

Essays on Models with Mixed Frequency Data and Time Varying Parameters

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A thesis submitted for the degree of

Doctor of Philosophy

July, 2022

I would like to dedicate this thesis to my family.

Declaration

I declare that the work presented in this thesis is, to the best of my knowledge and belief, original and my own work. The material has not been submitted, either in whole or in part, for a degree at this, or any other university.

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Abstract

This thesis focuses on two statistical challenges in time-series modelling. The first is when variables' observations are available at different frequencies. The second is when the coefficients of a model are time-varying with stochastic volatility. The impact of these challenges and the value of the suggested remedies are assessed in empirical financial-economic applications.

In addressing the first statistical challenge, disaggregation from the low- to the highfrequency domain is one of the methods that has long been used in several pieces of literature. The first chapter evaluates the existing disaggregation methods with thorough comparisons to provide comprehensive guidance for an empirical user. The second chapter builds on these results to examine the value-added in forecasting the volatility of financial stock prices by incorporating information from variables with mixed frequencies such as market sentiment indicators, economic variables, and activity measures. A representative factor(s) of all potential predictors from both frequencies results in significant forecast gains in predicting long-term financial volatility even during the 2007-08 financial crisis. The third chapter proposes a state-space model to incorporate features of timevarying coefficients to reflect the dynamic relationship between the dependent and the explanatory variables. Mainly, the model consists of two hierarchical states. First, the time-varying coefficients follow an autoregressive (AR) process with heteroskedastic innovations. Second, the log-transformation of the conditional variance of these innovations is also modelled as an AR process. In an empirical study, we utilize the proposed methodology to forecast the volatility of financial stock prices. We find that the proposed features consistently and significantly enhance the forecasting accuracy compared to a benchmark model and its existing variants.

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Contents

	Eva	luation of interpolation and extrapolation methods of low-	
	freq	uency series to the high-frequency domain	1
	1.1	Introduction	2
	1.2	Chow-Lin Framework: a Linear Transformation Function	5
	1.3	Theoretical Properties of the Linear Transformations	15
	1.4	Estimation	17
	1.5	Simulation Study	19
	1.6	Conclusion	26
	App	endices	27
	1.A		27
		1.A.1 Proof of Proposition 1.2.1	27
		1.A.2 Proof of Lemma 1.2.1.1	33
		1.A.3 Proof of Proposition 1.2.2	38
	1.B	Monte Carlo Study - Tables	44
-			
2		e role of economic and financial indicators in forecasting stock	
		atility 6	67
	2.1	tility 6 Introduction	68
		tility 6 Introduction 6 Predictors of the financial volatility 7	58 71
	2.1	tility 6 Introduction 6 Predictors of the financial volatility 7 Measures and Models 7	58 71 74
	$2.1 \\ 2.2$	tility 6 Introduction 6 Predictors of the financial volatility 7 Measures and Models 7	58 71
	$2.1 \\ 2.2 \\ 2.3$	tility 6 Introduction 6 Predictors of the financial volatility 7 Measures and Models 7 Data 7	58 71 74 79
	2.1 2.2 2.3 2.4	tility 6 Introduction 6 Predictors of the financial volatility 7 Measures and Models 7 Data 7 Empirical Results 8	58 71 74 79 81
	2.1 2.2 2.3 2.4	tility 6 Introduction 6 Predictors of the financial volatility 7 Measures and Models 7 Data 7 Empirical Results 8 2.5.1 In-sample Estimation	58 71 74 79 81 81
	2.1 2.2 2.3 2.4	tility6Introduction6Predictors of the financial volatility7Measures and Models7Data7Empirical Results82.5.1In-sample Estimation2.5.2Out-of-Sample Analysis	58 71 74 79 81 81
	2.1 2.2 2.3 2.4	tility6Introduction6Predictors of the financial volatility7Measures and Models7Data7Empirical Results82.5.1In-sample Estimation2.5.2Out-of-Sample Analysis2.5.3Forecasting Significance	58 71 74 79 81 81 83
	 2.1 2.2 2.3 2.4 2.5 	tility6Introduction6Predictors of the financial volatility7Measures and Models7Data7Empirical Results72.5.1In-sample Estimation2.5.2Out-of-Sample Analysis2.5.3Forecasting SignificanceConclusion6	58 71 74 79 81 81 83 88

3	\mathbf{SH}	ARP: A State-Space HAR model using Particle Gibbs Sampling	121
	3.1	Introduction	122
	3.2	Volatility Measure and HARL Family of Models	126
	3.3	Methodology	129
	3.4	Empirical Study	133
		3.4.1 Data	133
		3.4.2 Empirical Estimation of the SHARP and SHARP-sv models .	135
		3.4.3 Out-of-sample Comparative analyses	139
	3.5	Conclusion	150
	App	endices	151
	3.A	Derivations of the first and second moments	151
	3.B	Posterior Derivation	153
	$3.\mathrm{C}$	Particle filtering within MCMC	157
	3.D	HARSL estimation using Kalman Filter	160
D.	oforo	neos	169

References

162

Chapter 1

Evaluation of interpolation and extrapolation methods of low-frequency series to the high-frequency domain

Abstract

We revisit the temporal disaggregation methods in the Chow-Lin (CL) framework. First, we extend the dynamic Chow-Lin model of Poissonnier (2018) to estimate high-frequency extrapolated and interpolated data points. Second, we show that CL methods can be expressed as a linear transformation function in a matrix form. Specifically, the estimated high-frequency series is the sum of two signals. The underlying assumptions drive the first signal. In contrast, the second signal combines the latter and the chosen indicator variable(s). Such decomposition reveals which factors influence the accuracy of both interpolated and extrapolated estimates. Lastly, we perform a Global Monte Carlo study to compare the models under different statistical scenarios. The study guides practitioners in choosing the most appropriate temporal disaggregation method for their application.

Keywords: Chow-Lin interpolation Extrapolation Mixed Frequency Temporal Disaggregation

1.1 Introduction

It is often the case that the financial and macroeconomic series are observed at different frequencies. Researchers use the mixed data sampling (MIDAS) model by Ghysels, Santa-Clara, and Valkanov (2004) to incorporate high-frequency (HF) variables to estimate a low-frequency (LF) variable. In contrast, Ghysels (2016) and Foroni, Guérin, and Marcellino (2018) propose the mixed-frequency VAR and Reverse Unrestricted MIDAS model, respectively, to incorporate LF information to predict a HF variable.

This paper's interest lies in the frequency conversion from low to high, including monthly to daily, which permits the study of models at a HF level. It is particularly interesting to practitioners who believe that HF forecasting accuracy is enhanced when LF variables are available on a HF level. Interpolation refers to generating HF estimates of a specific LF variable between two of its LF observations. Whereas generating HF values beyond a LF observation is expressed as extrapolation (Chow and Lin, 1971). By definition, extrapolation estimates future HF values of a LF variable by assuming that its unobserved HF values will continue to follow a particular hypothesised model.

The most prominent interpolation-extrapolation method in the literature is the one proposed by Chow and Lin (1971). More recently, a number of researchers extended the Chow-Lin (CL) methodology to dynamic models. However, their primary interest lay in temporal disaggregation (interpolation) rather than inter-and extrapolation. For example, Poissonnier (2018) mention that extrapolation can be obtained in a similar approach to the interpolation method but do not illustrate the extrapolation adaption to their proposed models. Furthermore, CL methodology is popular in applications with constant frequency ratios such as quarterly to monthly or yearly to quarterly. Researchers suggest that it can also be implemented, in a similar fashion, for non-constant frequency ratios such as monthly to daily. However, such implementation in the case of non-constant and large frequency ratio was neither demonstrated nor assessed.

There are multiple methods introduced in the literature. While all have common grounds, the difference is in the proposed relationship between the unobserved HF values of the LF variable and the HF indicator variable(s). In particular, Chow and Lin (1971), Fernández (1981), and Litterman (1983) assume a static model with a first-order auto-regression, unit root, and ARIMA(1,1,0) stochastic disturbance terms, respectively. Poissonnier (2018) presents a dynamic model with a general data generating process (DGP), such that the stochastic disturbance terms may follow either an AR, unit root or ARIMA process.

According to the authors above, each of these models is deemed helpful depending on the setting of the application. For example, the traditional CL method is more suited when the data are generated from a static model. On the other hand, it is more sensible to apply dynamic models when it is assumed that the observations possess some stagnant characteristics. Although the intuition behind these arguments is sound, to the best of our knowledge, there is a lack of theoretical justification and thorough empirical comparison of these models in the literature. Hence, the purpose of this paper is to address the following points. i) The algebraic difference between the methods. ii) The consequences of using a misspecified model to obtain in-sample and out-of-sample estimates of unobserved HF values under the CL framework. iii) The recommendation for practitioners interested in estimating unobserved HF in-sample and out-of-sample values of a LF variable.

In Section 1.2, we summarize the main models and their estimation procedures within the CL framework. Then, we present the theoretical properties of the hypothetical estimates in Section 1.3. We show that all the HF estimates obtained by the temporal disaggregation methods in the CL framework can be decomposed into the sum of two series: 1) one that relies on the assumptions underlying the model and 2) another that depends on the assumptions as well as the observed HF indicator variable(s). Exploring the properties of each signal unveils the consequence of each input on the estimation. Section 1.4 illustrates the feasible estimates of the unobserved HF estimates for each method. Section 1.5 provides an extensive simulation study based on multiple pre-assigned DGP, then obtains the feasible HF estimates of each simulated series using all the mentioned models in the Chow-Lin framework. The estimated series are compared with the true simulated ones to evaluate the models under various specifications. Section 1.6 concludes.

1.2 Chow-Lin Framework: a Linear Transformation Function

To apply the timescale switching techniques from low to high, the practitioner needs to consider the following ingredients: the first, second, third, and fourth, respectively, as below.

Ingredients for time scale switching techniques

- 1. Observe the LF observations, \mathbf{x}_{lf} , of the variable of interest. (N×1)
- 2. Design the temporal aggregation method that defines the relationship between HF values, $\mathbf{x}_{(c\times 1)}$, and its LF counterparts, \mathbf{x}_{lf} . Where c represents the number of HF data points in the sample period where the LF observations, \mathbf{x}_{lf} , occurred. Note that c > N since there are more HF than the LF data points in a given period.
- 3. Specify the indicator variables, $\underset{(m \times k)}{\mathbf{S}}$, where k is the number of indicator variables used, m is the number of HF observations in the period of interest such that $m \ge c$ and m c is the number of extrapolated HF data points.
- 4. Specify a hypothesized relationship model between $\mathbf{z}_{(m \times 1)}$ and $\mathbf{S}_{(m \times k)}$, where $\{\mathbf{z}_i\}_{i=1}^c := \mathbf{x}$ are the interpolated data points and $\{\mathbf{z}_i\}_{i=c+1}^m$ are the extrapolated data points.

The first and third ingredients are self-explanatory. The second ingredient is not usually specified for series when the frequency ratio of low- and high-frequency variables (periodicity of the flows) is not constant. There are various aggregation setup schemes based on the assumed relationship between the HF and LF observations of the variable of interest. Let \mathbf{C} be the aggregation matrix that transforms the unknown interpolated HF vector \mathbf{x} to the LF observed vector \mathbf{x}_{lf} such that $\mathbf{x}_{lf} := \mathbf{C}\mathbf{x}$. For example, if the observed LF vector, \mathbf{x}_{lf} , and the unknown interpolated HF vector, \mathbf{x} , are defined as:

$$\mathbf{x}_{lf} := (X_{f_0}, X_{f_1}, \cdots, X_{f_{N-1}})', \quad \text{where} \quad f_j = f_{j-1} + \kappa_j = 1 + \sum_{i=1}^j \kappa_i, \quad (1.1)$$

where $\{\kappa_i\}_{i=1}^N$ is the number of HF periods in each LF period *i* (i.e. the number of days in a month¹) and $f_0 := 1$.

$$\mathbf{x} := (X_1, X_2, \cdots, X_c)', \quad \text{where} \quad c := f_{N-1}$$

then, in this case, the aggregation matrix, $\mathbf{C_1}$, is:

$$\mathbf{C_1}_{(N \times c)} := \begin{bmatrix} 1^{(1)} & 0^{(2)} \cdots & 0 & \cdots & 0^{(f_{N-1})} \\ 0^{(1)} & \cdots & 1^{(f_1)} \cdots & 0 & \cdots & 0 \\ 0^{(1)} & 0^{(2)} \cdots & 1^{(f_2)} \cdots & 0 & \cdots & 0 \\ \vdots & \ddots & \ddots & \ddots & \vdots \\ 0^{(1)} & \cdots & 0 & \cdots & 1^{(f_{N-1})} \end{bmatrix}$$
(1.2)

Whereas, if the observed LF variable, \mathbf{x}_{lf} , is assumed as follows:

$$\{\mathbf{x}_{lf}\}_i := \frac{1}{\kappa_i} \sum_{j=f_{i-1}}^{f_i-1} X_j, \quad \text{where} \quad i = 1, \cdots, N$$
 (1.3)

¹For example, if i = 1 represents January 1999 this means that there are $\kappa_1 = 20$ working days in this month.

then, the aggregation matrix, C_2 , is:

$$\{\mathbf{C}_{2}\}_{i,j} := \kappa_{i}^{-1} \mathbb{1}_{(f_{i-1} \le j \le f_{i} - 1)}; \quad i = 1, \cdots, N \quad \text{and} \quad j = 1, \cdots, c.$$
(1.4)

where $\mathbb{1}_{(.)}$ is the indicator function.

The fourth ingredient depends on the practitioner's statistical understanding and assumptions of the series. CL presents the best linear unbiased estimate (BLUE) for interpolated as well as extrapolated HF observations of the LF variable based on the LF observations, the HF indicator variable(s), and the static relation model proposed (eq. 1.5). Poissonnier (2018) points out that high autocorrelation of the residuals limits the impact of the unexplained component on the HF profile. Therefore, they introduce a dynamic relationship model (eq. 1.6) in the CL framework and is assumed to be superior, in some circumstances², to the traditional static relationship model of Chow and Lin (1971).

The static model is defined as:

$$X_t = \sum_{j=1}^k \varphi_j s_{t,j} + e_t \tag{1.5}$$

The dynamic model is defined as:

$$X_{t} = \sum_{i=1}^{p} \rho_{i} X_{t-i} + \sum_{j=1}^{k} \varphi_{j} s_{t,j} + e_{t}$$
(1.6)

where X_t is the HF observation at time t, $s_{t,j}$ is the indicator variable, j, at time t, and k is the number of indicator variables. The stochastic disturbance terms in either

 $^{^{2}}$ For example, when the HF observations exhibit a temporal time dependency feature.

models³ may follow:

- An AR(1) process $e_t = \mu_{CL} e_{t-1} + \omega_t$ (Chow and Lin, 1971).
- A Random walk (RW) process $e_t = e_{t-1} + \omega_t$ (Fernández, 1981).
- An ARIMA(1,1,0) process $\Delta e_t = \mu_{litt} \Delta e_{t-1} + \omega_t$ (Litterman, 1983).

where ω_t _{t=1} is a series of white noise with mean zero and variance σ^2 .

Hence, we denote the six combination of models studied in this paper as follows:

1. CLCL model (Chow and Lin, 1971):

$$\begin{split} X_t &= \sum_{j=1}^k \varphi_j s_{t,j} + e_t \\ e_t &= \mu e_{t-1} + \omega_t \ ; \quad \omega_t \sim i.i.d. \ N(0,\sigma^2) \end{split}$$

2. CLL model (Chow and Lin, 1971; Litterman, 1983)

$$X_t = \sum_{j=1}^k \varphi_j s_{t,j} + e_t$$
$$\Delta e_t = \mu \Delta e_{t-1} + \omega_t \; ; \quad \omega_t \sim i.i.d. \; N(0, \sigma^2)$$

3. CLF model (Chow and Lin, 1971; Fernández, 1981)

$$X_t = \sum_{j=1}^k \varphi_j s_{t,j} + e_t$$

 $^{^{3}}$ We intentionally do not refer to these terms as error terms of the model in order to distinguish between them and the error that we will define later in the sequel.

$$e_t = e_{t-1} + \omega_t$$
; $\omega_t \sim i.i.d. \ N(0, \sigma^2)$

4. PCL model (Chow and Lin, 1971; Poissonnier, 2018)

$$X_t = \rho X_{t-i} + \sum_{j=1}^k \varphi_j s_{t,j} + e_t$$
$$e_t = \mu e_{t-1} + \omega_t \; ; \quad \omega_t \sim i.i.d. \; N(0, \sigma^2)$$

5. PL model (Litterman, 1983; Poissonnier, 2018)

$$X_t = \rho X_{t-i} + \sum_{j=1}^k \varphi_j s_{t,j} + e_t$$
$$\Delta e_t = \mu \Delta e_{t-1} + \omega_t \; ; \quad \omega_t \sim i.i.d. \; N(0, \sigma^2)$$

6. PF model (Fernández, 1981; Poissonnier, 2018)

$$\begin{split} X_t &= \rho X_{t-i} + \sum_{j=1}^k \varphi_j s_{t,j} + e_t \\ e_t &= e_{t-1} + \omega_t \ ; \ \ \omega_t \sim i.i.d. \ N(0,\sigma^2) \end{split}$$

where $|\mu| < 1$ and $|\rho| < 1$.

In what follows, we demonstrate how the HF estimates produced by the above models can be expressed as a linear transformation of the LF observations and the HF indicator series. In particular, Proposition 1.2.1 shows that the BLUE estimates of the unknown HF observations for models CLCL, CLL, and CF can be written in the form of a slope matrix pre-multiplied by the column vector of the observed LF data points.

Proposition 1.2.1 Given the LF vector \mathbf{x}_{lf} , the BLUE (interpolated and extrapolated) estimates of the unobserved HF vector, $\mathbf{z}_{(m \times 1)}$, of the static model (1.5) defined in matrix form, $\mathbf{z} := \mathbf{S}\varphi + \mathbf{e}$, can be found using the following linear transformation

$$\hat{\mathbf{z}} = f(\mathbf{x}_{lf}) = \underset{(m \times N)_{(N \times 1)}}{\mathbf{A}} \mathbf{x}_{lf}$$
(1.7)

• In the case of the CLCL model, $\mathbf{A} := \mathbf{A}_{CL}$

$$\mathbf{A}_{CL}_{(m \times N)} = \mathbf{A}_{CL,1} + \mathbf{A}_{CL,2}_{(m \times N)}$$

where

$$\begin{split} \mathbf{A}_{CL,1} &:= \mathbf{\Omega} \phi_{TCL} \\ \mathbf{A}_{CL,2} &:= \left(\mathbf{I}_m - \mathbf{A}_{CL,1} \mathbf{C}_z \right) \mathbf{S} \delta_{TCL} \\ \delta_{TCL} &:= \left(\mathbf{S}' \phi_{TCL} \mathbf{C}_z \mathbf{S} \right)^{-1} \mathbf{S}' \phi_{TCL} \\ \phi_{TCL} &:= \left(\mathbf{C}_z \left(\mathbf{C}_z \mathbf{\Omega} \mathbf{C}_z' \right)^{-1} \right) \\ \mathbf{C}_z &:= \mathbf{C}_z \mathbf{F} \\ {}_{(N \times m)} &:= {}_{(N \times c)(\mathbf{C} \times m)} \\ \end{split}$$

Where **F** is an identity matrix, **C** is the aggregation matrix (see eq. 1.2 and 1.4), and $\Omega_{(m \times m)}$ is the variance-covariance matrix of the stochastic disturbance

terms $\mathop{\mathbf{e}}_{(m \times 1)}$:

$$\Omega_{(m \times m)} := \frac{\sigma^2}{1 - \mu^2} \begin{pmatrix} 1 & \mu & \cdots & \mu^{m-1} \\ \mu & 1 & \cdots & \mu^{m-2} \\ \vdots & \vdots & \ddots & \vdots \\ \mu^{m-1} & \mu^{m-2} & \cdots & 1 \end{pmatrix}$$

 \bullet In the case of the CLL model, $\mathbf{A}:=\mathbf{A}_{\mathit{litt}}$

$$\mathbf{A}_{litt}_{(m \times N)} = \mathbf{A}_{litt,1} + \mathbf{A}_{litt,2}_{(m \times N)}$$

where

 $\mathbf{A}_{litt,1} := \mathbf{D}^{-1} \mathbf{\Omega} \phi_{litt}$

$$\mathbf{A}_{litt,2} := (\mathbf{I}_m - \mathbf{A}_{litt,1}\mathbf{C}_z) \, \mathbf{S}\delta_{litt}$$
$$\phi_{litt}_{(m \times N)} := \mathbf{D}^{'-1}\mathbf{C}_z^{'} \left(\mathbf{C}_z \left(\mathbf{D}^{'} \boldsymbol{\Omega}^{-1} \mathbf{D}\right)^{-1} \mathbf{C}_z^{'}\right)^{-1}$$
$$\delta_{litt}_{(k \times N)} := \left(\mathbf{S}^{'} \mathbf{D}^{'} \phi_{litt} \mathbf{C}_z \mathbf{S}\right)^{-1} \mathbf{S}^{'} \mathbf{D}^{'} \phi_{litt}$$

Where **C** is the aggregation matrix (see eq. 1.2 and 1.4), Ω is the variancecovariance matrix of Δe_t defined as above, and **D** is difference matrix:

$$\mathbf{D}_{(m \times m)} := \begin{pmatrix} 1 & 0 & \cdots & 0 \\ -1 & 1 & \cdots & 0 \\ 0 & -1 & \ddots & 0 \\ 0 & \cdots & \ddots & 1 \end{pmatrix}$$

• In the case of the CLF model:

 $\mathbf{A} := \mathbf{A}_{Fer}$ which is a special case of \mathbf{A}_{litt} where $\mathbf{\Omega} = \sigma^2 \mathbf{I}_m$.

Observe that the slope matrix, **A**, defined in equation (1.7) for each of the proposed methods, is invariant to the variance of stochastic disturbance terms, σ^2 . Furthermore, equation (1.7) presents a direct way of computing $\hat{\mathbf{z}}$ using the CL framework.

For the dynamic models, Poissonnier (2018) use partial recursive substitutions to obtain the last recorded LF observation and then, under some assumptions, formulate a constrained optimization problem to find these estimates in a closed-form expression. In particular, as illustrated in Lemma 1.2.1.1 below, the slope matrix, \mathbf{A}_{pois} , and the transition vector, \mathbf{b}_{pois} , of this linear temporal disaggregation defined in equation (1.8) are a function of Model (1.6) parameters, the underlying structure of the stochastic disturbance terms, and the HF indicator series.

Lemma 1.2.1.1 Given the LF vector, \mathbf{x}_{lf} , to generate interpolated estimates of the unobserved HF vector, $\mathbf{x}_{(\mathbf{C}\times 1)}$, from Model (1.6), Poissonnier (2018) suggests the following linear transformation

$$\hat{\mathbf{x}} = f(\mathbf{x}_{lf}) = \mathbf{A}_{pois} \mathbf{x}_{lf} + \mathbf{b}_{pois} \tag{1.8}$$

• In the case of the PCL model:

$$\mathbf{A}_{pois,1} = \mathbf{\Omega}_{\mathbf{x}} \mathbf{C}' (\mathbf{C} \mathbf{\Omega}_{\mathbf{x}} \mathbf{C}')^{-1}$$

(C×N)

$$\mathbf{b}_{pois,1} = (\mathbf{I}_c - \mathbf{A}_{pois}\mathbf{C})(\mathbf{I}_c - \mathbf{M}_1)^{-1}(\mathbf{M}_2\mathbf{x}^{init} + \mathbf{S}_{inter}\varphi)$$
(**C**×1)

 $\mathbf{r} \mathbf{r}' > -1$

where

$$\mathbf{M}_{\mathbf{x}} := (\mathbf{I}_{c} - \mathbf{M}_{1})^{-\mathbf{F}\mathbf{M}\mathbf{F}} (\mathbf{I}_{c} - \mathbf{M}_{1})^{-1}$$
$$\begin{bmatrix} 0 & 0 & \cdots & \cdots & 0 \\ \rho_{1} & 0 & \ddots & \ddots & \ddots & \vdots \\ \rho_{2} & \rho_{1} & \cdots & \ddots & \ddots & \vdots \\ \rho_{3} & \rho_{2} & \cdots & \ddots & \ddots & \vdots \\ \vdots & \ddots & \ddots & \ddots & \ddots & \vdots \\ 0 & \cdots & \rho_{p} & \cdots & \rho_{1} & 0 \end{bmatrix} \quad and \quad \mathbf{M}_{2} := \begin{bmatrix} \rho_{1} & \rho_{2} & \cdots & \rho_{p} \\ 0 & \rho_{1} & \cdots & \rho_{p-1} \\ \vdots & \ddots & \ddots & \vdots \\ \vdots & \ddots & \ddots & \rho_{1} \\ \vdots & \ddots & \ddots & \rho_{1} \\ \vdots & \ddots & \ddots & 0 \\ \vdots & \ddots & \ddots & \vdots \\ 0 & \cdots & 0 \end{bmatrix}$$

 $\mathbf{N} \mathbf{f} = 1 \mathbf{D} \mathbf{O} \mathbf{D}' \mathbf{T}$

 $\begin{array}{l} \Omega_{(m\times m)} \text{ is the variance-covariance matrix of } \underset{(m\times 1)}{\mathbf{e}} \text{ as defined in Proposition 1.2.1,} \\ \mathbf{F} \text{ is an identity matrix, } \mathbf{C} \text{ is the aggregation matrix (see (1.2)), } \mathbf{S}_{inter} \coloneqq \{\mathbf{S}_i\}_{i=1}^c \\ \text{where inter stands for interpolation, and } \mathbf{x}^{init} \text{ is the initial value of } \mathbf{x} \end{array}$

• In the case of the PL model:

$$\mathbf{A}_{pois,2} := \mathbf{\Omega}_{\mathbf{x},\mathbf{2}} \mathbf{C}' \mathbf{\Omega}_{\mathbf{x},\mathbf{1}}^{-1} \qquad and \qquad \mathbf{b}_{pois,2} = \lambda_{pois} \left(\mathbf{M}_2 \mathbf{x}^{init} + \mathbf{F} \mathbf{S}_{inter} \varphi \right)$$

where

$$\lambda_{pois} := (\mathbf{I}_c - \mathbf{M}_1)^{-1} - \boldsymbol{\Omega}_{\mathbf{x},\mathbf{2}} \mathbf{C}' \boldsymbol{\Omega}_{x,1}^{-1} \mathbf{C} (\mathbf{I}_c - \mathbf{M}_1)^{-1}$$
$$\boldsymbol{\Omega}_{\mathbf{x},\mathbf{1}} = \mathbf{C} \boldsymbol{\Omega}_{\mathbf{x},\mathbf{2}} \mathbf{C}' , \quad \boldsymbol{\Omega}_{\mathbf{x},\mathbf{2}} := (\mathbf{I}_c - \mathbf{M}_1)^{-1} \left(\mathbf{D}_1' \left(\mathbf{F} \boldsymbol{\Omega} \mathbf{F}' \right)^{-1} \mathbf{D}_1 \right)^{-1} (\mathbf{I}_c - \mathbf{M}_1')^{-1}$$

 \mathbf{D}_1 is the $(c \times c)$ difference matrix and $\mathbf{\Omega}$ is the variance covariance matrix Δe as defined in Proposition 1.2.1, \mathbf{M}_1 and \mathbf{M}_2 are defined as in the PCL case.

• In the case of the PF model:

The $(c \times N)$ matrix $\mathbf{A}_{pois,3}$ and the c vector $\mathbf{b}_{pois,3}$ are exactly as $\mathbf{A}_{pois,2}$ and $\mathbf{b}_{pois,2}$, respectively. The only difference is that, in this case, $\mathbf{\Omega} = \sigma^2 \mathbf{I}_m$.

Unlike the static model, the dynamic models were proposed and demonstrated for interpolation only and not extrapolation. Hence, the application of the linear transformation stated in Lemma 1.2.1.1 generates only interpolated HF data estimates of the LF variable. Proposition 1.2.2 generalizes this Lemma to produce extrapolated and interpolated HF data estimates of the LF variable.

Proposition 1.2.2 Given the LF vector, \mathbf{x}_{lf} , to generate the best linear unbiased interpolated as well as extrapolated HF estimates from Model (1.6), the linear transformation becomes:

$$\hat{\mathbf{z}} = f(\mathbf{x}_{lf}) = \mathbf{A}_{epois} \mathbf{x}_{lf} + \mathbf{b}_{epois} \tag{1.9}$$

where the subscript epois stands for extended Poissonnier (2018).

• In the case of the PCL model:

$$\mathbf{A}_{epois,1}_{(m \times N)} = \mathbf{\Omega}_{\mathbf{x},\mathbf{2}} \mathbf{C}_{z}' \mathbf{\Omega}_{\mathbf{x},\mathbf{1}}^{-1}$$

$$\mathbf{b}_{epois,1} = (\mathbf{I}_m - \mathbf{A}_{epois,1} \mathbf{C}_z) (\mathbf{I}_m - \mathbf{M}_{1,z})^{-1} (\mathbf{M}_{2,z} \mathbf{z}^{init} + \mathbf{S}\varphi)$$
_(m×1)

where

$$\Omega_{{f x},{f 1}} = {f C}_z \Omega_{{f x},{f 2}} {f C}'_z \;, \;\;\; \Omega_{{f x},{f 2}} := ({f I}_m - {f M}_{1,z})^{-1} \Omega ({f I}_m - {f M}'_{1,z})^{-1}$$

 \mathbf{C}_z is the N $\times\,m$ matrix, defined exactly in the same manner as the N $\times\,c$ matrix

C defined in either set-ups. $\Omega_{(m \times m)}$ is the variance-covariance matrix of $\mathbf{e}_{(m \times 1)}$.

• In the case of the PL model:

 $\mathbf{A}_{epois,2}$ and $\mathbf{b}_{epois,2}$ have the same structure as $\mathbf{A}_{epois,1}$ and $\mathbf{b}_{epois,1}$, respectively, with variance-covariance matrix $\mathbf{D}^{-1}\mathbf{\Omega}\mathbf{D}'^{-1}$ instead of $\mathbf{\Omega}$ in $\mathbf{\Omega}_{\mathbf{x},2}$.

• In the case of the PF model:

 $\mathbf{A}_{epois,3}$ and $\mathbf{b}_{epois,3}$ are exactly as $\mathbf{A}_{epois,2}$ and $\mathbf{b}_{epois,2}$ respectively, with $\mathbf{\Omega} = \sigma^2 \mathbf{I}_m$.

1.3 Theoretical Properties of the Linear Transformations

Firstly, observe that both linear transformations depicted in equations (1.7) and (1.9) generate interpolated as well as extrapolated estimates, $\hat{\mathbf{z}}$, of the unobserved HF variable, \mathbf{z} . Note that the slope coefficient or/and the transition vector in these transformations depend on unknown coefficients to be estimated. For example, \mathbf{A}_{CL} is a function of the unknown autoregressive parameter μ . We denote by $\tilde{\mathbf{z}}$ the feasible estimate of the unobserved HF vector \mathbf{z} . One needs to distinguish between hypothetical errors, feasible errors, and residuals. In this framework, a hypothetical error term is a difference between the unobserved HF variable \mathbf{z} and its theoretical estimate $\hat{\mathbf{z}}$, i.e. $u_z := \mathbf{z} - \hat{\mathbf{z}}$. A feasible error term is the difference between the unobserved HF variable \mathbf{z} interpolated estimate $\tilde{\mathbf{z}}$ i.e. $u_{\tilde{z}} := \mathbf{z} - \tilde{\mathbf{z}}$. Residuals, by definition, are the difference between an actual and estimated value. In this context,

actual values (observed values) only exist at a LF (i.e. \mathbf{x}_{lf}). Hence, to avoid confusion, we restrict the definition of residuals, in this framework, to the LF timescale.⁴

Secondly, the slope matrix \mathbf{A} and \mathbf{A}_{epois} in equations (1.7) and (1.9) are independent of the variance scale of the stochastic disturbance terms, σ^2 . Although this seems like an encouraging feature in an econometric method – i.e. no need to estimate the variance scale parameter–, this has a detrimental effect on the mean squared hypothetical error. The theoretical estimate $\hat{\mathbf{z}}$ may not capture (and hence mimic) the variation of the actual unobserved variable \mathbf{z} . For a given two sets of HF data points, \mathbf{z} , generated from the same pre-determined model with different variance scales each, the theoretical estimates of the data with a larger variance scale have a higher average squared hypothetical error. Hence, the CL method is less accurate for data with high variances, such as disaggregating monthly historical stock returns to a higher frequency.

Thirdly, the estimated HF data points using the models discussed earlier are the sum of two signals. Precisely, the estimation represented by equation (1.7) of the static models, can be rewritten as $\hat{\mathbf{z}} = \mathbf{A}_{\pi,1}\mathbf{x}_{lf} + \mathbf{A}_{\pi,2}\mathbf{x}_{lf}$ where π refers to the disaggregation method: CLCL, CLL, or CLF. $\mathbf{A}_{\pi,1}\mathbf{x}_{lf}$ and $\mathbf{A}_{\pi,2}\mathbf{x}_{lf}$ represent the trajectory and perturbation signal, respectively. Model parameters and the LF observations determine the trajectory signal. It passes through the LF data points similar to the naive interpolation of connecting LF observations. Because of the model parameters involved, these segments are bent. In the out-of-sample period, where there are no observed LF data points, the path of the trajectory signal will rely solely on the model's parameters. Hence, the estimation accuracy of the in-

 $^{{}^{4}}$ In Poissonnier (2018)'s paper, the residual is similar to the feasible error term in this paper.

sample dataset outperforms the relative out-of-sample estimation accuracy. The perturbation signal, independent of the trajectory, represents the fluctuations driven by the model's parameters and the observed series of the LF variable and the HF indicator variable(s). Therefore, if the indicator variable(s) variance is high, the perturbation signal fluctuations tend to be substantial.

The above discussions provide an insightful understanding of the disaggregation characteristics. However, in reality, a practitioner is often unaware of the defined model and parameters of the data generating process (DGP). Therefore, it is reasonable to ask whether assuming a different model from the DGP might lead to a substantial difference and if there is a particular model that out-stands its peers regardless of the DGP. To answer these questions, we first outline the estimation process in each of the mentioned models. Second, we generate artificial series from each of the discussed models with various parameters' values as DGPs to evaluate the in- and out-of-sample mean squared feasible errors (MSFE) in each case for each model.

1.4 Estimation

Once the DGP model is hypothesised, unknown parameters need to be estimated to find a feasible estimate of the HF points. Hence, one needs to distinguish between parameters: Parameters Estimated using data available (PEDA) and Parameters that require estimation (PRE). For example, using the CLCL model, Chow and Lin (1971) considered both φ_j 's in equation (1.5) and the AR(1) coefficient, μ , to be unknown parameters that need to be estimated. However, based on Proposition 1.2.1, only μ needs to be estimated using a likelihood function. In such a setting, the φ_j 's are PEDA, and μ is a PRE.

