of the *E* field amplitude in this region [*Gurevich et al.*, 2001; *Gondarenko et al.*, 2003]. Self-focusing of this nature results in further nonlinear enhancement of the electron temperature and causes the evolution of a hierarchy of density-depleted structures around the *O*-mode reflection height, from the small-transverse-scale (1–10s of meters) striations associated with the resonance (thermal parametric) instability to self-organized bunches of striations forming duct-like density-depleted structures due to the self-focusing instability spanning up to many kilometers in transverse scale [*Gurevich et al.*, 1998]. Close-packed duct-like structures of this nature are known to form large-scale semipermanent density-depleted regions or duct-like channels which may affect wave propagation.

Observations have revealed that the magnitude of such artificial heating-induced perturbations to the plasma density depends strongly on the inclination angle of the RF pump beam, with a greater modification to the plasma observed when the heating beam is directed away from the vertical direction toward the magnetic zenith, as reported, for example, by *Honary et al.* [2011], *Kosch et al.* [2000], and *Pedersen et al.* [2003]. This magnetic zenith effect has been observed as an increase in electron temperature enhancement by a factor of 2 or greater compared to off-zenith operation [*Rietveld et al.*, 2003; *Dhillon and Robinson*, 2005] and has also been seen to correspond to a greater enhancement in the characteristic ion and electron plasma lines seen in UHF incoherent scatter radar spectra and an increase in the backscatter power due to field-aligned irregularities.

To explain theoretically the unexpected electron temperature and optical emission enhancements arising due to heating directed along the magnetic zenith direction, the self-focusing instability was first suggested as the mechanism behind this effect [*Gurevich et al.*, 2002, 2005]. The timescale for this process to manifest is known to be long (of the order of 1 min). The experiments recently performed by *Honary et al.* [2011] demonstrated that the rise of electron temperature occurs on a much more rapid timescale (of the order of a few seconds; comparative to the rise time of field-aligned striations). Instead, *Honary et al.* [2011] proposed that this aspect-angle dependence of heating is in part due to the *O*-mode to *Z*-mode (slow branch of the extraordinary mode) conversion process that can occur in the *F* region of the ionosphere for a narrow window of pump wave inclination angles. The *Z* mode is able to penetrate to altitudes above the *O*-mode reflection height, where the conversion of EM waves trapped in field-aligned irregularities to plasma modes is more efficient. This leads to a more rapid development of the resonance instability and potentially a greater magnitude of electron temperature enhancement.

To fully understand the fundamental wave interaction process involved in the *O*-mode to *Z*-mode conversion process, it is useful to augment experimental findings with accurate numerical simulations. Ionospheric plasma processes have previously benefited from investigation using full-wave numerical simulation codes, for example, as reported by *Gondarenko et al.* [2003] who used an alternating direction implicit model to simulate linear mode conversion processes, *Eliasson and Stenflo* [2008] who studied the parametric decay instability and the generation of Langmuir turbulence around the interaction height using a generalized Zakharov model, or *Leyser and Nordblad* [2009] and *Nordblad and Leyser* [2010] who simulated the *Z*-mode pumping of a nonlinear density ducts such as that formed via the self-focusing instability using ray tracing coupled with a Runge-Kutta scheme. In this paper, a recently developed full-wave finite-difference time-domain (FDTD) code [*Cannon and Honary*, 2015] is used to study this conversion process in two spatial dimensions, allowing the variety of background density scenarios to be investigated. Time-dependent on-grid updates to the plasma density and temperature included in the FDTD algorithm meant that the impact of *O*- to *Z*-mode conversion on the growth of thermal-scale plasma perturbations could be simulated.

The paper is structured as follows. In section 2, the numerical simulation algorithm is introduced. In section 3 the simulation code is used to calculate the radio window for which conversion from *O*-mode to *Z* mode is favorable for a simplified linear density gradient, with the results shown to be in good agreement with the theoretical predictions of [*Mjølhus*, 1984]. The code is then used to demonstrate how the radio window is

modified by the presence of plasma density features of varying scale sizes, such as density-depleted ducts or striations. Section 4 presents the results of full-wave plasma simulations demonstrating that the excitation and subsequent reflection of the *Z*-mode wave has a strong influence on the magnitude of heating-induced effects such as temperature enhancement developed around the interaction region. Simulations of plasma heating experiments for a range of pump wave launch angles show that the *O*-mode to *Z*-mode conversion process is a key mechanism contributing to the observed magnetic zenith effect. In section 5 the effects of adding density-depleted plasma irregularities are investigated, with the influence of this process on the enhancement of electron temperature demonstrated to be heavily dependent on both pump wave inclination angle and irregularity characteristics.

#### 2. Methodology

To numerically simulate an ionospheric modification experiment, this study uses the GPU-accelerated FDTD code described and validated in Cannon and Honary [2015]. The FDTD method, first proposed in Yee [1966], has an advantage over many other numerical simulation techniques as it deals with complex, nonlinear, and impulsive interactions in a natural and time-explicit manner, avoiding complex and computationally intensive linear algebra calculations. FDTD is a particularly useful tool for geophysical modeling due to the fact that the grid-based structure allows medium properties such as plasma density or temperature to be defined separately at each point. The algorithm used here assumes a multifluid description of a dynamic, anisotropic, collisional plasma, in which electron or charged ion species are treated as individual fluids of continuous mass and charge. The effect on wave propagation due to the presence of plasma is introduced through the coupling of Maxwell's wave equations with the Lorentz equations of motion for each constituent plasma species, with anisotropy introduced through inclusion of a static externally applied magnetic field in a manner similar to that described by Young [1994] or Yu and Simpson [2010]. The time-dependent variation of plasma temperature and density are treated by the inclusion of expressions for the dynamic behaviors of small perturbations of the plasma fluid temperature and density [Gurevich, 1978]. Together, these form a set of coupled first-order partial differential equations which govern the time-dependent behavior of the EM wave and plasma medium:

$$\nabla \times \mathbf{E} = -\mu_0 \frac{\partial \mathbf{H}}{\partial t} \tag{1}$$

$$\nabla \times \mathbf{H} - \sum_{a} N_{a} e_{a} \mathbf{U}_{a} = \varepsilon_{0} \frac{\partial \mathbf{E}}{\partial t}$$
<sup>(2)</sup>

$$N_a m_a \frac{\partial \mathbf{U}_a}{\partial t} = N_a e_a (\mathbf{E} + \mathbf{U}_a \times \mathbf{B}) - N_a m_a v_a \mathbf{U}_a - \nabla (k_B N_a T_a)$$
(3)

$$\frac{\partial N_a}{\partial t} + \nabla \cdot (N_a \mathbf{U}_a) = 0 \tag{4}$$

$$\frac{3}{2}k_{\beta}\frac{\partial}{\partial t}(N_{a}T_{a}) + \nabla \cdot \mathbf{Q}_{a} - N_{a}e_{a}\mathbf{E} \cdot \mathbf{U}_{a} - \triangle \varepsilon_{a} = 0.$$
(5)

In these expressions, subscript *a* refers to plasma component species. **U** is the time-varying fluid bulk velocity vector, **B** =  $B\hat{\mathbf{b}}$  is the static background magnetic field, *T* and *N* are the plasma temperature and number density, respectively,  $\nabla \cdot \mathbf{Q}$  describes the heat flux transport,  $\Delta \varepsilon$  is a collisional heating term,  $e_a$  and  $m_a$  refer to particle charge and mass, respectively,  $\varepsilon_0$  and  $\mu_0$  are the permittivity and permeability of free space respectively, and  $k_B$  is the Boltzmann constant. v represents the effective collision frequency given by equation (6) [Gondarenko et al., 2005]:

$$\nu(N, T) = \nu_0 \left(\frac{N}{N_0}\right) \left(\frac{T}{T_0}\right)^{-3/2},\tag{6}$$

where  $v_0$ ,  $N_0$ , and  $T_0$  are the unperturbed effective collision frequency, number density, and temperature, respectively. The effective collision frequency incorporates both collisions with other plasma particles and

with the neutral background species, such that  $v_e = v_{ei} + v_{en}$  and  $v_i = v_{ie} + v_{in}$ , where  $v_{ei}$ ,  $v_{ie}$ ,  $v_{en}$ , and  $v_{in}$  are the electron-ion, ion-electron, electron-neutral, and ion-neutral collision frequencies, respectively. In the simulation studies described below, the undisturbed collision frequencies  $v_0$  were chosen to represent typical conditions in the *F* region at Tromsø, and updated every time step via equation (6) to reflect simulated perturbations to the plasma medium. Simulated fields are updated sequentially via a leapfrog time-stepping scheme, following the cyclical update pattern:  $\mathbf{E}^q \rightarrow T^q \rightarrow \mathbf{H}^{q+\frac{1}{2}} \rightarrow \mathbf{U}^{q+\frac{1}{2}} \rightarrow N^{q+1} \rightarrow \mathbf{E}^{q+1} \rightarrow \dots$  Here **E**, **H**, **U**, *T*, and *N* represent the simulated electric field, magnetic field, fluid velocity, plasma temperature, and density, respectively. The index *q* refers to the simulation time step number. The inclusion in the time-stepping algorithm of on-grid updates to perturbations of the plasma medium allows nonlinear plasma processes such as the resonance instability or self-focusing instability to develop naturally in the simulation volume.

