

Demonstration of new MeV-scale capabilities in large neutrino LArTPCs using ambient radiogenic and cosmogenic activity in MicroBooNE

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Large neutrino liquid argon time projection chamber (LArTPC) experiments can broaden their physics reach by reconstructing and interpreting MeV-scale energy depositions, or blips, present in their data. We demonstrate new calorimetric and particle discrimination capabilities at the MeV energy scale using reconstructed blips in data from the MicroBooNE LArTPC at Fermilab. We observe a concentration of low energy (<3 MeV) blips around fiberglass mechanical support struts along the TPC edges with energy spectrum features consistent with the Compton edge of 2.614 MeV ^{208}Tl decay γ rays. These features are used to verify proper calibration of electron energy scales in MicroBooNE's data to few percent precision and to measure the specific activity of ^{208}Tl in the fiberglass composing these struts, $(11.7 \pm 0.2$ (stat) ± 3.1 (syst)) Bq/kg. Cosmogenically-produced blips above 3 MeV in reconstructed energy are used to showcase the ability of large LArTPCs to distinguish between low-energy proton and electron energy depositions. An enriched sample of low-energy protons selected using this new particle discrimination technique is found to be smaller in data than in dedicated CORSIKA cosmic ray simulations, suggesting either incorrect CORSIKA modeling of incident cosmic fluxes or particle transport modeling issues in Geant4.

I. INTRODUCTION

Liquid argon time projection chambers (LArTPCs) are unique among existing deployed neutrino detector technologies in combining millimeter-scale position resolution with sub-MeV energy detection thresholds. In a LArTPC, MeV-scale and sub-MeV charged particles generate ionization electron clouds with μm to cm spatial extents, which efficiently drift through the liquid argon environment in a uniform electric field towards planes of closely-spaced conducting wires. After processing the wire signals, these electron clouds appear as isolated compact features, or ‘blips’, in a LArTPC event, contrasting the extended shower-like or track-like features produced by higher-energy charged particles. The position resolution of neutrino LArTPCs is defined by the spacing of these sense wires (3 to 5 mm in current LArTPCs [1–4]), while detection thresholds are defined by the noise levels achieved by sense wire readout electronics (as low as $300 e^-$ in equivalent noise charge per ADC sample in the MicroBooNE LArTPC, for example [5]).

The combination of precise position resolution and low energy thresholds can be exploited to expand the scope of particle physics that can be performed by GeV-scale accelerator neutrino LArTPC experiments. Existing neutrino LArTPCs can use these capabilities to identify neutrino-argon (ν -Ar) interactions with important yet under-studied attributes. For example, ν -Ar final-

state neutrons, visible primarily as proton- or electron-induced low-energy activity [6–8], will play an important role in defining energy reconstruction biases for neutrinos and antineutrinos in future leptonic CP-violation measurements [9–11]. Low-activity ν -Ar vertices, such as those generated by low momentum-transfer neutral-current scatters, can be overlooked by standard reconstruction tools despite their potential value in probing the nature of the long-standing MiniBooNE neutrino experiment anomaly [12–15]. More generally, the definition of ‘zero-proton’ ($0p$) and ‘one or more protons’ (Np) neutrino final-state topologies in recent LArTPC studies [13–20] is based on a LArTPC’s effective threshold for proton reconstruction (20 or 35 MeV on kinetic energy, depending on the analysis). However, low-energy reconstruction tools may enable reliable reconstruction of proton activity below these thresholds, increasing precision in knowledge of the hadronic content of final-state ν -Ar interactions. MeV-scale activity can also enable searches for low-energy muon or meson decay-at-rest neutrino interactions [21–23] or new physics beyond the Standard Model (BSM) [24–26]. In the coming decade, application of low-energy LArTPC capabilities in the Deep Underground Neutrino Experiment (DUNE) far detector will be crucial to calibrating detector response [11] and to achieving sensitivity to supernova and solar neutrino ν -Ar interactions [27–29]. A range of potential physics benefits are surveyed in Refs. [24, 26].

The exploration of the MeV-scale physics capabilities of neutrino LArTPCs is a relatively recent effort. Following early studies of ~ 5 -50 MeV cosmic muon decay electrons in the ICARUS and MicroBooNE detectors [30, 31],

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data from the ArgoNeuT LArTPC was used to detect isolated blips produced by final-state photons and neutrons from ν -Ar interactions [6], and to search for similar features from hypothetical beam-produced millicharged particles [32]. The MeV-scale capabilities of the MicroBooNE detector were first studied in [33–35]. The MicroBooNE LArTPC was used to measure the properties of decay radiation from ^{222}Rn progeny in its liquid argon bulk [36], including ^{214}Po decay α -particles generating ~ 100 -200 keV worth of electron-equivalent ionization [37]. That analysis also used ^{214}Bi decay electrons to demonstrate other MeV-scale calorimetric capabilities, showing data-simulation energy scale agreement at $<5\%$ precision and a modeled energy resolution better than 10% above 1 MeV. A first study of blip-based particle discrimination was recently performed by LArIAT using purified samples of stopping μ^- and π^- [38]. With one exception [39], MeV-scale measurements to date have focused on low-energy capabilities of neutrino LArTPC ionization charge collection systems, bypassing consideration of associated scintillation light signals.

In this study, we extend the range of demonstrated neutrino LArTPC capabilities and tools available at the MeV scale using ambient radiogenic and cosmogenic blips reconstructed in MicroBooNE data. Leveraging an intrinsic detector source of ^{208}Tl decay γ rays, we perform a percent-level precision energy scale calibration for reconstructed blips, providing a template for blip-based energy scale calibrations in DUNE and other future LArTPC experiments. A byproduct of this calibration is a measurement of the specific activity of ^{208}Tl in MicroBooNE’s fiberglass TPC support struts that highlights the importance of radio-purity screening in DUNE and other next-generation LArTPCs.

Next, we develop a new particle identification (PID) metric for blips based on size-to-energy comparisons, capable of distinguishing low-energy charged particles of differing stopping powers in argon, such as electrons and protons. We demonstrate the value of this PID capability by identifying, for the first time in a neutrino LArTPC, a sub-dominant population of isolated cosmogenic protons in MicroBooNE. Comparisons to CORSIKA cosmogenic simulations reveal an excess of measured protons in CORSIKA, indicating potential cosmogenic flux or cosmic ray particle transport issues in the CORSIKA or Geant4 toolkits used by MicroBooNE.

Section II summarizes the MicroBooNE detector, while Sec. III describes the analysis tools used to process MicroBooNE data and reconstruct MeV-scale blips. Datasets and Monte Carlo (MC) simulations used in this analysis are described in Sec. IV. After a general description of all selected blips in Sec. V, Sec. VI provides a detailed analysis of the ^{208}Tl -produced 2.614 MeV γ -ray Compton edge. Section VII describes the developed PID metric and cosmogenic proton analysis, with a summary of study outcomes provided in Sec. VIII.

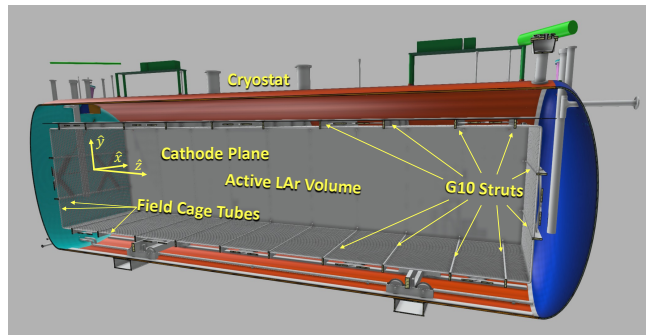


FIG. 1. A labeled isometric model of MicroBooNE’s cryostat and TPC interior. The model shows the TPC viewed from its anode plane wire side. The coordinate system ($\hat{x}, \hat{y}, \hat{z}$) is indicated, with the neutrino beam oriented along the $+z$ direction.

II. THE MICROBOONE DETECTOR

MicroBooNE was a single-phase LArTPC detector located in the Booster Neutrino Beamline (BNB) at Fermi National Accelerator Laboratory that operated from 2015 to 2021. The TPC, its cryostat, and all supporting detector infrastructure were located in the Liquid Argon Test Facility, an on-surface building providing minimal overburden shielding the LArTPC from cosmic rays. While details of the MicroBooNE detector and support systems are presented in Ref. [2], we will briefly overview MicroBooNE design elements relevant to this work.

The primary component of the MicroBooNE detector was a 2.56 m width \times 2.33 m height \times 10.37 m length TPC containing 85 metric tons of purified LAr. As illustrated in Fig. 1, the TPC formed a rectangular prism, with a uniform 274 V/cm electric field between the two vertically-oriented long faces of the TPC. This electric field, oriented in the x direction (indicated in Fig. 1) in MicroBooNE’s coordinate system, was generated by a stainless steel cathode plane charged to -70 kV at $x = 256$ cm and three anode planes of conducting read-out wires at $x \approx 0$ cm, spaced 3 mm apart, with a 3 mm separation gap between planes. A series of 64 stainless steel field cage tubes of 2.54 cm diameter surrounded the active LAr volume, running perpendicular to the electric field along the walls of the TPC. A series of voltage divider circuits were used to step the field down linearly along the field cage from the cathode to the anode to ensure a uniform electric field within the active volume.