In the six mentioned models, most of the literature assumes that the error term, ω , in the model of the stochastic disturbance term, e_{hf} or Δe_{hf} , is white noise. However, many authors assume normality with zero mean and unknown variance, σ^2 , to formulate a likelihood function. In particular, $\mathbf{e}_{hf} \sim N(0, \mathbf{\Omega})$ where \mathbf{e}_{hf} represents the vector of HF error terms in the CLCL or PCL models, or $\mathbf{De}_{hf} \sim N(0, \mathbf{\Omega})$ in the other models. Therefore, the LF vector of error terms, \mathbf{e}_{lf} , has a similar structure. Having said this, Bournay and Laroque (1979) and Sax and Steiner (2013), defined the log-likelihood function as:

$$\mathscr{L}(\text{PRE}) := -\frac{N}{2}\ln(2\pi) - \frac{1}{2}\ln\left|\mathbf{C}_{\mathbf{z}}\boldsymbol{\Omega}^{*}\mathbf{C}_{\mathbf{z}}^{'}\right| - \frac{1}{2}\hat{\mathbf{e}}_{lf}^{'}(\mathbf{C}_{z}\boldsymbol{\Omega}^{*}\mathbf{C}_{z}^{'})^{-1}\hat{\mathbf{e}}_{lf}$$
(1.10)

where $\Omega^* := \Omega$ in the case of CLCL and PCL, or $\Omega^* := \left(\mathbf{D}'(\Omega)^{-1}\mathbf{D}\right)^{-1}$ in the other cases.

However, Poissonnier (2018), instead, maximizes the log-likelihood function of the HF error terms directly for their proposed dynamic model:

$$\mathscr{L}(\text{PRE}) := -\frac{m}{2}\ln(2\pi) - \frac{1}{2}\ln|\mathbf{\Omega}^*| - \frac{1}{2}\hat{\mathbf{e}}'_{hf}(\mathbf{\Omega}^*)^{-1}\hat{\mathbf{e}}_{hf}$$
(1.11)

This likelihood function, in fact, can be used for all the models in this paper including the static ones and it can be rewritten as:

$$\mathscr{L}(\text{PRE}) := -\frac{m}{2}\ln(2\pi) - \frac{1}{2}\ln|\mathbf{\Omega}^*| - \frac{1}{2}\hat{\mathbf{e}}_{lf}'(\mathbf{C}_z\mathbf{\Omega}^*\mathbf{C}_z')^{-1}\hat{\mathbf{e}}_{lf}$$
(1.12)

Hint: $\hat{\mathbf{e}}_{hf} = \mathbf{\Omega}^* \mathbf{C}'_z (\mathbf{C}_z \mathbf{\Omega}^* \mathbf{C}'_z)^{-1} \hat{\mathbf{e}}_{lf}$ and equivalently $\hat{\mathbf{e}}_{lf} = \mathbf{C} \hat{\mathbf{e}}_{hf}$.

Note that Ω and $\hat{\mathbf{e}}_{hf}$ are a function of the unknown parameters. In particular, for the static models, $\hat{\mathbf{e}}_{hf} = (\mathbf{A} - \mathbf{S}\delta)x_{lf}$ where the expressions of \mathbf{A} and δ depend on the model as outlined in Proposition 1.2.1 (see Appendix 1.A.1). Similarly, for the dynamic models, $\hat{\mathbf{e}}_{hf} = (I_m - \mathbf{M}_{1,z})(\mathbf{A}\mathbf{x}_{lf} + \mathbf{b}) - \mathbf{M}_{2,z}\mathbf{z}^{init} - \mathbf{S}\varphi$ where $\mathbf{M}_{1,z}$ and $\mathbf{M}_{2,z}$ are defined in Lemma 1.2.1.1, while the expressions of \mathbf{A} and \mathbf{b} depend on the specified model as outlined in Proposition 1.2.2 (see Appendix 1.A.3).

The log-likelihood can be used to estimate the PREs in the hypothesised model using maximum likelihood estimation. However, not all the PREs are required to find feasible HF estimates, $\tilde{\mathbf{z}}$. For example, in the CLCL model, there are two PREs, the autoregressive parameter, μ , and the variance σ^2 . The only parameter that needs to be estimated using the likelihood function is μ since the matrix \mathbf{A}_{CL} in equation (1.5) is a function of it, $\mathbf{A}_{CL} = \mathbf{A}_{CL}(\mu)$. Thus, the feasible HF estimate $\tilde{\mathbf{z}}$ is computed by premultiplying the LF vector \mathbf{x}_{lf} by the estimated transformation matrix $\hat{\mathbf{A}}_{CL} = \mathbf{A}_{CL}(\hat{\mu})$ as suggested by Proposition 1.2.1. However, in the case of the CLF model, the slope transformation matrix does not depend on the PREs, so there is no need to estimate the likelihood function.

1.5 Simulation Study

We implement a Monte Carlo simulation to examine the performance of the models covered in this paper. The simulation compares the models' estimation of the "unobserved" HF daily series using the observed LF monthly series. To this end, firstly, we choose two consecutive calendar years to determine the number of working days in each calendar month; this leads to around 500 HF observations. Secondly, in every simulation, we use the same simulated daily indicator series from an AR(1) process for all the DGPs, for comparison purposes. Thirdly, we generate the error term, ω_t , from a standard normal distribution with zero mean and variance of σ^2 . Fourthly, the stochastic disturbance term, e_t , and the HF series, \mathbf{x} , are generated according to the specified model and parameters (μ , ρ , φ) as illustrated in Table (1.1). Fifthly, the LF series, $\mathbf{x}_{lf} := \mathbf{C}\mathbf{x}$, are computed using the specified aggregation matrix (eq. 1.2 or 1.4). Finally, using the LF series and the indicator series, we estimate the HF series using each of the six models as illustrated in sections 1.2 and 1.4. We repeat the process 1000 times to obtain 1000 artificial samples for each DGP. We use the six models for each simulated sample to estimate the HF series given the simulated LF series and the HF indicator. We estimate the models with the following constraints on the parameters: $0 < \mu < 1$, $0 < \rho < 1$, and $0 < \varphi < 1$.

Table 1.1: DGP Parameters

	σ^2	μ	φ	ρ	Total # DGPs
CLCL	0.1, 1, 5, or 10	0.2 or 0.8	0.2 or 0.8		16
\mathbf{CLL}	0.1, 1, 5, or 10	0.2 or 0.8	0.2 or 0.8		16
\mathbf{CLF}	0.1, 1, 5, or 10		0.2 or 0.8		8
\mathbf{PCL}	0.1, 1, 5, or 10	0.2 or 0.8	0.2 or 0.8	0.2 or 0.8	32
\mathbf{PL}	0.1, 1, 5, or 10	0.2 or 0.8	0.2 or 0.8	0.2 or 0.8	32
\mathbf{PF}	0.1, 1, 5, or 10		0.2 or 0.8	$0.2~{\rm or}~0.8$	16

Many studies are focused on evaluating and comparing models in estimation and forecasting; see, for example, Poon and Granger (2003) for an extensive review. Most apply a loss function, where model-based predictions of the series of interest are compared to its estimates. We compare the estimated HF series with the true simulated one to evaluate the models. We use the Mean Squared Error (MSE) as a model evaluation method (Andersen, Bollerslev, and Lange, 1999):

$$MSE = \frac{1}{N} \Sigma_{t=1}^{N} [Xt - \hat{X}_t]^2$$

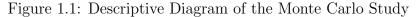
where N is the total number of observations in a given series.

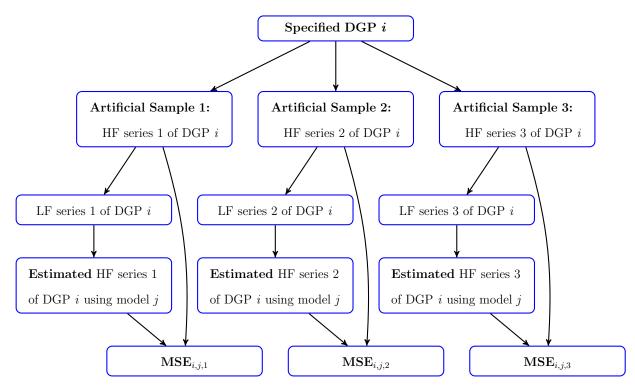
After computing the MSE for each simulation, we report for every model of estimation the average of the MSEs using the 1000 simulations of each DGP. As illustrated in figure 1.1, we compute the $MSE_{i,j,sim}$ for simulations sim = 1, ..., 1000, then we get the average mean squared error (AMSE) over the simulations:

$$AMSE_{i,j} = \frac{1}{1000} \Sigma_{sim=1}^{1000} MSE_{i,j,sim}$$

where i and j refer to the DGP and model of estimation respectively. We also consider the mean absolute error as a robustness check and found no significant differences in the findings. Hence, we only report the results using the MSE as a loss function for conciseness.

Table 1.1 summarises the DGPs (models and parameters) analysed in the simulation study. We include a range of magnitudes for each parameter to unveil their consequences on the estimated data points, ceteris paribus. In particular, it helps the practitioners understand what features they need to check before choosing the "right" model for a particular exercise. For the AR parameters, μ and ρ , we consider low (0.2) and high (0.8) values. The range of ρ allows us to study whether the dynamic models are better than their static counterparts if an autoregressive relationship exists in the HF series. At the same time, the range of μ helps us examine the influence of its magnitude on the overall AMSE and the choice of the "best" model. We also consider a high correlation, $\varphi = 0.8$, as well as a low correlation, $\varphi = 0.2$, between the "unobserved" HF series and the indicator variable. We can thus examine the importance of choosing the "right" indicator variable. Finally, we consider different magnitudes of the variance, σ^2 , to determine if different models are preferred based on the magnitude of variation of the series of interest. Finally, we repeat the study twice, once using the aggregation matrix \mathbf{C}_1 (eq. 1.2) and another using \mathbf{C}_2 (eq. 1.4).

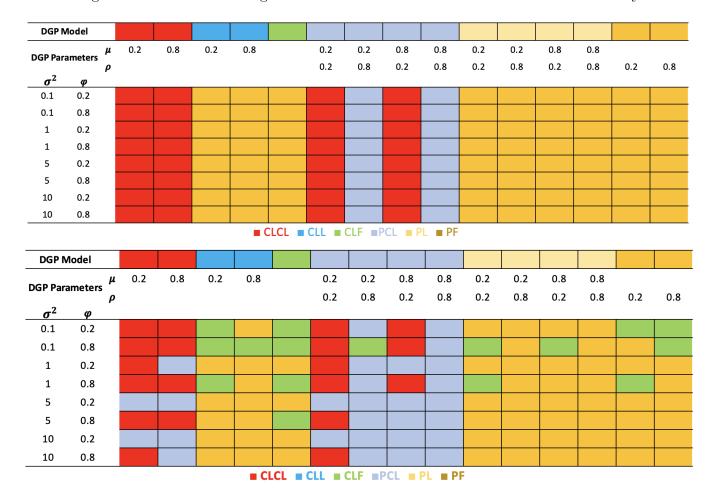




For each DGP, i, we generate 1000 artificial samples. Each artificial sample consists of a HF and its associated LF series. We use the LF series along with the indicator variable to estimate the HF series using each of the six models.

The Monte Carlo study reveals essential features in the performance of the models. We divide the analysis of the findings into two. In the first part of the analysis, we comment on the influence of the chosen DGP model and parameters on the AMSE. The second part examines the best-performing model under various DGPs.

We report the AMSE and its standard deviation using the 1000 simulations from each model, j, in each DGP, i, in Appendix 1.B for the interpolation as well as extrapolation sample. Our findings can be summarized as follows. Firstly, while the AMSE of extrapolation is higher than that of interpolation, there is no significant difference in the ranking of models between interpolation and extrapolation. Secondly, we note that the performance of the models is sensitive to some parameters of the underlying DGP and the relationship between the HF observations and its corresponding LF series (i.e. the aggregation matrix). Thirdly, the AMSE is overall lower when the DGP is CLCL or PCL. In other words, temporal disaggregation is generally more accurate for series where the e_t is an AR(1) rather than ARIMA(1,1,0) or RW. Fourthly, better estimates can be found for series with a defined aggregation matrix as in eq. (1.4) (i.e. when the LF observation represents the average of the HF values within the period) than eq. (1.2) (i.e. when the LF observation is one particular observed datapoint of the HF values). Fifthly, temporal disaggregation is more accurate when the defined DGP is static compared to dynamic. Finally, we do not find a significant impact of the relevancy of the indicator variable on the AMSE.



The tables represent the best estimating model in each DGP of the GMC. The upper table represents the results for aggregation matrix one, and the lower represents the results for aggregation matrix two. The colour of the cells in the True model row represents the model used to generate the true HF series. The model and parameters determine the DGP is. Hence, each cell in the inner triangle represents one case of the 120 DGPs considered in the MC. The colour of the cell represents the estimating model with the lowest average MSE over the 1000 simulations for a particular DGP.

Figure 1.2 presents the results of the best model with the lowest AMSE for each DGP considered. In the case of the first aggregation matrix (eq. 1.2), the findings are consistent regardless of the magnitudes of σ^2 and φ . In particular, the CLCL model is always dominant when the DGP is CLCL or PCL with low ρ . In contrast, the PCL model is only dominant (by a small margin) when the DGP is PCL with high ρ . In other words, the CLCL model is preferred when the stochastic disturbance terms, e_t , follow an AR(1) process, except when the model is dynamic with a high autoregressive parameter, ρ . The PF model is dominant in all cases where the stochastic disturbance terms, e_t is superior to the ARIMA(1,1,0) or RW. In particular, the RW specification for the e_t is superior to the ARIMA(1,1,0) in static and dynamic estimated models under all DGPs.

There are a couple of instances where the results are different when the aggregation matrix is C_2 instead of C_1 . For example, the PCL model moderately dominates, with few exceptions, when the DGP is CLCL or PCL with high variance, σ^2 . It means that under the second aggregation matrix, the PCL model performs better when the stochastic disturbance terms follow an AR process with high variance. Further, the CLF ranks moderately better in a few instances, especially when the DGP is CLL, CLF, or PL (with low ρ) with high φ and low variance, σ^2 . That is to say, using the second aggregation matrix, the CLF may outperform under three conditions: a relevant indicator is added, the stochastic disturbance terms follow an ARIMA(1,1,0) or RW process with low variance, and the autoregressive parameter is low under the dynamic model.

Therefore, based on the findings of the Monte Carlo, our recommendation is as follows. When the stochastic disturbance terms are assumed to follow an AR(1)

process, the static model (CLCL) is the best model to use unless the AR parameter of the unobserved HF values is assumed to be high. If the AR parameter is low, the relationship seems to be captured by the AR(1) process of the stochastic disturbance terms. On the other hand, when the stochastic disturbance terms follow an ARIMA(1,1,0) or RW, the dynamic model by Poissonnier (2018) would be more suitable since it depicts the autocorrelation component in the HF series. Specifically, the RW assumption is recommended in this case because although it yields only moderate improvement over its counterparts, it also has fewer parameters to estimate. Further, as noted by Poissonnier (2018), using the RW specification minimizes the residual variations from period to period and hence reduces the impact of the unexplained component of the model on the HF estimates.

1.6 Conclusion

This essay revisits temporal disaggregation models under the Chow-Lin framework: the static (Chow and Lin, 1971) and dynamic (Poissonnier, 2018) models, each with three different variants of the stochastic disturbance terms process. We show that the HF estimates found by these models are a linear transformation of their corresponding LF series. We also illustrate the necessary modification for a non-constant frequency ratio such as monthly to daily series. Further, we present the extended model of Poissonnier (2018) to estimate extrapolated and interpolated HF series.

We discuss the theoretical features of the temporal disaggregation methods. Specifically, we show that the HF estimates consist of a trajectory and a perturbation signal. We examine the assumptions and parameters that influence each signal. It gives a closer look to the practitioner on how and which assumptions and parameters drive the estimated values. We also illustrate the log-likelihood estimation and clarify which parameters need to be estimated.

Finally, we perform a global Monte Carlo study whereby we analyse the performance of the models under various combinations of different representative magnitudes of the parameters. The study examines which model provides the highest accuracy in the HF estimates under various data generating processes. We find that the preferred model generally depends on the process of the stochastic disturbance terms in the DGP rather than whether the true model is static or dynamic. The traditional static model by Chow and Lin (1971) dominates when the DGP error term follows an AR(1) process. In contrast, the dynamic model of Poissonnier (2018) with RW error term process (Fernández, 1981) is preferred when the DGP error terms follow an ARIMA(1,1,0) or RW process.

Appendices

1.A Proofs

1.A.1 Proof of Proposition **1.2.1**

Observe that, in matrix form, the static model is defined by: $\mathbf{x}_{lf} = \mathbf{C}\mathbf{x}$, and the true high-frequency observation \mathbf{z} is given by the model: $\mathbf{z} = \mathbf{S}\varphi + \mathbf{e}$. \mathbf{S} is the $(m \times k)$ indicator matrix that includes both interpolated data-points and extrapolated datapoints (if m > c). \mathbf{C} is the $(N \times c)$ aggregation matrix defined in either (1.2) or (1.4). \mathbf{e} is a $m \times 1$ vector of stochastic disturbances terms (both interpolated and extrapolated) and is defined by $\mathbf{e} := \begin{pmatrix} e_1 & \cdots & e_m \end{pmatrix}$.

Note: $\mathbf{z} = \mathbf{x}$ if m = c, hence, when $m \neq c$, $\mathbf{x}_{lf} = \mathbf{C}\mathbf{x} = \mathbf{C}\mathbf{F}\mathbf{z}$ where \mathbf{F} is a $c \times m$ non-square identity matrix defined as $\mathbf{F} = \mathbf{I}_{c \times m}$.

The proof consists of three parts for: CLCL, CLL, and CLF.

• Let the estimated high-frequency observation of \mathbf{z} be denoted by $\hat{\mathbf{z}}$ where $\hat{\mathbf{z}} = \mathbf{A}\mathbf{x}_{lf} = \mathbf{A}\mathbf{CF}\mathbf{x} = \mathbf{A}\mathbf{CF}\mathbf{S}\varphi + \mathbf{A}\mathbf{CF}\mathbf{e}$. Define the error of the high-frequency estimation by

$$egin{aligned} \mathbf{u}_z &= \mathbf{z} - \mathbf{\hat{z}} \ &= \mathbf{S}arphi - \mathbf{ACFS}arphi + \mathbf{e} - \mathbf{ACFe} \end{aligned}$$

For $\hat{\mathbf{z}}$ to be an unbiased estimator of \mathbf{z} in the case of Chow and Lin (1971), $\mathbb{E}(\mathbf{u}_z) = 0$. Hence, $\mathbf{S} = \mathbf{ACFS}$ and $var(\mathbf{u}_z) = (\mathbf{I}_m - \mathbf{ACF}) \mathbf{\Omega} (\mathbf{I}_m - \mathbf{ACF})'$ where $\mathbf{\Omega}^e$ is the $m \times m$ variance covariance matrix of \mathbf{e} .

Our objective is to minimize the var (\mathbf{u}_z) s.t. $\mathbf{S} = \mathbf{ACFS}$ hence, the Lagrangian expression is given by

$$\mathscr{L} = \frac{1}{2} tr(var(\mathbf{u}_z)) - tr(M'(\mathbf{ACFS} - \mathbf{S}))$$

where M is a $m \times k$ matrix of Lagrange multipliers.

$$\frac{\partial \mathscr{L}}{\partial \mathbf{A}} = (\mathbf{A}\mathbf{C}\mathbf{F} - \mathbf{I}_m)\mathbf{\Omega}\mathbf{F}'\mathbf{C}' - \mathbf{M}\mathbf{S}'\mathbf{F}'\mathbf{C}' = 0$$

Hence,

$$\mathbf{A} = \mathbf{\Omega} \mathbf{F}' \mathbf{C}' (\mathbf{C} \mathbf{F} \mathbf{\Omega} \mathbf{F}' \mathbf{C}')^{-1} + \mathbf{M} \mathbf{S}' \mathbf{F}' \mathbf{C}' (\mathbf{C} \mathbf{F} \mathbf{\Omega} \mathbf{F}' \mathbf{C}')^{-1}$$
(1.13)

since $\mathbf{S} = \mathbf{ACFS}$, then

$$\mathbf{M} = \left(\mathbf{S} - \mathbf{\Omega}\mathbf{F}'\mathbf{C}'\left(\mathbf{C}\mathbf{F}\mathbf{\Omega}\mathbf{F}'\mathbf{C}'\right)^{-1}\mathbf{C}\mathbf{F}\mathbf{S}\right)\left(\mathbf{S}'\mathbf{F}'\mathbf{C}'\left(\mathbf{C}\mathbf{F}\mathbf{\Omega}\mathbf{F}'\mathbf{C}'\right)^{-1}\mathbf{C}\mathbf{F}\mathbf{S}\right)^{-1}$$

Replace the matrix M in equation (1.13), then denote $\phi_{TCL} = \mathbf{C}' \left(\mathbf{CF} \mathbf{\Omega} \mathbf{F}' \mathbf{C}' \right)^{-1}$ and $\delta_{TCL} := \left(\mathbf{S}' \mathbf{F}' \phi_{TCL} \mathbf{CFS} \right)^{-1} \mathbf{S}' \mathbf{F}' \phi_{TCL}$ then one can show that

$$\mathbf{A} = \mathbf{\Omega} \mathbf{F}' \phi_{TCL} + (\mathbf{I}_m - \mathbf{\Omega} \mathbf{F}' \phi_{TCL} \mathbf{C} \mathbf{F}) \mathbf{S} \delta_{TCL}$$
$$= \mathbf{\Omega} \mathbf{F}' \phi_{TCL} (\mathbf{I}_N - \mathbf{C} \mathbf{F} \mathbf{S} \delta_{TCL}) + \mathbf{S} \delta_{TCL}$$

Note that the GLS estimator of φ :

$$\begin{split} \hat{\varphi} &= (\mathbf{S}'\mathbf{C}'(\mathbf{C}\mathbf{\Omega}\mathbf{C}')^{-1}\mathbf{C}\mathbf{S})^{-1}\mathbf{S}'\mathbf{C}(\mathbf{C}\mathbf{\Omega}\mathbf{C}')^{-1}\mathbf{C}\mathbf{x} \\ &= (\mathbf{S}'\phi_{TCL}\mathbf{C}\mathbf{S})^{-1}\mathbf{S}'\phi_{TCL}\mathbf{x}_{\mathbf{l}\mathbf{f}} \\ &= \delta_{TCL}\mathbf{x}_{\mathbf{l}\mathbf{f}} \end{split}$$

Hint: $\mathbf{Cx} = (\mathbf{CS})\varphi + \mathbf{Ce}.$

Hence,

$$\mathbf{\hat{e}}_{lf} = \mathbf{x_{lf}} - \mathbf{CS}\delta_{TCL}\mathbf{x_{lf}}$$

Hint: $\hat{\mathbf{e}}_{hf} = \hat{\mathbf{z}} - \mathbf{S}\hat{\varphi}$

• Using similar notation, it is assume that the stochastic disturbances terms follow an AR(1) process or an ARIMA(1,1,0) process. Let Δ be the $N \times N$ difference matrix, defined by

$$\boldsymbol{\Delta} := \begin{pmatrix} 1 & 0 & \cdots & 0 \\ -1 & 1 & \cdots & 0 \\ 0 & -1 & \ddots & 0 \\ 0 & \cdots & \ddots & 1 \end{pmatrix}$$
(1.14)

Furthermore, let **D** be also a $m \times m$ difference matrix similar in structure to Δ . Let **Q** be $N \times c$ matrix such that **QFD** := ΔCF , then

$$\mathbf{\Delta x}_{lf} = \mathbf{\Delta Cx} = \mathbf{\Delta CFz} = \mathbf{QFDz}$$

The difference series $D\hat{z}$ and Dz can be presented by

$$\mathbf{D}\hat{\mathbf{z}} = \mathbf{R}\mathbf{\Delta}\mathbf{x}_{lf} = \mathbf{R}\mathbf{Q}\mathbf{F}\mathbf{D}\mathbf{z} = \mathbf{R}\mathbf{Q}\mathbf{F}(\mathbf{D}\mathbf{S}\varphi + \mathbf{D}\mathbf{e})$$
 and $\mathbf{D}\mathbf{z} = \mathbf{D}\mathbf{S}\varphi + \mathbf{D}\mathbf{e}$

the error in estimation is given by

$$u_z = \mathbf{D}\mathbf{z} - \mathbf{D}\mathbf{\hat{z}} = (\mathbf{I}_m - \mathbf{R}\mathbf{Q}\mathbf{F})\mathbf{D}\mathbf{S}\varphi + (\mathbf{I}_m - \mathbf{R}\mathbf{Q}\mathbf{F})\mathbf{D}\mathbf{e}$$

imposing $E(u_z) = 0$ for unbiasedness of estimator, we get **RQFDS** = **DS**. Furthermore,

$$var(u_z) = (\mathbf{I}_m - \mathbf{RQF}) \Omega \left(\mathbf{I}_m - \mathbf{F'Q'R'} \right)$$

where here $\mathbf{\Omega} := \mathbb{E}\left(\mathbf{Dee'D}\right)$, hence almost the variance covariance matrix of the $(m-1) \times 1$ vector of white noise $\mathbf{\Omega}$. However, this is not the case, this is due to the fact that \mathbf{D} is square matrix with an element 1 in the upper left corner.

It is assumed that $e_0 = 0$, if the sample size is not too small, this assumption should not constitute a serious problem. Hence, in case of Fernández (1981), $\Omega = \sigma^2 I_m$. In case of Litterman (1983),

$$\mathbf{\Omega} := \frac{\sigma^2}{1 - \mu^2} \begin{pmatrix} 1 & \mu & \cdots & \mu^{m-1} \\ \mu & 1 & \cdots & \mu^{m-2} \\ \vdots & \vdots & \ddots & \vdots \\ \mu^{m-1} & \mu^{m-2} & \cdots & 1 \end{pmatrix}$$

Our objective is to minimize the $var(u_z)$ s.t. $\mathbf{RQFDS} = \mathbf{DS}$ hence, the Lagrangian expression is given by

$$\mathscr{L} = \frac{1}{2} tr(var(u_z)) - tr(M'(\mathbf{RQFDS} - \mathbf{DS}))$$

where M is a $m \times k$ matrix of Lagrange multipliers

$$\frac{\partial \mathscr{L}}{\partial R} = (\mathbf{R}\mathbf{Q}\mathbf{F} - \mathbf{I}_m)\mathbf{\Omega}\mathbf{F}^{'}\mathbf{Q}^{'} - \mathbf{M}\mathbf{S}^{'}\mathbf{D}^{'}\mathbf{F}^{'}\mathbf{Q}^{'} = 0$$

Hence,

$$\mathbf{R} = \mathbf{\Omega} \mathbf{F}' \mathbf{Q}' (\mathbf{Q} \mathbf{F} \mathbf{\Omega} \mathbf{F}' \mathbf{Q}')^{-1} + \mathbf{M} \mathbf{S}' \mathbf{D}' \mathbf{F}' \mathbf{Q}' (\mathbf{Q} \mathbf{F} \mathbf{\Omega} \mathbf{F}' \mathbf{Q}')^{-1}$$
(1.15)

replacing \mathbf{R} in $\mathbf{RQFDS} = \mathbf{DS}$, we get the Lagrange matrix as

$$\mathbf{M} = (I_m - \mathbf{\Omega} \mathbf{F}' \mathbf{Q}' (\mathbf{Q} \mathbf{F} \mathbf{\Omega} \mathbf{F}' \mathbf{Q}')^{-1} \mathbf{Q} \mathbf{F}) \mathbf{D} \mathbf{S} (\mathbf{S}' \mathbf{D}' \mathbf{F}' \mathbf{Q}' (\mathbf{Q} \mathbf{F} \mathbf{\Omega} \mathbf{F}' \mathbf{Q}')^{-1} \mathbf{Q} \mathbf{F} \mathbf{D} \mathbf{S})^{-1}$$

Replacing M in equation (1.15), we get

$$\begin{split} \mathbf{R} &= \mathbf{\Omega} \mathbf{F}^{'} \mathbf{Q}^{'} \left(\mathbf{Q} \mathbf{F} \mathbf{\Omega} \mathbf{F}^{'} \mathbf{Q}^{'} \right)^{-1} + \left(\mathbf{I}_{m} - \mathbf{\Omega} \mathbf{F}^{'} \mathbf{Q}^{'} \left(\mathbf{Q} \mathbf{F} \mathbf{\Omega} \mathbf{F}^{'} \mathbf{Q}^{'} \right)^{-1} \mathbf{Q} \mathbf{F} \right) \mathbf{D} \mathbf{S} \\ & \left(\mathbf{S}^{'} \mathbf{D}^{'} \mathbf{F}^{'} \mathbf{Q}^{'} \left(\mathbf{Q} \mathbf{F} \mathbf{\Omega} \mathbf{F}^{'} \mathbf{Q}^{'} \right)^{-1} \mathbf{Q} \mathbf{F} \mathbf{D} \mathbf{S} \right)^{-1} \mathbf{S}^{'} \mathbf{D}^{'} \mathbf{F}^{'} \mathbf{Q}^{'} \left(\mathbf{Q} \mathbf{F} \mathbf{\Omega} \mathbf{F}^{'} \mathbf{Q}^{'} \right)^{-1} \end{split}$$

Let

$$\begin{split} \phi_1 &:= \mathbf{F}' \mathbf{Q}' (\mathbf{Q} \mathbf{F} \mathbf{\Omega}^{\omega} \mathbf{F}' \mathbf{Q}')^{-1} \\ &= \mathbf{D}'^{-1} \mathbf{F}' \mathbf{C}' \mathbf{\Delta}' (\mathbf{\Delta} \mathbf{C} \mathbf{F} \mathbf{D}^{-1} \mathbf{\Omega}^{\omega} \mathbf{D}'^{-1} \mathbf{F}' \mathbf{C}' \mathbf{\Delta}')^{-1} \\ &= \mathbf{D}'^{-1} \mathbf{F}' \mathbf{C}' (\mathbf{C} \mathbf{F} (\mathbf{D}' (\mathbf{\Omega}^{\omega})^{-1} \mathbf{D})^{-1} \mathbf{F}' \mathbf{C}')^{-1} \mathbf{\Delta}^{-1} \end{split}$$

Observe the fact that $\mathbf{QF} := \mathbf{\Delta CFD}^{-1}$ and define the $m \times N$ matrix

$$\phi := \mathbf{D}^{'-1} \mathbf{F}^{'} \mathbf{C}^{'} \left(\mathbf{C} \mathbf{F} \left(\mathbf{D}^{'} \left(\mathbf{\Omega}^{\omega} \right)^{-1} \mathbf{D} \right)^{-1} \mathbf{F}^{'} \mathbf{C}^{'} \right)^{-1}$$

Observe that $\phi_1 \mathbf{QF} = \phi \Delta^{-1} \Delta \mathbf{CFD}^{-1} = \phi \mathbf{CFD}^{-1}$. Furthermore, define

$$\begin{split} \delta_1 &:= (\mathbf{S}'\mathbf{D}'\phi_1\mathbf{QFDS})^{-1}\mathbf{S}'\mathbf{D}'\phi_1 \\ &= (\mathbf{S}'\mathbf{D}'\phi\mathbf{CFD}^{-1}\mathbf{DS})^{-1}\mathbf{S}'\mathbf{D}'\phi\mathbf{\Delta}^{-1} \end{split}$$

Let $\delta := \left(\mathbf{S}' \mathbf{D}' \phi \mathbf{CF} \mathbf{D}^{-1} \mathbf{DS} \right)^{-1} \mathbf{S}' \mathbf{D}' \phi$. Then,

$$\begin{split} \mathbf{R} &= \mathbf{\Omega}\phi_1 + (\mathbf{I}_m - \mathbf{\Omega}\phi_1 \mathbf{QF})\mathbf{DS}(\mathbf{S}'\mathbf{D}'\phi_1 \mathbf{QFDS})^{-1}\mathbf{S}'\mathbf{D}'\phi_1 \\ &= \mathbf{\Omega}\phi_1 + (\mathbf{I}_m - \mathbf{\Omega}\phi_1 \mathbf{QF})\mathbf{DS}\delta_1 \\ &= \mathbf{\Omega}\phi + (\mathbf{I}_m - \mathbf{\Omega}\phi\mathbf{CFD}^{-1})\mathbf{DS}\delta\mathbf{\Delta}^{-1} \end{split}$$

Recall that $\mathbf{D}\hat{\mathbf{z}} = \mathbf{R}\Delta\mathbf{x}_{lf}$, hence

$$\hat{\mathbf{z}} = \mathbf{D}^{-1} \mathbf{R} \Delta \mathbf{x}_{lf}$$

$$= \mathbf{A}_{litt} \mathbf{x}_{lf}$$

where

$$\begin{split} \mathbf{A}_{litt} &:= \mathbf{D}^{-1} \mathbf{R} \Delta \\ &= \mathbf{D}^{-1} (\mathbf{\Omega}^{\omega} \phi + (\mathbf{I}_m - \mathbf{\Omega}^{\omega} \phi \mathbf{C} \mathbf{F} \mathbf{D}^{-1}) \mathbf{D} \mathbf{S} \delta) \\ &= \mathbf{D}^{-1} (\mathbf{\Omega}^{\omega} \phi (\mathbf{I}_N - \mathbf{C} \mathbf{F} \mathbf{S} \delta) + \mathbf{D} \mathbf{S} \delta) \\ &= \mathbf{D}^{-1} \mathbf{\Omega}^{\omega} \phi (\mathbf{I}_N - \mathbf{C} \mathbf{F} \mathbf{S} \delta) + \mathbf{S} \delta \end{split}$$

Also, $\hat{\mathbf{e}}_{lf} = \mathbf{x}_{lf} - \mathbf{CS}' \delta \mathbf{x}_{lf}$

1.A.2 Proof of Lemma 1.2.1.1

The proof consists of three parts for: PCL, PL, and PF. The proof provided here is different than the one stated in Poissonnier (2018)—however, the same result.