#### 3. Numerical Simulation of Radio Window

A vital process in understanding the magnetic zenith effect is thought to be the transition from O-mode polarized pump wave to a Z-mode (slow branch X-mode) wave that may occur as the pump wave approaches the interaction height [Miølhus and Fla, 1984; Miølhus, 1990]. This process is highly dependent on the initial inclination angle of the pump wave with respect to the direction of the geomagnetic field near the O-mode reflection point. Conditions most favorable for conversion to Z-mode occur when the k vector of the incident wave is close to being parallel to the geomagnetic field direction at the height at which the transformation takes place ( $\omega_{ne} = \omega_0$ , where  $\omega_0$  is the frequency of the pump wave). This condition corresponds to the *L*-mode wave near the  $k_{\perp} = 0$  limit of the O-mode dispersion surface; this wave is able to continue beyond  $\omega_{pe} = \omega_0$  on the Z-mode branch of the X-mode dispersion surface [Leyser and Nordblad, 2009]. The angular dependence leads to the formation of a "radio window" as described in, for example, Mjølhus [1984]: a range of inclination angles for which the pump wave is partially or fully transmitted beyond the O-mode reflection height as a Z-mode wave. As the direction of the incident wave  $\mathbf{k}$ -vector is governed by the refractive properties of the local plasma the wave passes through, the spatial profile of ionospheric plasma around the interaction height has a significant influence on the angular position and characteristics of the radio window, and thus on the development of the magnetic zenith effect. A strength of a full-wave numerical code such as that described in section 2 is that it allows wave propagation through plasma to be accurately computed in scenarios for which analytical evaluation may be impossible without resorting to constraining regimes of approximation, for example, those employed by geometric optics or similar [Budden, 1961; Ginzburg, 1970]. The FDTD scheme used here automatically incorporates linear mode conversion processes such as the O-mode to Z-mode transition into the time-domain evolution of the simulated EM fields and the coupled on-grid updates to plasma density and temperature perturbation mean that the time-dependent and nonlinear plasma response to an incident wave is directly included in any calculation. This section presents the results of a series of simulations run using the FDTD code described in section 2 to investigate the effect of different plasma density profiles on the characteristics of the ionospheric radio window.

The simulation was set up as a two-dimensional scenario, with all wave propagation restricted to lie in the *xz* plane, and all medium properties assumed to be homogeneous in the *y* direction. A computational domain of dimensions 1280 × 1 × 1280 Yee cells [*Yee*, 1966] was used, with the discrete spatial step size ( $\Delta_x$ ) set as 5.50 m in all directions. To ensure numerical stability the Courant number [*Taflove and Hagness*, 2000] was chosen to be 0.5, which fixed the size of the discrete time step ( $\Delta_t$ ) to be 9.17 × 10<sup>-9</sup> s. The simulation was terminated on all sides by a 32 cell complex frequency-shifted perfectly matched layer (CFS-PML) designed to ensure that any waves leaving the domain were absorbed with the minimum possible numerical reflection [*Berenger*, 1994]. For the plasma temperature and density, edge termination was achieved using a Neumann boundary condition that set the gradient of each to zero along the direction of the background magnetic field. A wave launching layer based on the principles of the total field scattered-field implementation described in *Schneider* [2004] was used to introduce continuous electromagnetic waves of *O*-mode polarization into the domain to represent the heating pump wave. The wave frequency was chosen to be  $\omega_0 = 2\pi \times 4.54$  MHz. Background plasma parameters were chosen to represent typical *F* region conditions during a heating experiment at Tromsø. Typical plasma parameters were obtained using IRI-2012 [*Bilitza et al.*, 2014]. An idealized ionospheric plasma was represented by a 1-D linear vertical electron density profile of the form (7).

$$N_0(z) = N_{\rm crit} \left( 1 + \frac{z - z_{\rm crit}}{L_z} \right)$$
(7)

In this expression,  $N_{crit}$  is the electron density at  $X = \frac{\omega_{pe}^2}{\omega_0^2} = 1$ , where  $\omega_{pe}$  is the fundamental plasma frequency for electrons.  $z_{crit}$  is the distance above the lower edge of the simulation at which the critical density occurs, here set to be 4.11 km above the z = 0 plane.  $L_z$  is the scale height of the density gradient, set to be 20 km. The background electron temperature was set to be 1800 K. A population of ions of average mass 16 AMU (representing  $O^+$  ions) was included with a background density distribution equivalent to that of the electrons to preserve neutrality, and an initialized temperature of 900 K. The unperturbed collision frequencies for electrons and ions were set to be 500 s<sup>-1</sup> and 6 s<sup>-1</sup>, respectively. A static magnetic field of magnitude  $|\mathbf{B}| = 4.16 \times 10^{-5}$  T was initialized in the domain at an angle of  $\alpha = 12^\circ$  to the vertical. This was set to be uniform throughout the simulation domain, hence resulting in electron and ion gyrofrequencies of  $\omega_{ce} = 2\pi \times 1.17$  MHz and  $\omega_{ci} = 2\pi \times 39.7$  Hz, respectively.

To investigate the angular characteristics of the *O*- to *Z*-mode conversion process, a set of simulations were performed with the launch angle of the pump wave with respect to the vertical direction,  $\theta$ , varied from  $\theta = 0^{\circ}$  (pump wave **k** vector directed vertically and in this case along the direction of the electron density gradient) to  $\theta = 12^{\circ}$  (pump wave **k** vector along the direction of the simulated geomagnetic field) in  $\Delta \theta = 0.01$  rad (0.57°) steps. The simulations were run for  $1 \times 10^{5}$  time steps (equivalent to  $9.17 \times 10^{-4}$  s), which was a sufficiently long time for a steady *E* field pattern to develop, but short compared to the timescales required for electron plasma density and temperature perturbations due to the action of the radio wave to develop.

Each trace in Figure 1 (top) shows the time-averaged *E* field amplitude measured along the original direction of propagation for a selection of simulated launch angles. In each case, a swelling effect can be seen as the pump wave approaches the  $X^{num} = 1$  point at  $z_c$  where the wave is reflected ( $X^{num}$  is a numerical equivalent of variable *X* which takes into account the dispersive and dissipative effects of the discrete computational grid [*Cummer*, 1997]). At this point the *O*-mode pump wave is reflected downward, leading to the generation of the large-amplitude standing waves visible in each plot around the interaction height. In most cases it can be seen that little or no wave amplitude is able to penetrate beyond the *O*-mode reflection point; however, it is apparent that there is a window around 6° for which propagation to higher altitudes is viable. This window is the range of angles for which mode conversion from the initial ordinary wave to an extraordinary-polarized *Z*-mode wave is favorable and corresponds to a set of angles for which the **k** vector of the incident wave is close to parallel to the local geomagnetic field as the wave reaches the reflection height. The optimal angle for this process to occur in a vertically inhomogeneous ionosphere is given by the Spitze angle (8) and is a function of the plasma medium parameters, where  $Y = \frac{\omega_{ce}}{\omega_0}$ ,  $\omega_{ce}$  is the electron cyclotron frequency and  $\alpha$  is the angle between the local geomagnetic field and the vertical.

$$\theta_{\text{Spitze}} = \arcsin\left[\left(\frac{\gamma}{\gamma+1}\right)^{\frac{1}{2}}\sin(\alpha)\right]$$
(8)

In the scenario described here, the value of the Spitze angle was calculated to be  $5.5^{\circ}$ . The theoretical propagation paths for rays of varying inclination angle in the ionosphere can be seen, for example, in Figure 1 of *Isham et al.* [2005]. The average fraction of wave amplitude penetrating beyond the *O*-mode reflection height for each angle is indicated in Figure 1 (bottom) by the red markers and shows a Gaussian-shaped transmission window centered on  $5.3^{\circ}$  and with a full width at half maximum of  $3.1^{\circ}$  (Gaussian fit is indicated by the dashed black line). The theoretical window for transmission of the EM wave through the reflection barrier via *O*-to-*Z* conversion as calculated by Mjolhus [*Mjølhus*, 1984] for the density profile used here is indicated by the blue line.