The cathode and anode surfaces of the TPC were mechanically supported by metal structures, while field cage tubes were stabilized using struts of 2.5 cm thick G10, an electrically insulating epoxied fiberglass laminate [40]. The G10 struts also provide the only mechanical connection between anode and cathode structures, enforcing a fixed parallel orientation between the two planes. Ten (two) elongated 15.4 kg struts were present on each TPC top and bottom (front and back) surface, with each

strut’s long axis running the entire x length of the TPC active volume. As shown in Fig. 2, field cage tubes penetrated machined holes centered 2.5 cm from one side of the strut’s 14 cm height in y , leaving 1.2 cm of G10 material extending beyond the field cage tubes. Thus, G10 is the only non-LAr material present in the active volume of MicroBooNE’s TPC.

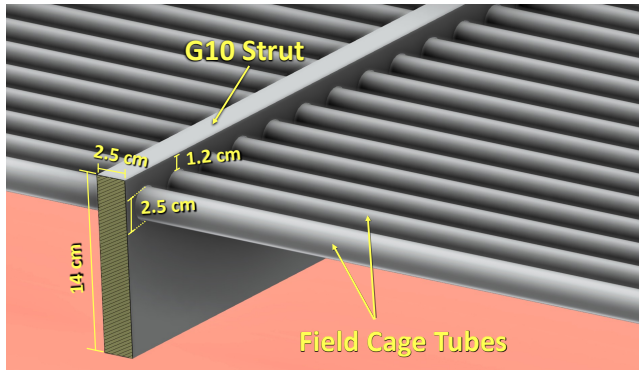


FIG. 2. A close-up labeled isometric drawing of the interface between the bottom TPC field cage and support strut elements. The pictured G10 strut provides mechanical support to the TPC while maintaining electrical insulation of all field cage elements.

Charged particle interactions within the active TPC volume generate ionized electrons which drift in the uniform electric field at a speed of $1.1 \text{ mm}/\mu\text{s}$ towards anode plane readout wires. Electrical signals generated on the conducting wires in response to nearby ionization charge are sensed, processed, and recorded by readout electronics with a sampling period of 500 ns as described in Refs. [5, 41, 42]. By collecting 6400 ADC samples (3.2 ms) per wire per detector trigger, all ionization charge present inside the TPC at the time of triggering can be recorded regardless of drift distance in x .

The 2,400 wires in each of the two planes closest to the active TPC volume are oriented at $\pm 60^\circ$ with respect to the vertical direction. These ‘induction’ planes are voltage-biased at -110 V and 0 V to ensure maximal electron transparency, causing charge to cross and induce bipolar signals on this plane. The 3,456 vertically-oriented wires in the third ‘collection’ plane are held at +220 V, and record unipolar ADC waveform pulses as ionization charge stops at this plane. Raw wire signals are further processed and noise-filtered to obtain charge waveforms upon which higher-level reconstruction and analysis is performed [5, 41, 42]. MicroBooNE’s charge collection system is sensitive to LArTPC energy depositions well below 1 MeV due to: (a) the achieved residual equivalent noise charge of $300 e^-$ and $400 e^-$ on the longest collection and induction wires, respectively, considered alongside the mean argon ionization energy of 23.6 eV per electron; (b) the electron-ion recombination survival fraction in MicroBooNE’s electric field of 60% for minimally-ionizing particles; and (c) MicroBooNE’s

long drift electron lifetime.

Scintillation light is also generated in a LArTPC over comparatively short (a few μs) timescales following ionization and excitation of the LAr by charged particles. While the MicroBooNE detector uses a photomultiplier-based light collection system to take advantage of scintillation signatures from higher-energy particle interactions, the magnitude of light collection in most detector regions is too low to enable sensitivity to MeV-scale signatures. For this reason, light is not used for the studies described in this paper. This results in complete ambiguity in the x position of MeV-scale energy depositions within the MicroBooNE TPC, since the time of energy deposition relative to the detector trigger time, and thus the drift time, cannot be known.

III. RECONSTRUCTING AND SELECTING MEV-SCALE ACTIVITY

MeV-scale charged particles in LAr leave ionization trails with lengths on the order of millimeters. For large LArTPCs with wire spacings on the order of MicroBooNE’s (3 mm), these ionization clouds will be sensed by only a few wires at most on each anode plane. Existing software targeted at reconstructing high-energy charged particles depositing 10s to 100s of MeV cannot properly identify or reconstruct the attributes of these low-energy particles. In this section, we will summarize how isolated MeV-scale energy depositions, or blips, are reconstructed in MicroBooNE, describe some important physics attributes of reconstructed blip objects, and outline the selection used to define a signal blip sample.

A. Blip Reconstruction

The reconstruction of blips in MicroBooNE is carried out using a LArSoft [43] custom-built toolkit called **BlipReco**, validated in previous MicroBooNE results and described in detail in Ref. [37]. Reconstructed blips are formed from TPC wire ‘hits’ which contain information about pulses on a wire’s filtered and deconvolved charge waveform. In this analysis, raw waveform deconvolution, noise filtering, and region-of-interest (ROI) identification is performed with the **WireCell** toolkit [41, 42]. Hit reconstruction is performed by the **GausHit** algorithm [44], which searches for threshold-crossing points and determines hit times, widths and amplitudes by applying Gaussian fits to these regions. All reconstructed hits in an event are then passed to **BlipReco**, which forms hit ‘clusters’ by grouping hits together on each plane based on their relative proximity to one another in wire and readout time. The proximity requirement for clustering scales with the width (in time) of each individual hit. Only hits from same or adjacent wires are grouped together. This clustering proceeds iteratively until no new hits on the plane can be added to the cluster. If an

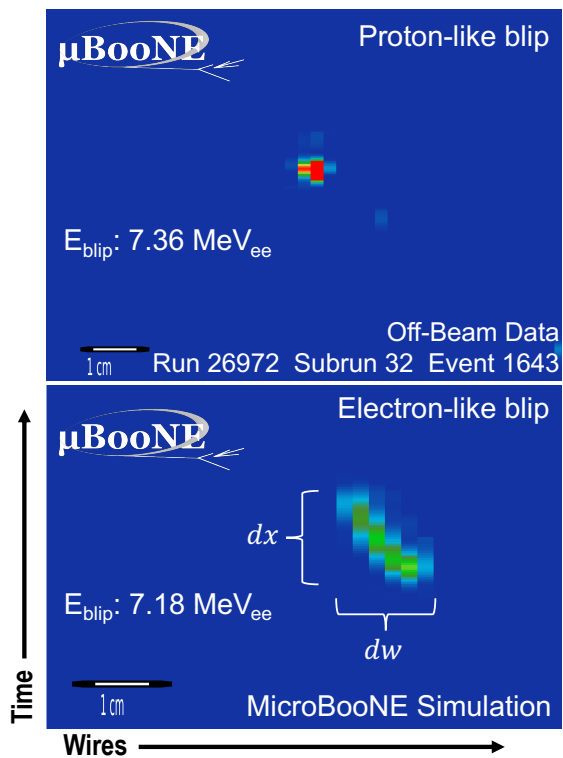


FIG. 3. A proton-like blip (top) from cosmic data and a low-energy electron blip (bottom) from simulation, as they appear in the MicroBooNE TPC event displays. Vertical columns represent individual wires on a wire plane, with the color scale indicating the relative charge collected at each ADC sample. This particular feature results in a cluster of hits spanning 4 (top) and 6 (bottom) wires respectively. The metrics dx and dw , described in the text, correspond to the measured lengths of the cluster along the drift direction and projected along the axis perpendicular to wires.

isolated hit is not accompanied by adjacent hits on its plane, it is also treated as a cluster.

Since MeV-scale depositions span lengths approaching the inherent limiting spatial resolution of the TPC, reconstructing their true physical extent is difficult. For this analysis, the length of the cluster projected onto the axis perpendicular to the orientation of the wires on the plane, termed dw , is determined by multiplying the number of wires by the wire spacing of 3 mm. Similarly, the cluster's extent along the drift direction, dx , is calculated using the overall time span of the cluster multiplied by the electron drift velocity in LAr. The full time span is calculated as $dt = [(t_1 + \text{RMS}_1) - (t_0 - \text{RMS}_0)]$, where t_0 (t_1) and RMS_0 (RMS_1) denote the hit time and Gaussian-fitted width of the earliest (latest) hit in the cluster. These two metrics dx and dw , depicted visually in Fig. 3, allow us to approximate the physical geometric extent of the contiguous trail of ionization that formed the cluster.