• The matrix representation of the proposed equation (1.6) is given by

$$\mathbf{x} = \mathbf{M}_1 \mathbf{x} + \mathbf{M}_2 \mathbf{x}^{init} + \mathbf{FS}\varphi + \mathbf{Fe}$$

where the $\mathbf{C} \times c$ and the $\mathbf{C} \times p$ matrices \mathbf{M}_1 and \mathbf{M}_2 are defined in the lemma. With some matrix algebra, one can reach

$$\mathbf{x} = (\mathbf{I}_c - \mathbf{M}_1)^{-1} (\mathbf{M}_2 \mathbf{x}^{init} + \mathbf{F} \mathbf{S} \varphi + \mathbf{F} \mathbf{e})$$
(1.16)

Observe that its interpolated estimate of \mathbf{x} , $\hat{\mathbf{x}}$, is given by the suggested linear transformation⁵

$$\hat{\mathbf{x}} = f(\mathbf{x}_{lf}) = \mathbf{A}\mathbf{x}_{lf} + \mathbf{b} = \mathbf{A}\mathbf{C}x + \mathbf{b}$$

the residual is given by

$$\mathbf{\hat{u}}_x = \mathbf{x} - \mathbf{\hat{x}} = (\mathbf{I}_c - \mathbf{AC})\mathbf{x} - \mathbf{b}$$

Observe that the matrix A and the vector b should satisfy the following three properties: 1) When you pre-multiply the interpolated estimate vector \hat{x} , by the conversion matrix C, you get the recorded observation vector x_{lf} , that is: $\mathbf{C}\hat{x} = x_{lf}$, 2) The estimate, \hat{x} , needs to be unbiased $\mathbb{E}(\hat{\mathbf{u}}_x) = 0$. 3) The estimate, \hat{x} , needs to be efficient subject to the other initial two constraints i.e.

$$\min_{\mathbf{A},\mathbf{b}} cov(\hat{\mathbf{u}}_x) \quad \text{s.t.} \quad \mathbf{C}\hat{\mathbf{x}} = \mathbf{x}_{lf} \quad \text{and} \quad \mathbb{E}(\hat{\mathbf{u}}_x) = 0$$

1. Solving the first constraint, we get

$$\mathbf{C}\hat{\mathbf{x}}_{inter} = \mathbf{C}\mathbf{A}\mathbf{x}_{lf} + \mathbf{C}\mathbf{b} = \mathbf{x}_{lf}$$

hence

$$\mathbf{C}\mathbf{A} = \mathbf{I}_N$$
 and $\mathbf{C}\mathbf{b} = \mathbf{0}_N$

⁵We chose to write it here as $\hat{\mathbf{x}}$ rather than $\hat{\mathbf{z}}$ purposely to distinguish between interpolated and extrapolated estimates.

2. Solving the second constraint, we get

$$\mathbf{b} = (\mathbf{I}_c - \mathbf{AC}) \mathbb{E}(\mathbf{x})$$

3. Define the variance covariance matrix of the vector \mathbf{x} , by the $(m \times m)$ matrix $\mathbf{\Omega}_x := Var(\mathbf{x}) = (\mathbf{I}_c - \mathbf{M}_1)^{-1} \mathbf{F} \mathbf{\Omega}^e \mathbf{F}' (\mathbf{I}_c - \mathbf{M}_1')^{-1}$. Furthermore, Observe that

$$cov(\hat{\mathbf{u}}_x) = (\mathbf{I}_c - \mathbf{AC})cov(x)(\mathbf{I}_c - \mathbf{AC})'$$
$$= (\mathbf{I}_c - \mathbf{AC})(\mathbf{I}_c - \mathbf{M}_1)^{-1}\mathbf{F}\mathbf{\Omega}^e\mathbf{F}'(\mathbf{I}_c - \mathbf{M}_1')^{-1}(\mathbf{I}_c - \mathbf{AC})'$$

We wish to minimize a scalar function $f_1(A)$

$$f_1(A) := \frac{1}{2} tr(cov(\hat{\mathbf{u}}_x))$$

= $\frac{1}{2} tr(\mathbf{\Omega}_x - \mathbf{A}\mathbf{C}\mathbf{\Omega}_x - \mathbf{\Omega}_x\mathbf{C}'\mathbf{A}' + \mathbf{A}\mathbf{C}\mathbf{\Omega}_x\mathbf{C}'\mathbf{A}')$

hence

$$\frac{\partial f_1}{\partial \mathbf{A}} = (\mathbf{A}\mathbf{C} - \mathbf{I}_m)\mathbf{\Omega}_x\mathbf{C}' = 0$$

This results

$$\mathbf{A} = \mathbf{\Omega}_x \mathbf{C}' (\mathbf{C} \mathbf{\Omega}_x \mathbf{C}')^{-1}$$

It is trivial to prove that the matrix **A** satisfies the other two constraints.

Using the matrix representation of equation (1.6), i.e. (1.16). Let R and b₁ be the C × N matrix and the C × 1 vector to be determined respectively, such that

the following equation holds:

$$\mathbf{D}_1 \hat{\mathbf{x}} = \mathbf{R} \Delta \mathbf{x}_{lf} + \mathbf{b}_1 = \mathbf{R} \Delta \mathbf{C} \mathbf{x} + \mathbf{b}_1 = \mathbf{R} \Delta \mathbf{C} (I_c - \mathbf{M}_1)^{-1} (\mathbf{M}_2 \mathbf{x}^{init} + \mathbf{F} \mathbf{S} \varphi + \mathbf{F} \mathbf{e}) + \mathbf{b}_1$$

where Δ is $N \times N$ matrix defined in 1.14, **D** is the $m \times m$ matrix defined in 1.2.1. **D**₁ is the **C** × c difference matrix of similar structure to **D** and Δ such that **FD** = **D**₁**F** where **F** is the **C** × m identity matrix.

$$\mathbf{D}_1 \mathbf{x} = (\mathbf{I}_c - \mathbf{M}_1)^{-1} \Lambda + (\mathbf{I}_c - \mathbf{M}_1)^{-1} \mathbf{F} \mathbf{D} \mathbf{e}$$

where Λ is a $\mathbf{C} \times 1$ vector defined as $\Lambda := \mathbf{D}_1 \mathbf{M}_2 \mathbf{x}^{init} + \mathbf{FDS}\varphi$. Let \mathbf{Q} be an $N \times c$ matrix defined by $\mathbf{QD}_1 := \mathbf{\Delta C} (\mathbf{I}_c - \mathbf{M}_1)^{-1}$, hence

$$\begin{split} \mathbf{D}_1 \mathbf{\hat{x}} = & \mathbf{R} \mathbf{Q} \mathbf{D}_1 (\mathbf{M}_2 \mathbf{x}^{init} + \mathbf{F} \mathbf{S} \varphi + \mathbf{F} \mathbf{e}) + \mathbf{b}_1 \\ &= & \mathbf{R} \mathbf{Q} (\mathbf{D}_1 \mathbf{M}_2 \mathbf{x}^{init} + \mathbf{D}_1 \mathbf{F} \mathbf{S} \varphi) + \mathbf{R} \mathbf{Q} \mathbf{D}_1 \mathbf{F} \mathbf{e} + \mathbf{b}_1 \\ &= & \mathbf{R} \mathbf{Q} \Lambda + \mathbf{R} \mathbf{Q} \mathbf{F} \mathbf{D} \mathbf{e} + \mathbf{b}_1 \end{split}$$

The residual is given by

$$\mathbf{\hat{u}}_x = \mathbf{D}_1 \mathbf{x} - \mathbf{D}_1 \mathbf{\hat{x}}$$

= $\lambda_1 (\Lambda + \mathbf{FDe}) - \mathbf{b}_1$

where $\lambda_1 := (\mathbf{I}_c - \mathbf{M}_1)^{-1} - \mathbf{R}\mathbf{Q}$. Observe that 1) When you pre-multiply the interpolated estimate vector $\hat{\boldsymbol{x}}$, by the conversion matrix \boldsymbol{C} , you get the recorded

observation vector x_{lf} , that is: $\mathbf{C}\hat{x} = x_{lf}$, hence

$$\mathbf{C}\mathbf{\hat{x}} = \mathbf{C}\mathbf{D}_{1}^{-1}\mathbf{R}\mathbf{\Delta}\mathbf{x}_{lf} + \mathbf{C}\mathbf{D}_{1}^{-1}\mathbf{b}_{1} = \mathbf{x}_{lf}$$

hence $\mathbf{C}\mathbf{D}_1^{-1}\mathbf{R}\boldsymbol{\Delta} = \mathbf{I}_N$ and $\mathbf{C}\mathbf{D}_1^{-1}\mathbf{b}_1 = \mathbf{0}_{N\times 1}$. Furthermore, 2) The estimate, $\mathbf{D}_1\hat{\boldsymbol{x}}$, needs to be unbiased $\mathbb{E}(\hat{\mathbf{u}}_x) = 0$. Hence, $\mathbf{b}_1 = \lambda_1 \Lambda$. Using simple algebra

$$\mathbf{C}\mathbf{D}_1^{-1}\mathbf{R}\mathbf{Q}\Lambda = \mathbf{C}\mathbf{D}_1^{-1}(\mathbf{I}_c - \mathbf{M}_1)^{-1}\Lambda$$

the variance covariance matrix of $\hat{\mathbf{u}}_x$ is

$$egin{aligned} \mathbb{E}(\mathbf{\hat{u}}_x\mathbf{\hat{u}}_x^{'}) &= \lambda_1\mathbf{F}\mathbb{E}(\mathbf{Dee}^{'}\mathbf{D}^{'})\mathbf{F}^{'}\lambda_1^{'} \ &= \lambda_1\mathbf{F}\mathbf{\Omega}^{\omega}\mathbf{F}^{'}\lambda_1^{'} \end{aligned}$$

where $\mathbf{\Omega}^{\omega}$ is the $m \times m$ variance matrix of the error terms in the stochastic disturbance equation. Our objective is to minimize the variance matrix of $\hat{\mathbf{u}}_x$ such that $\mathbf{CD}_1^{-1}\mathbf{R}\mathbf{\Delta} = \mathbf{I}_N$ holds

$$\begin{split} \mathscr{L} &= \frac{1}{2} tr(\mathbb{E}(\hat{\mathbf{u}}_x \hat{\mathbf{u}}_x')) - tr(\wp'(\mathbf{C} \mathbf{D}_1^{-1} \mathbf{R} \boldsymbol{\Delta} - \mathbf{I}_N)) \\ &\frac{\partial \mathscr{L}}{\partial \mathbf{R}} = -\lambda_1 \mathbf{F} \boldsymbol{\Omega}^{\omega} \mathbf{F}' \mathbf{Q}' - \mathbf{D}_1'^{-1} \mathbf{C}' \wp \boldsymbol{\Delta}' = 0 \end{split}$$

since $\mathbf{Q}' = \mathbf{D}_{1}^{'-1} (I_{c} - \mathbf{M}_{1}^{'})^{-1} \mathbf{C}' \mathbf{\Delta}'$ and $(\mathbf{QF} \mathbf{\Omega}^{\omega} \mathbf{F}' \mathbf{Q}')^{-1} = \mathbf{\Delta}^{'-1} \mathbf{\Omega}_{x,1}^{-1} \mathbf{\Delta}^{-1}$ where $\mathbf{\Omega}_{x,1} = \mathbf{C} \mathbf{\Omega}_{\mathbf{x},\mathbf{2}} \mathbf{C}'$ and $\mathbf{\Omega}_{x,2} := (\mathbf{I}_{c} - \mathbf{M}_{1})^{-1} (\mathbf{D}_{1}' (\mathbf{F} \mathbf{\Omega}^{\omega} \mathbf{F}')^{-1} \mathbf{D}_{1})^{-1} (\mathbf{I}_{c} - \mathbf{M}_{1})^{-1}$. One can show that,

$$\mathbf{R} = \mathbf{D}_{1}^{\prime-1}\mathbf{C}^{\prime} \wp \mathbf{\Omega}_{x,1}^{-1} \mathbf{\Delta}^{-1} + (\mathbf{I}_{c} - \mathbf{M}_{1})^{-1} \mathbf{F} \mathbf{\Omega}^{\omega} \mathbf{F}^{\prime} \mathbf{D}_{1}^{\prime-1} (I_{c} - \mathbf{M}_{1}^{\prime})^{-1} \mathbf{C}^{\prime} \mathbf{\Omega}_{x,1}^{-1} \mathbf{\Delta}^{-1}$$

since $\mathbf{C}\mathbf{D}_{1}^{-1}\mathbf{R}\boldsymbol{\Delta} = \mathbf{I}_{N}$

$$\mathbf{C}\mathbf{D}_{1}^{-1}\mathbf{D}_{1}^{'-1}\mathbf{C}'\wp = \mathbf{\Omega}_{x,1} - \mathbf{C}(\mathbf{I}_{c} - \mathbf{M}_{1})^{-1}\mathbf{D}_{1}^{-1}\mathbf{F}\mathbf{\Omega}^{\omega}\mathbf{F}'\mathbf{D}_{1}^{'-1}(I_{c} - \mathbf{M}_{1}^{'})^{-1}\mathbf{C}'$$

= 0

Hence,

$$\mathbf{R} = \mathbf{D_1} \mathbf{\Omega_{x,2}} \mathbf{C}' \mathbf{\Omega}_{x,1}^{-1} \mathbf{\Delta}^{-1}$$

and

$$\lambda_1 = (\mathbf{I}_c - \mathbf{M}_1)^{-1} - \mathbf{D}_1 \boldsymbol{\Omega}_{\mathbf{x}, \mathbf{2}} \mathbf{C}' \boldsymbol{\Omega}_{x, 1}^{-1} \mathbf{C} (I_c - \mathbf{M}_1)^{-1} \mathbf{D}_1^{-1}$$

Observe that $\mathbf{D}_1^{-1}\lambda_1 = \lambda_{pois,1}\mathbf{D}_1^{-1}$, where $\lambda_{pois,1} := (\mathbf{I}_c - \mathbf{M}_1)^{-1} - \mathbf{\Omega}_{\mathbf{x},2}\mathbf{C}'\mathbf{\Omega}_{x,1}^{-1}\mathbf{C}(I_c - \mathbf{M}_1)^{-1}$. Hence,

$$\mathbf{\hat{x}} = \mathbf{D_1^{-1}R\Delta x}_{lf} + \mathbf{D_1^{-1}}\lambda_1 \Lambda$$

$$\begin{split} \mathbf{b} &= \mathbf{D}_{1}^{-1} \lambda_{1} \Lambda \\ &= \lambda_{pois,1} \mathbf{D}_{1}^{-1} \Lambda \\ &= \lambda_{pois,1} (\mathbf{M}_{2} \mathbf{x}^{init} + \mathbf{FS} \varphi) \end{split} \qquad \qquad \mathbf{A}_{pois,2} := \mathbf{D}_{1}^{-1} \mathbf{R} \Delta \\ &= \mathbf{\Omega}_{\mathbf{x},2} \mathbf{C}' \mathbf{\Omega}_{x,1}^{-1} \end{split}$$

1.A.3 Proof of Proposition 1.2.2

The proof consists of three parts for: PCL, PL, and PF.

1. The matrix representation of the proposed equation (1.6) is given by

$$\mathbf{z} := \mathbf{M}_{1,z}\mathbf{z} + \mathbf{M}_{2,z}\mathbf{z}^{init} + \mathbf{S}\varphi + \mathbf{e}$$

where the $m \times m$ and the $m \times p$ matrices $\mathbf{M}_{1,z}$ and $\mathbf{M}_{2,z}$ are defined of similar structure to the \mathbf{M}_1 and \mathbf{M}_2 in the Lemma. In particular, $\mathbf{FM}_{2,z} = \mathbf{M}_2$ and $\mathbf{FM}_{1,\mathbf{z}}\mathbf{F}^{'} = \mathbf{M}_{1}.$ With some matrix algebra, one can reach

$$\mathbf{z} = (\mathbf{I}_m - \mathbf{M}_{1,z})^{-1} (\mathbf{M}_{2,z} \mathbf{z}^{init} + \mathbf{S}\varphi + \mathbf{e})$$

Observe that $\mathbf{x} = \mathbf{F}\mathbf{z} = \mathbf{F}\mathbf{M}_{1,z}\mathbf{z} + \mathbf{F}\mathbf{M}_{2,z}\mathbf{z}^{init} + \mathbf{F}\mathbf{S}\varphi + \mathbf{F}\mathbf{e} = \mathbf{M}_{1}\mathbf{F}\mathbf{z} + \mathbf{M}_{2}\mathbf{z}^{init} + \mathbf{F}\mathbf{S}\varphi + \mathbf{F}\mathbf{e}$. Hence, $\mathbf{x} = (\mathbf{I}_{c} - \mathbf{M}_{1})^{-1}(\mathbf{M}_{2}\mathbf{z}^{init} + \mathbf{F}\mathbf{S}\varphi + \mathbf{F}\mathbf{e})$. Furthermore, its estimate, $\hat{\mathbf{z}}$, is given by the suggested linear transformation

$$\hat{\mathbf{z}} = \mathbf{A}\mathbf{x}_{lf} + \mathbf{b}$$

where **A** is a $m \times N$ matrix and **b** is a $m \times 1$ vector to be determined. Observe that

$$\mathbf{z} = (\mathbf{I}_m - \mathbf{M}_{1,z})^{-1} (\mathbf{M}_{2,z} \mathbf{z}^{init} + \mathbf{S}\varphi + \mathbf{e}) = \Lambda_2 + (\mathbf{I}_m - \mathbf{M}_{1,z})^{-1} \mathbf{e}$$

Where Λ is a $\mathbf{C} \times 1$ vector defined as $\Lambda := (\mathbf{I}_c - \mathbf{M}_1)^{-1} (\mathbf{M}_2 \mathbf{x}^{init} + \mathbf{F} \mathbf{S} \varphi) = \mathbf{F} \Lambda_2,$ $\Lambda_2 := (\mathbf{I}_m - \mathbf{M}_{1,z})^{-1} (\mathbf{M}_{2,z} \mathbf{z}^{init} + \mathbf{S} \varphi).$

$$\hat{\mathbf{z}} = \mathbf{A}\mathbf{x}_{lf} + \mathbf{b} = \mathbf{A}\mathbf{CF}\mathbf{\Lambda}_2 + \mathbf{A}\mathbf{CF}(\mathbf{I}_m - \mathbf{M}_{1,z})^{-1}\mathbf{e} + b$$

The residuals is defined by

$$\hat{\mathbf{u}}_z = \mathbf{z} - \hat{\mathbf{z}} = (\mathbf{I}_m - \mathbf{ACF})\Lambda_2 + (\mathbf{I}_m - \mathbf{ACF})(\mathbf{I}_m - \mathbf{M}_{1,z})^{-1}e - b$$

Let \mathbf{C}_z be the $N \times m$ matrix, defined exactly in the same manner as the $N \times c$ matrix \mathbf{C} defined in either set-ups. In particular, when post-multiplied

by the vector $\hat{\mathbf{z}}$ one should get the vector \mathbf{x}_{lf} , that is: $\mathbf{C}_{z}\hat{\mathbf{z}} = \mathbf{x}_{lf}$. Hence, $\mathbf{C}_{z}\hat{\mathbf{z}} = \mathbf{C}_{z}\mathbf{A}\mathbf{x}_{lf} + \mathbf{C}_{z}\mathbf{b} = \mathbf{x}_{lf}$. This implies that $\mathbf{C}_{z}\mathbf{A} = \mathbf{I}_{N}$ and $\mathbf{C}_{z}\mathbf{b} = \mathbf{0}_{N}$. Furthermore, 2) The estimate, $\hat{\mathbf{z}}$, needs to be unbiased estimate of \mathbf{z} , $\mathbb{E}(\hat{\mathbf{u}}_{z}) = 0$. Hence, $\mathbf{b} = (\mathbf{I}_{m} - \mathbf{A}\mathbf{C}_{z})\Lambda_{2}$.

Observe that if $\mathbf{C}_{\mathbf{z}}\mathbf{A} = \mathbf{I}_N$, then **b** satisfies $\mathbf{C}_{\mathbf{z}}\mathbf{b} = \mathbf{0}_N$. So the only constraint that needs to be guaranteed is $\mathbf{C}_{\mathbf{z}}\mathbf{A} = \mathbf{I}_N$. Let $\lambda_1 := (\mathbf{I}_m - \mathbf{A}\mathbf{C}_{\mathbf{z}})(\mathbf{I}_m - \mathbf{M}_{\mathbf{1},\mathbf{z}})^{-1}$, then the variance covariance matrix of $\hat{\mathbf{u}}_x$ is

$$\mathbb{E}(\mathbf{\hat{u}}_{x}\mathbf{\hat{u}}_{x}^{'})=\lambda_{1}\mathbb{E}(\mathbf{ee}^{'})\lambda_{1}^{'}=\lambda_{1}\mathbf{\Omega}\lambda_{1}^{'}$$

where Ω^e is the $m \times m$ variance matrix of the error terms in the stochastic disturbance equation. Our target is to minimize this matrix such that the $\mathbf{C_z A} = \mathbf{I}_N$ holds. The Lagrangian expression is

$$\mathscr{L} = \frac{1}{2} tr(\mathbb{E}(\hat{\mathbf{u}}_{x} \hat{\mathbf{u}}_{x}^{'}))$$

The partial derivative is given by

$$\frac{\partial \mathscr{L}}{\partial \mathbf{A}} = -\lambda_1 \mathbf{\Omega}^e (\mathbf{I}_m - \mathbf{M}'_{1,z})^{-1} \mathbf{C}'_z = 0$$

Let $\Omega_{x,1} = \mathbf{C}_{\mathbf{z}} \Omega_{\mathbf{x},2} \mathbf{C}'_{z}$ and $\Omega_{x,2} := (\mathbf{I}_m - \mathbf{M}_{1,z})^{-1} \Omega^e (\mathbf{I}_m - \mathbf{M}'_{1,z})^{-1}$, hence solving for \mathbf{A} , we get

$$\mathbf{A} = \mathbf{\Omega}_{x,2} \mathbf{C}_{z}' \mathbf{\Omega}_{x,1}^{-1} \qquad \text{and} \qquad \mathbf{b} = (\mathbf{I}_{m} - \mathbf{A}\mathbf{C}_{z}) (\mathbf{I}_{m} - \mathbf{M}_{1,z})^{-1} (\mathbf{M}_{2,z} \mathbf{z}^{init} + \mathbf{S}\varphi)$$

It is easy to see that the constraint $C_z A = I_N$ holds for this choice of A.

2. Under the assumptions of ARIMA(1,1,0) stochastic disturbance terms, we will consider the following linear transformation equation

$$D\hat{z} = R\Delta x_{lf} + b_1 = R\Delta Cx + b_1 = R\Delta C_z z + b_1$$

where **D** and Δ is the $m \times m$ and the $N \times N$ difference matrix. **R** is a $m \times N$ matrix and **b**₁ is a $m \times 1$ vector to be determined. Observe that

$$\mathbf{D}\mathbf{z} = \mathbf{D}(I_m - \mathbf{M}_{1,z})^{-1}(\mathbf{M}_{2,z}\mathbf{z}^{init} + \mathbf{S}\varphi + \mathbf{e}) = \Lambda_2 + (I_m - \mathbf{M}_{1,z})^{-1}\mathbf{D}\mathbf{e}$$

where $\Lambda_2 := \mathbf{D}(I_m - \mathbf{M}_{1,z})^{-1}(\mathbf{M}_{2,z}\mathbf{z}^{init} + \mathbf{S}\varphi).$

Hence, $\mathbf{D}\hat{\mathbf{z}} = \mathbf{R}\boldsymbol{\Delta}\mathbf{C}_{\mathbf{z}}\mathbf{D}^{-1}\boldsymbol{\Lambda}_{2} + \mathbf{R}\boldsymbol{\Delta}\mathbf{C}_{\mathbf{z}}(\mathbf{I}_{\mathbf{m}} - \mathbf{M}_{\mathbf{1},\mathbf{z}})^{-1}\mathbf{e} + \mathbf{b}_{1}$.

The differenced residuals is given by

$$\hat{\mathbf{u}}_z = \mathbf{D}\mathbf{z} - \mathbf{D}\hat{\mathbf{z}} = (I_m - \mathbf{R}\mathbf{\Delta}\mathbf{C}_z\mathbf{D}^{-1})\Lambda_2 + \lambda_1\mathbf{D}\mathbf{e} - \mathbf{b}_1$$

where $\lambda_1 := (\mathbf{I_m} - \mathbf{M_{1,z}})^{-1} - \mathbf{R} \Delta \mathbf{C_z} (\mathbf{I_m} - \mathbf{M_{1,z}})^{-1} \mathbf{D}^{-1}$. When pre-multiplying the vector $\hat{\mathbf{z}}$ by \mathbf{C}_z one should get the vector \mathbf{x}_{lf} . In particular,

$$\mathbf{C_z} \mathbf{\hat{z}} = \mathbf{C_z} \mathbf{D}^{-1} \mathbf{R} \mathbf{\Delta} \mathbf{x_{lf}} + \mathbf{C_z} \mathbf{D}^{-1} \mathbf{b}_1 = \mathbf{x_{lf}}$$

Hence, $\mathbf{C}_{\mathbf{z}}\mathbf{D}^{-1}\mathbf{R}\boldsymbol{\Delta} = \mathbf{I}_N$ and $\mathbf{C}_{\mathbf{z}}\mathbf{D}^{-1}\mathbf{b}_1 = \mathbf{0}_N$ should be satisfied. Furthermore, the differenced estimate, $\mathbf{D}\hat{\mathbf{z}}$, needs to be unbiased estimate of $\mathbf{D}\mathbf{z}$, $\mathbb{E}(\hat{\mathbf{u}}_z) = \mathbf{0}$. Hence, $\mathbf{b}_1 = (I_m - \mathbf{R}\boldsymbol{\Delta}\mathbf{C}_{\mathbf{z}}\mathbf{D}^{-1})\Lambda_2$. Observe that if $\mathbf{C}_{\mathbf{z}}\mathbf{D}^{-1}\mathbf{R}\boldsymbol{\Delta} = \mathbf{I}_N$, then \mathbf{b}_1 satisfies $\mathbf{C}_{\mathbf{z}}\mathbf{D}^{-1}\mathbf{b}_1 = 0$. So the only constraint that need to be guaranteed is $\mathbf{C}_{\mathbf{z}}\mathbf{D}^{-1}\mathbf{R}\mathbf{\Delta} = \mathbf{I}_N$. The variance covariance matrix of $\hat{\mathbf{u}}_x$ is

$$\mathbb{E}(\hat{\mathbf{u}}_{x}\hat{\mathbf{u}}_{x}^{'}) = \lambda_{1}\mathbb{E}(\mathbf{Dee}^{'}\mathbf{D}^{'})\lambda_{1}^{'} = \lambda_{1}\mathbf{\Omega}^{\omega}\lambda_{1}^{'}$$

where Ω^{ω} is the $m \times m$ variance matrix of the error terms in the stochastic disturbance equation. Our target is to minimize this matrix such that the constraint holds. The Lagrangian expression is

$$\mathscr{L} = \frac{1}{2} tr(E(\hat{u}_x \hat{u}_x'))$$

The partial derivative is given by

$$\begin{aligned} \frac{\partial \mathscr{L}}{\partial \mathbf{R}} &= -\lambda_1 \mathbf{\Omega}^{\omega} \mathbf{D}^{\prime - 1} (\mathbf{I}_m - \mathbf{M}_{1,z}^{\prime})^{-1} \mathbf{F}^{\prime} \mathbf{C}^{\prime} \mathbf{\Delta}^{\prime} \\ &= (\mathbf{R} \mathbf{\Delta} \mathbf{C}_{\mathbf{z}} (\mathbf{I}_m - \mathbf{M}_{1,z})^{-1} \mathbf{D}^{-1} - (\mathbf{I}_m - \mathbf{M}_{1,z})^{-1}) \mathbf{\Omega}^{\omega} \mathbf{D}^{\prime - 1} (\mathbf{I}_m - \mathbf{M}_{1,z}^{\prime})^{-1} \mathbf{F}^{\prime} \mathbf{C}^{\prime} \mathbf{\Delta}^{\prime} \\ &= 0 \end{aligned}$$

Let $\Omega_{x,1} = \mathbf{C}_{\mathbf{z}} \Omega_{\mathbf{x},2} \mathbf{C}'_{z}$ and $\Omega_{x,2} := (\mathbf{I}_{m} - \mathbf{M}_{1,z})^{-1} \mathbf{D}^{-1} \Omega^{\omega} \mathbf{D}'^{-1} (\mathbf{I}_{m} - \mathbf{M}'_{1,z})^{-1}$, hence solving for \mathbf{R} , we get

$$\mathbf{R} = (\mathbf{I}_m - \mathbf{M}_{1,z})^{-1} \mathbf{\Omega}^{\omega} \mathbf{D}^{'-1} (\mathbf{I}_m - \mathbf{M}_{1,z}^{'})^{-1} \mathbf{C}_z^{\prime} (\mathbf{\Omega}_{x,1})^{-1} \mathbf{\Delta}^{-1}$$

It is easy to see that the constraint $C_z D^{-1} R \Delta = I_N$ holds for this choice of **R**. Recall that

$$\hat{\mathbf{z}} = \mathbf{D}^{-1} \mathbf{R} \mathbf{\Delta} \mathbf{x}_{lf} + \mathbf{D}^{-1} \mathbf{b}_1$$

Hence,

$$\mathbf{A}_{pois,2} = \mathbf{D}^{-1}\mathbf{R}\mathbf{\Delta} = \mathbf{\Omega}_{x,2}\mathbf{C}^{'}(\mathbf{\Omega}_{x,1})^{-1}$$

and

$$\mathbf{b}_{pois,2} := \mathbf{D}^{-1} \mathbf{b}_1 = (\mathbf{I}_m - \mathbf{A}_{pois,2} \mathbf{C}_{\mathbf{z}}) (\mathbf{I}_m - \mathbf{M}_{1,z})^{-1} (\mathbf{M}_{2,z} \mathbf{z}^{init} + \mathbf{S} \varphi)$$

1.B Monte Carlo Study - Tables

DGP reference #	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
Model	CLCL	CLCL	CLL	CLL	CLF	PCL	PCL	PCL	PCL	\mathbf{PL}	\mathbf{PL}	\mathbf{PL}	\mathbf{PL}	\mathbf{PF}	PF
μ	0.2	0.8	0.2	0.8		0.2	0.2	0.8	0.8	0.2	0.2	0.8	0.8		
ρ						0.2	0.8	0.2	0.8	0.2	0.8	0.2	0.8	0.2	0.8

Table 1.2: DGP Reference Key

Table 1.3: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 0.1$ with aggregation matrix C_1