A small *E* field swelling at the reflection height is seen in Figure 1 (top) for waves directed close to the Spitze angle despite the *Z*-mode conversion efficiency being close to 100%. This is due to a combination of effects: anisotropic divergence in wave direction due to numerical dispersion and the slight spread in propagation angle around the pump wave edges both lead to the reflection of a small fraction of the pump wave.

The linear density slope used in the simulation is only a simplified representation of the *F* region of the ionosphere. Experiments such as *Kosch et al.* [2011] have observed the magnetic zenith effect to manifest most strongly for inclination angles between the Spitze and field-aligned directions. *Isham et al.* [2005] have observed the maximum topside backscattered power to occur in this range, indicative of the radio window shifting toward the magnetic zenith. One mechanism that may be responsible for this shift is the presence of a horizontal density gradient in addition to the vertical gradient. As mentioned in section 1 above, a varied



**Figure 1.** Angular dependence of *O*-mode to *Z*-mode conversion for linear vertical density gradient. (top) *E* field amplitude averaged over  $9.17 \times 10^{-4}$  s ( $1 \times 10^{5}$  time steps) for a selection of inclination angles. Penetration of the *O*-mode reflection height 4.1 km above the lower plane of the computational domain can be seen to occur most effectively for angles close to the Spitze at 5.5°. (bottom) The simulated radio transmission window (red points, dashed black Gaussian fit) and the theoretical prediction from *Mjølhus* [1984] (blue).

hierarchy of density structures are likely to arise when the ionosphere is illuminated by high-power radiation due to the formation of plasma irregularities of different scales. In particular, the density gradient in the horizontal direction is unlikely to be homogeneous as assumed in the above analysis. *Honary et al.* [2011] obtained a theoretical expression for the shift in the radio window from the Spitze angle toward the magnetic zenith; however, this calculation assumed that the ionospheric inhomogeneity existed along one direction only (which in the general case can deviate from the vertical direction). The case of 2-D variation on the plasma density has not been discussed before and can only be investigated numerically.

To this end, the FDTD code was used to investigate the effect that density structures of various spatial scales have on the position of the *Z*-mode window. Three different inhomogeneity regimes were investigated: large-scale 2-D linear gradients, single field-aligned duct-like density depletions, and periodically arranged smaller-scale field-aligned irregularities.

To simulate a large-scale density-depleted patch, such as that caused by the bunching together of several nonlinear density structures in the heated volume, a linear horizontal slope of scale size  $L_x$  with density increasing from left to right in the computational domain was added to the background electron density profile. Examples of the 2-D background electron density profiles for a selection of  $L_x$  can be seen in Figure 2 (top). A computational domain of dimensions  $1920 \times 1 \times 1920$  was used, with the discrete step sizes set to



**Figure 2.** Simulated Z-mode window for varying horizontal slope scale size  $L_x$  when background density profile (9) is used as a background density profile (smaller  $L_x$  implies a steeper slope). It can be seen that increasing the steepness of the slope shifts the center of the radio window away from the Spitze direction and toward the magnetic zenith. The background plasma density, magnetic field direction (dashed line), vertical direction (dotted line) and position of the *O*-mode reflection height (solid black line) used for each  $L_x$  are shown in Figure 2 (top).

 $\Delta_t = 4.58 \times 10^{-9}$  s and  $\Delta_x = 2.75$  m. All other simulation parameters were set as described above. The full form of the background density profile in this case is given by equation (9).

$$N_0(x, z) = N_{\text{crit}} \left( 1 + \frac{x}{L_x} \right) \left( 1 + \frac{z - z_{\text{crit}}}{L_z} \right)$$
(9)

The vertical slope scale size was kept constant at  $L_z = 20$  km. The position and width of the Z-mode window was measured for different grades of horizontal slope, changed by varying  $L_x$ , with the results shown in Figure 2. The introduction of a horizontal density inhomogeneity of this form was found to shift the center of the radio window away from the Spitze direction and toward the magnetic zenith direction, with larger shifts experienced as the gradient was increased. The angular width of the window was also found to increase overall as the horizontal gradient increased, whereas the maximum transmitted *E* field amplitude was found to decrease. This result supports the theory that the observed angular position of peak magnetic zenith effect could be explained by the presence of a horizontal density slope. For the vertical plasma gradient used here, a peak in the Z-mode window at around  $\theta \simeq (8 - 10)^\circ$  would correspond to a horizontal inhomogeneity of around  $L_x \simeq 50$  km. In the auroral ionosphere, horizontal gradients in the opposing direction (increasing from right to left in the simulation domain) may also exist. In this case, the density gradient has the effect of shifting the center of the Z-mode window away from the Spitze and toward the vertical direction, again with a greater deviation occurring for a greater steepness of horizontal slope.

To simulate the effects of a duct-like horizontal density feature, a field-aligned density depletion of the form equation (10) was included in the background density profile.

$$N_{\rm irreg}(x', z') = N_0(x', z') \left[ 1 - A_{\rm irreg} \exp\left(-\frac{(x_0 - x')^2}{2L_{\rm width}^2}\right) \right],$$
(10)

where x' and z' are coordinates perpendicular and parallel to the simulated geomagnetic field,  $A_{irreg}$  and  $L_{width}$  are the amplitude and width of the irregularity respectively,  $x_0$  is a reference distance, and  $N_0(x', z')$  is the vertically inhomogeneous background density profile given by equation (7), with the vertical slope scale height kept at  $L_z = 20$  km as before.  $A_{irreg}$  was set to be 0.05 to represent a maximum density depletion of 5% of the background. A smaller computational domain of dimensions  $1280 \times 1 \times 1280$  was used to reduce computing time, with the discrete step sizes set to  $\Delta_t = 9.17 \times 10^{-9}$  s and  $\Delta_x = 5.50$  m.  $L_{width}$  was varied to investigate a



**Figure 3.** Simulated *Z*-mode window for varying duct width  $L_{width}$  when density perturbation of equation (10) is included in the background density profile. Initial amplitude of perturbation was set as 5% of the unperturbed background density given by equation (7). Smaller-width structures with  $L_{width} \le 0.1$  km can be seen to shift the distribution away from the Spitze direction toward vertical by  $(0.5-1)^\circ$ . As the irregularity width increases, the window broadens and loses its Gaussian shape. For widths  $\ge 0.2$  km, a significant fraction of incident wave amplitude was transmitted at all sampled angles, leading to a dramatic widening and flattening of the window. The background plasma density, magnetic field direction (dashed line), vertical direction (dotted line) and position of the *O*-mode reflection height (solid black line) used for each  $L_{width}$  are shown in Figure 3 (top).

range of irregularity scale sizes, from small-scale single striations with a width of tens of meters, up to larger ducts with widths of several kilometers. Examples of the 2-D background electron density profiles for a selection of  $L_{width}$  can be seen in Figure 3 (top). The effect of different perturbation widths on the position of the radio window is shown by Figure 3.

The presence of the irregularity was found to have a strong effect on the *O*- to *Z*-mode conversion process, with the resulting radio window falling into one of several broad regimes depending on the irregularity width. Smaller-scale irregularities with widths  $\leq 0.1$  km were found to have a minimal effect on the *Z*-mode window shape, resulting in only a slight shift the radio window away from the Spitze position and toward vertical by  $(0.5-1)^\circ$  measured. For duct widths in this range, the radio window was found to be broader than in the unperturbed case, with a full width at half maximum in the range  $(4.0-4.5)^\circ$ . Increasing the perturbation width in the range 0.1km  $\leq 2$  km was found to lead to an abrupt change in radio window shape, with a significant fraction of incident wave amplitude transmitted as a *Z*-mode for all sampled angles. This manifested as a broadening of the window as the scale size of the irregularity was increased, with the angular width of the window for a duct of 1 km width almost three times that of the 1-D slope case. Increasing the irregularity width also had the effect of decreasing the *Z*-mode amplitude transmitted of the *Z*-mode less than half that transmitted at the window central angle for the 1-D slope regime. As the irregularity width was increased beyond 2 km, the shape of the *Z*-mode window was found to tend toward the unperturbed form measured for the case of a linear vertical slope only.