Clusters generated on different planes by the same drifting electron cloud are then matched between planes

to form three-dimensional blip objects. The plane-matching procedure incorporates information about how closely the charge-weighted mean time of the clusters coincide, the overlap of their time spans, consistency in cluster charges, and physical wire crossing locations. While a blip is allowed to contain matched clusters from two or three planes, it must contain a collection plane cluster since this is the plane used for calorimetry.

Plane-matched blips are assigned location, energy, and size metrics based on the attributes of its clusters. The total size (ds) or extent of the blip in 3D space is determined by adding in quadrature dx from the collection plane cluster together with the largest dw from any of the matched planes, which represents an estimate of the blip's projected length onto the anode (y - z) plane. Three-dimensional blip location, determined from the wire crossing point and the projected distance along the drift axis, is given in the same (x, y, z) coordinates assigned to other reconstructed objects like tracks and showers. Length and location metrics in the y - z coordinate plane have resolutions limited by the 3 mm wire spacing.

Reconstructed blip energy, E_{blip} , is defined based on the number of electrons Q reconstructed on the collection plane, the mean argon ionization energy $W_{\text{ion}} = 23.6$ eV [45], and a linear charge-energy conversion based on recombination levels expected for electrons in this energy regime,

$$E_{\text{blip}}[\text{MeV}_{\text{ee}}] = \frac{Q}{0.584} \times W_{\text{ion}}. \quad (1)$$

As described in detail in Ref. [37], this linear charge-energy conversion approximation results in percent-level energy scale biases in E_{blip} with respect to deposited energy above 1.5 MeV, and 10%-level biases below 1 MeV. Blip energy resolution, characterized with MC simulations of uniformly distributed low-energy electrons, is estimated to be less than 10% above 1 MeV.

The algorithm matches truth-level MC information to each reconstructed blip, including the identity (PDG number) of the charged particle that produced a majority of a blip's ionization, the parents of that particle, the producing processes of that particle, its parents, and more.

B. Signal Blip Selection and Sorting

For this analysis, we apply minimal signal selection cuts to the full population of reconstructed blips in MicroBooNE. Blips formed from hits directly adjacent to cosmic muon tracks, such as those induced by low-energy δ rays, are identified and excluded from consideration. Blips reconstructed on wires directly adjacent to nonfunctional wires are also rejected to reduce inclusion of track fragments or blips with biased reconstructed physics quantities in the signal dataset.

The combined efficiency of blip reconstruction and signal blip selection was determined using a low-energy electron MC dataset with kinetic energies ranging from 0-12 MeV, similar to that described in Ref. [37]. Efficiency is strongly driven at low energies by chosen settings for the `WireCell` and `GausHit` modules used to generate reconstructed hits. As in Ref [37], this study uses specialized datasets processed using lowered settings in these modules that achieved roughly 50% blip reconstruction efficiency at 0.5 MeV. The efficiency approaches zero below a detected charge of 6,000 electrons (e^-), or an effective energy deposition of 250 keV, and reaches a maximum plateau above 80% around 35,000 e^- (1.5 MeV deposited energy). The upper limit of 80% is a consequence of nonfunctional TPC wires limiting the sensitive fiducial volume; when this effect is factored out, the maximal efficiency is $>95\%$.

In contrast to prior MicroBooNE blip studies or analyses in other large MeV-scale particle detectors, we refrain from applying selection cuts based on spatial proximity of blips to one another (*i.e.* single-site versus multi-site event topologies in $0\nu\beta\beta$ or dark matter experiments [46–50]) or performing reconstruction of larger multi-site neutrino LArTPC event topologies (*i.e.* multiple-scattering γ rays or neutron capture γ -cascades [24, 51]). To perform background subtractions or data-driven validations for some aspects of this study, we will segregate and/or compare signal blip datasets based on their location in the TPC, their proximity to cosmic ray tracks, and their energies.

IV. DATASETS AND MONTE CARLO SIMULATIONS

To perform our study, we use data collected over a 46-day period in June-July of 2018. For this dataset, readout of the detector was triggered randomly with a function generator during periods where the BNB was not delivering neutrinos to MicroBooNE (referred to in other MicroBooNE publications as ‘unbiased beam-external’ data). This is the same dataset used to measure the steady-state presence of radon daughters in MicroBooNE’s purified LAr bulk [37]. As mentioned above, this specially processed dataset achieves lower energy thresholds than other MicroBooNE production datasets due to altered signal processing and hit-finding settings. The dataset consists of 653,367 total triggered event readouts, equivalent to a total cumulative live time of roughly 35 minutes.

Portions of the analysis involving study of radiogenic energy depositions requires production of MC datasets in which monoenergetic γ rays were generated uniformly in the volume of a subset of the TPC’s G10 struts. Using `LArSoft`, three γ rays per strut were generated for each event and subsequently propagated through the detector using `Geant4`. This level of γ -ray activity per event, while not realistic, offered faster MC processing

times with negligible spatial overlap of blips from unrelated physics processes. Simulated LArTPC detector response attributes match those of other MicroBooNE physics analyses [52], with the exception of ionization charge diffusion, which was modeled using diffusion parameters reported in Ref. [53]; for further discussion of LArTPC response modeling, see Sec. VI 1. To ensure the inclusion of realistic wire noise features in addition to simulated MeV-scale energy depositions, MicroBooNE-standard “event overlay” procedures were used, in which the generated wire signatures of the simulated photons mentioned above are added to waveforms from existing unbiased beam-external data events in order to incorporate data-realistic noise in the wire signals. To guarantee only simulated γ -produced blips were studied in MC datasets, reconstructed blips’ truth-level variables were used to discard all real blips present in overlaid events. A total of 109,268 overlaid monoenergetic γ -ray event displays are generated using these methods.

Studies of cosmogenically-produced MeV-scale energy depositions were aided by production of MC datasets containing simulated cosmic rays. The use of the CORSIKA cosmic ray generator to simulate cosmogenic activity in MicroBooNE is discussed in detail in Ref. [54]. For our study, we sample from the contents of CORSIKA showers generated for this previous study. Since that study showed consistency between data and simulation for observed muon rates when CORSIKA simulated only proton shower generation in the upper atmosphere, we maintain this CORSIKA setting for our study. The shower content is generated on a surface 18 m above the cryostat that extends 10 m beyond the LArTPC cryostat in the x and z directions. Initial and secondary particles are propagated from this surface through the detector by `Geant4`. Cosmic particles are generated randomly in time ± 2.8 ms around the readout trigger to account for ionization produced prior to the trigger that can still be collected by the drift readout. As with monoenergetic γ -ray MC samples described above, CORSIKA MC samples are also overlaid with beam-off external data, and truth-level variables are used to discard consideration of real blips and tracks present in overlaid events.

V. GENERAL DESCRIPTION OF MEV-SCALE SIGNALS

A total of 69,427,354 candidate signal blips were reconstructed, corresponding to an average of 106 blips per event, or roughly 0.39 blips per kg of LAr inside the TPC per second of active detector readout. The position distribution of signal blips within the TPC in the y - z plane, shown in Fig. 4, exhibits a few visible artifacts due to the presence of non-functional or noisy wires, resulting in diagonal and vertical regions containing more or less blip activity than usual. Distinct regions of elevated blip activity, termed “hot spots”, are visible along the edges of the detector. These closely correspond to

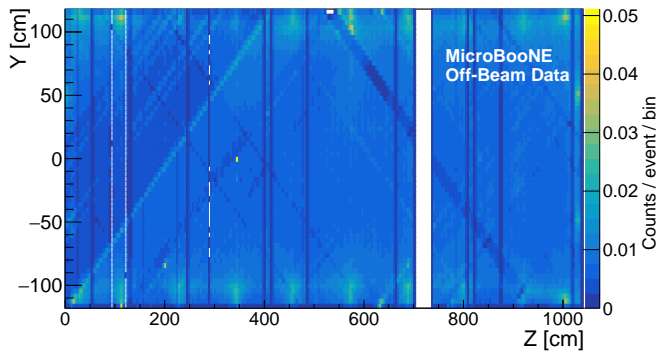


FIG. 4. Distribution of reconstructed blip positions per event display in the TPC volume projected onto the y - z plane. As described in the text, y and z coordinates correspond to the vertical and neutrino beam directions, respectively.