	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.024	2.694	11.288	42.992	9.319	2.282	8.971	3.718	26.608	17.858	262.318	68.127	1028.365	14.232	14.106
	(0.111)	(0.193)	(5.619)	(22.671)	(4.611)	(0.134)	(1.041)	(0.296)	(3.538)	(9.136)	(143.080)	(36.514)	(536.710)	(7.044)	(7.083)
\mathbf{CLL}	2.529	3.435	5.155	14.279	4.313	2.886	11.785	4.823	33.178	7.663	71.198	22.210	233.326	6.266	7.052
	(0.362)	(0.589)	(1.120)	(4.466)	(0.961)	(0.427)	(2.653)	(0.824)	(8.755)	(1.857)	(25.460)	(7.082)	(95.650)	(1.497)	(1.997)
\mathbf{CLF}	2.307	3.016	4.205	11.016	3.528	2.607	9.475	4.134	26.034	6.127	52.620	16.956	174.798	5.061	5.464
	(0.207)	(0.301)	(0.500)	(1.853)	(0.419)	(0.244)	(1.274)	(0.426)	(3.853)	(0.808)	(9.775)	(2.890)	(36.496)	(0.660)	(0.877)
\mathbf{PCL}	2.257	2.949	4.509	12.190	3.778	2.560	8.893	4.034	23.403	6.638	60.666	18.814	213.441	5.452	5.517
	(0.164)	(0.237)	(0.710)	(2.868)	(0.601)	(0.202)	(0.984)	(0.348)	(2.956)	(1.139)	(18.943)	(4.671)	(74.874)	(0.931)	(0.914)
\mathbf{PL}	2.342	3.048	4.101	9.887	3.447	2.644	9.380	4.182	24.930	5.927	42.875	15.008	130.640	4.902	4.973
	(0.172)	(0.258)	(0.424)	(1.296)	(0.359)	(0.217)	(1.096)	(0.381)	(3.330)	(0.665)	(7.152)	(2.048)	(23.620)	(0.524)	(0.538)
\mathbf{PF}	2.340	3.043	4.057	9.650	3.418	2.642	9.217	4.169	24.208	5.832	41.941	14.641	128.912	4.837	4.910
	(0.172)	(0.256)	(0.409)	(1.242)	(0.348)	(0.216)	(1.059)	(0.378)	(3.122)	(0.644)	(6.737)	(1.955)	(23.237)	(0.506)	(0.520)
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	1 2.022	2 2.709	3 11.400	4 43.769	5 9.228	6 2.287	7 9.108	8 3.739	9 26.532	10 17.710	11 270.441	12 67.247	13 1041.455	14 19.097	15 19.624
CLCL	<u> </u>														
CLCL CLL	2.022	2.709	11.400	43.769	9.228	2.287	9.108	3.739	26.532	17.710	270.441	67.247	1041.455	19.097	19.624
	2.022 (0.114)	2.709 (0.196)	11.400 (5.811)	43.769 (23.902)	9.228 (4.593)	2.287 (0.136)	9.108 (1.000)	3.739 (0.302)	26.532 (3.689)	17.710 (9.082)	270.441 (142.413)	67.247 (36.870)	$1041.455 \\ (588.701)$	19.097 (9.302)	19.624 (9.657)
	2.022 (0.114) 3.507	2.709 (0.196) 4.287	$ \begin{array}{c} 11.400 \\ (5.811) \\ 5.793 \end{array} $	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \end{array}$	$9.228 \\ (4.593) \\ 5.006$	2.287 (0.136) 4.140	9.108 (1.000) 13.091	3.739 (0.302) 5.779	$26.532 \\ (3.689) \\ 34.092$	$ \begin{array}{c c} 17.710 \\ (9.082) \\ 8.342 \end{array} $	$270.441 \\ (142.413) \\ 71.785$	67.247 (36.870) 22.797	$\begin{array}{c} 1041.455 \\ (588.701) \\ 234.955 \end{array}$	19.097 (9.302) 8.838	$19.624 \\ (9.657) \\ 9.273$
CLL CLF	2.022 (0.114) 3.507 (0.708)	2.709 (0.196) 4.287 (1.015)	$ \begin{array}{c} 11.400 \\ (5.811) \\ 5.793 \\ (1.596) \end{array} $	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \end{array}$	$\begin{array}{c} 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \end{array}$	2.287 (0.136) 4.140 (0.849)	$9.108 \\ (1.000) \\ 13.091 \\ (3.468)$	3.739 (0.302) 5.779 (1.380)	$26.532 \\ (3.689) \\ 34.092 \\ (9.352)$	$ \begin{array}{c c} 17.710 \\ (9.082) \\ 8.342 \\ (2.346) \end{array} $	$270.441 \\ (142.413) \\ 71.785 \\ (26.265)$	$\begin{array}{c} 67.247 \\ (36.870) \\ 22.797 \\ (7.094) \end{array}$	$1041.455 \\ (588.701) \\ 234.955 \\ (91.480)$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$19.624 \\ (9.657) \\ 9.273 \\ (2.308)$
CLL	$\begin{array}{ c c c } \textbf{2.022} \\ (0.114) \\ 3.507 \\ (0.708) \\ 2.905 \end{array}$	2.709 (0.196) 4.287 (1.015) 3.508	$ \begin{array}{c c} 11.400 \\ (5.811) \\ 5.793 \\ (1.596) \\ 4.553 \end{array} $	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \\ 11.210 \end{array}$	$\begin{array}{c c} 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \\ 3.922 \end{array}$	2.287 (0.136) 4.140 (0.849) 3.333	$9.108 \\ (1.000) \\ 13.091 \\ (3.468) \\ 10.227$	3.739 (0.302) 5.779 (1.380) 4.663	26.532 (3.689) 34.092 (9.352) 26.327	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 270.441 \\ (142.413) \\ 71.785 \\ (26.265) \\ 53.200 \end{array}$	67.247 (36.870) 22.797 (7.094) 17.193 (2.880) 18.736	$\begin{array}{c} 1041.455 \\ (588.701) \\ 234.955 \\ (91.480) \\ 173.333 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380$
CLL CLF	$\begin{array}{c c} 2.022 \\ (0.114) \\ 3.507 \\ (0.708) \\ 2.905 \\ (0.370) \end{array}$	2.709 (0.196) 4.287 (1.015) 3.508 (0.505)	$ \begin{vmatrix} 11.400 \\ (5.811) \\ 5.793 \\ (1.596) \\ 4.553 \\ (0.713) \end{vmatrix} $	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \\ 11.210 \\ (1.843) \end{array}$	$\begin{array}{c c} 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \\ 3.922 \\ (0.590) \end{array}$	$\begin{array}{c c} \textbf{2.287} \\ (0.136) \\ 4.140 \\ (0.849) \\ 3.333 \\ (0.438) \end{array}$	$\begin{array}{c} 9.108 \\ (1.000) \\ 13.091 \\ (3.468) \\ 10.227 \\ (1.624) \end{array}$	$\begin{array}{c} \textbf{3.739} \\ (0.302) \\ 5.779 \\ (1.380) \\ 4.663 \\ (0.671) \end{array}$	$\begin{array}{c} 26.532 \\ (3.689) \\ 34.092 \\ (9.352) \\ 26.327 \\ (4.182) \end{array}$	$\begin{array}{c c} 17.710 \\ (9.082) \\ 8.342 \\ (2.346) \\ 6.496 \\ (1.042) \end{array}$	$\begin{array}{c} 270.441 \\ (142.413) \\ 71.785 \\ (26.265) \\ 53.200 \\ (10.171) \end{array}$	$\begin{array}{c} 67.247\\ (36.870)\\ 22.797\\ (7.094)\\ 17.193\\ (2.880) \end{array}$	$\begin{array}{c} 1041.455 \\ (588.701) \\ 234.955 \\ (91.480) \\ 173.333 \\ (34.899) \end{array}$	$ \begin{vmatrix} 19.097 \\ (9.302) \\ 8.838 \\ (2.047) \\ 7.161 \\ (0.900) \end{vmatrix} $	$19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380 \\ (1.024)$
CLL CLF	$\begin{array}{c c} 2.022 \\ (0.114) \\ 3.507 \\ (0.708) \\ 2.905 \\ (0.370) \\ 2.396 \end{array}$	2.709 (0.196) 4.287 (1.015) 3.508 (0.505) 3.072	$ \begin{vmatrix} 11.400 \\ (5.811) \\ 5.793 \\ (1.596) \\ 4.553 \\ (0.713) \\ 4.617 \end{vmatrix} $	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \\ 11.210 \\ (1.843) \\ 12.312 \end{array}$	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c c} \textbf{2.287} \\ (0.136) \\ 4.140 \\ (0.849) \\ 3.333 \\ (0.438) \\ 2.703 \end{array}$	9.108 (1.000) 13.091 (3.468) 10.227 (1.624) 8.922	$\begin{array}{c} \textbf{3.739} \\ (0.302) \\ 5.779 \\ (1.380) \\ 4.663 \\ (0.671) \\ 4.151 \end{array}$	26.532 (3.689) 34.092 (9.352) 26.327 (4.182) 23.370	$\begin{array}{c c} 17.710 \\ (9.082) \\ 8.342 \\ (2.346) \\ 6.496 \\ (1.042) \\ 6.660 \end{array}$	$\begin{array}{c} 270.441 \\ (142.413) \\ 71.785 \\ (26.265) \\ 53.200 \\ (10.171) \\ 62.317 \end{array}$	67.247 (36.870) 22.797 (7.094) 17.193 (2.880) 18.736	$\begin{array}{c} 1041.455 \\ (588.701) \\ 234.955 \\ (91.480) \\ 173.333 \\ (34.899) \\ 214.648 \end{array}$	$ \begin{vmatrix} 19.097 \\ (9.302) \\ 8.838 \\ (2.047) \\ 7.161 \\ (0.900) \\ 7.602 \end{vmatrix} $	$\begin{array}{c} 19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380 \\ (1.024) \\ 7.696 \end{array}$
CLL CLF PCL PL	$\begin{array}{c c} \textbf{2.022} \\ (0.114) \\ 3.507 \\ (0.708) \\ 2.905 \\ (0.370) \\ 2.396 \\ (0.179) \\ 2.475 \\ (0.188) \end{array}$	$\begin{array}{c} \textbf{2.709} \\ (0.196) \\ 4.287 \\ (1.015) \\ 3.508 \\ (0.505) \\ 3.072 \\ (0.241) \\ 3.170 \\ (0.264) \end{array}$	$ \begin{array}{c c} 11.400 \\ (5.811) \\ 5.793 \\ (1.596) \\ 4.553 \\ (0.713) \\ 4.617 \\ (0.748) \end{array} $	$\begin{array}{c} 43.769\\ (23.902)\\ 14.700\\ (4.519)\\ 11.210\\ (1.843)\\ 12.312\\ (2.998)\\ 9.998\\ (1.342)\end{array}$	$ \begin{vmatrix} 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \\ 3.922 \\ (0.590) \\ 3.865 \\ (0.578) \end{vmatrix} $	2.287 (0.136) 4.140 (0.849) 3.333 (0.438) 2.703 (0.201) 2.781 (0.215)	$\begin{array}{c} 9.108 \\ (1.000) \\ 13.091 \\ (3.468) \\ 10.227 \\ (1.624) \\ \textbf{8.922} \\ (0.954) \\ 9.400 \\ (1.062) \end{array}$	$\begin{array}{c} \textbf{3.739} \\ (0.302) \\ 5.779 \\ (1.380) \\ 4.663 \\ (0.671) \\ 4.151 \\ (0.353) \end{array}$	26.532 (3.689) 34.092 (9.352) 26.327 (4.182) 23.370 (3.097) 25.007 (3.334)		$\begin{array}{c} 270.441 \\ (142.413) \\ 71.785 \\ (26.265) \\ 53.200 \\ (10.171) \\ 62.317 \\ (19.290) \end{array}$	$\begin{array}{c} 67.247\\ (36.870)\\ 22.797\\ (7.094)\\ 17.193\\ (2.880)\\ 18.736\\ (4.774)\\ 15.085\\ (2.059)\end{array}$	$\begin{array}{c} 1041.455\\ (588.701)\\ 234.955\\ (91.480)\\ 173.333\\ (34.899)\\ 214.648\\ (81.083)\end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 19.624\\ (9.657)\\ 9.273\\ (2.308)\\ 7.380\\ (1.024)\\ 7.696\\ (1.210)\end{array}$
CLL CLF PCL	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} \textbf{2.709} \\ (0.196) \\ 4.287 \\ (1.015) \\ 3.508 \\ (0.505) \\ 3.072 \\ (0.241) \\ 3.170 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \\ 11.210 \\ (1.843) \\ 12.312 \\ (2.998) \\ 9.998 \end{array}$	$\begin{array}{c} 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \\ 3.922 \\ (0.590) \\ 3.865 \\ (0.578) \\ 3.529 \end{array}$	2.287 (0.136) 4.140 (0.849) 3.333 (0.438) 2.703 (0.201) 2.781	$\begin{array}{c} 9.108 \\ (1.000) \\ 13.091 \\ (3.468) \\ 10.227 \\ (1.624) \\ \textbf{8.922} \\ (0.954) \\ 9.400 \end{array}$	$\begin{array}{c} \textbf{3.739} \\ (0.302) \\ 5.779 \\ (1.380) \\ 4.663 \\ (0.671) \\ 4.151 \\ (0.353) \\ 4.302 \end{array}$	26.532 (3.689) 34.092 (9.352) 26.327 (4.182) 23.370 (3.097) 25.007	$ \begin{vmatrix} 17.710 \\ (9.082) \\ 8.342 \\ (2.346) \\ 6.496 \\ (1.042) \\ 6.660 \\ (1.132) \\ 5.985 \end{vmatrix} $	$\begin{array}{c} 270.441 \\ (142.413) \\ 71.785 \\ (26.265) \\ 53.200 \\ (10.171) \\ 62.317 \\ (19.290) \\ 43.245 \end{array}$	$\begin{array}{c} 67.247\\ (36.870)\\ 22.797\\ (7.094)\\ 17.193\\ (2.880)\\ 18.736\\ (4.774)\\ 15.085\end{array}$	$\begin{array}{c} 1041.455\\ (588.701)\\ 234.955\\ (91.480)\\ 173.333\\ (34.899)\\ 214.648\\ (81.083)\\ 129.433\end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 19.624\\ (9.657)\\ 9.273\\ (2.308)\\ 7.380\\ (1.024)\\ 7.696\\ (1.210)\\ 7.017\end{array}$

Table 1.4: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 1$ with aggregation matrix \mathbf{C}_1

	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.870	3.821	16.362	63.419	12.808	3.225	12.606	5.257	37.596	24.771	370.621	95.229	1522.353	19.097	296.528
	(0.162)	(0.274)	(8.435)	(33.590)	(6.253)	(0.184)	(1.388)	(0.440)	(5.012)	(12.297)	(206.809)	(49.425)	(851.675)	(9.302)	(154.906)
\mathbf{CLL}	3.621	4.922	7.416	20.537	6.174	4.092	16.599	6.824	47.161	10.892	100.482	30.959	333.804	8.838	77.948
	(0.522)	(0.822)	(1.714)	(6.168)	(1.337)	(0.548)	(3.837)	(1.179)	(12.145)	(2.631)	(36.510)	(9.628)	(132.133)	(2.047)	(27.929)
\mathbf{CLF}	3.246	4.255	5.991	15.677	5.000	3.651	13.347	5.820	36.819	8.682	74.243	23.826	247.499	7.161	59.456
	(0.284)	(0.431)	(0.743)	(2.497)	(0.579)	(0.304)	(1.744)	(0.612)	(5.584)	(1.112)	(13.982)	(3.773)	(48.612)	(0.900)	(10.813)
\mathbf{PCL}	3.220	4.176	6.420	17.562	5.329	3.628	12.496	5.703	32.947	9.335	86.076	26.419	309.639	7.602	69.630
	(0.235)	(0.342)	(1.053)	(4.284)	(0.826)	(0.259)	(1.330)	(0.503)	(4.286)	(1.551)	(27.639)	(6.566)	(117.513)	(1.261)	(20.836)
\mathbf{PL}	3.313	4.320	5.851	14.024	4.896	3.738	13.266	5.921	35.144	8.400	60.315	21.208	183.631	6.952	48.890
	(0.255)	(0.378)	(0.640)	(1.860)	(0.495)	(0.283)	(1.524)	(0.552)	(4.770)	(0.904)	(9.555)	(2.858)	(33.494)	(0.766)	(7.939)
\mathbf{PF}	3.309	4.307	5.762	13.667	4.837	3.732	12.951	5.889	34.092	8.229	59.064	20.704	181.261	6.835	47.884
	(0.254)	(0.374)	(0.609)	(1.776)	(0.478)	(0.281)	(1.426)	(0.544)	(4.508)	(0.871)	(9.108)	(2.734)	(32.427)	(0.735)	(7.642)
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	1 2.865	2 3.814	3 16.091	4 59.530	5	6 3.222	7 12.741	8 5.243	<mark>9</mark> 37.336	10 24.469	11 375.206	12 94.898	13 1482.696	14 19.624	15 304.011
CLCL	1					-									
CLCL	2.865	3.814	16.091	59.530	12.765	3.222	12.741	5.243	37.336	24.469	375.206	94.898	1482.696	19.624	304.011
	2.865 (0.154)	3.814 (0.285)	16.091 (7.839)	59.530 (31.016)	12.765 (6.228)	3.222 (0.198)	12.741 (1.362)	5.243 (0.438)	37.336 (5.097)	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	375.206 (210.825)	94.898 (51.295)	$1482.696 \\ (781.425)$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	304.011 (162.312)
	2.865 (0.154) 4.378	3.814 (0.285) 5.461	16.091 (7.839) 7.809	59.530 (31.016) 20.595	$ \begin{array}{c c} 12.765 \\ (6.228) \\ 6.585 \end{array} $	3.222 (0.198) 5.060	$ \begin{array}{r} 12.741 \\ (1.362) \\ 17.562 \end{array} $	5.243 (0.438) 7.422	37.336 (5.097) 47.176	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	375.206 (210.825) 100.570	$94.898 \\ (51.295) \\ 31.620$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 304.011 \\ (162.312) \\ 81.110 \end{array}$
CLL	2.865 (0.154) 4.378 (0.837)	3.814 (0.285) 5.461 (1.105)	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$59.530 \\ (31.016) \\ 20.595 \\ (6.431)$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	3.222 (0.198) 5.060 (1.028)	$ \begin{array}{r} 12.741 \\ (1.362) \\ 17.562 \\ (4.323) \end{array} $	5.243 (0.438) 7.422 (1.581)	$\begin{array}{c} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \end{array}$	$94.898 \\ (51.295) \\ 31.620 \\ (10.019)$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \\ (130.849) \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 304.011 \\ (162.312) \\ 81.110 \\ (28.582) \end{array}$
CLL	$\begin{array}{ c c c } 2.865 \\ (0.154) \\ 4.378 \\ (0.837) \\ 3.686 \end{array}$	$\begin{array}{c} \textbf{3.814} \\ (0.285) \\ 5.461 \\ (1.105) \\ 4.559 \end{array}$	$ \begin{array}{c c} 16.091 \\ (7.839) \\ 7.809 \\ (2.079) \\ 6.172 \end{array} $	$59.530 \\ (31.016) \\ 20.595 \\ (6.431) \\ 15.765$	$ \begin{array}{c cccc} 12.765 \\ (6.228) \\ 6.585 \\ (1.705) \\ 5.264 \end{array} $	3.222 (0.198) 5.060 (1.028) 4.183	$12.741 \\ (1.362) \\ 17.562 \\ (4.323) \\ 13.876$	5.243 (0.438) 7.422 (1.581) 6.137	$\begin{array}{r} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \\ 36.795 \end{array}$	$\begin{array}{c c} & 24.469 \\ (12.186) \\ 11.299 \\ (2.889) \\ 8.906 \end{array}$	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \\ 74.440 \end{array}$	$94.898 \\ (51.295) \\ 31.620 \\ (10.019) \\ 23.934$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \\ (130.849) \\ 247.097 \end{array}$	$\begin{array}{ c c c c c } & 19.624 \\ & (9.657) \\ & 9.273 \\ & (2.308) \\ & 7.380 \end{array}$	$\begin{array}{r} 304.011 \\ (162.312) \\ 81.110 \\ (28.582) \\ 59.996 \end{array}$
CLL CLF	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	3.814 (0.285) 5.461 (1.105) 4.559 (0.558)	$ \begin{vmatrix} 16.091 \\ (7.839) \\ 7.809 \\ (2.079) \\ 6.172 \\ (0.933) \end{vmatrix} $	$59.530 \\ (31.016) \\ 20.595 \\ (6.431) \\ 15.765 \\ (2.564)$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$12.741 \\ (1.362) \\ 17.562 \\ (4.323) \\ 13.876 \\ (2.022)$	5.243 (0.438) 7.422 (1.581) 6.137 (0.771)	$\begin{array}{r} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \\ 36.795 \\ (5.850) \end{array}$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \\ 74.440 \\ (13.577) \end{array}$	$\begin{array}{c} 94.898 \\ (51.295) \\ 31.620 \\ (10.019) \\ 23.934 \\ (3.961) \end{array}$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \\ (130.849) \\ 247.097 \\ (49.183) \end{array}$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} 304.011 \\ (162.312) \\ 81.110 \\ (28.582) \\ 59.996 \\ (11.196) \end{array}$
CLL CLF	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} \textbf{3.814} \\ (0.285) \\ 5.461 \\ (1.105) \\ 4.559 \\ (0.558) \\ 4.245 \\ (0.346) \\ 4.393 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$59.530 \\ (31.016) \\ 20.595 \\ (6.431) \\ 15.765 \\ (2.564) \\ 17.119 \\ (3.960) \\ 14.027 \\ \end{cases}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} 12.741 \\ (1.362) \\ 17.562 \\ (4.323) \\ 13.876 \\ (2.022) \\ \textbf{12.534} \\ (1.331) \\ 13.299 \end{array}$	5.243 (0.438) 7.422 (1.581) 6.137 (0.771) 5.728	$\begin{array}{c} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \\ 36.795 \\ (5.850) \\ \textbf{32.882} \\ (4.196) \\ 35.082 \end{array}$	$ \begin{vmatrix} 24.469 \\ (12.186) \\ 11.299 \\ (2.889) \\ 8.906 \\ (1.247) \\ 9.331 \\ (1.572) \\ 8.428 \end{vmatrix} $	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \\ 74.440 \\ (13.577) \\ 87.071 \end{array}$	$\begin{array}{c} 94.898\\(51.295)\\31.620\\(10.019)\\23.934\\(3.961)\\26.487\\(6.791)\\20.906\end{array}$	$\begin{array}{c} 1482.696\\ (781.425)\\ 335.536\\ (130.849)\\ 247.097\\ (49.183)\\ 306.880\\ (109.301)\\ 183.361\end{array}$	$\begin{vmatrix} 19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380 \\ (1.024) \\ 7.696 \end{vmatrix}$	$\begin{array}{c} 304.011\\ (162.312)\\ 81.110\\ (28.582)\\ 59.996\\ (11.196)\\ 69.908\\ (21.598)\\ 48.860\end{array}$
CLL CLF PCL PL	2.865 (0.154) 4.378 (0.837) 3.686 (0.429) 3.317 (0.227) 3.411 (0.248)	$\begin{array}{c} \textbf{3.814} \\ (0.285) \\ 5.461 \\ (1.105) \\ 4.559 \\ (0.558) \\ 4.245 \\ (0.346) \\ 4.393 \\ (0.380) \end{array}$		$59.530 \\ (31.016) \\ 20.595 \\ (6.431) \\ 15.765 \\ (2.564) \\ 17.119 \\ (3.960)$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c c} \textbf{3.222} \\ (0.198) \\ 5.060 \\ (1.028) \\ 4.183 \\ (0.518) \\ 3.719 \\ (0.281) \\ 3.824 \\ (0.305) \end{array}$	$\begin{array}{c} 12.741 \\ (1.362) \\ 17.562 \\ (4.323) \\ 13.876 \\ (2.022) \\ 12.534 \\ (1.331) \end{array}$	$\begin{array}{c} \textbf{5.243} \\ (0.438) \\ 7.422 \\ (1.581) \\ 6.137 \\ (0.771) \\ 5.728 \\ (0.515) \end{array}$	$\begin{array}{c} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \\ 36.795 \\ (5.850) \\ \textbf{32.882} \\ (4.196) \end{array}$	$ \begin{vmatrix} 24.469 \\ (12.186) \\ 11.299 \\ (2.889) \\ 8.906 \\ (1.247) \\ 9.331 \\ (1.572) \\ 8.428 \\ (0.897) \end{vmatrix} $	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \\ 74.440 \\ (13.577) \\ 87.071 \\ (28.228) \end{array}$	$\begin{array}{c} 94.898 \\ (51.295) \\ 31.620 \\ (10.019) \\ 23.934 \\ (3.961) \\ 26.487 \\ (6.791) \end{array}$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \\ (130.849) \\ 247.097 \\ (49.183) \\ 306.880 \\ (109.301) \end{array}$	$ \begin{vmatrix} 19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380 \\ (1.024) \\ 7.696 \\ (1.210) \end{vmatrix} $	$\begin{array}{c} 304.011\\ (162.312)\\ 81.110\\ (28.582)\\ 59.996\\ (11.196)\\ 69.908\\ (21.598)\\ 48.860\\ (7.994) \end{array}$
CLL CLF PCL	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} \textbf{3.814} \\ (0.285) \\ 5.461 \\ (1.105) \\ 4.559 \\ (0.558) \\ 4.245 \\ (0.346) \\ 4.393 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$59.530 \\ (31.016) \\ 20.595 \\ (6.431) \\ 15.765 \\ (2.564) \\ 17.119 \\ (3.960) \\ 14.027 \\ \end{cases}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} 12.741 \\ (1.362) \\ 17.562 \\ (4.323) \\ 13.876 \\ (2.022) \\ \textbf{12.534} \\ (1.331) \\ 13.299 \end{array}$	$\begin{array}{c} \textbf{5.243} \\ (0.438) \\ 7.422 \\ (1.581) \\ 6.137 \\ (0.771) \\ 5.728 \\ (0.515) \\ 5.946 \end{array}$	$\begin{array}{c} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \\ 36.795 \\ (5.850) \\ \textbf{32.882} \\ (4.196) \\ 35.082 \end{array}$	$ \begin{vmatrix} 24.469 \\ (12.186) \\ 11.299 \\ (2.889) \\ 8.906 \\ (1.247) \\ 9.331 \\ (1.572) \\ 8.428 \end{vmatrix} $	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \\ 74.440 \\ (13.577) \\ 87.071 \\ (28.228) \\ 60.479 \end{array}$	$\begin{array}{c} 94.898\\(51.295)\\31.620\\(10.019)\\23.934\\(3.961)\\26.487\\(6.791)\\20.906\end{array}$	$\begin{array}{c} 1482.696\\ (781.425)\\ 335.536\\ (130.849)\\ 247.097\\ (49.183)\\ 306.880\\ (109.301)\\ 183.361\end{array}$	$ \begin{vmatrix} 19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380 \\ (1.024) \\ 7.696 \\ (1.210) \\ 7.017 \end{vmatrix} $	$\begin{array}{c} 304.011\\ (162.312)\\ 81.110\\ (28.582)\\ 59.996\\ (11.196)\\ 69.908\\ (21.598)\\ 48.860\end{array}$

Table 1.5: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 5$ with aggregation matrix \mathbf{C}_1

	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.024	2.694	11.288	42.992	9.319	2.282	8.971	3.718	26.608	17.858	262.318	68.127	1028.365	14.232	14.106
	(0.111)	(0.193)	(5.619)	(22.671)	(4.611)	(0.134)	(1.041)	(0.296)	(3.538)	(9.136)	(143.080)	(36.514)	(536.710)	(7.044)	(7.083)
\mathbf{CLL}	2.529	3.435	5.155	14.279	4.313	2.886	11.785	4.823	33.178	7.663	71.198	22.210	233.326	6.266	7.052
	(0.362)	(0.589)	(1.120)	(4.466)	(0.961)	(0.427)	(2.653)	(0.824)	(8.755)	(1.857)	(25.460)	(7.082)	(95.650)	(1.497)	(1.997)
\mathbf{CLF}	2.307	3.016	4.205	11.016	3.528	2.607	9.475	4.134	26.034	6.127	52.620	16.956	174.798	5.061	5.464
	(0.207)	(0.301)	(0.500)	(1.853)	(0.419)	(0.244)	(1.274)	(0.426)	(3.853)	(0.808)	(9.775)	(2.890)	(36.496)	(0.660)	(0.877)
\mathbf{PCL}	2.257	2.949	4.509	12.190	3.778	2.560	8.893	4.034	23.403	6.638	60.666	18.814	213.441	5.452	5.517
	(0.164)	(0.237)	(0.710)	(2.868)	(0.601)	(0.202)	(0.984)	(0.348)	(2.956)	(1.139)	(18.943)	(4.671)	(74.874)	(0.931)	(0.914)
\mathbf{PL}	2.342	3.048	4.101	9.887	3.447	2.644	9.380	4.182	24.930	5.927	42.875	15.008	130.640	4.902	4.973
	(0.172)	(0.258)	(0.424)	(1.296)	(0.359)	(0.217)	(1.096)	(0.381)	(3.330)	(0.665)	(7.152)	(2.048)	(23.620)	(0.524)	(0.538)
\mathbf{PF}	2.340	3.043	4.057	9.650	3.418	2.642	9.217	4.169	24.208	5.832	41.941	14.641	128.912	4.837	4.910
	(0.172)	(0.256)	(0.409)	(1.242)	(0.348)	(0.216)	(1.059)	(0.378)	(3.122)	(0.644)	(6.737)	(1.955)	(23.237)	(0.506)	(0.520)
	1 4														
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	1 2.022	2 2.709	3 11.400	4 43.769	5 9.228	6 2.287	7 9.108	8 3.739	9 26.532	10 17.710	11 270.441	12 67.247	13 1041.455	14 19.097	15 19.624
CLCL	-													<u> </u>	
CLCL CLL	2.022	2.709	11.400	43.769	9.228	2.287	9.108	3.739	26.532	17.710	270.441	67.247	1041.455	19.097	19.624
	2.022 (0.114)	2.709 (0.196)	11.400 (5.811)	43.769 (23.902)	$ \begin{array}{c} 9.228 \\ (4.593) \end{array} $	2.287 (0.136)	9.108 (1.000)	3.739 (0.302)	26.532 (3.689)	17.710 (9.082)	270.441 (142.413)	67.247 (36.870)	1041.455 (588.701)	$ \begin{array}{c c} 19.097 \\ (9.302) \end{array} $	19.624 (9.657)
	2.022 (0.114) 3.507	2.709 (0.196) 4.287	$ \begin{array}{c} 11.400 \\ (5.811) \\ 5.793 \end{array} $	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \end{array}$	$ \begin{array}{c c} 9.228 \\ (4.593) \\ 5.006 \end{array} $	2.287 (0.136) 4.140	9.108 (1.000) 13.091	3.739 (0.302) 5.779	$26.532 \\ (3.689) \\ 34.092$	$ \begin{array}{c c} 17.710 \\ (9.082) \\ 8.342 \end{array} $	$270.441 \\ (142.413) \\ 71.785$	67.247 (36.870) 22.797	$\begin{array}{c} 1041.455 \\ (588.701) \\ 234.955 \end{array}$	19.097 (9.302) 8.838	$19.624 \\ (9.657) \\ 9.273$
CLL	2.022 (0.114) 3.507 (0.708)	2.709 (0.196) 4.287 (1.015)	$ \begin{array}{c} 11.400 \\ (5.811) \\ 5.793 \\ (1.596) \end{array} $	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \end{array}$	$ \begin{array}{ c c c c c } 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \end{array} $	2.287 (0.136) 4.140 (0.849)	$9.108 \\ (1.000) \\ 13.091 \\ (3.468)$	3.739 (0.302) 5.779 (1.380)	$26.532 \\ (3.689) \\ 34.092 \\ (9.352)$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$270.441 \\ (142.413) \\ 71.785 \\ (26.265)$	$\begin{array}{c} 67.247 \\ (36.870) \\ 22.797 \\ (7.094) \end{array}$	$1041.455 \\ (588.701) \\ 234.955 \\ (91.480)$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$19.624 \\ (9.657) \\ 9.273 \\ (2.308)$
CLL	$\begin{array}{ c c c } \textbf{2.022} \\ (0.114) \\ 3.507 \\ (0.708) \\ 2.905 \end{array}$	2.709 (0.196) 4.287 (1.015) 3.508	$ \begin{array}{c c} 11.400 \\ (5.811) \\ 5.793 \\ (1.596) \\ 4.553 \end{array} $	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \\ 11.210 \end{array}$	$\begin{array}{ c c c } 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \\ 3.922 \end{array}$	2.287 (0.136) 4.140 (0.849) 3.333	$9.108 \\ (1.000) \\ 13.091 \\ (3.468) \\ 10.227$	3.739 (0.302) 5.779 (1.380) 4.663	26.532 (3.689) 34.092 (9.352) 26.327	$ \begin{array}{c c} 17.710 \\ (9.082) \\ 8.342 \\ (2.346) \\ 6.496 \end{array} $	$\begin{array}{c} 270.441 \\ (142.413) \\ 71.785 \\ (26.265) \\ 53.200 \end{array}$	$\begin{array}{c} 67.247\\ (36.870)\\ 22.797\\ (7.094)\\ 17.193 \end{array}$	$\begin{array}{c} 1041.455 \\ (588.701) \\ 234.955 \\ (91.480) \\ 173.333 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$ \begin{array}{r} 19.624\\(9.657)\\9.273\\(2.308)\\7.380\end{array} $
CLL CLF	$\begin{array}{c c} 2.022 \\ (0.114) \\ 3.507 \\ (0.708) \\ 2.905 \\ (0.370) \end{array}$	2.709 (0.196) 4.287 (1.015) 3.508 (0.505)	$ \begin{vmatrix} 11.400 \\ (5.811) \\ 5.793 \\ (1.596) \\ 4.553 \\ (0.713) \end{vmatrix} $	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \\ 11.210 \\ (1.843) \end{array}$	$\begin{array}{ c c c } 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \\ 3.922 \\ (0.590) \end{array}$	$\begin{array}{ } \textbf{2.287} \\ (0.136) \\ 4.140 \\ (0.849) \\ 3.333 \\ (0.438) \end{array}$	$\begin{array}{c} 9.108 \\ (1.000) \\ 13.091 \\ (3.468) \\ 10.227 \\ (1.624) \end{array}$	$\begin{array}{c} \textbf{3.739} \\ (0.302) \\ 5.779 \\ (1.380) \\ 4.663 \\ (0.671) \end{array}$	$\begin{array}{c} 26.532 \\ (3.689) \\ 34.092 \\ (9.352) \\ 26.327 \\ (4.182) \end{array}$	$\begin{array}{c c} 17.710 \\ (9.082) \\ 8.342 \\ (2.346) \\ 6.496 \\ (1.042) \end{array}$	$\begin{array}{c} 270.441 \\ (142.413) \\ 71.785 \\ (26.265) \\ 53.200 \\ (10.171) \end{array}$	$\begin{array}{c} 67.247\\ (36.870)\\ 22.797\\ (7.094)\\ 17.193\\ (2.880) \end{array}$	$\begin{array}{c} 1041.455 \\ (588.701) \\ 234.955 \\ (91.480) \\ 173.333 \\ (34.899) \end{array}$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380 \\ (1.024)$
CLL CLF	$\begin{array}{c c} \textbf{2.022} \\ (0.114) \\ 3.507 \\ (0.708) \\ 2.905 \\ (0.370) \\ 2.396 \end{array}$	2.709 (0.196) 4.287 (1.015) 3.508 (0.505) 3.072	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \\ 11.210 \\ (1.843) \\ 12.312 \end{array}$	$\begin{array}{ c c c } 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \\ 3.922 \\ (0.590) \\ 3.865 \end{array}$	$\begin{array}{ c c c c } \hline \textbf{2.287} \\ (0.136) \\ 4.140 \\ (0.849) \\ 3.333 \\ (0.438) \\ 2.703 \end{array}$	9.108 (1.000) 13.091 (3.468) 10.227 (1.624) 8.922	$\begin{array}{c} \textbf{3.739} \\ (0.302) \\ 5.779 \\ (1.380) \\ 4.663 \\ (0.671) \\ 4.151 \end{array}$	26.532 (3.689) 34.092 (9.352) 26.327 (4.182) 23.370	$\begin{array}{c c} 17.710 \\ (9.082) \\ 8.342 \\ (2.346) \\ 6.496 \\ (1.042) \\ 6.660 \end{array}$	$\begin{array}{c} 270.441 \\ (142.413) \\ 71.785 \\ (26.265) \\ 53.200 \\ (10.171) \\ 62.317 \end{array}$	67.247 (36.870) 22.797 (7.094) 17.193 (2.880) 18.736	$\begin{array}{c} 1041.455 \\ (588.701) \\ 234.955 \\ (91.480) \\ 173.333 \\ (34.899) \\ 214.648 \end{array}$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} 19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380 \\ (1.024) \\ 7.696 \end{array}$
CLL CLF PCL	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} \textbf{2.709} \\ (0.196) \\ 4.287 \\ (1.015) \\ 3.508 \\ (0.505) \\ 3.072 \\ (0.241) \end{array}$	$ \begin{array}{c c} 11.400 \\ (5.811) \\ 5.793 \\ (1.596) \\ 4.553 \\ (0.713) \\ 4.617 \\ (0.748) \end{array} $	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \\ 11.210 \\ (1.843) \\ 12.312 \\ (2.998) \end{array}$	$ \begin{vmatrix} 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \\ 3.922 \\ (0.590) \\ 3.865 \\ (0.578) \end{vmatrix} $	$\begin{array}{ } \textbf{2.287} \\ (0.136) \\ 4.140 \\ (0.849) \\ 3.333 \\ (0.438) \\ 2.703 \\ (0.201) \end{array}$	$\begin{array}{c} 9.108 \\ (1.000) \\ 13.091 \\ (3.468) \\ 10.227 \\ (1.624) \\ \textbf{8.922} \\ (0.954) \end{array}$	$\begin{array}{c} \textbf{3.739} \\ (0.302) \\ 5.779 \\ (1.380) \\ 4.663 \\ (0.671) \\ 4.151 \\ (0.353) \end{array}$	26.532 (3.689) 34.092 (9.352) 26.327 (4.182) 23.370 (3.097)	$ \begin{vmatrix} 17.710 \\ (9.082) \\ 8.342 \\ (2.346) \\ 6.496 \\ (1.042) \\ 6.660 \\ (1.132) \end{vmatrix} $	$\begin{array}{c} 270.441 \\ (142.413) \\ 71.785 \\ (26.265) \\ 53.200 \\ (10.171) \\ 62.317 \\ (19.290) \end{array}$	$\begin{array}{c} 67.247\\ (36.870)\\ 22.797\\ (7.094)\\ 17.193\\ (2.880)\\ 18.736\\ (4.774)\end{array}$	$\begin{array}{c} 1041.455\\ (588.701)\\ 234.955\\ (91.480)\\ 173.333\\ (34.899)\\ 214.648\\ (81.083)\end{array}$	$ \begin{vmatrix} 19.097 \\ (9.302) \\ 8.838 \\ (2.047) \\ 7.161 \\ (0.900) \\ 7.602 \\ (1.261) \end{vmatrix} $	$\begin{array}{c} 19.624\\ (9.657)\\ 9.273\\ (2.308)\\ 7.380\\ (1.024)\\ 7.696\\ (1.210)\end{array}$
CLL CLF PCL	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} \textbf{2.709} \\ (0.196) \\ 4.287 \\ (1.015) \\ 3.508 \\ (0.505) \\ 3.072 \\ (0.241) \\ 3.170 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 43.769 \\ (23.902) \\ 14.700 \\ (4.519) \\ 11.210 \\ (1.843) \\ 12.312 \\ (2.998) \\ 9.998 \end{array}$	$ \begin{vmatrix} 9.228 \\ (4.593) \\ 5.006 \\ (1.318) \\ 3.922 \\ (0.590) \\ 3.865 \\ (0.578) \\ 3.529 \end{vmatrix} $	$\begin{array}{c c} \textbf{2.287} \\ (0.136) \\ 4.140 \\ (0.849) \\ 3.333 \\ (0.438) \\ 2.703 \\ (0.201) \\ 2.781 \end{array}$	$\begin{array}{c} 9.108 \\ (1.000) \\ 13.091 \\ (3.468) \\ 10.227 \\ (1.624) \\ \textbf{8.922} \\ (0.954) \\ 9.400 \end{array}$	$\begin{array}{c} \textbf{3.739} \\ (0.302) \\ 5.779 \\ (1.380) \\ 4.663 \\ (0.671) \\ 4.151 \\ (0.353) \\ 4.302 \end{array}$	26.532 (3.689) 34.092 (9.352) 26.327 (4.182) 23.370 (3.097) 25.007	$ \begin{vmatrix} 17.710 \\ (9.082) \\ 8.342 \\ (2.346) \\ 6.496 \\ (1.042) \\ 6.660 \\ (1.132) \\ 5.985 \end{vmatrix} $	$\begin{array}{c} 270.441 \\ (142.413) \\ 71.785 \\ (26.265) \\ 53.200 \\ (10.171) \\ 62.317 \\ (19.290) \\ 43.245 \end{array}$	$\begin{array}{c} 67.247\\ (36.870)\\ 22.797\\ (7.094)\\ 17.193\\ (2.880)\\ 18.736\\ (4.774)\\ 15.085\end{array}$	$\begin{array}{c} 1041.455\\ (588.701)\\ 234.955\\ (91.480)\\ 173.333\\ (34.899)\\ 214.648\\ (81.083)\\ 129.433\end{array}$	$ \begin{vmatrix} 19.097 \\ (9.302) \\ 8.838 \\ (2.047) \\ 7.161 \\ (0.900) \\ 7.602 \\ (1.261) \\ 6.952 \end{vmatrix} $	$\begin{array}{c} 19.624\\ (9.657)\\ 9.273\\ (2.308)\\ 7.380\\ (1.024)\\ 7.696\\ (1.210)\\ 7.017\end{array}$