Figure 4 shows the *E*-filed amplitude developed in the simulation domain and averaged over  $9.17 \times 10^{-4}$  s  $(1 \times 10^{5}$  time steps) for pump wave inclination angles of 0.6°, 2.9°, 5.2°, 7.4°, 9.7°, and 12.0° and for the cases of irregularities of width 0.08, 0.10, 0.20, 0.50, and 0.80 km. The color range of this plot has been logarithmically normalized to more clearly show the transmitted *Z*-mode waves. The position of the X = 1 *O*-mode reflection contour is marked by a solid line. The *Z*-mode can be identified as signal which is able to penetrate beyond the *O*-mode reflection height. Around the region of the duct it can be seen that the *Z* mode is transmitted as one or more narrow beams only, with the point(s) of transmission varying greatly according to the pump wave angle and duct geometry. For irregularities of width > 0.1 km, it can be seen that a strong *Z*-mode beam



**Figure 4.** Time-averaged *E* field amplitude for pump wave inclination angles of  $0.6^{\circ}$ ,  $2.9^{\circ}$ ,  $5.2^{\circ}$ ,  $7.4^{\circ}$ ,  $9.7^{\circ}$ , and  $12.0^{\circ}$  when density perturbations of the form equation (10) and with  $L_{width}$ s of 0.08, 0.10, 0.20, 0.50, and 0.80 km are included in the background density profile. The position of the X = 1 *O*-mode reflection contour marked by a solid line. The color range has been logarithmically normalized to more clearly show the transmitted *Z*-mode waves passing beyond the *O*-mode reflection height. Figure 4 (top) shows the background electron density, position of the *O*-mode reflection layer (solid black line), vertical direction (dotted black line) and magnetic field direction (dashed black line) for each  $L_{width}$  condition.

is transmitted for all angles shown (albeit emerging from very different points of the duct), leading to the flat, broad window profiles shown in Figure 3. Away from the region around the duct, behavior reverts to that of the unperturbed case described above, with optimal *O*- to *Z*-mode conversion occurring at the Spitze angle. Only waves transmitted within a full width of the irregularity central axis were considered when calculating the shape of the window.

The effect of including multiple periodically arranged field-aligned density depletions was investigated by including density irregularities of the form (11) to the background density profile.

$$N_{\text{irreg}}(x',z') = N_0(x',z') \left\{ 1 - \frac{1}{2}A_{\text{irreg}} \left[ 1 - \cos\left(\frac{2\pi x'}{L_{\text{width}}}\right) \right] \right\}$$
(11)

Similarly to the duct-like irregularity simulation above, computational domain dimensions of  $1280 \times 1 \times 1280$ were used, with the discrete step parameters set to  $\Delta_t = 9.17 \times 10^{-9}$  s and  $\Delta_x = 5.50$  m. The vertical slope scale size was set to be  $L_z = 20$  km as before. Figure 5 shows the *E*-filed amplitude developed in the simulation domain and averaged over  $9.17 \times 10^{-4}$  s ( $1 \times 10^{5}$  time steps) for pump wave inclination angles of  $0.6^{\circ}$ ,  $2.9^{\circ}$ ,  $5.2^{\circ}$ , 7.4°, 9.7°, and 12.0° and for the cases of periodic irregularities of  $L_{\text{width}} = 0.08, 0.10, 0.20, 0.50, \text{ and } 0.80 \text{ km}$ . As with the single-irregularity case above, the X = 1 contour is shown by a solid line and the color bar has been logarithmically normalized to more clearly show any transmitted or trapped waves. The 2-D background electron density profiles for each  $L_{width}$  can be seen in Figure 5 (top). In the case of a single duct the transmitted Z-mode wave can emerge from either side of the structure, or even multiple points simultaneously, depending on the inclination angle of the source wave, as shown by Figure 4. In the case of periodic structures, this results in a converted Z-mode potentially entering multiple adjacent irregularities before emerging. This has a strong effect on the propagation of the converted Z-mode, with transmitted waves multiply scattered by adjacent irregularities making it difficult to measure the Z-mode window shape. In general the scattering effect of periodic irregularities with spatial scales comparable to or smaller than the pump wave wavelength  $(\leq 0.08 \text{ km})$  had the effect of attenuating the transmitted Z-mode amplitude at all angles while maintaining the angular window shape measured for the unperturbed case. As the irregularity width increased, multiple scattering of the Z-mode wave between adjacent irregularities resulted in the presence of high-amplitude E field waves beyond the O-mode reflection height for all sampled pump wave angles. Scattered field beyond the reflection height was found to have amplitudes comparable to that found in the Airy field below the reflection height and in some cases was well in excess of the maximum Z-mode amplitude obtained beyond X = 1 for the unperturbed 1-D slope case (shown in Figure 1).

#### 4. Numerical Simulation of Heating Effects

In the previous section, the FDTD code was used to show that the *O*- to *Z*-mode conversion process is strongly dependent on the electron plasma profile around the interaction height. In this section, simulation results demonstrating that such behavior can directly lead to an angular dependence in heating-induced effects such as electron temperature enhancement and the growth of density-depleted field-aligned irregularities are presented.

In an idealized heating experiment with a 1-D linear vertical electron density gradient only, an *O*-mode wave launched at the Spitze angle will be almost fully converted to the *Z*-mode as it approaches the interaction height. The converted *Z*-mode wave passes beyond the *O*-mode reflection barrier and, provided the ionospheric peak density is sufficiently high, is reflected at X = Y+1. The reflected *Z*-mode wave can then freely propagate downward, back toward the interaction region. As it approaches the point where  $X = (1-Y^2)/(1-Y^2\cos^2\alpha)$ , close to the O-mode reflection height, the reflected *Z*-mode wave encounters the *X*-mode resonance layer. This resonance is normally inaccessible to ground-launched fast *X*-mode waves which are reflected at X = 1-Y but when excited results in the very efficient conversion of the EM wave to upper hybrid plasma waves [*Gondarenko et al.*, 2002]. This process leads to a large increase in the *E* field amplitude measured along the direction of propagation is shown for a selection of inclination angles (panels 6.0–6.4). The simulation parameters described in section 3 for the case of an undisturbed 1-D vertical density gradient were used, along with simple linear background electron and ion density profiles as given by equation (7), with  $L_z = 20$  km. The blue curve shows *E* field amplitude for the case when a perfectly matched layer (PML) was used to absorb the *Z*-mode wave as it reaches the upper boundary of the simulation, suppressing



**Figure 5.** Time-averaged *E* field amplitude for pump wave inclination angles of  $0.6^{\circ}$ ,  $2.9^{\circ}$ ,  $5.2^{\circ}$ ,  $7.4^{\circ}$ ,  $9.7^{\circ}$ , and  $12.0^{\circ}$  when periodic density perturbations of the form (11) and with  $L_{width}$ s of 0.08, 0.10, 0.20, 0.50, and 0.80 km are included in the background density profile. The position of the X = 1 *O*-mode reflection contour is marked by a solid line. The color range has been logarithmically normalized to more clearly show the transmitted and scattered *Z*-mode waves beyond the *O*-mode reflection height. Figure 5 (top) shows the background electron density, position of the *O*-mode reflection layer (solid black line), vertical direction (dotted black line) and magnetic field direction (dashed black line) for each  $L_{width}$  condition.



**Figure 6.** (top) Comparison of the *E* field amplitude averaged over  $1 \times 10^5$  time steps ( $9.17 \times 10^{-4}$  s) for the case that the upward boundary of the computational domain is terminated with an absorbing PML (blue) or a reflecting layer (red-dashed). Traces are shown for a selection of initial pump wave inclination angles. (bottom) The variation of maximum averaged *E* field amplitude with inclination angle. It can be seen that allowing the *Z*-mode wave to reflect enhances the amplification of *E* field around the interaction region, particularly for waves directed toward the center of the radio window around 5.5°.

its reflection. The dashed-red curve shows the same signal, but for the case that the upper boundary of the simulation was set to be a reflecting boundary to simulate the reflection of the *Z*-mode wave at higher altitudes. A partial reflector was used in preference to a perfect reflector as the *Z*-mode wave would expect to experience partial attenuation due to collisional damping, particularly as it approaches X = Y + 1. The variation in the peak *E* field amplitude in the interaction region with pump wave inclination angle between the *Z*-mode reflection-allowed and *Z*-mode reflection-suppressed cases is shown in Figure 6 (bottom). Allowing the *Z*-mode wave to reflect and propagate back toward the resonance layer was found to lead to a significantly enhanced amplification of the *E* field around the interaction region. This effect was particularly pronounced for waves directed toward the center of the radio window at around 5.5° from vertical, with the maximum magnitude of *E* field amplitude found to be greater than 6 times that developed under similar conditions with *Z*-mode reflection suppressed.