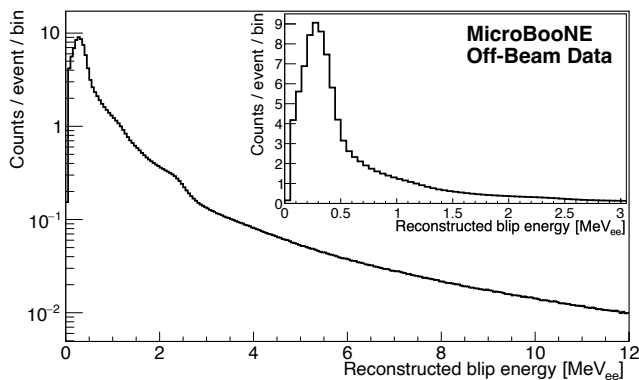


FIG. 5. Spectrum of reconstructed energies, E_{blip} , for all signal blips in MicroBooNE cosmic data. The inset shows the same spectrum from 0-3 MeV_{ee} plotted with a linear scale.

the locations of G10 field cage support struts described in Sec. II and pictured in Fig. 1. This observation supports the hypothesis that the fiberglass laminate composing the strut contains radioactive impurities. Further insight into ambient blip activity and the observed hot spots is provided by their reconstructed energy (E_{blip}) spectrum given in Fig. 5. Highest counts appear at the lowest accessible E_{blip} in MicroBooNE, with a kink in the spectrum occurring near the 0.58 MeV β -decay endpoint of ^{39}Ar , a naturally-occurring isotope present in LAr at (~ 1 Bq/kg) [55, 56]. The region of the E_{blip} spectrum below 0.58 MeV_{ee} , containing on average 65 blips per event (roughly 0.24 blips per kg of LAr inside the TPC per second of active detector readout), is heavily sculpted by changes in blip reconstruction efficiency described in Sec. IV, making it difficult to perform other studies in the absence of precise threshold characterization, modeling, and calibration. For this reason, we leave the lower-energy blip population as a topic for future study.

Another obvious energy spectrum feature is present between 2 and 3 MeV_{ee} , a regime well above the blip reconstruction threshold. Roughly 4.8 signal blips per

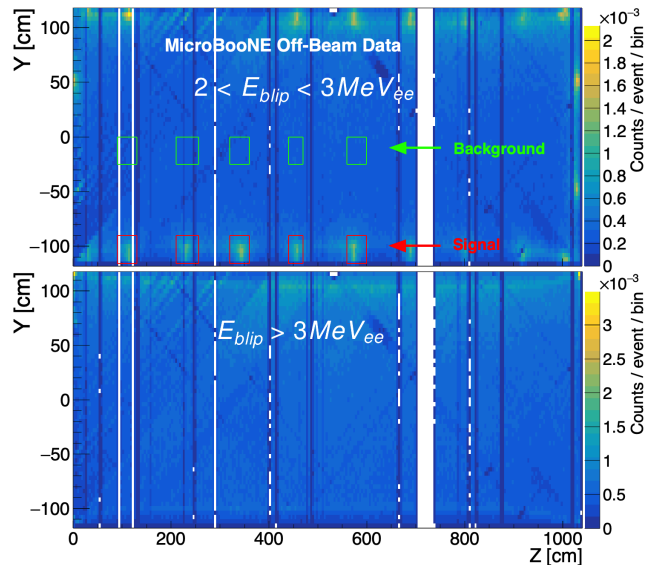


FIG. 6. Distribution of reconstructed y - z positions for blips consisting of matched hit clusters on all three planes, with E_{blip} between 2 and 3 MeV_{ee} (top) and above 3 MeV_{ee} (bottom). Red and green boxes in the top panel indicate signal and background regions for blip hot spot studies described in Sec. VI.

event fall into this energy range. This spectral feature is generally consistent with a prominent monoenergetic 2.614 MeV γ -ray associated with β -decay of ^{208}Tl , a daughter of the long-lived ^{232}Th radioisotope [57]. Figure 6 shows the y - z distribution of blips occurring in this 2-3 MeV_{ee} range. Hot spot regions are still present in the narrower energy range dominated by this spectral feature. This suggests that the G10 struts are a source of ^{208}Tl γ rays, and that ^{208}Tl and its ancestor ^{232}Th may be present in this material. This strong spectral feature in MicroBooNE's blip energy spectrum, combined with its apparent localization in specific MicroBooNE components, offers unique possibilities for LArTPC response calibration that will be studied in detail in the following section.

Above 3 MeV_{ee} in Fig. 5, a featureless and exponentially falling energy spectrum is observed, with around 7.8 blips per event in this energy range (4.4 blips with $E_{\text{blip}} > 5 \text{ MeV}_{ee}$). As shown in the bottom panel of Fig. 6, these higher energy blips are not clustered around the G10 struts, but instead are spread evenly throughout the detector, with a greater concentration near the top. This position distribution suggests a cosmogenic origin for these blips, which will be explored further in Sec. VII.

Reconstructed blip size is visualized in Fig. 7, showing a clear positive correlation between a blip's estimated length ds and its energy. The secondary population at high ds and low E_{blip} are blips featuring plane-matched clusters generated by different physics processes (or 'bad plane matches'), which represent roughly 1% of the total

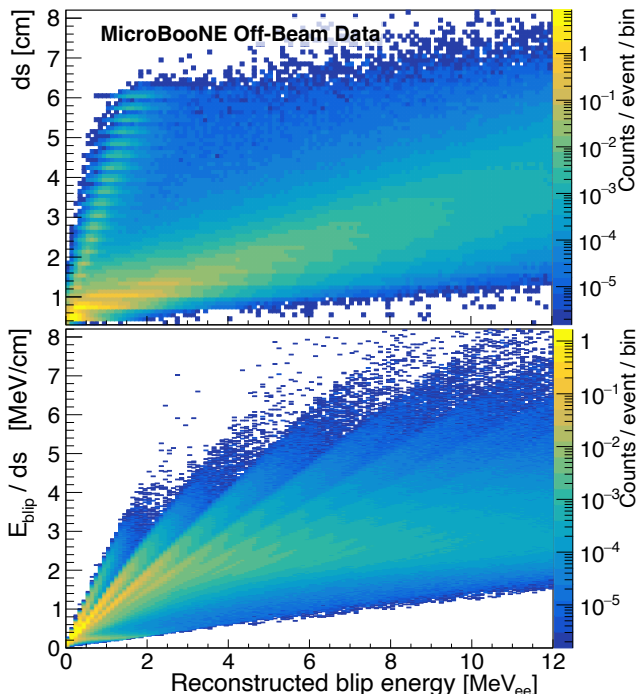


FIG. 7. Distributions of reconstructed blip size ds (top), and average blip energy per unit length E/ds (bottom), plotted against the reconstructed blip energy. Striped patterns present in these distributions are artifacts of the discrete spacing of MicroBooNE’s charge collection elements. The smaller secondary population extending to high ds in the top panel corresponds to blips featuring incorrectly plane-matched clusters.

reconstructed blip population. This figure also depicts blips’ E/ds , which is analogous to the stopping power (dE/dx) typically reconstructed for extended tracks in LArTPCs. Artifacts due to the use of discretized charge readout elements are present in both distributions in Fig. 7, an unavoidable consequence of studying small energy depositions with physical lengths on the order of the wire-to-wire separation. The substantial spread is also caused by the meandering, stochastic nature of low-mass charged particle trajectories in matter (well illustrated in, for example, Ref. [58]) and of ionization charge-sharing between wires on a plane.

In Sec. VII, we will study further whether particles with differing average dE/dx , such as protons and electrons, can be distinguished using blip size variables.

VI. STUDY OF RADIOGENIC HOT SPOTS

The 2-3 MeV_{ee} spectral feature present in the vicinity of MicroBooNE’s G10 struts calls for further study. In this section, we use this feature to perform the first percent-level-precision energy scale calibration of a neutrino LArTPC in the MeV regime and to measure the

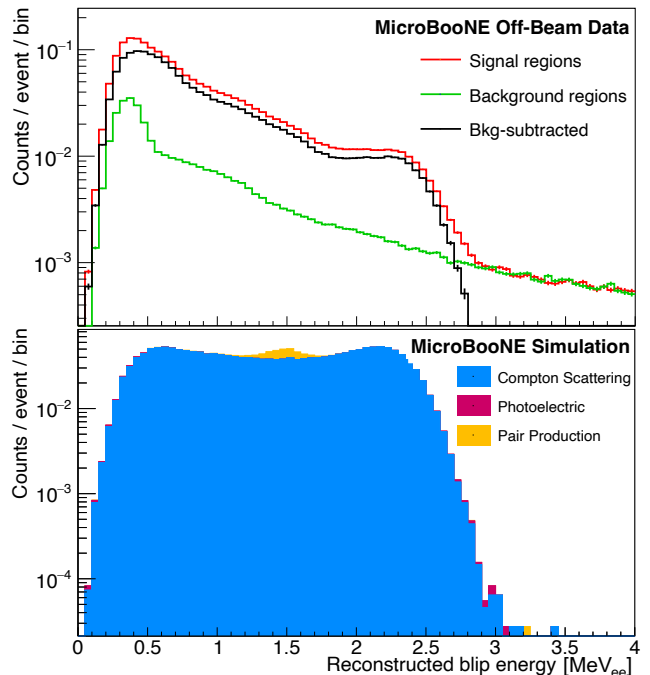


FIG. 8. E_{blip} spectra (top) for data blips located in the signal and background regions defined in Fig. 6, as well as their difference. Background-subtracted E_{blip} spectrum (bottom) in the same TPC regions produced by simulated monoenergetic 2.614 MeV γ rays in nearby G10 struts with stacked components by process.

specific activity of ^{208}Tl in MicroBooNE’s G10 struts. Prior to carrying out these studies, our hypothesis regarding the physics origin of this spectral feature was further verified using the single- γ MC datasets described in Sec. IV.