Table 1.6: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 10$ with aggregation matrix \mathbf{C}_1

	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.870	3.821	16.362	63.419	12.808	3.225	12.606	5.257	37.596	24.771	370.621	95.229	1522.353	19.097	296.528
	(0.162)	(0.274)	(8.435)	(33.590)	(6.253)	(0.184)	(1.388)	(0.440)	(5.012)	(12.297)	(206.809)	(49.425)	(851.675)	(9.302)	(154.906)
\mathbf{CLL}	3.621	4.922	7.416	20.537	6.174	4.092	16.599	6.824	47.161	10.892	100.482	30.959	333.804	8.838	77.948
	(0.522)	(0.822)	(1.714)	(6.168)	(1.337)	(0.548)	(3.837)	(1.179)	(12.145)	(2.631)	(36.510)	(9.628)	(132.133)	(2.047)	(27.929)
\mathbf{CLF}	3.246	4.255	5.991	15.677	5.000	3.651	13.347	5.820	36.819	8.682	74.243	23.826	247.499	7.161	59.456
	(0.284)	(0.431)	(0.743)	(2.497)	(0.579)	(0.304)	(1.744)	(0.612)	(5.584)	(1.112)	(13.982)	(3.773)	(48.612)	(0.900)	(10.813)
\mathbf{PCL}	3.220	4.176	6.420	17.562	5.329	3.628	12.496	5.703	32.947	9.335	86.076	26.419	309.639	7.602	69.630
	(0.235)	(0.342)	(1.053)	(4.284)	(0.826)	(0.259)	(1.330)	(0.503)	(4.286)	(1.551)	(27.639)	(6.566)	(117.513)	(1.261)	(20.836)
\mathbf{PL}	3.313	4.320	5.851	14.024	4.896	3.738	13.266	5.921	35.144	8.400	60.315	21.208	183.631	6.952	48.890
	(0.255)	(0.378)	(0.640)	(1.860)	(0.495)	(0.283)	(1.524)	(0.552)	(4.770)	(0.904)	(9.555)	(2.858)	(33.494)	(0.766)	(7.939)
\mathbf{PF}	3.309	4.307	5.762	13.667	4.837	3.732	12.951	5.889	34.092	8.229	59.064	20.704	181.261	6.835	47.884
	(0.254)	(0.374)	(0.609)	(1.776)	(0.478)	(0.281)	(1.426)	(0.544)	(4.508)	(0.871)	(9.108)	(2.734)	(32.427)	(0.735)	(7.642)
	1	~													
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	$\begin{array}{c c} 1 \\ \hline 2.865 \end{array}$	2 3.814	3 16.091	4 59.530	5 12.765	6 3.222	7 12.741	8 5.243	9 37.336	10 24.469	11 375.206	12 94.898	13 1482.696	14 19.624	15 304.011
CLCL	-						-								
CLCL	2.865	3.814	16.091	59.530	12.765	3.222	12.741	5.243	37.336	24.469	375.206	94.898	1482.696	19.624	304.011
	2.865 (0.154)	3.814 (0.285)	16.091 (7.839)	59.530 (31.016)	$ \begin{array}{c} 12.765 \\ (6.228) \end{array} $	3.222 (0.198)	12.741 (1.362)	5.243 (0.438)	37.336 (5.097)	$ \begin{array}{c c} 24.469 \\ (12.186) \end{array} $	375.206 (210.825)	94.898 (51.295)	$1482.696 \\ (781.425)$	$ \begin{array}{c c} 19.624 \\ (9.657) \end{array} $	304.011 (162.312)
	2.865 (0.154) 4.378	3.814 (0.285) 5.461	$ \begin{array}{c} 16.091 \\ (7.839) \\ 7.809 \end{array} $	59.530 (31.016) 20.595	$ \begin{array}{c c} 12.765 \\ (6.228) \\ 6.585 \end{array} $	3.222 (0.198) 5.060	$ \begin{array}{c} 12.741 \\ (1.362) \\ 17.562 \end{array} $	5.243 (0.438) 7.422	37.336 (5.097) 47.176	$ \begin{array}{c c} 24.469 \\ (12.186) \\ 11.299 \end{array} $	375.206 (210.825) 100.570	$94.898 \\ (51.295) \\ 31.620$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 304.011 \\ (162.312) \\ 81.110 \end{array}$
CLL	2.865 (0.154) 4.378 (0.837)	3.814 (0.285) 5.461 (1.105)	$ \begin{array}{c c} 16.091 \\ (7.839) \\ 7.809 \\ (2.079) \end{array} $	$59.530 \\ (31.016) \\ 20.595 \\ (6.431)$	$ \begin{array}{c c} 12.765 \\ (6.228) \\ 6.585 \\ (1.705) \end{array} $	3.222 (0.198) 5.060 (1.028)	$ \begin{array}{r} 12.741 \\ (1.362) \\ 17.562 \\ (4.323) \end{array} $	5.243 (0.438) 7.422 (1.581)	$\begin{array}{c} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \end{array}$	$94.898 \\ (51.295) \\ 31.620 \\ (10.019)$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \\ (130.849) \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 304.011 \\ (162.312) \\ 81.110 \\ (28.582) \end{array}$
CLL	$\begin{array}{ c c c } \textbf{2.865} \\ (0.154) \\ 4.378 \\ (0.837) \\ 3.686 \end{array}$	3.814 (0.285) 5.461 (1.105) 4.559	$ \begin{array}{c c} 16.091 \\ (7.839) \\ 7.809 \\ (2.079) \\ 6.172 \end{array} $	$59.530 \\ (31.016) \\ 20.595 \\ (6.431) \\ 15.765$	$ \begin{array}{c c} 12.765 \\ (6.228) \\ 6.585 \\ (1.705) \\ 5.264 \end{array} $	3.222 (0.198) 5.060 (1.028) 4.183	$12.741 \\ (1.362) \\ 17.562 \\ (4.323) \\ 13.876$	5.243 (0.438) 7.422 (1.581) 6.137	$\begin{array}{r} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \\ 36.795 \end{array}$	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \\ 74.440 \end{array}$	$94.898 \\ (51.295) \\ 31.620 \\ (10.019) \\ 23.934$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \\ (130.849) \\ 247.097 \end{array}$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{r} 304.011 \\ (162.312) \\ 81.110 \\ (28.582) \\ 59.996 \end{array}$
CLL CLF	$\begin{array}{c c} 2.865 \\ (0.154) \\ 4.378 \\ (0.837) \\ 3.686 \\ (0.429) \end{array}$	3.814 (0.285) 5.461 (1.105) 4.559 (0.558)	$ \begin{vmatrix} 16.091 \\ (7.839) \\ 7.809 \\ (2.079) \\ 6.172 \\ (0.933) \end{vmatrix} $	$59.530 \\ (31.016) \\ 20.595 \\ (6.431) \\ 15.765 \\ (2.564)$	$\begin{array}{c c} 12.765 \\ (6.228) \\ 6.585 \\ (1.705) \\ 5.264 \\ (0.772) \end{array}$	$\begin{array}{c c} \textbf{3.222} \\ (0.198) \\ 5.060 \\ (1.028) \\ 4.183 \\ (0.518) \end{array}$	$12.741 \\ (1.362) \\ 17.562 \\ (4.323) \\ 13.876 \\ (2.022)$	5.243 (0.438) 7.422 (1.581) 6.137 (0.771)	$\begin{array}{c} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \\ 36.795 \\ (5.850) \end{array}$	$\begin{array}{c c} & 24.469 \\ (12.186) \\ 11.299 \\ (2.889) \\ 8.906 \\ (1.247) \end{array}$	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \\ 74.440 \\ (13.577) \end{array}$	$\begin{array}{c} 94.898 \\ (51.295) \\ 31.620 \\ (10.019) \\ 23.934 \\ (3.961) \end{array}$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \\ (130.849) \\ 247.097 \\ (49.183) \end{array}$	$ \begin{vmatrix} 19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380 \\ (1.024) \end{vmatrix} $	$\begin{array}{c} 304.011 \\ (162.312) \\ 81.110 \\ (28.582) \\ 59.996 \\ (11.196) \end{array}$
CLL CLF	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} \textbf{3.814} \\ (0.285) \\ 5.461 \\ (1.105) \\ 4.559 \\ (0.558) \\ 4.245 \end{array}$		$\begin{array}{c} 59.530\\ (31.016)\\ 20.595\\ (6.431)\\ 15.765\\ (2.564)\\ 17.119\\ (3.960)\\ 14.027\end{array}$	$ \begin{vmatrix} 12.765 \\ (6.228) \\ 6.585 \\ (1.705) \\ 5.264 \\ (0.772) \\ 5.392 \\ (0.794) \\ 4.961 \end{vmatrix} $	$\begin{array}{c c} \textbf{3.222} \\ (0.198) \\ 5.060 \\ (1.028) \\ 4.183 \\ (0.518) \\ 3.719 \\ (0.281) \\ 3.824 \end{array}$	$\begin{array}{c} 12.741 \\ (1.362) \\ 17.562 \\ (4.323) \\ 13.876 \\ (2.022) \\ \textbf{12.534} \\ (1.331) \\ 13.299 \end{array}$	5.243 (0.438) 7.422 (1.581) 6.137 (0.771) 5.728	$\begin{array}{c} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \\ 36.795 \\ (5.850) \\ \textbf{32.882} \\ (4.196) \\ 35.082 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \\ 74.440 \\ (13.577) \\ 87.071 \\ (28.228) \\ 60.479 \end{array}$	$\begin{array}{c} 94.898 \\ (51.295) \\ 31.620 \\ (10.019) \\ 23.934 \\ (3.961) \\ 26.487 \\ (6.791) \\ 20.906 \end{array}$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \\ (130.849) \\ 247.097 \\ (49.183) \\ 306.880 \\ (109.301) \\ 183.361 \end{array}$	$ \begin{vmatrix} 19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380 \\ (1.024) \\ 7.696 \\ (1.210) \\ 7.017 \end{vmatrix} $	$\begin{array}{c} 304.011 \\ (162.312) \\ 81.110 \\ (28.582) \\ 59.996 \\ (11.196) \\ 69.908 \\ (21.598) \\ 48.860 \end{array}$
CLL CLF PCL PL	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} \textbf{3.814} \\ (0.285) \\ 5.461 \\ (1.105) \\ 4.559 \\ (0.558) \\ 4.245 \\ (0.346) \\ 4.393 \\ (0.380) \end{array}$		$\begin{array}{c} 59.530\\ (31.016)\\ 20.595\\ (6.431)\\ 15.765\\ (2.564)\\ 17.119\\ (3.960)\\ 14.027\\ (1.823)\end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c c} \textbf{3.222} \\ (0.198) \\ 5.060 \\ (1.028) \\ 4.183 \\ (0.518) \\ 3.719 \\ (0.281) \\ 3.824 \\ (0.305) \end{array}$	$\begin{array}{c} 12.741 \\ (1.362) \\ 17.562 \\ (4.323) \\ 13.876 \\ (2.022) \\ \textbf{12.534} \\ (1.331) \\ 13.299 \\ (1.525) \end{array}$	$\begin{array}{c} \textbf{5.243} \\ (0.438) \\ 7.422 \\ (1.581) \\ 6.137 \\ (0.771) \\ 5.728 \\ (0.515) \\ 5.946 \\ (0.561) \end{array}$	$\begin{array}{c} 37.336\\ (5.097)\\ 47.176\\ (12.823)\\ 36.795\\ (5.850)\\ \textbf{32.882}\\ (4.196)\\ 35.082\\ (4.777) \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \\ 74.440 \\ (13.577) \\ 87.071 \\ (28.228) \\ 60.479 \\ (10.069) \end{array}$	$\begin{array}{c} 94.898\\ (51.295)\\ 31.620\\ (10.019)\\ 23.934\\ (3.961)\\ 26.487\\ (6.791)\\ 20.906\\ (2.832) \end{array}$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \\ (130.849) \\ 247.097 \\ (49.183) \\ 306.880 \\ (109.301) \\ 183.361 \\ (32.677) \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 304.011\\ (162.312)\\ 81.110\\ (28.582)\\ 59.996\\ (11.196)\\ 69.908\\ (21.598)\\ 48.860\\ (7.994) \end{array}$
CLL CLF PCL	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} \textbf{3.814} \\ (0.285) \\ 5.461 \\ (1.105) \\ 4.559 \\ (0.558) \\ 4.245 \\ (0.346) \\ 4.393 \end{array}$		$\begin{array}{c} 59.530\\ (31.016)\\ 20.595\\ (6.431)\\ 15.765\\ (2.564)\\ 17.119\\ (3.960)\\ 14.027\end{array}$	$ \begin{vmatrix} 12.765 \\ (6.228) \\ 6.585 \\ (1.705) \\ 5.264 \\ (0.772) \\ 5.392 \\ (0.794) \\ 4.961 \end{vmatrix} $	$\begin{array}{c c} \textbf{3.222} \\ (0.198) \\ 5.060 \\ (1.028) \\ 4.183 \\ (0.518) \\ 3.719 \\ (0.281) \\ 3.824 \end{array}$	$\begin{array}{c} 12.741 \\ (1.362) \\ 17.562 \\ (4.323) \\ 13.876 \\ (2.022) \\ \textbf{12.534} \\ (1.331) \\ 13.299 \end{array}$	$\begin{array}{c} \textbf{5.243} \\ (0.438) \\ 7.422 \\ (1.581) \\ 6.137 \\ (0.771) \\ 5.728 \\ (0.515) \\ 5.946 \end{array}$	$\begin{array}{c} 37.336 \\ (5.097) \\ 47.176 \\ (12.823) \\ 36.795 \\ (5.850) \\ \textbf{32.882} \\ (4.196) \\ 35.082 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 375.206 \\ (210.825) \\ 100.570 \\ (35.550) \\ 74.440 \\ (13.577) \\ 87.071 \\ (28.228) \\ 60.479 \end{array}$	$\begin{array}{c} 94.898 \\ (51.295) \\ 31.620 \\ (10.019) \\ 23.934 \\ (3.961) \\ 26.487 \\ (6.791) \\ 20.906 \end{array}$	$\begin{array}{c} 1482.696 \\ (781.425) \\ 335.536 \\ (130.849) \\ 247.097 \\ (49.183) \\ 306.880 \\ (109.301) \\ 183.361 \end{array}$	$ \begin{vmatrix} 19.624 \\ (9.657) \\ 9.273 \\ (2.308) \\ 7.380 \\ (1.024) \\ 7.696 \\ (1.210) \\ 7.017 \end{vmatrix} $	$\begin{array}{c} 304.011 \\ (162.312) \\ 81.110 \\ (28.582) \\ 59.996 \\ (11.196) \\ 69.908 \\ (21.598) \\ 48.860 \end{array}$

Table 1.7: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 0.1$ with aggregation matrix \mathbf{C}_2

	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	0.250	0.352	0.779	2.525	0.641	0.289	1.355	0.496	3.647	1.196	14.446	4.043	53.490	0.961	0.982
	(0.009)	(0.028)	(0.235)	(0.915)	(0.186)	(0.013)	(0.243)	(0.048)	(0.781)	(0.359)	(5.968)	(1.558)	(24.565)	(0.295)	(0.292)
CLL	0.273	0.364	0.520	1.456	0.437	0.309	1.244	0.499	3.449	0.770	6.958	2.235	22.434	0.630	0.779
	(0.012)	(0.028)	(0.064)	(0.273)	(0.055)	(0.016)	(0.143)	(0.042)	(0.550)	(0.101)	(1.674)	(0.445)	(6.049)	(0.083)	(0.127)
\mathbf{CLF}	0.273	0.363	0.516	1.385	0.434	0.309	1.228	0.498	3.283	0.760	6.605	2.114	21.333	0.624	0.773
	(0.012)	(0.028)	(0.062)	(0.229)	(0.054)	(0.016)	(0.137)	(0.042)	(0.475)	(0.096)	(1.393)	(0.367)	(4.972)	(0.079)	(0.123)
PCL	0.315	0.393	0.693	1.926	0.583	0.360	1.202	0.522	3.085	1.016	8.936	2.828	30.924	0.833	1.259
	(0.011)	(0.023)	(0.175)	(0.590)	(0.134)	(0.015)	(0.129)	(0.037)	(0.423)	(0.260)	(3.364)	(0.853)	(13.670)	(0.212)	(0.198)
\mathbf{PL}	0.318	0.400	0.530	1.319	0.457	0.363	1.248	0.535	3.210	0.759	5.989	1.979	18.888	0.636	1.078
	(0.011)	(0.025)	(0.054)	(0.196)	(0.042)	(0.016)	(0.141)	(0.041)	(0.461)	(0.087)	(1.147)	(0.307)	(4.137)	(0.068)	(0.068)
\mathbf{PF}	0.319	0.400	0.530	1.319	0.457	0.364	1.248	0.535	3.210	0.759	5.975	1.978	18.599	0.636	1.078
	(0.011)	(0.025)	(0.054)	(0.196)	(0.042)	(0.016)	(0.141)	(0.041)	(0.461)	(0.087)	(1.142)	(0.306)	(3.989)	(0.068)	(0.068)
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	1 0.251	2 0.352	3 0.773	4 2.583	5 0.629	6 0.330	7 2.552	8 0.523	9 4.215	10 1.215	11 15.256	12 4.045	13 55.606	14 3.059	15 3.040
CLCL	1						-								
CLCL CLL	0.251	0.352	0.773	2.583	0.629	0.330	2.552	0.523	4.215	1.215	15.256	4.045	55.606	3.059	3.040
	0.251 (0.009)	0.352 (0.029)	0.773 (0.232)	2.583 (0.992)	$0.629 \\ (0.178)$	0.330 (0.016)	2.552 (0.394)	0.523 (0.050)	4.215 (1.021)	1.215 (0.368)	15.256 (6.423)	4.045 (1.538)	55.606 (26.120)	$\begin{vmatrix} 3.059\\ (0.919) \end{vmatrix}$	3.040 (0.908)
	0.251 (0.009) 0.500	0.352 (0.029) 0.557	$\begin{array}{c c} 0.773 \\ (0.232) \\ 0.671 \end{array}$	2.583 (0.992) 1.512	$\begin{array}{c} 0.629 \\ (0.178) \\ 0.602 \end{array}$	0.330 (0.016) 0.558	$2.552 \\ (0.394) \\ 1.861$	0.523 (0.050) 0.676	$\begin{array}{c} 4.215 \\ (1.021) \\ 3.765 \end{array}$	$ \begin{array}{c c} 1.215 \\ (0.368) \\ 0.892 \end{array} $	$ \begin{array}{c} 15.256 \\ (6.423) \\ 7.210 \end{array} $	$\begin{array}{c} 4.045 \\ (1.538) \\ 2.268 \end{array}$	$55.606 \\ (26.120) \\ 22.684$	$\begin{array}{ c c c } 3.059 \\ (0.919) \\ 2.031 \end{array}$	$3.040 \\ (0.908) \\ 2.061$
CLL	0.251 (0.009) 0.500 (0.018)	0.352 (0.029) 0.557 (0.047)	$ \begin{array}{c c} 0.773 \\ (0.232) \\ 0.671 \\ (0.108) \end{array} $	$\begin{array}{c} 2.583 \\ (0.992) \\ 1.512 \\ (0.332) \end{array}$	$\begin{array}{c c} 0.629 \\ (0.178) \\ 0.602 \\ (0.092) \end{array}$	0.330 (0.016) 0.558 (0.027)	$2.552 \\ (0.394) \\ 1.861 \\ (0.158)$	0.523 (0.050) 0.676 (0.066)	$\begin{array}{c} 4.215 \\ (1.021) \\ 3.765 \\ (0.570) \end{array}$	$ \begin{array}{c c} 1.215 \\ (0.368) \\ 0.892 \\ (0.151) \end{array} $	$15.256 \\ (6.423) \\ 7.210 \\ (1.699)$	$\begin{array}{c} 4.045 \\ (1.538) \\ 2.268 \\ (0.477) \end{array}$	$55.606 \\ (26.120) \\ 22.684 \\ (6.172)$	$\begin{vmatrix} 3.059 \\ (0.919) \\ 2.031 \\ (0.289) \end{vmatrix}$	$\begin{array}{c} 3.040 \\ (0.908) \\ 2.061 \\ (0.312) \end{array}$
CLL	$\begin{array}{ c c c } \textbf{0.251} \\ (0.009) \\ 0.500 \\ (0.018) \\ 0.497 \end{array}$	0.352 (0.029) 0.557 (0.047) 0.553	0.773 (0.232) 0.671 (0.108) 0.666	2.583 (0.992) 1.512 (0.332) 1.450	0.629 (0.178) 0.602 (0.092) 0.596	0.330 (0.016) 0.558 (0.027) 0.554	2.552 (0.394) 1.861 (0.158) 1.831	$\begin{array}{c} \textbf{0.523} \\ (0.050) \\ 0.676 \\ (0.066) \\ 0.671 \end{array}$	$\begin{array}{c} 4.215 \\ (1.021) \\ 3.765 \\ (0.570) \\ 3.575 \end{array}$	1.215 (0.368) 0.892 (0.151) 0.886	$15.256 \\ (6.423) \\ 7.210 \\ (1.699) \\ 6.807$	4.045 (1.538) 2.268 (0.477) 2.168	$55.606 \\ (26.120) \\ 22.684 \\ (6.172) \\ 21.652$	$\begin{array}{c c} & 3.059 \\ (0.919) \\ 2.031 \\ (0.289) \\ 1.954 \end{array}$	3.040 (0.908) 2.061 (0.312) 2.002
CLL CLF	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	0.352 (0.029) 0.557 (0.047) 0.553 (0.047)		$\begin{array}{c} 2.583 \\ (0.992) \\ 1.512 \\ (0.332) \\ 1.450 \\ (0.280) \end{array}$	0.629 (0.178) 0.602 (0.092) 0.596 (0.090)	$\begin{array}{ c c c } \textbf{0.330} \\ (0.016) \\ 0.558 \\ (0.027) \\ 0.554 \\ (0.027) \end{array}$	2.552 (0.394) 1.861 (0.158) 1.831 (0.146)	$\begin{array}{c} \textbf{0.523} \\ (0.050) \\ 0.676 \\ (0.066) \\ 0.671 \\ (0.067) \end{array}$	$\begin{array}{r} 4.215\\(1.021)\\3.765\\(0.570)\\3.575\\(0.476)\end{array}$	$ \begin{vmatrix} 1.215 \\ (0.368) \\ 0.892 \\ (0.151) \\ 0.886 \\ (0.144) \end{vmatrix} $	$\begin{array}{c} 15.256 \\ (6.423) \\ 7.210 \\ (1.699) \\ 6.807 \\ (1.378) \end{array}$	$\begin{array}{c} 4.045 \\ (1.538) \\ 2.268 \\ (0.477) \\ \textbf{2.168} \\ (0.395) \end{array}$	$55.606 \\ (26.120) \\ 22.684 \\ (6.172) \\ 21.652 \\ (5.177)$	$\begin{array}{c c} 3.059 \\ (0.919) \\ 2.031 \\ (0.289) \\ 1.954 \\ (0.238) \end{array}$	$\begin{array}{c} 3.040 \\ (0.908) \\ 2.061 \\ (0.312) \\ \textbf{2.002} \\ (0.269) \end{array}$
CLL CLF	$\begin{array}{ c c c c } \textbf{0.251} \\ (0.009) \\ 0.500 \\ (0.018) \\ 0.497 \\ (0.018) \\ 0.808 \end{array}$	$\begin{array}{c} \textbf{0.352} \\ (0.029) \\ 0.557 \\ (0.047) \\ 0.553 \\ (0.047) \\ 0.841 \end{array}$		$\begin{array}{c} 2.583 \\ (0.992) \\ 1.512 \\ (0.332) \\ 1.450 \\ (0.280) \\ 2.121 \end{array}$		$\begin{array}{ c c c } \textbf{0.330} \\ (0.016) \\ 0.558 \\ (0.027) \\ 0.554 \\ (0.027) \\ 0.924 \end{array}$	$\begin{array}{c} 2.552\\ (0.394)\\ 1.861\\ (0.158)\\ \textbf{1.831}\\ (0.146)\\ 2.121 \end{array}$	$\begin{array}{c} \textbf{0.523} \\ (0.050) \\ 0.676 \\ (0.066) \\ 0.671 \\ (0.067) \\ 0.995 \end{array}$	$\begin{array}{r} 4.215\\(1.021)\\3.765\\(0.570)\\3.575\\(0.476)\\\textbf{3.542}\end{array}$	$ \begin{vmatrix} 1.215 \\ (0.368) \\ 0.892 \\ (0.151) \\ 0.886 \\ (0.144) \\ 1.403 \end{vmatrix} $	$\begin{array}{c} 15.256 \\ (6.423) \\ 7.210 \\ (1.699) \\ 6.807 \\ (1.378) \\ 9.569 \end{array}$	$\begin{array}{r} 4.045\\(1.538)\\2.268\\(0.477)\\\textbf{2.168}\\(0.395)\\3.043\end{array}$	$55.606 \\ (26.120) \\ 22.684 \\ (6.172) \\ 21.652 \\ (5.177) \\ 31.978$	$\begin{vmatrix} 3.059 \\ (0.919) \\ 2.031 \\ (0.289) \\ 1.954 \\ (0.238) \\ 2.494 \end{vmatrix}$	$\begin{array}{c} 3.040 \\ (0.908) \\ 2.061 \\ (0.312) \\ \textbf{2.002} \\ (0.269) \\ 2.670 \end{array}$
CLL CLF PCL	$\begin{array}{ c c c c c } \textbf{0.251} \\ (0.009) \\ 0.500 \\ (0.018) \\ 0.497 \\ (0.018) \\ 0.808 \\ (0.014) \end{array}$	$\begin{array}{c} \textbf{0.352} \\ (0.029) \\ 0.557 \\ (0.047) \\ 0.553 \\ (0.047) \\ 0.841 \\ (0.027) \end{array}$		$\begin{array}{c} 2.583\\ (0.992)\\ 1.512\\ (0.332)\\ \textbf{1.450}\\ (0.280)\\ 2.121\\ (0.601) \end{array}$	$ \begin{vmatrix} 0.629 \\ (0.178) \\ 0.602 \\ (0.092) \\ 0.596 \\ (0.090) \\ 0.994 \\ (0.128) \end{vmatrix} $	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 2.552\\ (0.394)\\ 1.861\\ (0.158)\\ \textbf{1.831}\\ (0.146)\\ 2.121\\ (0.153) \end{array}$	$\begin{array}{c} \textbf{0.523} \\ (0.050) \\ 0.676 \\ (0.066) \\ 0.671 \\ (0.067) \\ 0.995 \\ (0.041) \end{array}$	$\begin{array}{c} 4.215\\(1.021)\\3.765\\(0.570)\\3.575\\(0.476)\\\textbf{3.542}\\(0.435)\end{array}$	$ \begin{vmatrix} 1.215 \\ (0.368) \\ 0.892 \\ (0.151) \\ 0.886 \\ (0.144) \\ 1.403 \\ (0.258) \end{vmatrix} $	$\begin{array}{c} 15.256 \\ (6.423) \\ 7.210 \\ (1.699) \\ 6.807 \\ (1.378) \\ 9.569 \\ (3.441) \end{array}$	$\begin{array}{c} 4.045\\ (1.538)\\ 2.268\\ (0.477)\\ \textbf{2.168}\\ (0.395)\\ 3.043\\ (0.855)\end{array}$	$\begin{array}{c} 55.606\\ (26.120)\\ 22.684\\ (6.172)\\ 21.652\\ (5.177)\\ 31.978\\ (13.961)\end{array}$	$ \begin{vmatrix} 3.059 \\ (0.919) \\ 2.031 \\ (0.289) \\ 1.954 \\ (0.238) \\ 2.494 \\ (0.610) \end{vmatrix} $	$\begin{array}{c} 3.040 \\ (0.908) \\ 2.061 \\ (0.312) \\ \textbf{2.002} \\ (0.269) \\ 2.670 \\ (0.602) \end{array}$
CLL CLF PCL	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} \textbf{0.352} \\ (0.029) \\ 0.557 \\ (0.047) \\ 0.553 \\ (0.047) \\ 0.841 \\ (0.027) \\ 0.855 \end{array}$		$\begin{array}{c} 2.583\\ (0.992)\\ 1.512\\ (0.332)\\ 1.450\\ (0.280)\\ 2.121\\ (0.601)\\ 1.532\end{array}$		$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 2.552\\ (0.394)\\ 1.861\\ (0.158)\\ \textbf{1.831}\\ (0.146)\\ 2.121\\ (0.153)\\ 2.109\end{array}$	$\begin{array}{c} \textbf{0.523} \\ (0.050) \\ 0.676 \\ (0.066) \\ 0.671 \\ (0.067) \\ 0.995 \\ (0.041) \\ 1.015 \end{array}$	$\begin{array}{c} 4.215\\(1.021)\\3.765\\(0.570)\\3.575\\(0.476)\\\textbf{3.542}\\(0.435)\\3.671\end{array}$	$ \begin{vmatrix} 1.215 \\ (0.368) \\ 0.892 \\ (0.151) \\ 0.886 \\ (0.144) \\ 1.403 \\ (0.258) \\ 1.157 \end{vmatrix} $	$\begin{array}{c} 15.256 \\ (6.423) \\ 7.210 \\ (1.699) \\ 6.807 \\ (1.378) \\ 9.569 \\ (3.441) \\ 6.287 \end{array}$	$\begin{array}{c} 4.045\\ (1.538)\\ 2.268\\ (0.477)\\ \textbf{2.168}\\ (0.395)\\ 3.043\\ (0.855)\\ 2.188\end{array}$	$\begin{array}{c} 55.606\\ (26.120)\\ 22.684\\ (6.172)\\ 21.652\\ (5.177)\\ 31.978\\ (13.961)\\ 19.100 \end{array}$	$ \begin{vmatrix} 3.059 \\ (0.919) \\ 2.031 \\ (0.289) \\ 1.954 \\ (0.238) \\ 2.494 \\ (0.610) \\ 1.899 \end{vmatrix} $	$\begin{array}{c} 3.040 \\ (0.908) \\ 2.061 \\ (0.312) \\ \textbf{2.002} \\ (0.269) \\ 2.670 \\ (0.602) \\ 2.098 \end{array}$

Table 1.8: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 1$ with aggregation matrix \mathbf{C}_2