The ability of the reflected Z-mode to excite large-amplitude plasma waves makes a clear difference to the magnitude of *E* field standing wave that forms around the interaction region and can drive the formation of artificially induced perturbations to the plasma medium. To investigate the impact of this on the growth of

temperature and density perturbations in the heated region, two parallel simulations were performed: the first with Z-mode reflection suppressed using a PML layer at the upper boundary of the computational domain and the second with Z-mode reflection allowed through the inclusion of a reflecting boundary as described above. The simulation domain was set up with the smaller dimensions of  $896 \times 1 \times 640$  Yee cells. The smaller size reduced the computing time required to calculate each simulation cycle and allowed longer-duration runs to be performed. The time step size was set to be  $1.83 \times 10^{-8}$  s. The Courant number was maintained at 0.5 to preserve stability, and the discrete spatial step size set to be 11.0 m in all directions. A vertical gradient (7) was used as the background density profile with  $L_z = 20$  km, and  $z_c = 4.46$  km above the bottom edge of the simulation box, as indicated the 2-D background density plots shown in Figure 7 (panel 7.0) and Figure 8 (panel 8.0). Initially, no horizontal inhomogeneity was included. For this background density profile, the upper hybrid resonance point occurred at 3.16 km above the bottom edge of the simulation. This domain setup was designed to include the important regions of interaction such as the reflection and upper hybrid resonance heights, while keeping the overall domain size small to reduce the computational run time. An O-mode pump wave of frequency  $\omega_0 = 2\pi \times 4.54$  MHz was launched at an angle of 5.16° from the lower edge of the computational domain. As this inclination angle is close to the center of the radio window, it was expected to lead to efficient conversion of the pump wave to the Z-mode. The beam was given a Gaussian profile to keep most of the source wave energy away from the domain boundaries. All other domain parameters were set as described in section 3 for the case of the undisturbed 1-D vertical density gradient.

Figures 7 and 8 illustrate the variation of E field, electron density perturbation (expressed as a fraction of the background electron density,  $N_{e0}$ , fractional electron density irregularity amplitude  $\delta N/N_{e0}$ , and temperature perturbation evolved in the simulation after  $1.83 \times 10^{-4}$  s,  $1.83 \times 10^{-2}$  s,  $9.17 \times 10^{-2}$  s,  $1.83 \times 10^{-1}$  s,  $2.93 \times 10^{-1}$  s,  $1.83 \times 10^{-1}$  s, and 1.10 s for the cases of Z-mode reflection suppressed (Figure 7) and allowed (Figure 8). The irregularity amplitude  $\delta N$  is calculated by subtracting a local average (taken along the horizontal x direction) from the electron density perturbation, such that  $\delta N = N_e - \langle N_e \rangle_x$ . Both figures show a rapid rise of the E field amplitude at around  $z_c = 4.46$  km corresponding to the high-amplitude reflected standing wave, which results in significant rise of the E field amplitude and an increase (due to the wave dissipation) of the electron temperature, which saturates after  $\sim 0.3$  s. Allowing the Z-mode to reflect had a significant effect on the E field strength developed in the simulation, with the saturation E field amplitude achieved with reflection allowed over 10 times greater than with reflection suppressed. For the case of suppressed reflection, the highest field amplitudes were located around the reflection height at  $z_c$  and remained relatively steady for the duration of the simulation. By contrast, with reflection allowed, E field amplitudes of up to 6 V/m were developed and located in a narrow band initially around z<sub>c</sub> corresponding to the location of the resonance. This can be seen to move upward in altitude with time as the simulation progresses, as shown by panels 8.1-6a of Figure 8. The Z-mode reflection allowed scenario was found to produce around double the increase in temperature compared to the reflection suppressed scenario. For both reflection suppressed and allowed cases, a strong electron density depletion is observed around the region of interaction. In the reflection-suppressed scenario, this takes the form of a pair of density-depleted patches located on the shoulders of the Z-mode window. The density depletions are correlated with the regions where the pump wave is reflected and the reflection-induced standing wave forms (toward the edges of the beam rather than toward the beam center where the O-mode wave is transmitted as a Z-mode). This suggests that the formation of the density cavities may, in part, be due to a ponderomotive interaction or the excitation of a parametric instability. Around the peak of the window, the EM wave is mostly transmitted as the Z-mode and the swollen-amplitude reflected standing wave does not form; hence, the density in this region is comparatively unperturbed. In the reflection-allowed scenario, the density perturbation takes the form of a single density-depleted cavity with a maximum depth of  $|N_{m}^{\text{ini}}| \sim 0.02N_{e0}$ . This is greater by almost a factor of 4 than the maximum magnitude of perturbation observed in the reflection-suppressed simulation. In this case, the site of maximum depletion can be seen to rise in altitude as the simulation progresses. This is caused by the downward traveling Z-mode wave encountering a region of plasma that matches the extraordinary wave resonance condition  $X = (1 - Y^2)/(1 - Y^2 \cos^2 \alpha)$ as it enters the depleted density patch. Due to the presence of the depletion, the resonance is encountered at a higher altitude than if there had been no density perturbation. High E field amplitudes are excited as a result, which accelerates the depletion of density around that region and leads to a gradual rise in altitude of the region of maximum density depletion as the simulation progresses. This process also explains why the band of maximum E field amplitude moves upward in altitude with time.



**Figure 7.** Change in the simulated *E* field amplitude, electron density perturbation (expressed as a fraction of the background density  $N_{e0}$ ), density irregularity amplitude ( $N_e - \langle N_e \rangle_x$ ), and electron temperature perturbation with time, when Z-mode reflection was suppressed. Spatial snapshots of each quantity are shown for times  $1.83 \times 10^{-4}$  s,  $1.83 \times 10^{-2}$  s,  $9.17 \times 10^{-2}$  s,  $1.83 \times 10^{-1}$  s,  $2.93 \times 10^{-1}$  s, and 1.10 s. The background density profile, geomagnetic field orientation (solid black line), *O*-mode reflection height (dot-dashed black line) and upper hybrid resonance height (dotted black line) are shown in panel 7.0. Field-aligned density irregularities can be seen to grow with time around the UHR height, as indicated in panel 7.4c.



**Figure 8.** Change in the simulated *E* field amplitude, electron density perturbation (expressed as a fraction of the background density  $N_{e0}$ ), density irregularity amplitude ( $N_e - \langle N_e \rangle_X$ ), and electron temperature perturbation with time, when Z-mode reflection was allowed. Spatial snapshots of each quantity are shown for times  $1.83 \times 10^{-4}$  s,  $1.83 \times 10^{-2}$  s,  $9.17 \times 10^{-2}$  s,  $1.83 \times 10^{-1}$  s,  $2.93 \times 10^{-1}$  s, and 1.10 s. The background density profile, geomagnetic field orientation (solid black line), O-mode reflection height (dot-dashed black line) and upper hybrid resonance height (dotted black line) are shown in panel 8.0.

An interesting feature to note is the spontaneous formation of small-scale field-aligned density structures visible at after  $\sim 0.1$  s in the simulation. These are particularly clear in the plots of the density irregularity amplitude shown by Figure 7, panels 7.4c (in which the features are indicated by an arrow) and 7.5c, and can also be seen for the reflection-allowed case in Figure 8, particularly panels 8.4c and 8.5c. These appear to evolve naturally in the simulation, first appearing in a narrow band located just below the point at which the background upper hybrid frequency matches with the pump wave frequency, and subsequently increase in both depth and elongation with time. At later times they can be seen to bunch together into larger-scale structures. Due to the location and orientation of these structures, it is clear that they are formed in the simulation due to the conversion of the pump wave to upper hybrid waves; the altitude at which the structures initially form is located just below the upper hybrid resonance height where the pump wave frequency matches the upper hybrid frequency, and the structures are exactly aligned along the geomagnetic field direction, consistent with a wave propagating in a direction perpendicular to this magnetic field. Changing the position of the upper hybrid resonance height within the computational domain was found to change the position of this feature, with the initial location of the irregularity perturbation consistently found to originate in a band  $\sim$  400 m below the UHR height before extending with time to both higher and lower altitudes along the magnetic field direction.

Resolvable field-aligned density irregularities were visible after as little as 0.1 s; this was unexpected as similar structures have been observed experimentally to require timescales of a few seconds or more to develop. This discrepancy may be partly due to the low amplitudes of the simulated irregularities: the maximum density irregularity amplitude developed over the simulation run was only  $|N_e - \langle N_e \rangle_x| < 0.001 N_{e0}$ , which may be difficult to detect experimentally in a real heating experiment. The early onset of irregularity development may also be linked to the numerical noise level in the simulation. The linear mode conversion process understood to govern UH generation from an *O*-mode pump wave requires small-amplitude seed density irregularities to be present before it may proceed; however, no such seed irregularities were included in the initial simulation density profiles. It is likely here that the seed for this process is provided by the low-amplitude numerical noise present in the simulated ion and electron densities. This is an artifact of the computational process and not physical in nature; however, it may play an equivalent role to the seeding effect of low-amplitude density variations in the real ionosphere. If the numerical noise is of a higher amplitude or grows at a faster rate than seed perturbations in the real ionosphere, it may explain the unexpectedly early appearance of the field-aligned structures in the simulation.