Potential issues related to dead wires were reduced by examining and simulating γ -ray emissions from a subset of five struts on the detector bottom, highlighted in Fig. 6. To study characteristics of only blips related to the G10 struts, we exclude all signal blips within 15 cm of a track, and then implement a background-subtraction procedure, with reconstructed signal and background blip locations indicated by the red and green boxes in Fig. 6. For both data and simulation, the energy spectrum of the green boxed sample far from the strut is then subtracted from that of the red boxed sample adjacent to the strut. Figure 8 depicts the reconstructed energy spectrum for data blips in the signal and background boxes as well as the subtracted result. The result of the subtraction is an increase in the prominence of the edge feature between 2 and 3 MeV_{ee} E_{blip} , mostly due to the reduction in content above roughly 2.5 MeV_{ee} .

The background-subtracted blip energy spectrum from the γ -ray MC dataset is also depicted in Fig. 8. In the 2-3 MeV_{ee} range, the simulated spectrum contains a very similar feature to the data spectrum, *i.e.* the falling edge. This feature in MC is due almost entirely to Compton-

scattered electrons produced by the simulated 2.614 MeV ^{208}Tl decay γ rays from the G10 struts. Thus, the feature of interest found in the data's E_{blip} spectrum is likely a Compton edge produced by interaction of 2.614 MeV γ rays generated by ^{208}Tl radioactive decays in MicroBooNE's G10 struts.

It should be noted in Fig. 9, that the background-subtracted E_{blip} spectrum contains substantial content below the edge feature, and also that data and MC spectra diverge widely below roughly 2 MeV. This is likely due to the presence of other radioisotopes in the decay chains of thorium and uranium that are also present in the struts. A closer analysis of this lower-energy portion of the E_{blip} spectrum requires accurate simulation of these isotopes and their various decay branches, a task that is beyond the scope of the current study. Fortunately, as demonstrated in a wide range of large MeV-scale particle detectors [59–62], the 2.614 MeV ^{208}Tl γ -ray's uniquely high energy and high intensity (produced in >99% of ^{208}Tl decays) allows its Compton edge and full-energy peak to appear in measured ambient background spectra with little interference from other radioisotopes.

1. MeV-Scale Energy Calibration Demonstration

With simulated and measured ^{208}Tl γ -ray Compton edges now clearly defined, we perform a calibration of reconstructed blip energy scales by testing the alignment of this feature between data and MC. Alignment is achieved by simultaneously applying an energy scale correction and a normalization scaling to the simulated spectrum, with best alignment judged by calculating the χ^2 between the data and adjusted simulated spectrum between 2.2 and 2.6 MeV_{ee}. As shown in Fig. 9, we scale the energy of simulated events up by 3.12% in order to best match the data (χ^2/ndf of 2.02/6), ndf being the number of bins minus the number of parameters used for the fit. Marginalizing over the normalization parameter, a $\Delta\chi^2$ below 2 was observed for an energy scaling parameter between 2.47% and 3.57%, defining the 1σ statistical uncertainty for that parameter.

A variety of systematic uncertainties, listed in Table I, have the potential to bias the simulated E_{blip} spectrum relative to data. The impacts of most parameter uncertainties are benchmarked by performing simulations with that parameter adjusted to the edge of its 1σ allowed region, refitting the data, and using the new best-fit value to define its 1σ systematic uncertainty contribution. Some parameters, such those related to diffusion, recombination, and space charge detector response parameters, have been well-established by previous MicroBooNE analyses [52, 63]. For diffusion, the longitudinal diffusion parameter D_L and its uncertainty were taken from Ref. [63], while the transverse diffusion parameter D_T , as done in Ref. [37], was assigned a large 30% uncertainty given the lack of existing measurements; D_L and

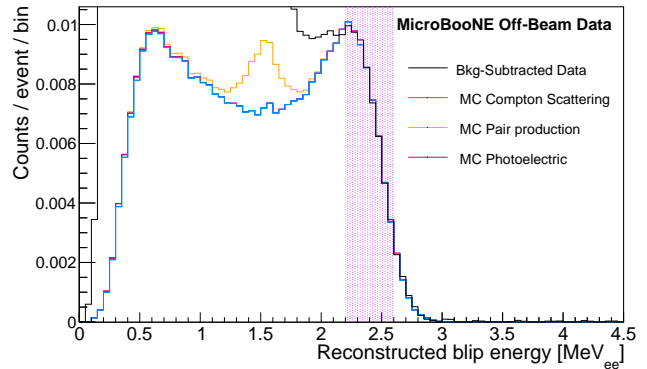


FIG. 9. Comparison of the measured Compton edge from 2.6 MeV ^{208}Tl decay γ rays to MC simulations of this feature with best-fit energy scale shift and specific activity parameters applied. Fit range indicated by the pink shaded region, see the text for further details about fit parameters and approach.

D_T uncertainties were assumed to be fully correlated. Recombination uncertainty was assessed by comparing results between the Birks and modified box models of ionization quenching [64, 65]. For space charge, we use maps corresponding to 1σ uncertainty ranges as defined in Refs. [52, 66]. Due to the high LAr purity in MicroBooNE during the period used for this analysis, ionization electron drift losses and associated uncertainties were negligible. Other uncertainties, such as that of the size and placement of struts, are conservatively estimated using mechanical drawings and photographs from the period of MicroBooNE detector construction. Background subtraction uncertainties were conservatively benchmarked using the case where no background subtraction is applied. Finally, minimization is performed with modest adjustments to the fit ranges to benchmark uncertainties related to the specific fit approach. The quadrature-summed total systematic uncertainty from all sources is also given in Table I. The approach followed to evaluate uncertainties is tailored to this MeV-scale analysis, and differs in several points to the data-driven detector systematic estimation detailed in [52] and used in MicroBooNE's higher energy measurements.

Thus we find that the MC energy scale correction parameter providing the best fit between MicroBooNE data and MC simulations is $(3.12 \pm 0.2 \text{ (stat)} \pm 1.2 \text{ (syst)})\%$. This MeV-scale calibration using ambient radiological activity is the first of its kind for a large neutrino LArTPC. The general level of correspondence between data and simulation indicates that MeV energy scales in MicroBooNE are well-modeled to the few-percent level. This result improves upon previous demonstrations in MicroBooNE performed using ^{214}Bi β -decays [37] by improving blip-based energy calibration precision to the percent level.

A blip-based calibration scheme similar to the one demonstrated here, potentially using the same ^{208}Tl Compton edge feature, could serve as an attractive option

Parameter	Value and Uncertainty	Deviation from CV	
		Energy Scale (absolute, %)	G10 Activity (relative, %)
Strut mass	Nominal strut dimensions versus ± 1 mm	<0.2	7.1
Strut extent into active volume	(1.2 ± 1.0) cm	<0.2	18.1
Recombination	Modified box versus Birks	<0.2	<2.0
Space charge	Default versus $\pm 1\sigma$ charge map	0.3	<2.0
Diffusion	$D_L = 3.74_{-0.29}^{+0.28}$, $D_T = 5.9_{-1.8}^{+1.8}$ (cm ² /s)	<0.2	<2.0
Bin width	0.05 MeV versus 0.1 MeV	0.02	0.1
Fit range	$2.2_{-0.1}^{+0.3} - 2.6_{-0.4}^{+0.2}$ MeV	0.6	12.4
Background subtraction	Default versus no subtraction	0.91	13.5
Total		1.2	26.9

TABLE I. Summary of systematic uncertainties considered in E_{blip} energy scale calibration and ^{208}Tl specific activity measurements. For each uncertainty source, the central value (CV) and bounding scenario cases are given, along with the absolute change in the energy scale correction and fractional change in activity when the parameter is varied in the specified manner. Statistical uncertainties associated with the MC-derived energy scale and activity changes are 0.2% and 2.0%, respectively. The extent of the G10 strut protruding past the field cage profiles into the active LAr volume is visualized in Fig. 2. See the text for more description of each systematic error source and its applied adjustment.

for detector-wide, region-specific, or wire-by-wire energy scale calibrations in future large LArTPCs, particularly those deployed underground in low cosmic ray environments, such as DUNE. The non-uniform distribution of ^{208}Tl inside the detector may limit its attractiveness as a calibration source: for example, the present analysis has provided a calibration of five detector regions comprising a small fraction of MicroBooNE’s active TPC volume. In addition, the lack of knowledge of blip true x locations represents a current limitation of blip-based calibration methods more generally.

2. Measurement of G10 Specific Activity

The same fit procedure used to evaluate energy scale agreement between data and MC in the last section can also be used to measure the specific activity of ^{208}Tl in MicroBooNE’s G10 struts.