	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	0.794	1.112	2.481	8.157	2.031	0.908	3.913	1.559	11.366	3.760	47.377	12.584	172.990	3.059	37.805
	(0.030)	(0.092)	(0.740)	(3.119)	(0.572)	(0.041)	(0.622)	(0.155)	(2.349)	(1.169)	(20.002)	(4.800)	(76.639)	(0.919)	(15.962)
\mathbf{CLL}	0.801	1.104	1.661	4.611	1.368	0.914	3.877	1.545	10.924	2.466	21.908	7.058	71.039	2.031	17.720
	(0.030)	(0.076)	(0.223)	(0.894)	(0.176)	(0.039)	(0.498)	(0.129)	(1.823)	(0.346)	(5.184)	(1.353)	(19.098)	(0.289)	(4.246)
\mathbf{CLF}	0.800	1.100	1.605	4.362	1.333	0.913	3.712	1.535	10.385	2.368	20.838	6.681	67.981	1.954	16.829
	(0.030)	(0.075)	(0.187)	(0.745)	(0.149)	(0.039)	(0.431)	(0.126)	(1.533)	(0.290)	(4.332)	(1.091)	(16.251)	(0.238)	(3.579)
\mathbf{PCL}	0.810	1.087	2.056	5.645	1.717	0.923	3.535	1.505	9.683	3.038	29.145	8.600	98.808	2.494	23.153
	(0.030)	(0.067)	(0.504)	(1.772)	(0.412)	(0.039)	(0.380)	(0.116)	(1.323)	(0.762)	(11.442)	(2.722)	(41.937)	(0.610)	(8.695)
\mathbf{PL}	0.817	1.116	1.571	4.080	1.305	0.935	3.689	1.549	10.175	2.313	19.051	6.249	60.808	1.899	15.278
	(0.030)	(0.075)	(0.172)	(0.614)	(0.134)	(0.041)	(0.431)	(0.127)	(1.501)	(0.270)	(3.820)	(0.949)	(13.538)	(0.212)	(2.944)
\mathbf{PF}	0.818	1.116	1.571	4.078	1.305	0.935	3.689	1.549	10.173	2.313	18.836	6.240	58.780	1.899	15.142
	(0.030)	(0.075)	(0.172)	(0.613)	(0.134)	(0.041)	(0.431)	(0.127)	(1.500)	(0.270)	(3.740)	(0.945)	(12.559)	(0.212)	(2.890)
	1	-													
	L 1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	0.795	2 1.118	2.445	4 8.320	5 2.018	6 0.923	7 4.475	8 1.580	9 11.780	10 3.769	11 47.182	12 12.899	13 171.951	14 3.040	15 36.896
CLCL	· -						-								
CLCL CLL	0.795	1.118	2.445	8.320	2.018	0.923	4.475	1.580	11.780	3.769	47.182	12.899	171.951	3.040	36.896
	0.795 (0.028)	1.118 (0.090)	2.445 (0.743)	8.320 (3.158)	2.018 (0.563)	0.923 (0.042)	4.475 (0.858)	1.580 (0.155)	11.780 (2.682)	3.769 (1.140)	47.182 (19.593)	12.899 (5.017)	171.951 (76.603)	$\begin{vmatrix} 3.040 \\ (0.908) \end{vmatrix}$	36.896 (14.953)
	0.795 (0.028) 0.905	1.118 (0.090) 1.185	$ \begin{array}{c c} 2.445 \\ (0.743) \\ 1.691 \end{array} $	8.320 (3.158) 4.639	$ \begin{array}{c c} 2.018 \\ (0.563) \\ 1.418 \end{array} $	0.923 (0.042) 1.024	$ \begin{array}{r} 4.475 \\ (0.858) \\ 4.116 \end{array} $	1.580 (0.155) 1.613	$ \begin{array}{c} 11.780 \\ (2.682) \\ 11.118 \end{array} $	$3.769 \\ (1.140) \\ 2.516$	$\begin{array}{c} 47.182 \\ (19.593) \\ 22.259 \end{array}$	$ \begin{array}{r} 12.899\\(5.017)\\7.155\end{array} $	$171.951 \\ (76.603) \\ 71.403$	$\begin{array}{ c c c } 3.040 \\ (0.908) \\ 2.061 \end{array}$	36.896 (14.953) 17.990
CLL	0.795 (0.028) 0.905 (0.038)	1.118 (0.090) 1.185 (0.098)	$ \begin{array}{c c} 2.445 \\ (0.743) \\ 1.691 \\ (0.246) \end{array} $	$8.320 \\ (3.158) \\ 4.639 \\ (0.930)$	$ \begin{array}{c c} 2.018 \\ (0.563) \\ 1.418 \\ (0.212) \end{array} $	0.923 (0.042) 1.024 (0.054)	$\begin{array}{c} 4.475 \\ (0.858) \\ 4.116 \\ (0.508) \end{array}$	1.580 (0.155) 1.613 (0.146)	$ \begin{array}{c} 11.780 \\ (2.682) \\ 11.118 \\ (1.809) \end{array} $	$\begin{array}{c} 3.769 \\ (1.140) \\ 2.516 \\ (0.399) \end{array}$	$\begin{array}{c} 47.182 \\ (19.593) \\ 22.259 \\ (5.209) \end{array}$	$12.899 \\ (5.017) \\ 7.155 \\ (1.413)$	$171.951 \\ (76.603) \\ 71.403 \\ (19.441)$	$\begin{vmatrix} 3.040 \\ (0.908) \\ 2.061 \\ (0.312) \end{vmatrix}$	$36.896 \\ (14.953) \\ 17.990 \\ (4.084)$
CLL	0.795 (0.028) 0.905 (0.038) 0.903	1.118 (0.090) 1.185 (0.098) 1.180	$ \begin{vmatrix} 2.445 \\ (0.743) \\ 1.691 \\ (0.246) \\ \textbf{1.648} \end{vmatrix} $	$8.320 \\ (3.158) \\ 4.639 \\ (0.930) \\ 4.399$	2.018 (0.563) 1.418 (0.212) 1.399	$\begin{array}{c c} 0.923 \\ (0.042) \\ 1.024 \\ (0.054) \\ 1.021 \end{array}$	$\begin{array}{c} 4.475\\(0.858)\\4.116\\(0.508)\\3.949\end{array}$	1.580 (0.155) 1.613 (0.146) 1.604	11.780 (2.682) 11.118 (1.809) 10.502	3.769 (1.140) 2.516 (0.399) 2.429	$\begin{array}{c} 47.182 \\ (19.593) \\ 22.259 \\ (5.209) \\ 21.026 \end{array}$	$12.899 \\ (5.017) \\ 7.155 \\ (1.413) \\ 6.773$	$171.951 \\ (76.603) \\ 71.403 \\ (19.441) \\ 68.259$	3.040 (0.908) 2.061 (0.312) 2.002	$\begin{array}{r} 36.896 \\ (14.953) \\ 17.990 \\ (4.084) \\ 16.985 \end{array}$
CLL CLF	$\begin{array}{c c} 0.795 \\ (0.028) \\ 0.905 \\ (0.038) \\ 0.903 \\ (0.038) \end{array}$	$\begin{array}{c} \textbf{1.118} \\ (0.090) \\ 1.185 \\ (0.098) \\ 1.180 \\ (0.097) \end{array}$	$\begin{array}{c c} 2.445 \\ (0.743) \\ 1.691 \\ (0.246) \\ \textbf{1.648} \\ (0.209) \end{array}$	$\begin{array}{c} 8.320 \\ (3.158) \\ 4.639 \\ (0.930) \\ 4.399 \\ (0.751) \end{array}$	2.018 (0.563) 1.418 (0.212) 1.399 (0.188)	$\begin{array}{c c} 0.923 \\ (0.042) \\ 1.024 \\ (0.054) \\ 1.021 \\ (0.054) \end{array}$	$\begin{array}{c} 4.475\\ (0.858)\\ 4.116\\ (0.508)\\ 3.949\\ (0.436)\end{array}$	$\begin{array}{c} \textbf{1.580} \\ (0.155) \\ 1.613 \\ (0.146) \\ 1.604 \\ (0.142) \end{array}$	$11.780 \\ (2.682) \\ 11.118 \\ (1.809) \\ 10.502 \\ (1.535)$	3.769 (1.140) 2.516 (0.399) 2.429 (0.330)	$\begin{array}{c} 47.182 \\ (19.593) \\ 22.259 \\ (5.209) \\ 21.026 \\ (4.357) \end{array}$	$\begin{array}{c} 12.899\\ (5.017)\\ 7.155\\ (1.413)\\ 6.773\\ (1.151)\end{array}$	$\begin{array}{c} 171.951 \\ (76.603) \\ 71.403 \\ (19.441) \\ 68.259 \\ (16.641) \end{array}$	3.040 (0.908) 2.061 (0.312) 2.002 (0.269)	$\begin{array}{r} 36.896 \\ (14.953) \\ 17.990 \\ (4.084) \\ 16.985 \\ (3.371) \end{array}$
CLL CLF	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	1.118 (0.090) 1.185 (0.098) 1.180 (0.097) 1.323	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 8.320 \\ (3.158) \\ 4.639 \\ (0.930) \\ 4.399 \\ (0.751) \\ 5.806 \end{array}$	2.018 (0.563) 1.418 (0.212) 1.399 (0.188) 1.895	$\begin{array}{c c} \textbf{0.923} \\ (0.042) \\ 1.024 \\ (0.054) \\ 1.021 \\ (0.054) \\ 1.255 \end{array}$	$\begin{array}{c} 4.475\\ (0.858)\\ 4.116\\ (0.508)\\ 3.949\\ (0.436)\\ \textbf{3.943}\end{array}$	$\begin{array}{c} \textbf{1.580} \\ (0.155) \\ 1.613 \\ (0.146) \\ 1.604 \\ (0.142) \\ 1.731 \end{array}$	11.780 (2.682) 11.118 (1.809) 10.502 (1.535) 9.770	$\begin{array}{c} 3.769 \\ (1.140) \\ 2.516 \\ (0.399) \\ \textbf{2.429} \\ (0.330) \\ 3.164 \end{array}$	$\begin{array}{r} 47.182\\(19.593)\\22.259\\(5.209)\\21.026\\(4.357)\\28.594\end{array}$	$\begin{array}{c} 12.899\\ (5.017)\\ 7.155\\ (1.413)\\ 6.773\\ (1.151)\\ 8.705\end{array}$	$\begin{array}{c} 171.951 \\ (76.603) \\ 71.403 \\ (19.441) \\ 68.259 \\ (16.641) \\ 98.241 \end{array}$	$ \begin{vmatrix} 3.040 \\ (0.908) \\ 2.061 \\ (0.312) \\ 2.002 \\ (0.269) \\ 2.670 \end{vmatrix} $	$\begin{array}{r} 36.896 \\ (14.953) \\ 17.990 \\ (4.084) \\ 16.985 \\ (3.371) \\ 23.062 \end{array}$
CLL CLF PCL	$\begin{array}{c c} \textbf{0.795} \\ (0.028) \\ 0.905 \\ (0.038) \\ 0.903 \\ (0.038) \\ 1.103 \\ (0.036) \end{array}$	$\begin{array}{c} \textbf{1.118} \\ (0.090) \\ 1.185 \\ (0.098) \\ 1.180 \\ (0.097) \\ 1.323 \\ (0.076) \end{array}$	$\begin{array}{c c} 2.445 \\ (0.743) \\ 1.691 \\ (0.246) \\ \textbf{1.648} \\ (0.209) \\ 2.194 \\ (0.481) \end{array}$	$\begin{array}{c} 8.320 \\ (3.158) \\ 4.639 \\ (0.930) \\ 4.399 \\ (0.751) \\ 5.806 \\ (1.749) \end{array}$	$ \begin{vmatrix} 2.018 \\ (0.563) \\ 1.418 \\ (0.212) \\ 1.399 \\ (0.188) \\ 1.895 \\ (0.393) \end{vmatrix} $	$\begin{array}{c c} \textbf{0.923} \\ (0.042) \\ 1.024 \\ (0.054) \\ 1.021 \\ (0.054) \\ 1.255 \\ (0.050) \end{array}$	$\begin{array}{c} 4.475\\ (0.858)\\ 4.116\\ (0.508)\\ 3.949\\ (0.436)\\ \textbf{3.943}\\ (0.394)\end{array}$	$\begin{array}{c} \textbf{1.580} \\ (0.155) \\ 1.613 \\ (0.146) \\ 1.604 \\ (0.142) \\ 1.731 \\ (0.117) \end{array}$	$\begin{array}{c} 11.780\\ (2.682)\\ 11.118\\ (1.809)\\ 10.502\\ (1.535)\\ \textbf{9.770}\\ (1.356)\end{array}$	$\begin{array}{c} 3.769 \\ (1.140) \\ 2.516 \\ (0.399) \\ \textbf{2.429} \\ (0.330) \\ 3.164 \\ (0.712) \end{array}$	$\begin{array}{c} 47.182 \\ (19.593) \\ 22.259 \\ (5.209) \\ 21.026 \\ (4.357) \\ 28.594 \\ (10.854) \end{array}$	$\begin{array}{c} 12.899\\ (5.017)\\ 7.155\\ (1.413)\\ 6.773\\ (1.151)\\ 8.705\\ (2.782)\end{array}$	$\begin{array}{c} 171.951 \\ (76.603) \\ 71.403 \\ (19.441) \\ 68.259 \\ (16.641) \\ 98.241 \\ (41.673) \end{array}$	$ \begin{vmatrix} 3.040 \\ (0.908) \\ 2.061 \\ (0.312) \\ 2.002 \\ (0.269) \\ 2.670 \\ (0.602) \end{vmatrix} $	$\begin{array}{c} 36.896 \\ (14.953) \\ 17.990 \\ (4.084) \\ 16.985 \\ (3.371) \\ 23.062 \\ (8.425) \end{array}$
CLL CLF PCL	$\begin{array}{c c} & 0.795 \\ (0.028) \\ 0.905 \\ (0.038) \\ 0.903 \\ (0.038) \\ 1.103 \\ (0.036) \\ 1.116 \end{array}$	$\begin{array}{c} \textbf{1.118} \\ (0.090) \\ 1.185 \\ (0.098) \\ 1.180 \\ (0.097) \\ 1.323 \\ (0.076) \\ 1.353 \end{array}$	$ \begin{vmatrix} 2.445 \\ (0.743) \\ 1.691 \\ (0.246) \\ \textbf{1.648} \\ (0.209) \\ 2.194 \\ (0.481) \\ 1.735 \end{vmatrix} $	$\begin{array}{c} 8.320\\ (3.158)\\ 4.639\\ (0.930)\\ 4.399\\ (0.751)\\ 5.806\\ (1.749)\\ 4.176\end{array}$	$ \begin{vmatrix} 2.018 \\ (0.563) \\ 1.418 \\ (0.212) \\ 1.399 \\ (0.188) \\ 1.895 \\ (0.393) \\ 1.519 \end{vmatrix} $	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 4.475\\ (0.858)\\ 4.116\\ (0.508)\\ 3.949\\ (0.436)\\ \textbf{3.943}\\ (0.394)\\ 4.061\end{array}$	$\begin{array}{c} \textbf{1.580} \\ (0.155) \\ 1.613 \\ (0.146) \\ 1.604 \\ (0.142) \\ 1.731 \\ (0.117) \\ 1.776 \end{array}$	$\begin{array}{c} 11.780\\ (2.682)\\ 11.118\\ (1.809)\\ 10.502\\ (1.535)\\ \textbf{9.770}\\ (1.356)\\ 10.205\end{array}$	$\begin{array}{c} 3.769 \\ (1.140) \\ 2.516 \\ (0.399) \\ \textbf{2.429} \\ (0.330) \\ 3.164 \\ (0.712) \\ 2.479 \end{array}$	$\begin{array}{c} 47.182\\(19.593)\\22.259\\(5.209)\\21.026\\(4.357)\\28.594\\(10.854)\\19.134\end{array}$	$\begin{array}{c} 12.899\\ (5.017)\\ 7.155\\ (1.413)\\ 6.773\\ (1.151)\\ 8.705\\ (2.782)\\ 6.373\end{array}$	$\begin{array}{c} 171.951\\ (76.603)\\ 71.403\\ (19.441)\\ 68.259\\ (16.641)\\ 98.241\\ (41.673)\\ 60.974 \end{array}$	$ \begin{vmatrix} 3.040 \\ (0.908) \\ 2.061 \\ (0.312) \\ 2.002 \\ (0.269) \\ 2.670 \\ (0.602) \\ 2.098 \end{vmatrix} $	$\begin{array}{c} 36.896 \\ (14.953) \\ 17.990 \\ (4.084) \\ 16.985 \\ (3.371) \\ 23.062 \\ (8.425) \\ 15.509 \end{array}$

Table 1.9: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 5$ with aggregation matrix \mathbf{C}_2

	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	1.772	2.499	5.449	18.449	4.525	2.032	8.704	3.516	25.825	8.299	104.273	28.350	383.173	6.769	6.795
	(0.064)	(0.208)	(1.593)	(6.985)	(1.288)	(0.094)	(1.320)	(0.343)	(5.581)	(2.549)	(44.026)	(10.424)	(167.648)	(2.054)	(2.052)
\mathbf{CLL}	1.773	2.488	3.740	10.320	3.088	2.032	8.594	3.533	24.771	5.498	49.146	15.863	161.116	4.527	4.532
	(0.063)	(0.180)	(0.500)	(2.039)	(0.402)	(0.089)	(1.091)	(0.304)	(3.951)	(0.812)	(11.585)	(3.185)	(44.060)	(0.629)	(0.656)
\mathbf{CLF}	1.773	2.458	3.596	9.766	2.981	2.029	8.246	3.446	23.419	5.279	46.432	14.989	152.345	4.347	4.367
	(0.063)	(0.168)	(0.416)	(1.675)	(0.340)	(0.089)	(0.943)	(0.268)	(3.359)	(0.664)	(9.569)	(2.576)	(36.380)	(0.520)	(0.542)
\mathbf{PCL}	1.770	2.402	4.359	12.531	3.675	2.020	7.847	3.347	21.816	6.430	63.748	19.359	221.670	5.258	5.364
	(0.062)	(0.151)	(0.945)	(4.006)	(0.819)	(0.086)	(0.862)	(0.240)	(3.006)	(1.505)	(24.598)	(6.083)	(93.908)	(1.206)	(1.225)
\mathbf{PL}	1.783	2.468	3.500	9.168	2.915	2.044	8.190	3.459	22.803	5.121	42.994	14.076	138.351	4.210	4.302
	(0.064)	(0.170)	(0.388)	(1.437)	(0.314)	(0.090)	(0.959)	(0.272)	(3.337)	(0.621)	(8.397)	(2.292)	(31.881)	(0.465)	(0.470)
\mathbf{PF}	1.783	2.468	3.500	9.146	2.915	2.044	8.189	3.459	22.785	5.119	41.986	14.008	132.771	4.209	4.301
	(0.064)	(0.170)	(0.387)	(1.429)	(0.313)	(0.090)	(0.959)	(0.272)	(3.331)	(0.620)	(7.997)	(2.263)	(28.832)	(0.464)	(0.470)
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	1 1.776	2 2.487	3 5.522	4 18.232	5 4.496	6 2.040	7 9.027	8 3.481	9 25.603	10 8.424	11 102.861	12 28.329	13 384.419	14 9.575	15 9.557
CLCL							-								
CLCL	1.776	2.487	5.522	18.232	4.496	2.040	9.027	3.481	25.603	8.424	102.861	28.329	384.419	9.575	9.557
	1.776 (0.062)	2.487 (0.197)	5.522 (1.705)	18.232 (6.883)	$ 4.496 \\ (1.235)$	2.040 (0.095)	9.027 (1.563)	3.481 (0.341)	25.603 (5.547)	8.424 (2.654)	102.861 (45.064)	28.329 (10.586)	384.419 (175.072)	9.575 (2.851)	9.557 (2.806)
	1.776 (0.062) 1.828	2.487 (0.197) 2.514	$ \begin{array}{c c} 5.522 \\ (1.705) \\ 3.721 \end{array} $	$18.232 \\ (6.883) \\ 10.304$	$ \begin{array}{ c c c c c } 4.496 \\ (1.235) \\ 3.092 \\ \end{array} $	2.040 (0.095) 2.087	9.027 (1.563) 8.730	$3.481 \\ (0.341) \\ 3.514$	25.603 (5.547) 24.713	$ \begin{array}{c c} 8.424 \\ (2.654) \\ 5.569 \end{array} $	$102.861 \\ (45.064) \\ 48.838$	$28.329 \\ (10.586) \\ 15.979$	$384.419 \\ (175.072) \\ 160.465$	$ \begin{array}{c c} 9.575 \\ (2.851) \\ 6.434 \end{array} $	9.557 (2.806) 6.399
CLL	$\begin{array}{ c c c } 1.776 \\ (0.062) \\ 1.828 \\ (0.067) \end{array}$	2.487 (0.197) 2.514 (0.190)	$ \begin{array}{c c} 5.522 \\ (1.705) \\ 3.721 \\ (0.498) \end{array} $	$18.232 \\ (6.883) \\ 10.304 \\ (2.033)$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	2.040 (0.095) 2.087 (0.102)	$9.027 \\ (1.563) \\ 8.730 \\ (1.131)$	$\begin{array}{c} 3.481 \\ (0.341) \\ 3.514 \\ (0.311) \end{array}$	$25.603 \\ (5.547) \\ 24.713 \\ (4.094)$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$102.861 \\ (45.064) \\ 48.838 \\ (11.661)$	$28.329 \\ (10.586) \\ 15.979 \\ (3.089)$	$\begin{array}{c} 384.419 \\ (175.072) \\ 160.465 \\ (44.769) \end{array}$	$ \begin{array}{ c c c c } 9.575 \\ (2.851) \\ 6.434 \\ (0.897) \end{array} $	$9.557 \\ (2.806) \\ 6.399 \\ (0.899)$
CLL	$\begin{array}{ c c c } & 1.776 \\ & (0.062) \\ & 1.828 \\ & (0.067) \\ & 1.826 \end{array}$	2.487 (0.197) 2.514 (0.190) 2.487	$ \begin{array}{c c} 5.522 \\ (1.705) \\ 3.721 \\ (0.498) \\ 3.583 \end{array} $	$18.232 \\ (6.883) \\ 10.304 \\ (2.033) \\ 9.762$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	2.040 (0.095) 2.087 (0.102) 2.083	$9.027 \\ (1.563) \\ 8.730 \\ (1.131) \\ 8.366$	$\begin{array}{r} 3.481 \\ (0.341) \\ 3.514 \\ (0.311) \\ 3.441 \end{array}$	$25.603 \\ (5.547) \\ 24.713 \\ (4.094) \\ 23.328$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$102.861 \\ (45.064) \\ 48.838 \\ (11.661) \\ 46.212$	$\begin{array}{c} 28.329 \\ (10.586) \\ 15.979 \\ (3.089) \\ 15.082 \end{array}$	$\begin{array}{c} 384.419 \\ (175.072) \\ 160.465 \\ (44.769) \\ 151.584 \end{array}$	$ \begin{array}{ c c c c c } 9.575 \\ (2.851) \\ 6.434 \\ (0.897) \\ 6.163 \end{array} $	$9.557 \\ (2.806) \\ 6.399 \\ (0.899) \\ 6.152$
CLL CLF	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} \textbf{2.487} \\ (0.197) \\ 2.514 \\ (0.190) \\ 2.487 \\ (0.178) \end{array}$	$\begin{array}{c c} 5.522 \\ (1.705) \\ 3.721 \\ (0.498) \\ 3.583 \\ (0.416) \end{array}$	$18.232 \\ (6.883) \\ 10.304 \\ (2.033) \\ 9.762 \\ (1.658)$	$\begin{array}{c c} & 4.496 \\ (1.235) \\ 3.092 \\ (0.400) \\ \textbf{2.993} \\ (0.338) \end{array}$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} 9.027 \\ (1.563) \\ 8.730 \\ (1.131) \\ 8.366 \\ (0.971) \end{array}$	$\begin{array}{c} 3.481 \\ (0.341) \\ 3.514 \\ (0.311) \\ 3.441 \\ (0.280) \end{array}$	$\begin{array}{c} 25.603 \\ (5.547) \\ 24.713 \\ (4.094) \\ 23.328 \\ (3.441) \end{array}$	$ \begin{vmatrix} 8.424 \\ (2.654) \\ 5.569 \\ (0.825) \\ 5.332 \\ (0.662) \end{vmatrix} $	$102.861 \\ (45.064) \\ 48.838 \\ (11.661) \\ 46.212 \\ (9.638)$	$\begin{array}{c} 28.329 \\ (10.586) \\ 15.979 \\ (3.089) \\ 15.082 \\ (2.498) \end{array}$	$\begin{array}{r} 384.419 \\ (175.072) \\ 160.465 \\ (44.769) \\ 151.584 \\ (36.610) \end{array}$	$\begin{array}{c c} 9.575 \\ (2.851) \\ 6.434 \\ (0.897) \\ 6.163 \\ (0.730) \end{array}$	$\begin{array}{c} 9.557 \\ (2.806) \\ 6.399 \\ (0.899) \\ 6.152 \\ (0.760) \end{array}$
CLL CLF	$\begin{array}{ c c c } & 1.776 \\ (0.062) \\ 1.828 \\ (0.067) \\ 1.826 \\ (0.067) \\ 1.927 \end{array}$	$\begin{array}{c} \textbf{2.487} \\ (0.197) \\ 2.514 \\ (0.190) \\ 2.487 \\ (0.178) \\ 2.513 \end{array}$	$\begin{array}{c c} 5.522 \\ (1.705) \\ 3.721 \\ (0.498) \\ 3.583 \\ (0.416) \\ 4.467 \end{array}$	$18.232 \\ (6.883) \\ 10.304 \\ (2.033) \\ 9.762 \\ (1.658) \\ 12.520$	4.496 (1.235) 3.092 (0.400) 2.993 (0.338) 3.746	$\begin{array}{ c c c } \hline \textbf{2.040} \\ (0.095) \\ 2.087 \\ (0.102) \\ 2.083 \\ (0.102) \\ 2.195 \end{array}$	$\begin{array}{c} 9.027\\(1.563)\\8.730\\(1.131)\\8.366\\(0.971)\\\textbf{8.033}\end{array}$	$\begin{array}{c} 3.481 \\ (0.341) \\ 3.514 \\ (0.311) \\ 3.441 \\ (0.280) \\ 3.420 \end{array}$	$\begin{array}{c} 25.603 \\ (5.547) \\ 24.713 \\ (4.094) \\ 23.328 \\ (3.441) \\ \textbf{21.732} \end{array}$	$ \begin{vmatrix} 8.424 \\ (2.654) \\ 5.569 \\ (0.825) \\ 5.332 \\ (0.662) \\ 6.516 \end{vmatrix} $	$\begin{array}{c} 102.861 \\ (45.064) \\ 48.838 \\ (11.661) \\ 46.212 \\ (9.638) \\ 63.392 \end{array}$	$\begin{array}{c} 28.329 \\ (10.586) \\ 15.979 \\ (3.089) \\ 15.082 \\ (2.498) \\ 19.297 \end{array}$	$\begin{array}{c} 384.419 \\ (175.072) \\ 160.465 \\ (44.769) \\ 151.584 \\ (36.610) \\ 220.483 \end{array}$	$\begin{array}{ c c c } 9.575 \\ (2.851) \\ 6.434 \\ (0.897) \\ 6.163 \\ (0.730) \\ 7.409 \end{array}$	$\begin{array}{c} 9.557 \\ (2.806) \\ 6.399 \\ (0.899) \\ 6.152 \\ (0.760) \\ 7.464 \end{array}$
CLL CLF PCL	$\begin{vmatrix} 1.776 \\ (0.062) \\ 1.828 \\ (0.067) \\ 1.826 \\ (0.067) \\ 1.927 \\ (0.066) \end{vmatrix}$	$\begin{array}{c} \textbf{2.487} \\ (0.197) \\ 2.514 \\ (0.190) \\ 2.487 \\ (0.178) \\ 2.513 \\ (0.157) \end{array}$	$\begin{vmatrix} 5.522 \\ (1.705) \\ 3.721 \\ (0.498) \\ 3.583 \\ (0.416) \\ 4.467 \\ (1.002) \end{vmatrix}$	$18.232 \\ (6.883) \\ 10.304 \\ (2.033) \\ 9.762 \\ (1.658) \\ 12.520 \\ (4.025) \\ \end{cases}$	$ \begin{vmatrix} 4.496 \\ (1.235) \\ 3.092 \\ (0.400) \\ 2.993 \\ (0.338) \\ 3.746 \\ (0.759) \end{vmatrix} $	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 9.027\\(1.563)\\8.730\\(1.131)\\8.366\\(0.971)\\\textbf{8.033}\\(0.873)\end{array}$	$\begin{array}{c} 3.481 \\ (0.341) \\ 3.514 \\ (0.311) \\ 3.441 \\ (0.280) \\ \textbf{3.420} \\ (0.253) \end{array}$	$\begin{array}{c} 25.603 \\ (5.547) \\ 24.713 \\ (4.094) \\ 23.328 \\ (3.441) \\ \textbf{21.732} \\ (2.984) \end{array}$	$ \begin{vmatrix} 8.424 \\ (2.654) \\ 5.569 \\ (0.825) \\ 5.332 \\ (0.662) \\ 6.516 \\ (1.515) \end{vmatrix} $	$\begin{array}{c} 102.861 \\ (45.064) \\ 48.838 \\ (11.661) \\ 46.212 \\ (9.638) \\ 63.392 \\ (24.063) \end{array}$	$\begin{array}{c} 28.329 \\ (10.586) \\ 15.979 \\ (3.089) \\ 15.082 \\ (2.498) \\ 19.297 \\ (5.997) \end{array}$	$\begin{array}{c} 384.419 \\ (175.072) \\ 160.465 \\ (44.769) \\ 151.584 \\ (36.610) \\ 220.483 \\ (98.903) \end{array}$	$ \begin{vmatrix} 9.575 \\ (2.851) \\ 6.434 \\ (0.897) \\ 6.163 \\ (0.730) \\ 7.409 \\ (1.670) \end{vmatrix} $	$\begin{array}{c} 9.557\\ (2.806)\\ 6.399\\ (0.899)\\ 6.152\\ (0.760)\\ 7.464\\ (1.576)\end{array}$
CLL CLF PCL	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} \textbf{2.487} \\ (0.197) \\ 2.514 \\ (0.190) \\ 2.487 \\ (0.178) \\ 2.513 \\ (0.157) \\ 2.581 \end{array}$	$ \begin{vmatrix} 5.522 \\ (1.705) \\ 3.721 \\ (0.498) \\ 3.583 \\ (0.416) \\ 4.467 \\ (1.002) \\ 3.548 \end{vmatrix} $	$\begin{array}{c} 18.232 \\ (6.883) \\ 10.304 \\ (2.033) \\ 9.762 \\ (1.658) \\ 12.520 \\ (4.025) \\ 9.230 \end{array}$	$ \begin{vmatrix} 4.496 \\ (1.235) \\ 3.092 \\ (0.400) \\ \textbf{2.993} \\ (0.338) \\ 3.746 \\ (0.759) \\ 2.996 \end{vmatrix} $	$\begin{array}{ c c c c c } \hline \textbf{2.040} \\ \hline (0.095) \\ \hline 2.087 \\ \hline (0.102) \\ \hline 2.083 \\ \hline (0.102) \\ \hline 2.195 \\ \hline (0.098) \\ \hline 2.223 \end{array}$	9.027 (1.563) 8.730 (1.131) 8.366 (0.971) 8.033 (0.873) 8.352	$\begin{array}{c} 3.481 \\ (0.341) \\ 3.514 \\ (0.311) \\ 3.441 \\ (0.280) \\ \textbf{3.420} \\ (0.253) \\ 3.525 \end{array}$	$\begin{array}{c} 25.603 \\ (5.547) \\ 24.713 \\ (4.094) \\ 23.328 \\ (3.441) \\ \textbf{21.732} \\ (2.984) \\ 22.792 \end{array}$	$ \begin{vmatrix} 8.424 \\ (2.654) \\ 5.569 \\ (0.825) \\ 5.332 \\ (0.662) \\ 6.516 \\ (1.515) \\ 5.215 \end{vmatrix} $	$\begin{array}{c} 102.861 \\ (45.064) \\ 48.838 \\ (11.661) \\ 46.212 \\ (9.638) \\ 63.392 \\ (24.063) \\ 42.959 \end{array}$	$\begin{array}{c} 28.329 \\ (10.586) \\ 15.979 \\ (3.089) \\ 15.082 \\ (2.498) \\ 19.297 \\ (5.997) \\ 14.179 \end{array}$	$\begin{array}{c} 384.419\\ (175.072)\\ 160.465\\ (44.769)\\ 151.584\\ (36.610)\\ 220.483\\ (98.903)\\ 137.539\end{array}$	$ \begin{vmatrix} 9.575 \\ (2.851) \\ 6.434 \\ (0.897) \\ 6.163 \\ (0.730) \\ 7.409 \\ (1.670) \\ 5.974 \end{vmatrix} $	$\begin{array}{c} 9.557\\ (2.806)\\ 6.399\\ (0.899)\\ 6.152\\ (0.760)\\ 7.464\\ (1.576)\\ 6.025\end{array}$

Table 1.10: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 10$ with aggregation matrix \mathbf{C}_2