To investigate this, the simulation was repeated for varying temporal and spatial step sizes, while keeping the Courant number constant. It was found that the field-aligned density structures become visible after approximately  $6 \times 10^{6} \Delta_{t}$ , regardless of the magnitude of  $\Delta_{t}$ , when test simulations were performed for  $\Delta_{t} = 9.17 \times 10^{-9}$  s,  $\Delta_{t} = 1.83 \times 10^{-8}$  s, and  $\Delta_{t} = 3.67 \times 10^{-8}$  s. Subsequent to this, the growth rate of the irregularities was consistent with development due to a thermal physical process such as the resonance instability and scaled with the simulated time in seconds rather than with the number of time steps. This implies that the onset time of irregularity growth was numerical in nature, dependent on the growth of numerical noise in the simulated plasma density, and thus was determined by the simulation time step number. After the onset of the irregularities, their growth was physical in nature and progressed according to the physical timescale in the simulation rather than the numerical time step.

To investigate the angular dependence of heating-induced effects, the simulations were repeated for pump wave launch angles  $\theta$  of 0.1°, 1.7°, 3.4°, 5.2°, 6.9°, and 8.6°. The electron density and temperature perturbation developed after 1.1 s for each wave direction is shown in Figure 9, with the *Z*-mode reflection-allowed case on the left (panels 9.1–6a and 9.1–6c) and *Z*-mode reflection-suppressed case on the right (panels 9.1–6b and 9.1–6d). In the reflection-allowed scenario, the strongest electron temperature enhancement was found to occur for an angle of 5.2° (shown in panel 9.4c), which supports the argument that the pump wave that converts most efficiently to the *Z*-mode (in this case, the wave with an inclination angle closest to the Spitze) produces the greatest temperature perturbation. For the reflection-suppressed scenario, the wave launched in a near-vertical direction produced the greatest temperature enhancement (shown in panel 9.1d), as this wave lost the least energy to *Z*-mode transmission and thus resulted in a greater amplitude of standing wave below the *O*-mode reflection height. In the reflection-suppressed case, the Spitze-directed wave was almost fully transmitted through the reflection barrier and subsequently absorbed by the PML, hence causing the least temperature perturbation.



**Figure 9.** Variation of electron density and electron temperature perturbation with pump wave inclination angle, for the cases that *Z*-mode reflection is allowed and suppressed. In the reflection-allowed scenario, the strongest electron temperature enhancement was found to occur for an angle of 5.2° (close to the Spitze angle; shown in panel 9.4c), supporting the argument that the pump wave that converts most efficiently to the *Z*-mode produces the greatest temperature perturbation. In the reflection-suppressed scenario, the wave launched in a near-vertical direction produced the greatest temperature enhancement (shown in panel 9.1d). The Spitze-directed wave in the reflection-suppressed case was almost fully transmitted through the reflection barrier and subsequently absorbed by the PML, hence causing the least temperature perturbation (9.4d). The lower panels (9.7a–b) show the variation of the maximum value of electron temperature enhancement and electron density depletion in the simulation domain with time, for the *Z*-mode reflection allowed (solid lines) and suppressed (dashed lines) scenarios. Background conditions are shown in the upper panel (9.0).

Summary plots showing how the maximum value of electron temperature perturbation and the minimum value of electron density perturbation in the simulation domain changed with time can be seen in the lower panels of Figure 9 (panels 9.7a - 9.7b), for the cases of the pump wave directed above ( $8.6^\circ$ ), below ( $0.0^\circ$ ), and along the Spitze direction ( $5.2^\circ$ ). The Z-mode reflection-allowed scenario is shown by a solid line, with the reflection-suppressed case indicated by a dashed line. Allowing Z-mode reflection can be seen to lead to a higher eventual magnitude of heating-induced effects by as much as a factor of 2 for the case of the Spitze-directed wave. Pump waves with initial inclination angles closer to the Spitze direction (for which conversion to Z-mode was more favorable) were found to achieve a greater magnitude of temperature enhancement and a greater variation between the reflection-allowed and reflection-suppressed case; for example, when the incident wave was directed at  $5.2^\circ$ , the steady state temperature is achieved after 0.25 s in the reflection-allowed case, compared with 0.6 s in the reflection-suppressed case. This can be attributed to a greater rate of thermal diffusion in the reflection-allowed scenario caused by the much higher temperature, which allowed thermal equilibrium (steady state conditions) to be achieved more quickly.

The pump wave directed close to the Spitze direction was also found to be most effective at exciting small-scale field-aligned density irregularities around the upper hybrid resonance height. This is shown by Figure 10 which displays the change in the fractional electron density irregularity amplitude,  $(N_e - \langle N_e \rangle_x) / N_{e0}$ , with time in a narrow band of altitude around the UH resonance level. Results are shown for pump waves directed along 0.0° (panels 10.0–4b), 5.2° (10.0–4c), and 8.6° (10.0–4d) to the vertical, after times of 0.11 s (uppermost panels; 10.0b–d), 0.13 s, 0.15 s, 0.17 s, and 0.18 s (lowermost panels; 10.4b–d). The UH resonance height for which the local UH frequency matches the pump wave frequency ( $\omega_{UH} = \omega_0$ ) is indicated by a dashed line. In each case, clear field-aligned density perturbations can be seen to grow with time, first visible from ~ 0.1 s and persisting until they begin to collapse into less-ordered small-scale structures leaving a slight net density depletion. Initial growth of the irregularities in all cases occurs at an altitude ~ 0.4 km below the UH height, after which they extend with time to both higher and lower altitudes. The rate of growth of striations can be seen to vary with pump wave angle, with the wave directed closest to the Spitze angle (5.2°) causing the most rapid growth, both in terms of depletion depth and elongation along the magnetic field direction.

Panels 10.0–4a in Figure 10 show the 1-D spatial discrete Fourier transform (DFT) of the magnitude of the fractional density irregularity amplitude  $\delta N_e/N_{e0} = |N_e - \langle N_e \rangle_x|/N_{e0}$  for each angle, sampled along the horizontal axis of the simulation domain at the height at which the irregularities begin to develop. The DFT plots show that there is a clear increase in the perturbation amplitude for spatial frequencies around  $k_x \sim 0.18 \text{ m}^{-1}$ , which increases in amplitude with time as the irregularities grow. This range of spatial frequencies corresponds to irregularities with scale sizes of 30–40 m perpendicular to the geomagnetic field direction. The amplitude of density perturbation components in this scale-size range is highly dependent on pump wave angle, with the amplitude greater by a factor of 2 or more in the case of the Spitze-directed wave. From this it follows that the UH wave excitation process is dependent on wave angle, with UH production occurring most favorably at the same range of angles for which the *O*- to *Z*-mode conversion process was most favorable.

Production of UH waves and their dissipation in the associated field-aligned irregularities would make a significant contribution to the electron temperature enhancement via the resonance instability. As such, the dependence of irregularity production on pump wave angle demonstrated by these simulations could be expected to add a further contribution to the magnetic zenith effect. Small-scale density irregularities such as this may also affect the development of large-scale plasma perturbations through anomalous attenuation of the pump wave. The code used here treats absorption of the pump wave through inclusion of an effective collision term  $v_0$  in a manner similar to that described by *Gondarenko et al.* [2006]. The magnitude of irregularities developed over the simulation timescale is small (<  $0.001N_{e0}$ ); however, there is a small attenuation in the pump wave amplitude (< 1 dB for the vertically directed wave) which may be due to absorption caused by irregularities.

These results show that the *O*-mode to *Z*-mode conversion process could be an important mechanism behind the observed magnetic zenith effect, with the observed angular dependence arising from the shape of the *Z*-mode window and the resonant excitation of large-amplitude electric fields via the reflected *Z*-mode contributing to the enhanced plasma perturbations.



**Figure 10.** (right) The simulated growth of small-scale field-aligned density irregularities with time in a narrow band of altitude below the UH resonance level for pump waves directed along  $0.0^{\circ}$  (panels 10.0-4b),  $5.2^{\circ}$  (10.0-4c), and  $8.6^{\circ}$  (10.0-4d) to the vertical. The UH resonance height is indicated by a dashed line. (left) The 1-D spatial DFT of  $\delta N/N_{e0}$  sampled along the horizontal axis of the simulation domain at the height at which the irregularities begin to develop. Amplitude of spatial frequency components around this value increases with time corresponding to the growth of irregularities and is greater by a factor of 2 or more in the case of the  $5.2^{\circ}$ -directed wave.

#### 5. Inclusion of Density Structures

The previous section demonstrates that the *O*-mode to *Z*-mode conversion process can lead to a variation in the magnitude of artificial plasma perturbation with the inclination angle of an incident *O*-mode wave, for the case of a 1-D linear vertical plasma density gradient. The addition of a component of horizontal inhomogeneity has been shown in section 3 to significantly modify the *Z*-mode window; thus, it is to be expected that the presence of a horizontally inhomogeneous density feature will also affect the dependence of plasma perturbation on pump wave direction. This section presents the results of simulations performed for various pump wave directions and for the cases that a linear horizontal density gradient, a single field-aligned density-depleted duct, or a periodic distribution of field-aligned density-depleted irregularities are included in the background density profile.