In the MC, three isotropically-oriented γ rays originating at random points within each G10 strut’s volume were simulated during each 3.2 ms detector readout, equivalent to a G10 specific activity of 60.8 Bq/kg. As shown in Fig. 9, a best fit between measured and simulated Compton edges occurs when the level of simulated γ -induced blip content per event is scaled by a factor of 0.193, to 11.7 Bq/kg. An examination of nearby fit phase space similar to that described in the previous sub-section gives a 1σ statistical uncertainty range of ± 0.18 Bq/kg, and an identical systematic error determination is also performed, with results given in Table I. The resulting measured specific activity of the G10 struts is $(11.7 \pm 0.2$ (stat) ± 3.1 (syst)) Bq/kg, with the systematic error dominated by uncertainties in the strut’s exact location. This activity level is comparable to existing measurements of ^{208}Tl and ^{232}Th activity in fiberglass detector components collected in the community-standard radiopurity database [67], specifically fiberglass component entries from Refs. [68, 69]. This database

also includes measurements of ^{208}Tl and ^{232}Th activity well below that measured for MicroBooNE (for example, other fiberglass components in Ref. [68]), indicating a large potential variability in radiopurity. Thus, care should be taken in the selection of fiberglass components, such as structural beams, circuit boards, and other electrically insulating or resistive components, for future LArTPC experiments aiming to perform MeV-scale physics measurements. Radiopurity screening measurement campaigns may be advisable.

VII. STUDY OF COSMOGENIC BLIPS

We now consider the blip population in Fig. 5 with $E_{\text{blip}} > 3$ MeV_{ee}, which are distributed throughout the active volume but more concentrated at the top, suggesting a cosmogenic origin. As shown in Fig. 10, the energy spectrum shape in this range is similar between data and the CORSIKA cosmic ray MC described in Sec. IV. While the shape is similar, we observe a 10% offset in overall normalization between the two, contrasting with the few-percent-level agreement in measured and modeled cosmic muon track rates we previously reported [54]. We observe the same data-MC correspondence in muon rates when considering only long (>5cm) tracks, but we see a substantial data-MC offset for short (<5cm) tracks that mirrors the aforementioned offset in overall data and MC blip activity. We reserve further discussion of observed data-MC mismatches in the cosmic ray flux for Sec. VII 2.

Figure 10 also shows a distribution of track proximity for all blips, defined as the 3D distance of closest approach between each blip and the reconstructed cosmic track >5 cm in length nearest to it. The shape of the blip-track proximity distribution provides clues to the physics processes generating the cosmic blip population. The strong peak at short proximity and subsequent exponential decrease within 10 to 35 cm suggests that a

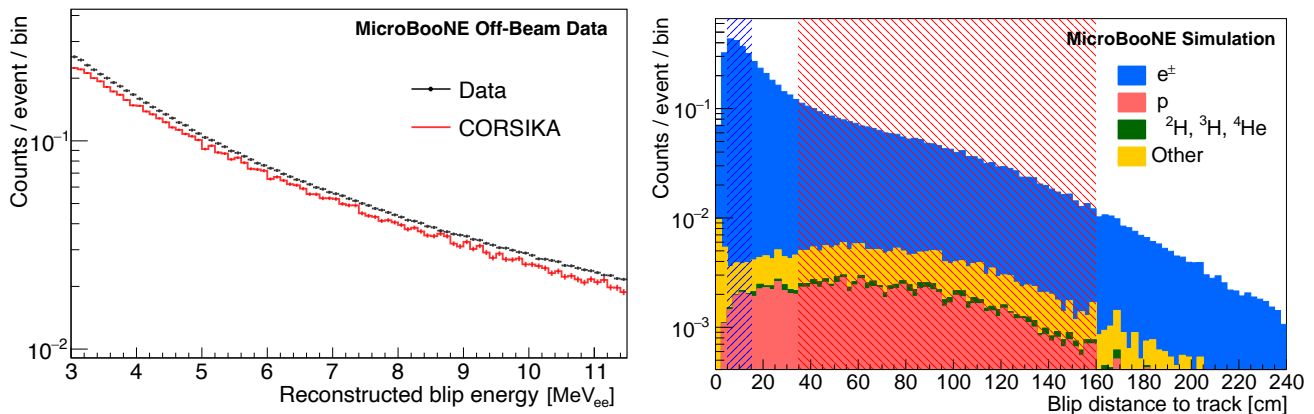


FIG. 10. (Left) reconstructed E_{blip} spectrum above 3 MeV for data (black points) and for CORSIKA MC (red). Error bars indicate statistical uncertainties. (Right) Blip-track proximity, or the distance of closest approach between blips with $E_{\text{blip}} > 3$ MeV and the nearest track longer than 5 cm for CORSIKA MC. Contributions from different blip parent particle ID classes are shown as stacked colored histograms. Besides p , and e^\pm , heavier nuclei also contribute to reconstructed blips. The *Other* category is comprised primarily by muons/pions indicating blips due to misreconstruction of tracks. Blue (red) transparent vertical bands indicate ranges selected for the short (long) blip-track proximity subsets described in the text.

majority of selected cosmic blips are the byproduct of bremsstrahlung γ rays generated following hard scattering of cosmic muons or production of high-energy δ electrons. This hypothesis is supported by probing the true parentage of CORSIKA MC blips at track proximities between 5 and 15 cm, which is also pictured in Fig. 10: an overwhelming majority of these blips ($\gg 99\%$) have electrons (from γ -ray Compton scattering or pair-production) or positrons (from pair production) as their direct producer, and $>85\%$ of these true electron parents descend from bremsstrahlung photons generated by μ^\pm δ rays.

According to Fig. 10, blips farther from tracks (long blip-track proximity) are also primarily produced by true electrons and positrons. Within the blip-track proximity window of 35 to 160 cm, 93% of blips have an e^\pm direct producer. Roughly 40% of these share the same δ -ray bremsstrahlung ancestry that dominated the short (5-15 cm) blip-track proximity sample, indicating that a substantial fraction of the diffuse, track-uncorrelated blips in MicroBooNE are produced by muons whose trajectories are largely or entirely outside the active TPC volume. Blips from γ rays from external cosmic muons appear randomly in time within the TPC readout period, so their measured x coordinate cannot be used for fiducial-based mitigation along the drift direction. The other primary parentage categories are cosmic neutron inelastic nuclear scattering (roughly 30%) and electromagnetic showers not associated with muon activity (roughly 25%). Absent further information from other detector systems, such as a muon tagger or an efficient light collection system, or cuts on other LArTPC variables, such as nearby shower objects or high adjacent blip activity, these blip categories represent potentially troublesome backgrounds to purely MeV-scale topologies for

future low-energy neutrino measurements in near-surface LArTPC experiments [22, 70–72].

Figure 10 also shows a sub-dominant but substantial population of blips in CORSIKA simulations that is directly produced by protons. While representing only 0.3% of the population at short (5-15 cm) blip-track proximity, proton-generated blips make up 5% of the predicted population at long blip-track proximity (35-160 cm). A vast majority of blip-generating protons, $>99\%$, have neutrons as their direct producer. This indicates that this population is generated by inelastic (n, p) interactions [73] between incident cosmic neutrons and argon nuclei in the active TPC.

1. Defining Particle Discrimination Criteria

Figure 11 shows reconstructed average energy deposition density, E_{blip}/ds , versus E_{blip} for the long blip-track proximity (35-160 cm) sample in CORSIKA MC, separated into two direct producer categories: p versus e^\pm . The distributions, while overlapping, are clearly offset from one another, indicating that MicroBooNE should be capable of exploiting the offset in stopping power of low-energy charged particles of differing masses. While an offset might also be expected between the two prominent electromagnetically-produced blip categories – pair-production (an e^+e^- pair) and Compton scattering (a single e^-) – we find this offset to be fairly small in reconstructed blip space and do not consider this comparison further.

To demonstrate p - e^\pm discrimination, we outline a particle identification (PID) cut based on the value of E_{blip}/ds below which 99% of all blips reside for each 0.5 MeV_{ee} bin. Discrete boundary values provided by the

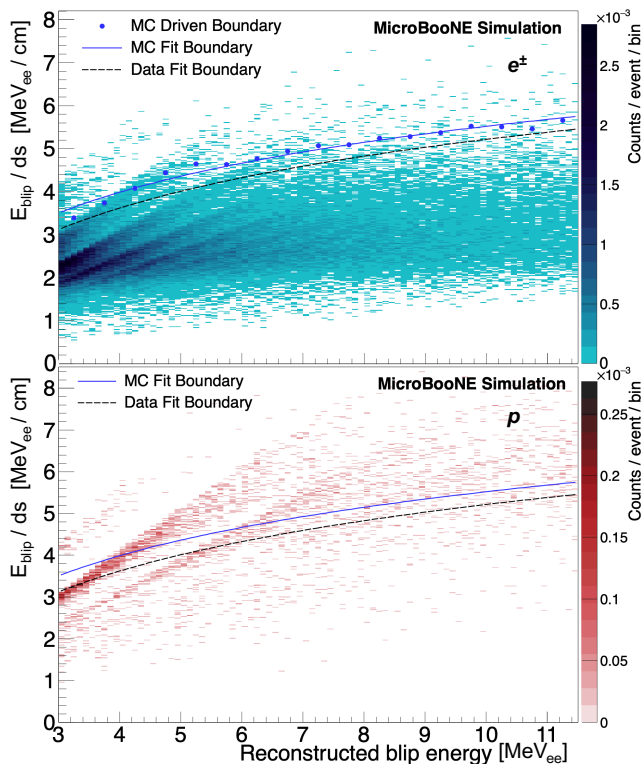


FIG. 11. Distribution of average energy deposition density, E_{blip}/ds , for blips of various reconstructed energies in CORSIKA MC simulations. Blips generated by true e^\pm parents are shown in blue (top), while those with p parents are shown in red (bottom). In the top panel, the blue points show the E_{blip}/ds value above which 1% of CORSIKA MC blips reside for each 0.5 MeV_{ee} bin. The blue and black lines show the best-fit log function (Eq. 2) for MC and data, respectively.