	5	6	13	14	19	29	30	31	32	45	46	47	48	57	58
CLCL	2.507	3.539	7.824	25.542	6.338	2.882	12.307	4.924	36.608	11.841	144.492	39.876	535.446	9.575	116.858
	(0.092)	(0.296)	(2.286)	(9.421)	(1.775)	(0.131)	(1.994)	(0.452)	(7.680)	(3.489)	(61.102)	(15.351)	(248.153)	(2.851)	(48.113)
\mathbf{CLL}	2.507	3.537	5.239	14.554	4.367	2.884	12.141	4.974	34.861	7.752	69.085	22.497	225.402	6.434	55.569
	(0.091)	(0.263)	(0.698)	(2.850)	(0.543)	(0.126)	(1.531)	(0.428)	(5.702)	(1.122)	(15.859)	(4.563)	(58.453)	(0.897)	(13.358)
\mathbf{CLF}	2.505	3.475	5.053	13.805	4.206	2.877	11.649	4.851	33.069	7.432	65.384	21.183	214.439	6.163	52.655
	(0.090)	(0.236)	(0.589)	(2.386)	(0.458)	(0.124)	(1.309)	(0.384)	(4.800)	(0.922)	(13.146)	(3.620)	(50.318)	(0.730)	(11.170)
\mathbf{PCL}	2.497	3.397	6.084	17.498	5.096	2.857	11.093	4.723	30.786	9.027	89.593	26.792	302.393	7.409	72.545
	(0.089)	(0.214)	(1.330)	(5.353)	(1.102)	(0.120)	(1.176)	(0.352)	(4.325)	(2.111)	(33.468)	(8.648)	(133.069)	(1.670)	(26.654)
\mathbf{PL}	2.514	3.484	4.919	13.008	4.114	2.891	11.564	4.881	32.160	7.218	61.113	19.943	194.459	5.974	48.765
	(0.091)	(0.240)	(0.563)	(2.029)	(0.427)	(0.127)	(1.295)	(0.395)	(4.823)	(0.875)	(11.786)	(3.138)	(42.523)	(0.692)	(10.173)
\mathbf{PF}	2.514	3.484	4.917	12.955	4.113	2.891	11.563	4.881	32.115	7.213	59.276	19.801	186.184	5.971	47.516
	(0.091)	(0.240)	(0.563)	(2.009)	(0.426)	(0.127)	(1.294)	(0.395)	(4.804)	(0.873)	(11.059)	(3.065)	(39.389)	(0.691)	(9.584)
	7	8	15	16	20	33	34	35	36	49	50	51	52	59	60
CLCL	7 2.515	8 3.516	15 7.787	16 25.714	20 6.398	33 2.879	34 12.549	35 4.936	<mark>36</mark> 36.293	49 11.995	50 146.923	51 40.520	52 547.736	59 9.557	60 117.372
CLCL			1			I									
CLCL	2.515	3.516	7.787	25.714	6.398	2.879	12.549	4.936	36.293	11.995	146.923	40.520	547.736	9.557	117.372
	2.515 (0.089)	3.516 (0.282)	7.787 (2.321)	25.714 (9.640)	6.398 (1.877)	2.879 (0.132)	$12.549 \\ (1.969)$	4.936 (0.481)	36.293 (8.057)	$ \begin{array}{c c} 11.995 \\ (3.862) \end{array} $	146.923 (63.184)	40.520 (14.891)	547.736 (246.160)	$ \begin{array}{c c} 9.557 \\ (2.806) \end{array} $	$ \begin{array}{c} 117.372 \\ (47.245) \end{array} $
	2.515 (0.089) 2.551	3.516 (0.282) 3.551	$ \begin{array}{c c} 7.787 \\ (2.321) \\ 5.281 \end{array} $	25.714 (9.640) 14.766	$ \begin{array}{c c} 6.398 \\ (1.877) \\ 4.408 \end{array} $	2.879 (0.132) 2.911	$12.549 \\ (1.969) \\ 12.353$	$\begin{array}{c} 4.936 \\ (0.481) \\ 4.978 \end{array}$	36.293 (8.057) 34.633	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$146.923 \\ (63.184) \\ 69.750$	$\begin{array}{c} 40.520 \\ (14.891) \\ 22.365 \end{array}$	$547.736 \\ (246.160) \\ 225.422$	$ \begin{array}{c c} 9.557 \\ (2.806) \\ 6.399 \end{array} $	$ \begin{array}{r} 117.372 \\ (47.245) \\ 55.847 \end{array} $
CLL	2.515 (0.089) 2.551 (0.094)	$\begin{array}{c} 3.516 \\ (0.282) \\ 3.551 \\ (0.263) \end{array}$	$\begin{array}{ c c c } 7.787 \\ (2.321) \\ 5.281 \\ (0.758) \end{array}$	$25.714 \\ (9.640) \\ 14.766 \\ (3.078)$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	2.879 (0.132) 2.911 (0.134)	$12.549 \\ (1.969) \\ 12.353 \\ (1.614)$	$\begin{array}{c} 4.936 \\ (0.481) \\ 4.978 \\ (0.439) \end{array}$	$\begin{array}{c} 36.293 \\ (8.057) \\ 34.633 \\ (5.651) \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$146.923 \\ (63.184) \\ 69.750 \\ (16.038)$	$\begin{array}{c} 40.520 \\ (14.891) \\ 22.365 \\ (4.622) \end{array}$	$547.736 \\ (246.160) \\ 225.422 \\ (59.690)$	$ \begin{array}{ c c c } 9.557 \\ (2.806) \\ 6.399 \\ (0.899) \end{array} $	$ \begin{array}{r} 117.372\\(47.245)\\55.847\\(12.825)\end{array} $
CLL	$\begin{array}{ c c c } \textbf{2.515} \\ (0.089) \\ 2.551 \\ (0.094) \\ 2.548 \end{array}$	$\begin{array}{c} 3.516 \\ (0.282) \\ 3.551 \\ (0.263) \\ 3.492 \end{array}$	$\begin{array}{ c c c } 7.787 \\ (2.321) \\ 5.281 \\ (0.758) \\ 5.068 \end{array}$	$\begin{array}{c} 25.714 \\ (9.640) \\ 14.766 \\ (3.078) \\ 13.940 \end{array}$	$ \begin{array}{c c} 6.398 \\ (1.877) \\ 4.408 \\ (0.610) \\ 4.240 \end{array} $	2.879 (0.132) 2.911 (0.134) 2.904	$12.549 \\ (1.969) \\ 12.353 \\ (1.614) \\ 11.819$	$\begin{array}{c} 4.936 \\ (0.481) \\ 4.978 \\ (0.439) \\ 4.855 \end{array}$	$\begin{array}{r} 36.293 \\ (8.057) \\ 34.633 \\ (5.651) \\ 32.805 \end{array}$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} 146.923 \\ (63.184) \\ 69.750 \\ (16.038) \\ 66.111 \end{array}$	$\begin{array}{r} 40.520 \\ (14.891) \\ 22.365 \\ (4.622) \\ 21.119 \end{array}$	$547.736 \\ (246.160) \\ 225.422 \\ (59.690) \\ 214.261$	$ \begin{array}{ c c c c c } 9.557 \\ (2.806) \\ 6.399 \\ (0.899) \\ 6.152 \\ \end{array} $	$\begin{array}{c} 117.372 \\ (47.245) \\ 55.847 \\ (12.825) \\ 53.115 \end{array}$
CLL CLF	$\begin{array}{ c c c } \textbf{2.515} \\ (0.089) \\ 2.551 \\ (0.094) \\ 2.548 \\ (0.094) \end{array}$	$\begin{array}{c} 3.516 \\ (0.282) \\ 3.551 \\ (0.263) \\ 3.492 \\ (0.236) \end{array}$	$\begin{array}{ c c c } 7.787 \\ (2.321) \\ 5.281 \\ (0.758) \\ 5.068 \\ (0.616) \end{array}$	$\begin{array}{c} 25.714 \\ (9.640) \\ 14.766 \\ (3.078) \\ 13.940 \\ (2.483) \end{array}$	$ \begin{array}{c c} 6.398 \\ (1.877) \\ 4.408 \\ (0.610) \\ 4.240 \\ (0.494) \end{array} $	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} 12.549 \\ (1.969) \\ 12.353 \\ (1.614) \\ 11.819 \\ (1.393) \end{array}$	$\begin{array}{c} 4.936 \\ (0.481) \\ 4.978 \\ (0.439) \\ 4.855 \\ (0.398) \end{array}$	$\begin{array}{c} 36.293 \\ (8.057) \\ 34.633 \\ (5.651) \\ 32.805 \\ (4.782) \end{array}$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c} 146.923 \\ (63.184) \\ 69.750 \\ (16.038) \\ 66.111 \\ (13.531) \end{array}$	$\begin{array}{c} 40.520 \\ (14.891) \\ 22.365 \\ (4.622) \\ 21.119 \\ (3.676) \end{array}$	$547.736 \\ (246.160) \\ 225.422 \\ (59.690) \\ 214.261 \\ (50.795)$	$\begin{array}{c c} 9.557 \\ (2.806) \\ 6.399 \\ (0.899) \\ 6.152 \\ (0.760) \end{array}$	$\begin{array}{c} 117.372 \\ (47.245) \\ 55.847 \\ (12.825) \\ 53.115 \\ (10.746) \end{array}$
CLL CLF	$\begin{array}{ c c c } \textbf{2.515} \\ (0.089) \\ 2.551 \\ (0.094) \\ 2.548 \\ (0.094) \\ 2.614 \end{array}$	$\begin{array}{c} 3.516 \\ (0.282) \\ 3.551 \\ (0.263) \\ 3.492 \\ (0.236) \\ 3.469 \end{array}$	$\begin{array}{ c c c } 7.787 \\ (2.321) \\ 5.281 \\ (0.758) \\ 5.068 \\ (0.616) \\ 6.165 \end{array}$	$\begin{array}{c} 25.714 \\ (9.640) \\ 14.766 \\ (3.078) \\ 13.940 \\ (2.483) \\ 17.680 \end{array}$	$\begin{vmatrix} 6.398 \\ (1.877) \\ 4.408 \\ (0.610) \\ 4.240 \\ (0.494) \\ 5.146 \end{vmatrix}$	$\begin{array}{ c c c } \hline \textbf{2.879} \\ (0.132) \\ 2.911 \\ (0.134) \\ 2.904 \\ (0.132) \\ 2.973 \end{array}$	12.549 (1.969) 12.353 (1.614) 11.819 (1.393) 11.279	$\begin{array}{c} 4.936\\ (0.481)\\ 4.978\\ (0.439)\\ 4.855\\ (0.398)\\ \textbf{4.780} \end{array}$	36.293 (8.057) 34.633 (5.651) 32.805 (4.782) 30.460	$ \begin{vmatrix} 11.995 \\ (3.862) \\ 7.899 \\ (1.202) \\ 7.574 \\ (0.990) \\ 9.168 \end{vmatrix} $	$\begin{array}{c} 146.923 \\ (63.184) \\ 69.750 \\ (16.038) \\ 66.111 \\ (13.531) \\ 90.537 \end{array}$	$\begin{array}{c} 40.520 \\ (14.891) \\ 22.365 \\ (4.622) \\ 21.119 \\ (3.676) \\ 27.176 \end{array}$	$547.736 \\ (246.160) \\ 225.422 \\ (59.690) \\ 214.261 \\ (50.795) \\ 314.050$	$\begin{array}{ c c c } 9.557 \\ (2.806) \\ 6.399 \\ (0.899) \\ 6.152 \\ (0.760) \\ 7.464 \end{array}$	$\begin{array}{c} 117.372 \\ (47.245) \\ 55.847 \\ (12.825) \\ 53.115 \\ (10.746) \\ 72.625 \end{array}$
CLL CLF PCL PL	$\begin{array}{ c c c c } \hline \textbf{2.515} \\ (0.089) \\ 2.551 \\ (0.094) \\ 2.548 \\ (0.094) \\ 2.614 \\ (0.093) \end{array}$	$\begin{array}{c} 3.516 \\ (0.282) \\ 3.551 \\ (0.263) \\ 3.492 \\ (0.236) \\ \textbf{3.469} \\ (0.211) \end{array}$	$ \begin{vmatrix} 7.787 \\ (2.321) \\ 5.281 \\ (0.758) \\ 5.068 \\ (0.616) \\ 6.165 \\ (1.349) \end{vmatrix} $	$\begin{array}{c} 25.714 \\ (9.640) \\ 14.766 \\ (3.078) \\ 13.940 \\ (2.483) \\ 17.680 \\ (5.424) \end{array}$	$ \begin{vmatrix} 6.398 \\ (1.877) \\ 4.408 \\ (0.610) \\ 4.240 \\ (0.494) \\ 5.146 \\ (1.079) \end{vmatrix} $	2.879 (0.132) 2.911 (0.134) 2.904 (0.132) 2.973 (0.129)	$\begin{array}{c} 12.549 \\ (1.969) \\ 12.353 \\ (1.614) \\ 11.819 \\ (1.393) \\ \textbf{11.279} \\ (1.285) \end{array}$	$\begin{array}{c} 4.936\\ (0.481)\\ 4.978\\ (0.439)\\ 4.855\\ (0.398)\\ \textbf{4.780}\\ (0.369) \end{array}$	$\begin{array}{c} 36.293 \\ (8.057) \\ 34.633 \\ (5.651) \\ 32.805 \\ (4.782) \\ \textbf{30.460} \\ (4.375) \end{array}$	$ \begin{vmatrix} 11.995 \\ (3.862) \\ 7.899 \\ (1.202) \\ 7.574 \\ (0.990) \\ 9.168 \\ (2.170) \end{vmatrix} $	$\begin{array}{c} 146.923 \\ (63.184) \\ 69.750 \\ (16.038) \\ 66.111 \\ (13.531) \\ 90.537 \\ (34.972) \end{array}$	$\begin{array}{c} 40.520 \\ (14.891) \\ 22.365 \\ (4.622) \\ 21.119 \\ (3.676) \\ 27.176 \\ (8.697) \end{array}$	$547.736 \\ (246.160) \\ 225.422 \\ (59.690) \\ 214.261 \\ (50.795) \\ 314.050 \\ (131.633)$	$ \begin{vmatrix} 9.557 \\ (2.806) \\ 6.399 \\ (0.899) \\ 6.152 \\ (0.760) \\ 7.464 \\ (1.576) \end{vmatrix} $	$\begin{array}{c} 117.372 \\ (47.245) \\ 55.847 \\ (12.825) \\ 53.115 \\ (10.746) \\ 72.625 \\ (27.447) \end{array}$
CLL CLF PCL	$\begin{array}{ c c c c } \hline \textbf{2.515} \\ (0.089) \\ 2.551 \\ (0.094) \\ 2.548 \\ (0.094) \\ 2.614 \\ (0.093) \\ 2.635 \end{array}$	$\begin{array}{c} 3.516\\ (0.282)\\ 3.551\\ (0.263)\\ 3.492\\ (0.236)\\ \textbf{3.469}\\ (0.211)\\ 3.564 \end{array}$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} 25.714\\ (9.640)\\ 14.766\\ (3.078)\\ 13.940\\ (2.483)\\ 17.680\\ (5.424)\\ 13.130\end{array}$	$ \begin{vmatrix} 6.398 \\ (1.877) \\ 4.408 \\ (0.610) \\ 4.240 \\ (0.494) \\ 5.146 \\ (1.079) \\ 4.176 \end{vmatrix} $	2.879 (0.132) 2.911 (0.134) 2.904 (0.132) 2.973 (0.129) 3.009	$\begin{array}{c} 12.549 \\ (1.969) \\ 12.353 \\ (1.614) \\ 11.819 \\ (1.393) \\ \textbf{11.279} \\ (1.285) \\ 11.740 \end{array}$	$\begin{array}{c} 4.936\\ (0.481)\\ 4.978\\ (0.439)\\ 4.855\\ (0.398)\\ \textbf{4.780}\\ (0.369)\\ 4.927\end{array}$	$\begin{array}{c} 36.293 \\ (8.057) \\ 34.633 \\ (5.651) \\ 32.805 \\ (4.782) \\ \textbf{30.460} \\ (4.375) \\ 31.860 \end{array}$	$ \begin{vmatrix} 11.995 \\ (3.862) \\ 7.899 \\ (1.202) \\ 7.574 \\ (0.990) \\ 9.168 \\ (2.170) \\ 7.376 \end{vmatrix} $	$\begin{array}{c} 146.923\\ (63.184)\\ 69.750\\ (16.038)\\ 66.111\\ (13.531)\\ 90.537\\ (34.972)\\ 61.281\end{array}$	$\begin{array}{c} 40.520\\(14.891)\\22.365\\(4.622)\\21.119\\(3.676)\\27.176\\(8.697)\\19.961\end{array}$	$547.736 \\ (246.160) \\ 225.422 \\ (59.690) \\ 214.261 \\ (50.795) \\ 314.050 \\ (131.633) \\ 196.006$	$ \begin{vmatrix} 9.557 \\ (2.806) \\ 6.399 \\ (0.899) \\ 6.152 \\ (0.760) \\ 7.464 \\ (1.576) \\ 6.025 \end{vmatrix} $	$\begin{array}{c} 117.372 \\ (47.245) \\ 55.847 \\ (12.825) \\ 53.115 \\ (10.746) \\ 72.625 \\ (27.447) \\ 48.781 \end{array}$

							INTE	RPOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.871	3.813	15.351	59.098	12.013	3.225	12.506	5.237	37.121	23.164	345.239	89.005	1415.381	17.921	275.293
	(0.163)	(0.278)	(7.787)	(31.036)	(5.774)	(0.189)	(1.416)	(0.444)	(5.053)	(11.378)	(191.233)	(45.782)	(789.588)	(8.590)	(144.075)
\mathbf{CLL}	3.592	4.863	7.254	19.964	6.058	4.054	16.266	6.723	46.035	10.658	97.441	30.070	321.774	8.650	75.220
	(0.517)	(0.815)	(1.692)	(6.071)	(1.324)	(0.543)	(3.799)	(1.158)	(12.040)	(2.606)	(36.498)	(9.576)	(132.287)	(2.030)	(27.846)
\mathbf{CLF}	3.226	4.215	5.849	15.121	4.899	3.625	13.110	5.752	35.981	8.471	70.939	22.961	233.947	6.985	56.653
	(0.280)	(0.430)	(0.728)	(2.402)	(0.574)	(0.303)	(1.744)	(0.601)	(5.537)	(1.094)	(13.767)	(3.695)	(46.761)	(0.889)	(10.516)
\mathbf{PCL}	3.220	4.165	5.680	14.254	4.761	3.627	12.354	5.678	32.348	8.148	66.320	21.604	225.117	6.740	53.232
	(0.237)	(0.349)	(0.622)	(2.257)	(0.499)	(0.264)	(1.351)	(0.510)	(4.265)	(0.927)	(14.359)	(3.617)	(61.881)	(0.769)	(11.097)
\mathbf{PL}	3.287	4.268	5.671	13.329	4.774	3.705	12.882	5.836	33.675	8.127	56.609	20.183	169.271	6.731	45.609
	(0.249)	(0.375)	(0.603)	(1.719)	(0.482)	(0.280)	(1.489)	(0.547)	(4.543)	(0.860)	(8.760)	(2.565)	(29.936)	(0.733)	(7.348)
\mathbf{PF}	3.284	4.257	5.598	13.052	4.722	3.700	12.639	5.806	32.950	7.987	55.641	19.792	167.448	6.636	44.907
	(0.248)	(0.372)	(0.581)	(1.658)	(0.467)	(0.279)	(1.414)	(0.539)	(4.375)	(0.829)	(8.443)	(2.489)	(28.785)	(0.710)	(7.140)
							EXTR	APOLAT	TION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.873	4.152	43.004	177.361	33.774	3.230	15.241	5.772	50.112	67.149	1039.918	259.334	4343.040	50.097	856.448
	(0.627)	(1.411)	(32.948)	(132.081)	(25.033)	(0.803)	(7.412)	(2.114)	(29.164)	(48.820)	(798.639)	(196.385)	(3263.325)	(38.261)	(603.817)
\mathbf{CLL}	5.494	7.332	11.703	35.624	9.228	5.089	25.371	9.482	76.871	17.051	180.655	54.397	651.026	13.802	149.875
	(2.544)	(3.789)	(6.916)	(22.937)	(5.267)	(2.433)	(14.801)	(5.181)	(48.446)	(9.886)	(117.088)	(34.777)	(432.520)	(8.214)	(100.089)
\mathbf{CLF}	4.359	5.837	9.733	30.359	7.664	4.327	19.584	7.616	58.934	14.259	161.351	46.632	604.840	11.813	133.352
	(1.797)	(2.770)	(5.530)	(19.125)	(4.260)	(1.838)	(10.870)	(3.811)	(36.437)	(8.234)	(108.983)	(29.988)	(438.496)	(6.829)	(90.227)
PCL	3.324	4.683	25.942	104.806	20.321	3.668	16.245	6.356	48.747	40.626	606.996	153.383	2538.333	30.341	502.018
	(1.059)	(1.926)	(19.319)	(76.734)	(14.714)	(1.249)	(8.425)	(2.721)	(30.587)	(28.721)	(468.732)	(114.807)	(1912.935)	(22.632)	(355.710)
\mathbf{PL}	4.111	5.965	10.597	32.349	8.117	4.608	23.378	8.177	73.891	15.622	158.057	48.218	562.277	12.778	135.384
	(1.892)	(3.154)	(6.380)	(21.710)	(4.786)	(2.187)	(14.347)	(4.369)	(51.431)	(9.505)	(111.802)	(32.991)	(410.817)	(7.803)	(95.265)
\mathbf{PF}	4.102	5.923	10.069	29.898	7.861	4.593	21.173	8.056	64.225	14.619	149.310	44.763	545.485	12.097	126.373
	(1.884)	(3.116)	(5.932)	(19.414)	(4.558)	(2.172)	(12.466)	(4.267)	(43.417)	(8.707)	(105.094)	(30.210)	(403.948)	(7.196)	(87.215)

Table 1.11: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 1$ and $\varphi = 0.2$ with aggregation matrix \mathbf{C}_1

							INTE	RPOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.865	3.802	15.031	55.591	11.973	3.222	12.619	5.222	36.893	22.964	349.056	88.084	1374.329	18.413	283.380
	(0.157)	(0.289)	(7.222)	(28.800)	(5.745)	(0.202)	(1.388)	(0.442)	(5.167)	(11.247)	(194.668)	(47.089)	(720.737)	(8.974)	(150.564)
\mathbf{CLL}	4.336	5.390	7.655	20.083	6.466	5.007	17.192	7.320	46.098	11.067	97.550	30.750	323.553	9.078	78.484
	(0.827)	(1.095)	(2.075)	(6.414)	(1.689)	(1.014)	(4.262)	(1.563)	(12.733)	(2.881)	(35.577)	(9.908)	(130.653)	(2.290)	(28.620)
\mathbf{CLF}	3.660	4.510	6.038	15.298	5.161	4.153	13.611	6.068	36.004	8.707	71.257	23.075	233.560	7.206	57.229
	(0.427)	(0.553)	(0.930)	(2.536)	(0.766)	(0.513)	(2.004)	(0.765)	(5.842)	(1.234)	(13.389)	(3.834)	(47.173)	(1.015)	(10.875)
\mathbf{PCL}	3.317	4.228	5.645	14.118	4.820	3.720	12.373	5.701	32.303	8.218	66.639	21.343	222.136	6.800	53.690
	(0.232)	(0.346)	(0.630)	(2.166)	(0.471)	(0.287)	(1.347)	(0.521)	(4.255)	(0.944)	(15.008)	(3.493)	(55.579)	(0.730)	(11.174)
\mathbf{PL}	3.384	4.334	5.659	13.386	4.833	3.800	12.908	5.857	33.675	8.170	56.803	19.854	169.276	6.798	45.700
	(0.246)	(0.369)	(0.583)	(1.727)	(0.462)	(0.305)	(1.488)	(0.554)	(4.582)	(0.856)	(9.357)	(2.651)	(30.195)	(0.696)	(7.245)
\mathbf{PF}	3.381	4.322	5.587	13.118	4.784	3.795	12.661	5.827	32.928	8.037	55.835	19.474	167.314	6.699	44.843
	(0.245)	(0.366)	(0.567)	(1.643)	(0.450)	(0.303)	(1.420)	(0.547)	(4.384)	(0.828)	(9.064)	(2.526)	(29.163)	(0.672)	(7.009)
							EXTR	APOLAT	TION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.873	4.152	44.040	163.380	33.643	3.223	15.963	5.813	49.012	64.160	1064.734	274.578	4340.173	51.557	848.033
	(0.627)	(1.411)	(31.894)	(123.237)	(25.297)	(0.764)	(7.724)	(2.306)	(28.825)	(48.986)	(826.948)	(206.547)	(3182.063)	(37.962)	(645.398)
\mathbf{CLL}	5.494	7.332	11.864	34.102	9.711	6.477	27.327	10.119	75.611	17.407	180.207	54.549	651.528	14.414	150.338
	(2.544)	(3.789)	(6.821)	(21.350)	(5.848)	(2.949)	(15.762)	(5.411)	(47.960)	(10.115)	(115.821)	(35.259)	(421.007)	(8.622)	(98.363)
\mathbf{CLF}	4.359	5.837	9.699	28.091	7.991	4.985	20.866	7.933	57.657	14.147	158.379	46.597	604.041	11.978	132.953
	(1.797)	(2.770)	(5.589)	(18.230)	(4.615)	(2.037)	(11.575)	(3.866)	(36.783)	(8.061)	(106.720)	(30.262)	(404.411)	(7.074)	(91.123)
\mathbf{PCL}	3.324	4.683	26.530	96.247	20.478	3.681	16.784	6.448	48.160	38.686	625.834	162.124	2541.434	31.332	497.564
	(1.059)	(1.926)	(18.587)	(72.313)	(14.813)	(1.170)	(8.879)	(2.812)	(29.877)	(28.683)	(478.240)	(120.487)	(1860.550)	(22.390)	(378.037)
\mathbf{PL}	4.111	5.965	10.201	30.934	8.322	4.458	23.624	8.291	72.165	15.222	157.410	48.641	554.770	12.774	132.170
	(1.892)	(3.154)	(6.131)	(20.224)	(4.872)	(1.923)	(14.639)	(4.386)	(49.940)	(9.481)	(109.686)	(32.428)	(385.011)	(7.946)	(92.187)
\mathbf{PF}	4.102	5.923	9.817	27.996	8.065	4.447	21.635	8.174	63.113	14.306	149.766	45.074	540.537	12.102	124.324
	(1.884)	(3.116)	(5.803)	(18.350)	(4.656)	(1.914)	(12.876)	(4.272)	(42.316)	(8.662)	(101.894)	(29.774)	(365.648)	(7.366)	(87.941)

Table 1.12: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 1$ and $\varphi = 0.8$ with aggregation matrix \mathbf{C}_1

							INTE	RPOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.024	2.687	10.559	40.112	8.735	2.281	8.889	3.701	26.290	16.703	245.106	63.568	955.957	13.310	13.194
	(0.113)	(0.197)	(5.196)	(21.094)	(4.258)	(0.136)	(1.053)	(0.299)	(3.577)	(8.435)	(133.125)	(33.754)	(497.320)	(6.513)	(6.521)
\mathbf{CLL}	2.508	3.393	5.055	13.892	4.226	2.859	11.551	4.756	32.414	7.485	69.204	21.631	224.445	6.135	6.909
	(0.358)	(0.590)	(1.116)	(4.455)	(0.960)	(0.420)	(2.627)	(0.814)	(8.706)	(1.837)	(25.322)	(7.020)	(95.410)	(1.495)	(1.981)
\mathbf{CLF}	2.292	2.986	4.114	10.621	3.451	2.588	9.311	4.085	25.454	5.970	50.417	16.368	165.000	4.938	5.346
	(0.206)	(0.306)	(0.491)	(1.809)	(0.419)	(0.241)	(1.267)	(0.419)	(3.865)	(0.793)	(9.654)	(2.806)	(35.410)	(0.652)	(0.870)
\mathbf{PCL}	2.258	2.940	3.980	9.972	3.353	2.558	8.790	4.013	22.984	5.792	47.086	15.298	156.277	4.775	4.858
	(0.166)	(0.243)	(0.424)	(1.522)	(0.354)	(0.204)	(0.992)	(0.349)	(3.005)	(0.689)	(10.254)	(2.508)	(39.454)	(0.547)	(0.538)
\mathbf{PL}	2.321	3.009	3.988	9.396	3.350	2.620	9.130	4.118	23.908	5.745	40.359	14.292	120.119	4.753	4.828
	(0.170)	(0.259)	(0.407)	(1.212)	(0.348)	(0.212)	(1.062)	(0.370)	(3.219)	(0.643)	(6.712)	(1.914)	(21.262)	(0.502)	(0.512)
\mathbf{PF}	2.320	3.005	3.949	9.215	3.324	2.618	8.997	4.106	23.406	5.664	39.667	13.999	118.871	4.697	4.774
	(0.170)	(0.258)	(0.395)	(1.166)	(0.338)	(0.211)	(1.034)	(0.368)	(3.091)	(0.626)	(6.428)	(1.852)	(21.025)	(0.488)	(0.496)
							EXTR	APOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.025	2.882	30.516	118.915	24.713	2.311	11.122	4.163	35.006	48.322	716.182	188.333	2937.658	38.524	38.161
	(0.423)	(0.931)	(22.275)	(87.785)	(18.447)	(0.564)	(5.443)	(1.603)	(20.236)	(36.230)	(553.026)	(141.838)	(2162.243)	(28.311)	(28.539)
\mathbf{CLL}	3.066	4.531	7.775	24.487	6.632	3.582	17.945	6.609	53.311	12.365	123.772	37.469	467.518	9.716	10.803
	(1.365)	(2.274)	(4.428)	(16.315)	(3.820)	(1.669)	(10.885)	(3.585)	(33.043)	(7.695)	(82.565)	(24.293)	(314.950)	(5.886)	(6.649)
\mathbf{CLF}	2.715	3.793	6.618	21.440	5.557	3.104	13.797	5.442	41.331	10.244	110.705	32.465	433.140	8.302	8.572
	(1.040)	(1.749)	(3.749)	(14.009)	(3.103)	(1.267)	(8.071)	(2.703)	(23.820)	(6.052)	(74.589)	(21.373)	(304.558)	(5.046)	(5.208)
\mathbf{PCL}	2.253	3.200	18.467	70.660	14.973	2.595	11.622	4.578	34.450	28.965	418.764	111.536	1720.759	23.300	22.881
	(0.642)	(1.280)	(12.951)	(51.647)	(10.730)	(0.841)	(6.264)	(2.036)	(19.939)	(20.999)	(319.821)	(83.314)	(1260.260)	(16.925)	(16.560)
\mathbf{PL}	2.877	4.068	7.094	22.831	6.001	3.274	15.980	5.874	51.871	10.732	109.209	33.888	408.065	8.811	8.794
	(1.223)	(2.101)	(4.150)	(15.069)	(3.456)	(1.518)	(10.132)	(3.168)	(34.147)	(6.631)	(74.445)	(22.530)	(287.379)	(5.449)	(5.443)
\mathbf{PF}	2.874	4.054	6.920	21.101	5.886	3.269	15.005	5.832	45.368	10.267	101.898	31.571	393.679	8.538	8.490
	(1.220)	(2.088)	(3.992)	(13.948)	(3.371)	(1.512)	(9.324)	(3.128)	(28.347)	(6.156)	(69.877)	(20.679)	(278.570)	(5.222)	(5.183)

Table 1.13: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 5$ and $\varphi = 0.2$ with aggregation matrix \mathbf{C}_1

	INTERPOLATION														
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.023	2.700	10.675	40.893	8.657	2.286	9.037	3.721	26.179	16.584	251.364	62.744	968.297	17.921	18.413
	(0.116)	(0.198)	(5.371)	(22.242)	(4.250)	(0.138)	(1.007)	(0.309)	(3.771)	(8.403)	(131.466)	(33.946)	(546.187)	(8.590)	(8.974)
\mathbf{CLL}	3.476	4.239	5.683	14.301	4.921	4.097	12.823	5.696	33.253	8.179	69.476	22.155	227.207	8.650	9.078
	(0.697)	(1.004)	(1.581)	(4.528)	(1.307)	(0.835)	(3.414)	(1.361)	(9.271)	(2.323)	(26.143)	(7.054)	(91.872)	(2.030)	(2.290)
\mathbf{CLF}	2.886	3.477	4.455	10.811	3.849	3.307	10.047	4.610	25.680	6.351	50.750	16.593	164.510	6.985	7.206
	(0.366)	(0.501)	(0.706)	(1.831)	(0.592)	(0.433)	(1.601)	(0.666)	(4.182)	(1.032)	(9.963)	(2.820)	(33.718)	(0.889)	(1.015)
\mathbf{PCL}	2.398	3.063	4.094	10.089	3.455	2.703	8.820	4.130	22.908	5.837	47.505	15.270	156.970	6.740	6.800
	(0.182)	(0.243)	(0.446)	(1.634)	(0.338)	(0.205)	(0.946)	(0.357)	(3.150)	(0.652)	(9.840)	(2.514)	(43.441)	(0.769)	(0.730)
\mathbf{PL}	2.460	3.134	4.086	9.521	3.450	2.763	9.140	4.237	23.929	5.812	40.450	14.347	119.650	6.731	6.798
	(0.186)	(0.259)	(0.420)	(1.269)	(0.329)	(0.214)	(1.012)	(0.381)	(3.304)	(0.619)	(6.532)	(1.926)	(21.495)	(0.733)	(0.696)
\mathbf{PF}	2.458	3.130	4.050	9.328	3.425	2.760	9.016	4.225	23.381	5.730	39.743	14.055	118.183	6.636	6.699
	(0.186)	(0.257)	(0.410)	(1.215)	(0.322)	(0.213)	(0.983)	(0.378)	(3.207)	(0.600)	(6.282)	(1.855)	(21.064)	(0.710)	(0.672)
							EXTR	APOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.005	2.926	30.521	119.607	24.279	2.327	10.975	4.197	35.837	47.421	773.489	185.978	2970.516	50.097	51.557
	(0.438)	(1.007)	(23.411)	(91.239)	(18.290)	(0.534)	(5.807)	(1.576)	(21.702)	(35.953)	(569.122)	(144.973)	(2236.619)	(38.261)	(37.962)
\mathbf{CLL}	4.337	5.549	8.682	25.221	7.243	5.294	20.159	7.970	56.215	12.635	132.675	39.719	439.259	13.802	14.414
	(1.934)	(2.846)	(4.943)	(15.833)	(4.113)	(2.307)	(12.008)	(4.334)	(35.000)	(7.602)	(87.015)	(25.384)	(292.107)	(8.214)	(8.622)
\mathbf{CLF}	3.406	4.321	7.125	21.729	5.825	4.019	14.992	6.075	43.395	10.329	117.797	32.996	405.983	11.813	11.978
	(1.334)	(1.983)	(3.997)	(13.930)	(3.194)	(1.576)	(8.558)	(3.006)	(26.288)	(6.089)	(76.995)	(22.582)	(287.697)	(6.829)	(7.074)
\mathbf{PCL}	2.341	3.292	18.421	70.912	14.691	2.699	11.623	4.703	35.563	28.374	452.904	110.130	1735.528	30.341	31.332
	(0.716)	(1.296)	(13.777)	(53.460)	(10.697)	(0.826)	(6.507)	(2.003)	(22.222)	(21.025)	(335.187)	(84.803)	(1304.200)	(22.632)	(22.390)
\mathbf{PL}	2.890	4.120	7.066	22.595	5.614	3.270	16.254	5.996	53.442	10.541	116.967	34.552	387.402	12.778	12.774
	(1.309)	(2.057)	(4.096)	(15.099)	(3.346)	(1.407)	(9.873)	(3.188)	(34.926)	(6.494)	(79.684)	(24.500)	(277.571)	(7.803)	(7.946)
\mathbf{PF}	2.887	4.106	6.918	21.307	5.522	3.266	15.297	5.952	46.965	10.088	110.008	31.726	370.893	12.097	12.102
	(1.305)	(2.043)	(3.952)	(13.818)	(3.279)	(1.402)	(9.153)	(3.145)	(29.746)	(6.059)	(73.121)	(22.607)	(268.563)	(7.196)	(7.366)

Table 1.14: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 5$ and $\varphi = 0.8$ with aggregation matrix \mathbf{C}_1