Simulations were performed for pump wave initial inclination angles 0.0°, 2.9°, 5.7°, 8.6°, and 11.5° with *Z*-mode reflection allowed. Computational domain dimensions were set to be  $1024 \times 1 \times 768$ . For increased accuracy, the time step size was reduced to  $1.22 \times 10^{-8}$  s, and the spatial step size to 7.33 m. All other simulation parameters were set as in section 4.

Figure 11 (top) shows the electron density and temperature perturbation developed in the simulation after 0.23 s when a 2-D density profile including a linear horizontal slope as given by equation (9) was used. Results for horizontal scale sizes  $L_x = 20$  km (11.1–5a and 11.1–5c) and  $L_x = 50$  km (11.1–5b and 11.1–5d) are shown. In both cases the vertical slope scale size was set to be  $L_z = 20$  km. The lowermost panels show the variation of minimum density perturbation (panel 11.6a) and maximum temperature perturbation (panel 11.6b) recorded in the simulation with time. The  $L_x = 20$  km case is indicated by a solid line, the  $L_x = 50$  km case by a dashed line.

The most striking difference between the results shown in Figure 11 and the unperturbed case with no horizontal inhomogeneity shown in Figure 9 is the shifting of the inclination angle responsible for the greatest temperature enhancement from the Spitze position toward the magnetic zenith direction with decreasing  $L_x$ . For  $L_x = 50$  km, the greatest temperature enhancement is produced by the wave launched at 8.6° (shown in panel 11.4d). This is consistent with the modification of the *Z*-mode window found by the simulations reported in section 3 for the case of a linear horizontal density inhomogeneity, as shown in Figure 2, which predicts maximum *Z*-mode transmission to occur for pump waves directed at an angle of 8.2° for the  $L_x = 50$  km case. Likewise, the analysis of section 3 shows *O*- to *Z*-mode conversion to be more favorable as the pump wave direction approaches the zenith direction for  $L_x = 20$  km, in good agreement with the simulations performed here which show that the temperature enhancement increases as the pump wave inclination angle increases, with maximum enhancement found for the pump wave directed at 11.5° (shown in panel 11.5c).

Similar plots are shown in Figures 12 and 13 for the cases of a single density-depleted field-aligned duct and a periodic distribution of such irregularities, respectively. In both scenarios, an inhomogeneities with scale parameters  $L_{width} = 0.08$  km and  $L_{width} = 0.8$  km are compared.

When the single, narrow (0.08 km) irregularity was included, the simulation behaved in a similar manner to the unperturbed case, as would be expected from Figure 3 which suggests that the *Z*-mode window would be relatively unmodified by irregularities of this scale, with only a slight (< 1°) shift toward vertical. In this case, the maximum temperature enhancement occurred for the pump waves directed close to the Spitze angle (5.7°) and between the Spitze and vertical directions (2.9°) consistent with the small shift in the *Z*-mode window. The maximum stationary-state temperature was similar to that shown for the unperturbed 1-D slope case in Figure 9. At all angles, the largest density modification occurred away from the position of the density duct. For the larger-scale (0.8 km) irregularity, the temperature perturbation at all angles was greater than in the narrow irregularity case at all sampled angles, due to the broadening of the window with increasing duct width (as shown in Figure 3). The greatest modifications to both temperature and density occurred for the wave directed between the Spitze angle and vertical (shown in panels 12.2b and 12.2d). Here the maximum steady state temperature achieved was more than 10% greater than the maximum temperature developed in the unperturbed 1-D slope scenario shown in Figure 9. At all sampled angles, the site of greatest density perturbation can be seen to occur within the duct, having the effect of deepening the depletion with time around the point at which the pump wave is reflected.

When multiple periodically distributed irregularities were included, the growth of plasma perturbations was very different. As expected from section 3, the smaller-scale (0.08 km) irregularities effectively attenuate the



**Figure 11.** Variation in electron density perturbation and electron temperature perturbation with pump wave inclination angle when horizontal density slopes of the form (9) with  $L_x = 20$  km and  $L_x = 50$  km were included in the background density profile. Uppermost panels (11.0a-b) show the background conditions for each  $L_x$  case. Lowermost panels (11.6a-b) show the variation of minimum density perturbation and maximum temperature perturbation recorded in the simulation with time for  $L_x = 20$  km (solid line) and  $L_x = 50$  km (dashed line). The inclination angle responsible for the greatest temperature enhancement shifts from the Spitze position toward the magnetic zenith direction with decreasing  $L_x$ , consistent with the modification of the *Z*-mode window simulated in section 3 (shown in Figure 2).



**Figure 12.** Variation in electron density perturbation and electron temperature perturbation with pump wave inclination angle when single density-depleted field-aligned irregularities of the form (10) with  $L_{width} = 0.08$  km and  $L_{width} = 0.8$  km were included in the background density profile. Uppermost panels (12.0a-b) show the background conditions for each  $L_{width}$  case. Lowermost panels (12.6a-b) show the variation of minimum density perturbation and maximum temperature perturbation recorded in the simulation with time for  $L_{width} = 0.08$  km (solid line) and  $L_{width} = 0.8$  km (dashed line).



**Figure 13.** Variation in electron density perturbation and electron temperature perturbation with pump wave inclination angle when periodic density-depleted field-aligned irregularities of the form (11) with  $L_{width} = 0.08$  km and  $L_{width} = 0.8$  km were included in the background density profile. Uppermost panels (13.0a-b) show the background conditions for each  $L_{width}$  case. Lowermost panels (13.6a-b) show the variation of minimum density perturbation and maximum temperature perturbation recorded in the simulation with time for  $L_{width} = 0.08$  km (solid line) and  $L_{width} = 0.8$  km (dashed line).

propagation of the *Z*-mode wave beyond the interaction region. As such, this narrow irregularity case behaves much like the unperturbed 1-D density slope scenario with the effects due the *Z*-mode suppressed. Accordingly, the maximum temperature enhancement was found to occur for the vertically directed pump wave (shown in panel 13.1c) and was found to decrease as the wave inclination angle was increased toward the field-aligned direction. For the larger-width (0.8 km) irregularities, as in the single-irregularity case, the greatest modification to the plasma was found to occur for the pump wave inclined at 2.9° (panel 13.2d). The greatest perturbations can be seen to be localized at certain points within each density depletion. These "hot spots" correspond to the pump wave reflection point within each irregularity case was greater than in the small-irregularity case for all pump wave angles aside from vertical (for which the steady state temperatures were approximately equal). By contrast, the magnitude of density depletions was greater in the case of narrow irregularities by a factor of 2 or greater at all sampled angles, corresponding to a more rapid growth of the depletions with time in this scenario.

In section 3 it was shown that addition of a 2-D variation to the electron density profile strongly affected both the process of conversion from the *O*-mode pump wave to *Z*-mode and the subsequent propagation of any *Z*-mode waves. The results presented in this section have demonstrated that this modification of the *Z*-mode window has a knock-on impact on the development of artificially-induced thermal plasma perturbations. Changing the shape or position of *Z*-mode window via the inclusion of a 2-D density inhomogeneity was found to lead to a corresponding variation in the dependence of electron temperature enhancement on pump wave inclination angle. This mechanism could explain why the greatest magnitudes of plasma heating have often been found to occur for non-Spitze directed waves during observations of the magnetic zenith effect.

#### 6. Summary and Conclusions

lonospheric heating experiments have observed that the magnitude of artificial heating-induced plasma perturbations depend strongly on the inclination angle of the pump beam, with a greater modification to the plasma observed when the heating beam is directed close to or along the magnetic zenith direction. *Honary et al.* [2011] proposed that this magnetic zenith effect is due to the *O*-mode to *Z*-mode conversion process that can occur in the *F* region for a narrow range of pump wave inclination angles. This conversion process has previously been investigated theoretically for the case of a 1-D variation in electron density; however, the case of a 2-D variation had not been discussed before, and in this paper was investigated numerically for the first time. A recently developed full-wave FDTD code [*Cannon and Honary*, 2015] was used to numerically explore the effect of an *O*-mode polarized EM pump wave on the plasma around the ionosphere critical interaction region. In particular, the behavior of the *O*- to *Z*-mode conversion process and magnetic zenith effect was investigated for a variety of density profiles. These simulations show that the presence of ionospheric plasma density inhomogeneities can explain many of the features of the observed magnetic zenith effect through modification of the angular window for *O*-mode to *Z*-mode conversion.