comparatively e^\pm -pure sample of CORSIKA MC blips with short (5-15 cm) blip-track proximities are illustrated by the blue points in the top panel of Fig. 11. A continuous PID selection boundary is then defined by a log fit applied to this set of per-sub-range boundary values:

$$f(E_{\text{blip}}) = a_1 \ln(a_2 E_{\text{blip}}). \quad (2)$$

Figure 11 also depicts the log function providing the best fit to the short blip-track proximity sample for data and for CORSIKA MC. For CORSIKA MC, the fit defines a continuous rejection cut boundary at $(a_1, a_2) = (1.69 [\text{MeV}_{ee}/\text{cm}], 2.64 [\text{MeV}_{ee}^{-1}])$. While 1σ statistically allowed ranges for the parameters, $(\pm 0.08 [\text{MeV}_{ee}/\text{cm}], \pm 0.35 [\text{MeV}_{ee}^{-1}])$, are wide, this is mostly due to fit degeneracy. If a_1 is held constant at its best-fit value, the 1σ range for a_2 is reduced to $\pm 0.04 [\text{MeV}_{ee}^{-1}]$.

In Fig. 12, we show the results of applying this PID selection to the long blip-track proximity CORSIKA MC sample shown in Fig. 11. The efficiency of the PID cut, defined as the number of reconstructed hadron-generated blips above the cut value over the total number of reconstructed blips, is shown for the p -like category as well as

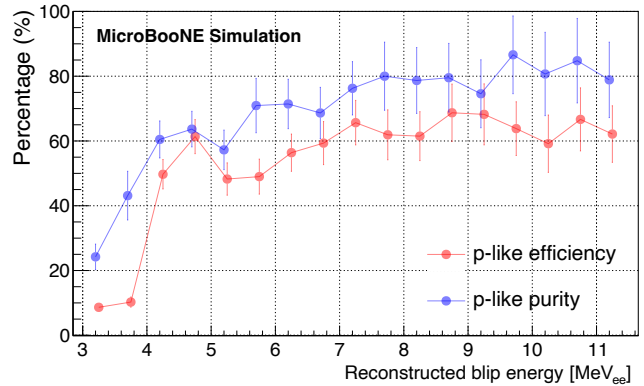


FIG. 12. Red: Efficiency for selecting reconstructed blips produced by true hadronic parents with the PID cut described in the text. Blue: composition of the reconstructed blip sample selected with this PID cut. The hadronic sample groups together contributions from p , ${}^2\text{H}$, ${}^3\text{H}$, and ${}^4\text{He}$, with the latter three comprising roughly 7% of the selected sample. Error bars represent statistical uncertainties in the CORSIKA MC sample.

its fractional contribution to the total selected sample. Blip contributions from other nuclear fragments (${}^2\text{H}$, ${}^3\text{H}$, and ${}^4\text{He}$) are included in the p -like sample, and comprise roughly 7% of the total. The efficiency of the proton-like PID cut starts at roughly 10% near 3 MeV_{ee} , but quickly rises to 50% around 4 MeV_{ee} and reaches a maximum of roughly 70% at higher E_{blip} . We note here that 4 MeV_{ee} E_{blip} corresponds to roughly 18 MeV true proton energy; this conversion is detailed more fully in subsequent sections, and in the supplementary materials accompanying this manuscript.

Sample purity, defined as the number of selected true hadron-generated blips divided by the total number of blips above the PID cut value, is also shown in Fig. 12. The applied cut yields a candidate proton blip sample that has $>50\%$ purity starting at around 3.75 MeV_{ee} E_{blip} , which increases to $>80\%$ at the highest considered blip energies. Even at the lowest considered E_{blip} , where proton selection efficiency is low, the cut nonetheless provides roughly an order of magnitude reduction in electron-like backgrounds relative to the proton-like sample. This suggests that meaningful p/e discrimination is possible at and even below 20 MeV in true proton energy.

When the demonstrated PID cut technique and boundary is applied to MicroBooNE data, statistical and systematic uncertainties related to these quoted efficiencies and background rejection factors must be considered and propagated. To propagate statistical uncertainties, we use the fit uncertainty in a_2 while fixing a_1 to address the fit degeneracy described above. A check of this procedure fixing a_2 and allowing a_1 to float yields similar allowable statistical variation in the PID boundary. We define systematic uncertainty contributions using similar techniques to that described in Sec. VI: the PID cut

boundary is rederived and applied using systematically varied MC datasets. Considered systematic effects include electron and proton recombination, diffusion, space charge, blip energy scales as benchmarked in Sec VI 1, and signal impurity. A blip energy scale uncertainty is defined using the difference in results between the CORSIKA dataset with 0% and the best-fit +3.12% applied energy scale shift from Sec. VI 1, while a signal impurity systematic is defined by the difference in PID cut boundary generated by the full short blip-track proximity dataset versus only the subset of true e^\pm -produced blips. Systematically varied recombination, diffusion, and space charge datasets are defined as described in Table I. Recombination and energy scale systematics provide the dominant uncertainty to particle identification measurements, while MC statistical uncertainties also serve as a non-negligible contributor. The impacts of systematic uncertainties on reported proton counts are described in the following subsection.

2. Identification of Cosmogenically Produced Protons in MicroBooNE

With the use and performance of the PID cut variable now defined using CORSIKA MC, we apply a data-derived PID cut to the MicroBooNE dataset described in Sec. IV. Using the short blip-track proximity dataset in data, we observe best-fit PID boundary parameters of $(a_1, a_2) = (1.72 \pm 0.03 \text{ [MeV}_{ee}/\text{cm}], 2.08 \pm 0.10 \text{ [MeV}_{ee}^{-1}])$. When setting a_1 to its best-fit value, we find an uncertainty on a_2 of $\pm 0.01 \text{ [MeV}_{ee}^{-1}]$. As shown in Fig. 11, substantial differences in PID boundary fit constants are observed between data and CORSIKA MC. These differences are likely related to the mismatch between modeled and measured wire pulse shapes, which are illustrated and discussed in Ref. [42]. While this offset is notable, its impact on PID systematic uncertainties is mitigated by using the data's e^\pm -rich sample at short blip-track proximity to define PID regions for other proton-rich blip subsets in data, as opposed to applying an MC-derived PID boundary directly to the data.

Applying the data-derived PID cut to all blips with energies greater than 3 MeV, we select a substantially higher fraction of proton-like blip candidates in the sample farther from tracks, $(2.06 \pm 0.01)\%$, compared to that selected in the close-proximity sample, which is $\sim 1\%$ by construction. This proton-like fraction, plotted as a function of E_{blip} in Fig. 13, increases with reconstructed energy, reaching an average of $(3.12 \pm 0.03)\%$ above 6 MeV. In regions that are less dominated by tracks and their associated bremsstrahlung radiation within MicroBooNE event readouts, an additional population of highly-ionizing, proton-like energy depositions is clearly visible.

As shown in Fig. 13, a comparatively high fraction of proton-like blips is also visible in the CORSIKA MC's long blip-track proximity subset. In the previous subsection,

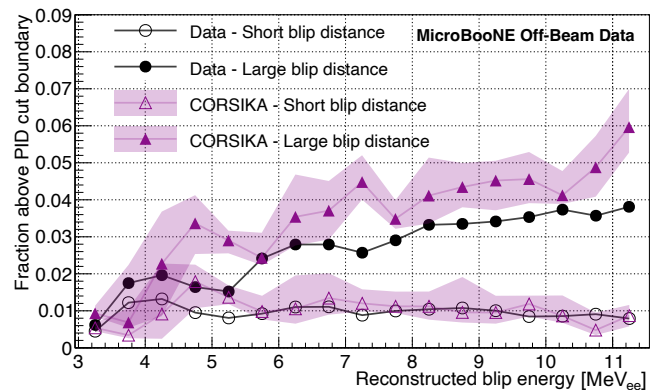


FIG. 13. Fraction of blips in the short (open markers) and long (solid markers) blip-track proximity categories passing the applied proton-like PID cut. Shaded bands on MC data points represent systematic uncertainties. While the fraction for short proximity blips is roughly 1% by construction, the fraction for the long blip-track proximity sample deviates strongly from this value, indicating the presence of true protons in this sample. The fractions shown for the CORSIKA MC simulations indicate an over-prediction of cosmogenically-produced protons.

tion, it was established that this higher proton-like blip fraction indeed derived from the presence of true proton energy depositions generated by cosmogenic neutrons. This strengthens confidence that the selected proton-like sample is largely composed of true cosmogenically-produced protons.

While these cosmogenically-generated proton blips are observable in both data and CORSIKA MC, the overall proportion is substantially lower in data. This difference is visible in Fig. 13, as well as in Fig. 14, which shows the predicted and measured E_{blip} spectrum of all selected proton-like blips in the long blip-track proximity subset. While spectral features are generally comparable, the overall normalization of the two does not agree.

Above 3 MeV, a total of $27,717 \pm 166$ (stat) proton-like blips are measured, while simulations predict $39,684^{+12714}_{-8073}$ (syst) ± 199 (stat), corresponding to 0.0425 ± 0.0003 proton-like blips per event for data and $0.0607^{+0.0195}_{-0.0124}$ for MC. Similarly, above 6 MeV, we expect $20,842^{+4057}_{-3140}$ (syst) ± 144 (stat) counts but we only detect $13,653 \pm 117$ (stat), that is, $0.0319^{+0.0062}_{-0.0048}$ and 0.0209 ± 0.0002 proton-like blips per event in MC and data respectively. Stated in terms of data-MC ratios, above 3 MeV, the ratio of proton-like blips in data compared to MC is only about 0.70 ± 0.18 , with the uncertainty dominated by systematic effects. Above 6 MeV, where our proton purity is expected to be higher, this data-to-MC ratio is 0.66 ± 0.11 . This deficit becomes substantially more pronounced when the MC-derived PID cuts are used on data, rather than the data-driven ones. The significance of the data-MC difference is limited by systematic uncertainties from electron-ion recombination

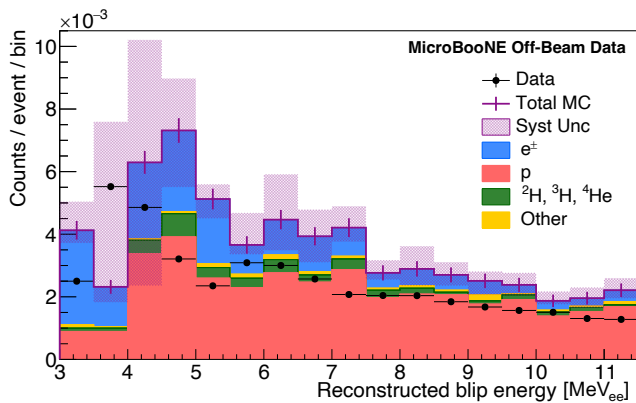


FIG. 14. Measured (black points) and simulated (stacked histograms) reconstructed energy spectrum of selected proton-like events in MicroBooNE’s long blip-track proximity sample. Purple vertical lines represent the statistical uncertainty in the MC dataset.

modeling and blip energy scales, with both producing substantial shifts in the fitted PID boundary.

Differences in measured and predicted cosmogenic protons could be due to incorrect predictions of cosmogenic neutron fluxes in the CORSIKA generator, and/or to incorrect neutron transport in the LArTPC and surrounding material by Geant4. As demonstrated in previous MicroBooNE cosmic ray studies [54], muon fluxes predicted at Earth’s surface differ by $\sim 40\%$ when considering incident cosmic protons versus the full array of nuclear fragments, with the former providing a better fit to MicroBooNE LArTPC track datasets. Even in cases where predicted muon rates in MicroBooNE are found to be relatively consistent between the different cosmic ray generator packages CORSIKA and CRY [74], predicted neutron fluxes incident on the TPC are still offset between these generators by more than a factor of 2 [75]. These past studies lend credence to the notion that incident neutron fluxes may be incorrectly predicted in the CORSIKA simulation used for this study. On the other hand, the inclusive interaction cross section of high-energy neutrons on argon nuclei has only recently been measured [76, 77], and no exclusive measurements of proton-producing interaction channels have been performed to date. Thus, it is also reasonable to attribute an excess of predicted cosmogenically-produced protons to mismodeling of cosmic neutron interactions within the MicroBooNE LArTPC by Geant4.

3. General Purpose Proton Selection

The pure sample of low-energy protons described in the previous section was selected with a blip PID metric and cut values tuned for a specific use case: isolating a small and sub-dominant cosmogenically-produced proton population amidst a dominant sea of low-energy electro-

magnetic cosmogenic backgrounds. In this case, an E_{blip} -independent 99% electron blip rejection factor was sufficient for achieving better than 50% purity for true proton signatures above 4 MeV_{ee} E_{blip} , a value above which a substantial fraction of cosmogenically-produced protons reside. Other use cases for p - e^{\pm} separation may call for differing levels of background rejection or efficiency than were implemented here. In an appendix that accompanies this manuscript we provide more general quantitative depictions and descriptions of blip energy response and low-energy particle discrimination capability for electrons and protons in the MicroBooNE detector. This appendix, alongside proton and electron response and PID data files, should enable the community to more broadly explore potential applications of low-energy p - e^{\pm} discrimination in neutrino LArTPCs. While LArTPCs operating at electric fields higher than MicroBooNE’s nominal 274 V/cm may achieve better charge-based detection efficiency for low-energy particles, the information provided with this manuscript offers a baseline level of performance, validated on data, from which future studies can build.

VIII. SUMMARY

Using ambient radiogenic and cosmogenic activity in the MicroBooNE detector, we have demonstrated new energy calibration and particle discrimination capabilities at the MeV scale in large neutrino LArTPCs. Isolated low-energy activity was identified and reconstructed in MicroBooNE data and MC simulations using a custom-built LArSoft toolkit, `BlipReco`, validated in previous MicroBooNE results. We performed detailed analysis on $>600,000$ MicroBooNE events recorded in the absence of BNB beam activity, with an average of roughly 106 blips observed per event.

The lowest-energy blips within the reconstructed sample exhibited multiple spectral features, including a prominent shoulder in the 2-3 MeV_{ee} E_{blip} range, which was found to primarily originate from 2.614 MeV γ rays released by ^{208}Tl present in G10 fiberglass mechanical struts that support the MicroBooNE TPC and field cage. This feature was used to perform a percent-level precision calibration of blip energy scales in MicroBooNE, with data and MC features found to agree within $(3.1 \pm 0.2 \text{ (stat)} \pm 1.2 \text{ (syst)})\%$. This demonstration is vital for performing reliable high-level physics analysis with MeV-scale blips in MicroBooNE. It also provides a useful blueprint that can be followed in the calibration of future large LArTPC detectors. In particular, the underground DUNE experiment will be unable to perform LArTPC response calibrations with large cosmic muon datasets, and will likely rely on MeV-scale calibration strategies akin to those demonstrated in this paper.

Study of the 2.614 MeV ^{208}Tl γ -ray Compton edge in MicroBooNE also enabled a measurement of the specific activity of ^{208}Tl in its fiber-

glass structural materials, which was found to be $(11.7 \pm 0.2 \text{ (stat)} \pm 3.1 \text{ (syst)}) \text{ Bq/kg}$. This measurement is relevant to other large neutrino LArTPC experiments, since all incorporate fiberglass structural materials in the immediate vicinity of their active TPC volume. While this background source has no direct impact on the primary MicroBooNE physics goals, its presence serves as a cautionary tale for future LArTPC efforts with ambitious MeV-scale physics programs. In particular, a comparable level of ^{208}Tl activity in the DUNE far detector may represent a potential challenge for the triggering and reconstruction of astrophysical neutrinos. Our result emphasizes the necessity for dedicated material screening campaigns for DUNE and other low-energy single-phase LArTPCs.

New particle discrimination capabilities at the MeV-scale were explored using cosmogenic blips with more than 3 MeV_{ee} of deposited energy. By making selections on the charge-to-size ratio of reconstructed blips, we identified a purified sample of topologically isolated low-energy protons generated by inelastic scatters of cosmic neutrons. With true proton kinetic energies ranging from $\sim 15\text{--}40 \text{ MeV}$, the dataset is the lowest-energy of its kind ever identified in a neutrino LArTPC. While typical MicroBooNE events contain more than a dozen reconstructed cosmic muon tracks, we identified roughly one reconstructed neutron-induced proton blip candidate per two MicroBooNE events. This rate of proton-like blip detection was found to be substantially lower than cosmic ray MC predictions, indicating neutron flux or transport issues in MicroBooNE's CORSIKA and Geant4 based cosmic simulation.

This analysis represents the first-ever exploration of particle discrimination capabilities at MeV energy scales in a large neutrino LArTPC, and it opens doors to a

range of future novel cross section and BSM physics measurements in these detectors. To further aid experimentalists and phenomenologists, we have provided an appendix and data files with descriptions of low-energy response and particle discrimination capabilities for the MicroBooNE LArTPC.

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