The fraction of the simulated pump wave field penetrating beyond the *O*-mode reflection height in the form of a *Z*-mode was measured for a range of launch angles and used to determine the angular shape of the window for which *O*- to *Z*-mode conversion was favorable. For the case of a 1-D vertical electron density gradient the simulation results (as shown in Figure 1) demonstrated good agreement with the predictions of *Mjølhus* [1984], with the conversion window centered on the Spitze angle (8). From this, one would expect that the magnetic zenith effect would manifest most strongly for pump waves directed at the Spitze angle; however, it has been reported that the maximum electron temperature enhancement at European Incoherent Scatter (EISCAT) can occur for inclination angles somewhere between the Spitze and field-aligned directions [for example, *Isham et al.*, 2005]. To explain this, *Honary et al.* [2011] proposed that the shape of the *Z*-mode window can be modified by the presence of horizontal inhomogeneities in the electron density, such as the large-scale depleted regions or field-aligned irregularities that may be excited during the course of an ionospheric heating experiment. This problem cannot be fully treated theoretically using the techniques of geometric optics; however, the numerical simulation techniques used here allowed 2-D plasma profiles to be fully taken into account.

A varied hierarchy of electron density structures has been observed during heating experiments due to the formation of plasma irregularities of different scales. The FDTD code was used to investigate how the inclusion

of 2-D density inhomogeneities could influence the shape of the Z-mode window. Here three distinct cases were investigated: large-scale horizontal density slopes, medium-scale single density-depleted ducts, and smaller-scale periodic field-aligned irregularities. The results demonstrate that the window is highly sensitive to the form of the density profile.

Adding a linear horizontal density gradient to represent a large-scale density-depleted patch was found to shift the center of the *Z*-mode window away from the Spitze direction, as shown in Figure 2, either toward the magnetic zenith direction or the vertical direction depending on the direction of the horizontal gradient. A steeper slope resulted in a greater shift of the window. Shifting the window all the way to the field-aligned direction required the presence of a horizontal density gradient of similar size to the vertical gradient (a similar gradient in the opposing direction would shift the window all the way to the vertical direction).

Inclusion of a single field-aligned density-depleted irregularity to represent a duct-like structure led to a slight shifting of the window toward the vertical direction for irregularities with widths < 0.1 km. Increasing the irregularity width in the range 0.1 km–2 km was found to increasingly broaden the *Z*-mode window, with the angular width of the window for a duct of 1 km width almost three times that of the 1-D slope case (see *Z*-mode window curves in Figure 3). For duct widths in this range, it could be seen that a narrow *Z*-mode beam was produced for every sampled pump wave direction; however, the location with respect to the irregularity geometry at which conversion took place varied between inclination angles, as can be seen in Figure 4. For irregularity widths beyond 2 km, this broadening effect was found to decrease, and with increasing width the shape of the *Z*-mode window tended toward the unperturbed form measured for the case of the 1-D vertical density gradient.

Inclusion of multiple duct-like structures in the form of periodic field-aligned density-depleted irregularities had a strong effect on the propagation of the converted *Z*-mode (as shown in Figure 5), with transmitted waves multiply scattered by adjacent irregularities, making it difficult to measure the window shape. In general the effect of periodic irregularities with spatial scales comparable to or smaller than the pump wave wavelength had the effect of attenuating the transmitted *Z*-mode amplitude at all angles while maintaining the angular window shape measured for the unperturbed case. As the irregularity scale size was increased, multiple scattering of the *Z*-mode wave between irregularities allowed significant *E* field amplitude to propagate beyond the *O*-mode reflection height for all sampled pump wave angles.

A novel feature of the FDTD code used here is that it allowed the time-explicit evolution of thermal perturbations to the plasma density and temperature to be simulated on the timescale of the EM wave. This feature was used to investigate how the O-mode to Z-mode conversion process may lead to an angularly-dependent modification of the plasma medium. In the F region of the ionosphere, the Z-mode wave is reflected at a higher altitude than the O-mode wave at a point corresponding to X = Y + 1. If reflected, the Z-mode wave is able to propagate back toward the interaction region where it can excite a resonance that is inaccessible to fast X-mode waves and results in the efficient conversion of the EM wave to upper hybrid plasma waves. As the Z-mode wave approaches the resonance height, its **k** vector approaches infinity and thus its group velocity decreases, leading to a sharp increase in electric field amplitude. The swelling in amplitude leads to the growth of local electron temperature, thus an increase in the pressure and causing the expansion of the local plasma. To investigate the impact of this process, simulations in which the Z-mode wave was allowed to reflect back toward the interaction height were compared with simulations in which the Z-mode was absorbed before reflection using an artificial PML, both for the case of a 1-D linear density variation (see Figures 7 and 8). For a Spitze-directed wave, the maximum E field amplitude measured when the Z-mode wave was allowed to reflect was more than 10 times greater than in the reflection-suppressed case. Correspondingly, the magnitude of thermal perturbations to the plasma were found to be greater in the reflection-allowed scenario, with electron temperature enhanced by factor of 2 and electron density depleted by a factor of 4 compared to the Z-mode reflection-suppressed results.

In the reflection-allowed simulations, the angular dependence in *O*- to *Z*-mode conversion translated to an angular dependence in plasma modification, demonstrated by the results presented in Figure 9. Pump waves with inclination angles more favorable for *Z*-mode conversion were found to achieve a greater magnitude of plasma perturbation, with the magnitude of steady state temperature enhancement increasing as the wave inclination angle was increased from 2780 K at vertical (0.0°) to 2940 K at the Spitze (5.2°), before falling off as the inclination angle was increased further. By contrast, without *Z*-mode reflection the greatest electron temperature enhancement was seen for the vertically directed (2750 K) wave and the level of

enhancement reduced as the inclination angle approached the Spitze (1440 K). In addition, excitement of small-scale field-aligned density irregularities was found to occur most favorably for pump beams directed along the Spitze direction, as shown in Figure 10. These results support the idea that resonant excitation of large-amplitude electric fields via the reflected *Z*-mode contributes to the enhanced plasma perturbations and pump wave inclination angle dependence associated with the magnetic zenith effect.

The addition of a horizontal component to the linear density profile was shown to modify the *Z*-mode window by shifting its center away from the Spitze direction, and here it was shown that this also had a significant impact on the development of thermal plasma perturbations. Simulations were run using the density profile given by equation (9) for horizontal slope scale sizes of 50 km and 20 km (see results presented in Figure 11). Maximum temperature enhancement found to occur for pump wave inclination angles 8.6° and 11.5°, respectively, corresponding closely to the center of the *Z*-mode window for each scenario (8.3° and 12.4°, respectively). Thus, by shifting the Z-mode window via a horizontal density inhomogeneity, the pump wave inclination angle resulting in the greatest heating was found to vary. This offers a potential mechanism to explain how maximal magnetic zenith effect heating has been observed to occur for non-Spitze-directed waves.

Simulations were run to compare the influence of duct-like irregularities of large (800 m) and smaller (80 m; close to the pump wavelength) scale sizes on the growth of thermal plasma perturbations (see results presented in Figure 12). As would be expected from the simulated Z-mode window curves shown in Figure 3, the narrow irregularity case behaved much like the unperturbed case with maximum temperature enhancement occurring at the Spitze angle. For the larger irregularity, greater stationary-state temperatures were achieved at all angles due to the broadening of the Z-mode window for irregularities of this type. Both the maximum temperature enhancement and the greatest depletion of density occurred for the wave directed between the Spitze angle and vertical. For a pump wave at this angle, the stationary temperature enhancement was 10% greater than in the unperturbed 1-D slope case shown in Figure 9.

For periodic irregularities of the same scale sizes, the growth of plasma perturbations was very different (as shown in Figure 13). The narrow irregularities ( $L_{width}$ =80~m) effectively attenuated conversion of the *O*-mode wave to the *Z* mode, leading to the greatest plasma perturbations occurring when the pump wave was directed vertically. For the larger irregularities ( $L_{width}$ =800~m), as in the single-irregularity case, the greatest modification to the plasma occurred for a pump wave inclination angle of 2.9°. The greatest perturbations can be seen to occur in highly-localized "hot spots" within the periodic density depletions. The magnitude of temperature in the large-irregularity case was greater than in the small-irregularity case for all pump wave angles aside from vertical (for which the steady state temperatures were approximately equal). By contrast, the magnitude of density depletions were greater in the case of narrow irregularities by a factor of 2 or more at all angles, corresponding to a more rapid growth of the depletions in this scenario.

The simulations presented in this paper demonstrate that the existence of 2-D electron density inhomogeneities around the *O*-mode reflection height can have a significant impact on the development of perturbation to the plasma during a heating experiment. Large-scale linear density gradients, medium-scale duct-like density depletions and small-scale field-aligned irregularities were all found to affect the *O*-mode to *Z*-mode conversion process and consequently modify the position of the *Z*-mode window. This was shown to have a knock-on effect on the growth of thermal plasma perturbations due to the interaction of the heating wave and offers a potential mechanism behind several of the observed features of the magnetic zenith effect.

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