							INTE	RPOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.871	3.813	15.351	59.098	12.013	3.225	12.506	5.237	37.121	23.164	345.239	89.005	1415.381	17.921	275.293
	(0.163)	(0.278)	(7.787)	(31.036)	(5.774)	(0.189)	(1.416)	(0.444)	(5.053)	(11.378)	(191.233)	(45.782)	(789.588)	(8.590)	(144.075)
\mathbf{CLL}	3.592	4.863	7.254	19.964	6.058	4.054	16.266	6.723	46.035	10.658	97.441	30.070	321.774	8.650	75.220
	(0.517)	(0.815)	(1.692)	(6.071)	(1.324)	(0.543)	(3.799)	(1.158)	(12.040)	(2.606)	(36.498)	(9.576)	(132.287)	(2.030)	(27.846)
\mathbf{CLF}	3.226	4.215	5.849	15.121	4.899	3.625	13.110	5.752	35.981	8.471	70.939	22.961	233.947	6.985	56.653
	(0.280)	(0.430)	(0.728)	(2.402)	(0.574)	(0.303)	(1.744)	(0.601)	(5.537)	(1.094)	(13.767)	(3.695)	(46.761)	(0.889)	(10.516)
\mathbf{PCL}	3.220	4.165	5.680	14.254	4.761	3.627	12.354	5.678	32.348	8.148	66.320	21.604	225.117	6.740	53.232
	(0.237)	(0.349)	(0.622)	(2.257)	(0.499)	(0.264)	(1.351)	(0.510)	(4.265)	(0.927)	(14.359)	(3.617)	(61.881)	(0.769)	(11.097)
\mathbf{PL}	3.287	4.268	5.671	13.329	4.774	3.705	12.882	5.836	33.675	8.127	56.609	20.183	169.271	6.731	45.609
	(0.249)	(0.375)	(0.603)	(1.719)	(0.482)	(0.280)	(1.489)	(0.547)	(4.543)	(0.860)	(8.760)	(2.565)	(29.936)	(0.733)	(7.348)
\mathbf{PF}	3.284	4.257	5.598	13.052	4.722	3.700	12.639	5.806	32.950	7.987	55.641	19.792	167.448	6.636	44.907
	(0.248)	(0.372)	(0.581)	(1.658)	(0.467)	(0.279)	(1.414)	(0.539)	(4.375)	(0.829)	(8.443)	(2.489)	(28.785)	(0.710)	(7.140)
							EXTR	APOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.853	4.033	43.004	177.361	33.774	3.230	15.241	5.772	50.112	67.149	1039.918	259.334	4343.040	50.097	856.448
	0.629	1.295	(32.948)	(132.081)	(25.033)	(0.803)	(7.412)	(2.114)	(29.164)	(48.820)	(798.639)	(196.385)	(3263.325)	(38.261)	(603.817)
\mathbf{CLL}	4.366	6.468	11.703	35.624	9.228	5.089	25.371	9.482	76.871	17.051	180.655	54.397	651.026	13.802	149.875
	2.058	3.296	(6.916)	(22.937)	(5.267)	(2.433)	(14.801)	(5.181)	(48.446)	(9.886)	(117.088)	(34.777)	(432.520)	(8.214)	(100.089)
\mathbf{CLF}	3.775	5.300	9.733	30.359	7.664	4.327	19.584	7.616	58.934	14.259	161.351	46.632	604.840	11.813	133.352
	1.533	2.376	(5.530)	(19.125)	(4.260)	(1.838)	(10.870)	(3.811)	(36.437)	(8.234)	(108.983)	(29.988)	(438.496)	(6.829)	(90.227)
PCL	3.205	4.448	25.942	104.806	20.321	3.668	16.245	6.356	48.747	40.626	606.996	153.383	2538.333	30.341	502.018
	1.007	1.663	(19.319)	(76.734)	(14.714)	(1.249)	(8.425)	(2.721)	(30.587)	(28.721)	(468.732)	(114.807)	(1912.935)	(22.632)	(355.710)
\mathbf{PL}	3.981	5.679	10.597	32.349	8.117	4.608	23.378	8.177	73.891	15.622	158.057	48.218	562.277	12.778	135.384
	1.801	2.775	(6.380)	(21.710)	(4.786)	(2.187)	(14.347)	(4.369)	(51.431)	(9.505)	(111.802)	(32.991)	(410.817)	(7.803)	(95.265)
\mathbf{PF}	3.973	5.643	10.069	29.898	7.861	4.593	21.173	8.056	64.225	14.619	149.310	44.763	545.485	12.097	126.373
	1.792	2.734	(5.932)	(19.414)	(4.558)	(2.172)	(12.466)	(4.267)	(43.417)	(8.707)	(105.094)	(30.210)	(403.948)	(7.196)	(87.215)

Table 1.15: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 10$ and $\varphi = 0.2$ with aggregation matrix \mathbf{C}_1

							INTE	RPOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.865	3.802	15.031	55.591	11.973	3.222	12.619	5.222	36.893	22.964	349.056	88.084	1374.329	18.413	283.380
	(0.157)	(0.289)	(7.222)	(28.800)	(5.745)	(0.202)	(1.388)	(0.442)	(5.167)	(11.247)	(194.668)	(47.089)	(720.737)	(8.974)	(150.564)
CLL	4.336	5.390	7.655	20.083	6.466	5.007	17.192	7.320	46.098	11.067	97.550	30.750	323.553	9.078	78.484
	(0.827)	(1.095)	(2.075)	(6.414)	(1.689)	(1.014)	(4.262)	(1.563)	(12.733)	(2.881)	(35.577)	(9.908)	(130.653)	(2.290)	(28.620)
\mathbf{CLF}	3.660	4.510	6.038	15.298	5.161	4.153	13.611	6.068	36.004	8.707	71.257	23.075	233.560	7.206	57.229
	(0.427)	(0.553)	(0.930)	(2.536)	(0.766)	(0.513)	(2.004)	(0.765)	(5.842)	(1.234)	(13.389)	(3.834)	(47.173)	(1.015)	(10.875)
PCL	3.317	4.228	5.645	14.118	4.820	3.720	12.373	5.701	32.303	8.218	66.639	21.343	222.136	6.800	53.690
	(0.232)	(0.346)	(0.630)	(2.166)	(0.471)	(0.287)	(1.347)	(0.521)	(4.255)	(0.944)	(15.008)	(3.493)	(55.579)	(0.730)	(11.174)
\mathbf{PL}	3.384	4.334	5.659	13.386	4.833	3.800	12.908	5.857	33.675	8.170	56.803	19.854	169.276	6.798	45.700
	(0.246)	(0.369)	(0.583)	(1.727)	(0.462)	(0.305)	(1.488)	(0.554)	(4.582)	(0.856)	(9.357)	(2.651)	(30.195)	(0.696)	(7.245)
\mathbf{PF}	3.381	4.322	5.587	13.118	4.784	3.795	12.661	5.827	32.928	8.037	55.835	19.474	167.314	6.699	44.843
	(0.245)	(0.366)	(0.567)	(1.643)	(0.450)	(0.303)	(1.420)	(0.547)	(4.384)	(0.828)	(9.064)	(2.526)	(29.163)	(0.672)	(7.009)
							EXTR	APOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.873	4.152	44.040	163.380	33.643	3.223	15.963	5.813	49.012	64.160	1064.734	274.578	4340.173	51.557	848.033
	(0.627)	(1.411)	(31.894)	(123.237)	(25.297)	(0.764)	(7.724)	(2.306)	(28.825)	(48.986)	(826.948)	(206.547)	(3182.063)	(37.962)	(645.398)
CLL	5.494	7.332	11.864	34.102	9.711	6.477	27.327	10.119	75.611	17.407	180.207	54.549	651.528	14.414	150.338
	(2.544)	(3.789)	(6.821)	(21.350)	(5.848)	(2.949)	(15.762)	(5.411)	(47.960)	(10.115)	(115.821)	(35.259)	(421.007)	(8.622)	(98.363)
\mathbf{CLF}	4.359	5.837	9.699	28.091	7.991	4.985	20.866	7.933	57.657	14.147	158.379	46.597	604.041	11.978	132.953
	(1.797)	(2.770)	(5.589)	(18.230)	(4.615)	(2.037)	(11.575)	(3.866)	(36.783)	(8.061)	(106.720)	(30.262)	(404.411)	(7.074)	(91.123)
PCL	3.324	4.683	26.530	96.247	20.478	3.681	16.784	6.448	48.160	38.686	625.834	162.124	2541.434	31.332	497.564
	(1.059)	(1.926)	(18.587)	(72.313)	(14.813)	(1.170)	(8.879)	(2.812)	(29.877)	(28.683)	(478.240)	(120.487)	(1860.550)	(22.390)	(378.037)
\mathbf{PL}	4.111	5.965	10.201	30.934	8.322	4.458	23.624	8.291	72.165	15.222	157.410	48.641	554.770	12.774	132.170
	(1.892)	(3.154)	(6.131)	(20.224)	(4.872)	(1.923)	(14.639)	(4.386)	(49.940)	(9.481)	(109.686)	(32.428)	(385.011)	(7.946)	(92.187)
\mathbf{PF}	4.102	5.923	9.817	27.996	8.065	4.447	21.635	8.174	63.113	14.306	149.766	45.074	540.537	12.102	124.324
	(1.884)	(3.116)	(5.803)	(18.350)	(4.656)	(1.914)	(12.876)	(4.272)	(42.316)	(8.662)	(101.894)	(29.774)	(365.648)	(7.366)	(87.941)

Table 1.16: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 10$ and $\varphi = 0.8$ with aggregation matrix \mathbf{C}_1

INTERPOLATION 1 $\mathbf{2}$ 6 7 8 9 3 4 5 10 11 1213 $\mathbf{14}$ 15CLCL 0.2500.3500.6041.8280.5040.288 1.3340.4913.5450.919 10.2562.98936.466 0.7480.772(0.009)(0.028)(0.159)(0.686)(0.134)(0.013)(0.245)(0.049)(0.783)(0.254)(4.567)(1.202)(18.292)(0.210)(0.218)CLL 0.2720.3581.3370.308 1.199 0.4893.2580.7276.2842.05519.812 0.5960.7470.4940.415(0.012)(0.027)(0.060)(0.257)(0.051)(0.016)(0.136)(0.042)(0.524)(0.091)(1.546)(0.414)(5.649)(0.076)(0.124)CLF 0.2720.3580.490 1.2680.4130.308 0.489 3.1190.7175.8881.93318.5750.589 0.7421.186(0.012)(0.027)(0.057)(0.201)(0.050)(0.016)(0.131)(0.042)(0.452)(0.086)(1.203)(0.324)(4.377)(0.072)(0.120)PCL 0.3150.3911.2620.443 0.3601.178 0.5182.9700.7365.5701.87617.2770.6151.0600.515(0.011)(0.023)(0.049)(0.179)(0.039)(0.016)(0.129)(0.037)(0.413)(0.081)(1.029)(0.270)(3.890)(0.064)(0.063) \mathbf{PL} 0.3150.3920.5031.1980.4340.3591.1950.5223.0080.7145.2861.79016.1910.6001.042(0.011)(0.023)(0.046)(0.167)(0.036)(0.015)(0.133)(0.038)(0.421)(0.074)(0.913)(0.247)(3.407)(0.059)(0.056) \mathbf{PF} 0.3150.3920.5031.197 0.4340.3591.1950.5223.008 0.714 5.2721.790 15.8890.600 1.042(0.011)(0.023)(0.046)(0.166)(0.036)(0.015)(0.133)(0.038)(0.421)(0.074)(0.906)(0.247)(3.223)(0.059)(0.056)EXTRAPOLATION 1 7 2 3 4 5 6 8 9 10 11 12 13 14 15CLCL 4.2571.913 0.6250.260 0.4275.40520.907 0.305 6.3278.485 124.91431.829 502.386 6.5946.515(0.048)(0.140)(4.045)(15.081)(3.100)(0.066)(0.913)(0.224)(3.406)(5.950)(95.170)(23.560)(383.878)(4.897)(4.674) \mathbf{CLL} 1.5310.3000.5091.2034.5850.9960.3502.4270.7498.4791.88724.7276.99691.5841.616(0.594)(0.061)(0.201)(0.756)(2.968)(0.090)(1.442)(0.357)(5.312)(1.216)(17.341)(4.727)(63.519)(0.913)(0.984)CLF 1.201 0.995 2.3391.5320.3000.5094.4670.3500.7467.5981.879 25.5286.888 94.0641.603(0.594)(0.911)(0.061)(0.200)(0.756)(2.992)(0.090)(1.372)(0.354)(4.598)(1.208)(17.690)(4.695)(65.629)(0.971)PCL 0.4525.39819.4384.2580.6396.1118.397 97.695 27.937390.757 6.5746.5210.3110.3601.836(4.030)(13.242)(3.094)(0.229)(5.850)(73.974)(19.109)(4.859)(0.056)(0.143)(0.078)(0.908)(3.401)(300.747)(4.614) \mathbf{PL} 0.3830.6071.2524.5301.0550.4602.6500.8798.522 1.94924.5256.94590.007 1.5922.039(0.270)(0.797)(2.978)(0.627)(0.144)(1.587)(0.459)(5.476)(1.268)(17.447)(4.750)(0.972)(0.095)(62.471)(1.152)PF 1.2524.5301.0558.522 24.52190.072 1.5920.4050.6140.4742.6500.8821.9496.9452.039(0.102)(0.274)(0.797)(2.978)(0.627)(0.148)(1.587)(0.459)(5.476)(1.268)(17.446)(4.750)(62.477)(0.972)(1.152)

Table 1.17: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 0.1$ and $\varphi = 0.2$ with aggregation matrix \mathbf{C}_2

Table 1.18: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 0.1$ and $\varphi = 0.8$ with aggregation matrix \mathbf{C}_2

	INTERPOLATION														
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	0.250	0.349	0.611	1.895	0.495	0.329	2.540	0.518	4.129	0.940	10.711	2.978	37.608	2.351	2.358
	(0.009)	(0.029)	(0.167)	(0.746)	(0.126)	(0.016)	(0.398)	(0.051)	(1.031)	(0.269)	(4.908)	(1.169)	(19.691)	(0.639)	(0.639)
\mathbf{CLL}	0.498	0.551	0.646	1.401	0.585	0.555	1.818	0.666	3.573	0.852	6.552	2.098	19.970	1.922	1.949
	(0.019)	(0.047)	(0.105)	(0.314)	(0.091)	(0.028)	(0.150)	(0.067)	(0.536)	(0.146)	(1.602)	(0.453)	(5.721)	(0.269)	(0.292)
\mathbf{CLF}	0.495	0.548	0.641	1.340	0.579	0.551	1.788	0.662	3.411	0.848	6.128	1.995	18.742	1.845	1.895
	(0.019)	(0.048)	(0.103)	(0.256)	(0.090)	(0.028)	(0.139)	(0.067)	(0.451)	(0.139)	(1.225)	(0.359)	(4.431)	(0.213)	(0.247)
\mathbf{PCL}	0.810	0.842	0.913	1.473	0.872	0.927	2.096	0.995	3.439	1.137	5.912	2.092	17.588	1.828	2.030
	(0.014)	(0.027)	(0.050)	(0.171)	(0.040)	(0.020)	(0.153)	(0.042)	(0.426)	(0.079)	(1.031)	(0.269)	(3.736)	(0.192)	(0.199)
\mathbf{PL}	0.805	0.839	0.900	1.419	0.861	0.920	2.065	0.993	3.469	1.116	5.602	2.014	16.288	1.785	1.985
	(0.014)	(0.027)	(0.044)	(0.154)	(0.036)	(0.020)	(0.152)	(0.041)	(0.432)	(0.072)	(0.918)	(0.253)	(3.214)	(0.182)	(0.185)
\mathbf{PF}	0.805	0.839	0.900	1.419	0.861	0.920	2.065	0.993	3.468	1.116	5.586	2.014	15.990	1.785	1.985
	(0.014)	(0.027)	(0.044)	(0.154)	(0.036)	(0.020)	(0.152)	(0.041)	(0.432)	(0.072)	(0.909)	(0.252)	(3.043)	(0.182)	(0.185)
							EXTRA	POLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	0.257	0.441	5.051	20.728	4.152	0.343	2.873	0.661	6.500	8.464	135.110	32.188	530.159	21.746	21.020
	(0.045)	(0.150)	(3.910)	(15.732)	(3.083)	(0.064)	(1.024)	(0.222)	(3.218)	(6.057)	(97.330)	(23.749)	(395.349)	(16.260)	(15.590)
\mathbf{CLL}	0.564	0.708	1.327	4.440	1.058	0.646	2.996	0.942	8.823	1.925	24.550	6.755	94.269	4.907	5.007
	(0.085)	(0.254)	(0.744)	(2.927)	(0.597)	(0.131)	(1.432)	(0.398)	(5.598)	(1.211)	(16.787)	(4.496)	(64.439)	(3.020)	(3.255)
\mathbf{CLF}	0.560	0.702	1.320	4.357	1.051	0.638	2.942	0.930	7.888	1.910	24.726	6.727	98.390	4.841	4.835
	(0.084)	(0.251)	(0.739)	(2.904)	(0.593)	(0.130)	(1.365)	(0.390)	(4.796)	(1.195)	(17.079)	(4.430)	(67.024)	(2.876)	(3.065)
PCL	0.736	0.814	5.064	19.218	4.196	0.851	2.770	0.990	6.254	8.419	105.985	28.131	411.429	20.062	19.565
	(0.070)	(0.163)	(3.844)	(13.700)	(3.039)	(0.098)	(0.988)	(0.261)	(3.226)	(5.943)	(76.030)	(19.135)	(309.178)	(13.953)	(13.558)
\mathbf{PL}	1.203	1.268	1.659	4.519	1.469	1.419	3.293	1.609	9.016	2.243	24.365	6.786	93.240	4.917	5.070
	(0.138)	(0.416)	(0.886)	(3.014)	(0.761)	(0.208)	(1.377)	(0.639)	(5.648)	(1.355)	(16.706)	(4.483)	(64.907)	(3.048)	(3.308)
\mathbf{PF}	1.203	1.268	1.659	4.519	1.469	1.419	3.293	1.609	9.016	2.243	24.357	6.786	93.018	4.917	5.070
	(0.138)	(0.416)	(0.886)	(3.014)	(0.761)	(0.208)	(1.377)	(0.639)	(5.648)	(1.355)	(16.703)	(4.483)	(64.481)	(3.048)	(3.308)

							INTE	ERPOLA	ΓION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	0.793	1.103	1.926	5.948	1.577	0.906	3.839	1.542	11.046	2.878	32.954	9.152	118.497	2.351	26.583
	(0.030)	(0.093)	(0.533)	(2.318)	(0.394)	(0.042)	(0.628)	(0.156)	(2.403)	(0.801)	(14.787)	(3.541)	(58.811)	(0.639)	(12.721)
\mathbf{CLL}	0.798	1.085	1.573	4.253	1.299	0.908	3.719	1.512	10.286	2.333	19.649	6.507	62.766	1.922	15.948
	(0.030)	(0.074)	(0.206)	(0.836)	(0.163)	(0.040)	(0.467)	(0.124)	(1.710)	(0.325)	(4.839)	(1.274)	(17.716)	(0.269)	(3.951)
\mathbf{CLF}	0.798	1.082	1.519	4.002	1.266	0.908	3.582	1.504	9.838	2.238	18.519	6.117	59.151	1.845	14.990
	(0.030)	(0.073)	(0.163)	(0.655)	(0.134)	(0.040)	(0.405)	(0.122)	(1.455)	(0.261)	(3.827)	(0.978)	(14.018)	(0.213)	(3.157)
\mathbf{PCL}	0.810	1.077	1.517	3.842	1.268	0.921	3.454	1.488	9.329	2.232	17.624	5.855	54.863	1.828	14.152
	(0.030)	(0.068)	(0.155)	(0.581)	(0.123)	(0.039)	(0.379)	(0.115)	(1.310)	(0.255)	(3.428)	(0.843)	(11.715)	(0.192)	(2.599)
\mathbf{PL}	0.814	1.092	1.478	3.708	1.235	0.927	3.519	1.509	9.503	2.174	16.713	5.673	52.221	1.785	13.443
	(0.030)	(0.071)	(0.143)	(0.516)	(0.113)	(0.040)	(0.389)	(0.119)	(1.351)	(0.233)	(3.087)	(0.787)	(10.927)	(0.182)	(2.383)
\mathbf{PF}	0.814	1.092	1.478	3.705	1.235	0.927	3.519	1.509	9.501	2.174	16.498	5.664	50.271	1.785	13.305
	(0.030)	(0.071)	(0.143)	(0.515)	(0.113)	(0.040)	(0.389)	(0.119)	(1.350)	(0.233)	(2.979)	(0.782)	(9.702)	(0.182)	(2.313)
							EXTI	RAPOLA	TION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	0.829	1.364	17.130	66.392	14.005	0.965	5.844	1.992	19.827	27.013	427.684	103.075	1609.887	21.746	333.716
	(0.149)	(0.470)	(12.550)	(50.394)	(9.943)	(0.217)	(2.724)	(0.744)	(10.567)	(20.408)	(322.850)	(77.606)	(1185.941)	(16.260)	(247.584)
\mathbf{CLL}	0.999	1.686	3.986	14.051	3.165	1.057	8.052	2.406	27.759	5.978	81.464	21.578	289.203	4.907	64.447
	(0.205)	(0.722)	(2.456)	(9.057)	(1.932)	(0.282)	(4.948)	(1.140)	(17.598)	(3.809)	(55.980)	(14.479)	(202.098)	(3.020)	(43.366)
\mathbf{CLF}	0.994	1.661	3.894	13.857	3.122	1.055	7.134	2.345	24.810	5.778	81.983	21.568	300.801	4.841	65.326
	(0.204)	(0.698)	(2.424)	(8.911)	(1.900)	(0.280)	(4.213)	(1.081)	(14.864)	(3.671)	(56.485)	(14.204)	(213.837)	(2.876)	(43.813)
\mathbf{PCL}	1.081	1.512	16.272	53.208	13.563	0.974	5.669	1.963	19.024	24.295	332.927	80.989	1257.589	20.062	260.499
	(0.182)	(0.481)	(11.345)	(38.871)	(9.379)	(0.218)	(2.779)	(0.742)	(10.810)	(16.848)	(252.104)	(60.487)	(924.251)	(13.953)	(193.253)
\mathbf{PL}	1.460	2.084	4.018	13.903	3.165	1.160	8.168	2.604	27.900	5.966	80.679	21.437	287.237	4.917	63.659
	(0.342)	(0.974)	(2.464)	(8.983)	(1.967)	(0.367)	(5.093)	(1.341)	(17.936)	(3.821)	(56.421)	(14.489)	(204.999)	(3.048)	(42.576)
\mathbf{PF}	1.460	2.084	4.018	13.902	3.165	1.163	8.168	2.604	27.896	5.966	80.473	21.435	283.165	4.917	63.580
	(0.342)	(0.974)	(2.464)	(8.982)	(1.967)	(0.368)	(5.093)	(1.341)	(17.932)	(3.821)	(56.229)	(14.485)	(202.675)	(3.048)	(42.511)

Table 1.19: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 1$ and $\varphi = 0.2$ with aggregation matrix \mathbf{C}_2

							INTE	ERPOLA	ΓΙΟΝ						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	0.794	1.109	1.916	6.016	1.591	0.920	4.403	1.565	11.456	2.913	33.493	9.474	118.053	2.358	25.956
	(0.029)	(0.091)	(0.533)	(2.317)	(0.404)	(0.043)	(0.867)	(0.157)	(2.718)	(0.834)	(15.059)	(3.754)	(58.079)	(0.639)	(11.331)
\mathbf{CLL}	0.901	1.166	1.610	4.277	1.352	1.018	3.959	1.583	10.477	2.375	19.978	6.608	62.928	1.949	16.220
	(0.039)	(0.098)	(0.234)	(0.868)	(0.195)	(0.055)	(0.488)	(0.145)	(1.733)	(0.374)	(4.874)	(1.328)	(18.207)	(0.292)	(3.811)
\mathbf{CLF}	0.899	1.162	1.569	4.038	1.332	1.015	3.818	1.576	9.958	2.290	18.685	6.218	59.180	1.895	15.157
	(0.039)	(0.097)	(0.198)	(0.669)	(0.171)	(0.055)	(0.416)	(0.141)	(1.462)	(0.303)	(3.816)	(1.033)	(14.345)	(0.247)	(2.932)
\mathbf{PCL}	1.103	1.316	1.684	3.934	1.477	1.256	3.861	1.718	9.403	2.384	17.553	5.964	54.764	2.030	14.333
	(0.036)	(0.076)	(0.157)	(0.549)	(0.122)	(0.051)	(0.388)	(0.118)	(1.323)	(0.246)	(3.192)	(0.906)	(11.801)	(0.199)	(2.579)
\mathbf{PL}	1.103	1.325	1.652	3.802	1.449	1.256	3.899	1.736	9.536	2.335	16.760	5.796	52.181	1.985	13.661
	(0.036)	(0.078)	(0.144)	(0.508)	(0.113)	(0.051)	(0.403)	(0.122)	(1.313)	(0.232)	(2.832)	(0.821)	(11.296)	(0.185)	(2.301)
\mathbf{PF}	1.103	1.325	1.652	3.799	1.449	1.256	3.899	1.736	9.534	2.335	16.562	5.786	50.225	1.985	13.531
	(0.036)	(0.078)	(0.144)	(0.506)	(0.113)	(0.051)	(0.403)	(0.122)	(1.312)	(0.232)	(2.751)	(0.816)	(10.071)	(0.185)	(2.238)
							EXTI	RAPOLA	TION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	0.829	1.364	16.404	69.083	13.290	0.982	6.374	1.976	20.339	26.344	408.146	103.234	1593.150	21.020	325.353
	(0.149)	(0.470)	(12.046)	(50.844)	(9.574)	(0.216)	(2.887)	(0.734)	(10.780)	(19.131)	(310.140)	(79.066)	(1180.896)	(15.590)	(239.316)
\mathbf{CLL}	0.999	1.686	3.830	14.203	3.174	1.188	8.239	2.415	28.012	6.247	82.398	21.575	294.880	5.007	64.672
	(0.205)	(0.722)	(2.343)	(9.489)	(2.037)	(0.319)	(4.851)	(1.139)	(17.623)	(3.853)	(55.741)	(14.478)	(202.833)	(3.255)	(43.454)
\mathbf{CLF}	0.994	1.661	3.744	13.904	3.171	1.179	7.396	2.342	24.852	6.100	82.739	21.421	307.673	4.835	65.173
	(0.204)	(0.698)	(2.273)	(9.236)	(1.992)	(0.313)	(4.167)	(1.070)	(15.381)	(3.734)	(56.732)	(14.226)	(215.500)	(3.065)	(44.017)
\mathbf{PCL}	1.081	1.512	15.664	55.149	12.902	1.244	6.102	2.071	19.431	23.728	319.748	80.978	1244.666	19.565	253.241
	(0.182)	(0.481)	(10.863)	(39.277)	(8.987)	(0.259)	(2.894)	(0.741)	(11.004)	(16.002)	(242.227)	(61.561)	(921.245)	(13.558)	(186.697)
\mathbf{PL}	1.460	2.084	3.910	14.037	3.342	1.767	8.340	2.833	27.852	6.289	81.712	21.579	292.838	5.070	64.246
	(0.342)	(0.974)	(2.412)	(9.580)	(2.137)	(0.535)	(4.917)	(1.422)	(18.027)	(3.877)	(55.287)	(14.582)	(203.555)	(3.308)	(43.093)
\mathbf{PF}	1.460	2.084	3.910	14.036	3.342	1.767	8.340	2.833	27.848	6.288	81.489	21.575	290.589	5.070	64.203
	(0.342)	(0.974)	(2.412)	(9.579)	(2.137)	(0.535)	(4.916)	(1.422)	(18.024)	(3.877)	(55.252)	(14.577)	(200.105)	(3.308)	(43.027)

Table 1.20: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 1$ and $\varphi = 0.8$ with aggregation matrix \mathbf{C}_2

							INTEF	RPOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	1.769	2.478	4.292	13.505	3.552	2.027	8.520	3.483	25.087	6.379	73.372	20.513	260.702	5.305	5.278
	(0.066)	(0.208)	(1.163)	(5.207)	(0.932)	(0.095)	(1.318)	(0.348)	(5.654)	(1.826)	(33.018)	(7.749)	(127.482)	(1.498)	(1.529)
\mathbf{CLL}	1.767	2.441	3.540	9.517	2.924	2.021	8.218	3.446	23.368	5.188	44.177	14.606	141.630	4.274	4.280
	(0.064)	(0.170)	(0.451)	(1.899)	(0.354)	(0.090)	(1.024)	(0.290)	(3.820)	(0.755)	(10.892)	(2.959)	(40.020)	(0.581)	(0.625)
\mathbf{CLF}	1.766	2.417	3.402	8.976	2.822	2.018	7.932	3.377	22.227	4.973	41.352	13.698	131.896	4.099	4.118
	(0.064)	(0.160)	(0.364)	(1.470)	(0.290)	(0.090)	(0.895)	(0.260)	(3.234)	(0.601)	(8.424)	(2.254)	(30.686)	(0.462)	(0.497)
\mathbf{PCL}	1.767	2.381	3.348	8.593	2.804	2.015	7.645	3.312	20.962	4.868	38.943	13.140	123.358	4.017	4.097
	(0.064)	(0.151)	(0.339)	(1.279)	(0.275)	(0.088)	(0.854)	(0.242)	(2.996)	(0.540)	(7.282)	(2.014)	(25.951)	(0.416)	(0.432)
\mathbf{PL}	1.774	2.413	3.294	8.341	2.748	2.028	7.794	3.366	21.326	4.797	37.850	12.771	117.919	3.950	4.038
	(0.065)	(0.157)	(0.322)	(1.185)	(0.253)	(0.090)	(0.882)	(0.253)	(3.080)	(0.536)	(7.091)	(1.905)	(25.312)	(0.393)	(0.405)
\mathbf{PF}	1.774	2.413	3.294	8.319	2.748	2.028	7.793	3.366	21.308	4.795	36.875	12.702	112.781	3.949	4.037
	(0.065)	(0.157)	(0.322)	(1.175)	(0.253)	(0.090)	(0.882)	(0.253)	(3.075)	(0.535)	(6.603)	(1.869)	(22.027)	(0.392)	(0.404)
							EXTRA	APOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	1.848	3.045	35.944	148.805	30.192	2.167	13.555	4.383	45.282	58.940	919.082	235.003	3612.535	45.365	46.796
	(0.322)	(0.979)	(26.610)	(114.005)	(22.615)	(0.472)	(6.584)	(1.672)	(24.396)	(43.831)	(704.500)	(172.903)	(2675.236)	(34.747)	(34.884)
\mathbf{CLL}	1.942	3.726	9.023	31.507	7.404	2.338	18.492	5.823	61.757	13.688	180.169	49.012	674.923	11.197	11.189
	(0.376)	(1.664)	(5.564)	(20.973)	(4.659)	(0.587)	(11.513)	(2.954)	(38.643)	(8.984)	(121.745)	(32.524)	(472.268)	(7.040)	(6.716)
\mathbf{CLF}	1.938	3.537	8.707	30.594	7.192	2.323	16.536	5.252	54.845	13.338	180.388	49.006	691.573	10.878	10.921
	(0.374)	(1.475)	(5.317)	(20.751)	(4.482)	(0.574)	(9.721)	(2.447)	(33.475)	(8.602)	(122.964)	(32.620)	(482.595)	(6.756)	(6.458)
PCL	1.844	2.965	31.022	116.371	26.632	2.147	13.169	4.278	44.334	47.617	717.827	183.343	2813.991	37.984	38.758
	(0.323)	(0.967)	(21.338)	(88.846)	(18.372)	(0.471)	(6.690)	(1.677)	(25.068)	(33.669)	(549.734)	(134.894)	(2095.935)	(27.149)	(26.992)
\mathbf{PL}	2.024	3.923	8.935	30.980	7.315	2.469	18.639	5.927	61.766	13.653	178.632	48.500	677.095	11.081	11.267
	(0.445)	(1.889)	(5.583)	(20.987)	(4.699)	(0.719)	(11.677)	(3.053)	(39.213)	(9.077)	(120.079)	(32.633)	(486.180)	(7.127)	(6.886)
\mathbf{PF}	2.024	3.923	8.935	30.966	7.315	2.469	18.638	(5.000) 5.927	(53.213) 61.713	13.652	176.767	48.439	659.881	11.080	(0.000) 11.266
	(0.446)	(1.889)	(5.583)	(20.972)	(4.699)	(0.719)	(11.676)	(3.053)	(39.170)	(9.076)	(118.942)	(32.578)	(466.937)	(7.126)	(6.886)

Table 1.21: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 5$ and $\varphi = 0.2$ with aggregation matrix \mathbf{C}_2

63

							INTEF	RPOLAT	[ON						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	1.774	2.466	4.307	13.364	3.536	2.035	8.854	3.448	24.859	6.511	72.416	20.712	262.635	7.405	7.429
	(0.063)	(0.200)	(1.183)	(5.175)	(0.906)	(0.096)	(1.569)	(0.345)	(5.605)	(1.989)	(33.468)	(7.966)	(131.470)	(2.025)	(2.133)
\mathbf{CLL}	1.821	2.467	3.537	9.505	2.942	2.075	8.365	3.432	23.236	5.245	44.056	14.768	142.824	6.068	6.046
	(0.068)	(0.185)	(0.465)	(1.927)	(0.367)	(0.102)	(1.073)	(0.298)	(3.866)	(0.771)	(10.767)	(2.941)	(41.951)	(0.831)	(0.840)
\mathbf{CLF}	1.819	2.446	3.407	8.967	2.848	2.072	8.066	3.375	22.075	5.017	41.298	13.858	133.016	5.808	5.805
	(0.068)	(0.175)	(0.373)	(1.487)	(0.299)	(0.102)	(0.925)	(0.272)	(3.213)	(0.596)	(8.407)	(2.238)	(31.830)	(0.654)	(0.678)
\mathbf{PCL}	1.925	2.492	3.421	8.641	2.889	2.192	7.839	3.386	20.868	4.954	39.035	13.238	122.931	5.667	5.740
	(0.068)	(0.158)	(0.339)	(1.341)	(0.261)	(0.100)	(0.864)	(0.254)	(2.839)	(0.537)	(7.382)	(1.986)	(26.771)	(0.592)	(0.596)
\mathbf{PL}	1.930	2.523	3.359	8.398	2.839	2.202	7.962	3.437	21.235	4.878	37.935	12.917	118.439	5.599	5.664
	(0.068)	(0.163)	(0.319)	(1.225)	(0.242)	(0.100)	(0.890)	(0.262)	(2.908)	(0.505)	(7.065)	(1.846)	(25.471)	(0.592)	(0.586)
\mathbf{PF}	1.930	2.523	3.359	8.375	2.839	2.202	7.962	3.437	21.218	4.877	37.025	12.855	113.050	5.597	5.661
	(0.068)	(0.163)	(0.319)	(1.213)	(0.242)	(0.100)	(0.890)	(0.262)	(2.903)	(0.504)	(6.548)	(1.822)	(22.116)	(0.590)	(0.585)
							EXTRA	APOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	1.835	3.033	37.549	146.575	29.795	2.166	13.609	4.339	45.222	58.846	905.648	229.177	3595.683	66.791	65.690
	(0.345)	(1.018)	(28.539)	(113.943)	(20.966)	(0.466)	(6.779)	(1.603)	(25.164)	(43.585)	(678.628)	(169.597)	(2811.117)	(49.508)	(46.032)
\mathbf{CLL}	1.998	3.770	8.562	31.379	7.045	2.407	18.364	5.681	63.674	14.115	174.928	47.902	625.629	16.068	15.708
	(0.409)	(1.730)	(5.244)	(21.193)	(4.186)	(0.618)	(10.979)	(2.873)	(40.071)	(9.047)	(123.466)	(32.451)	(440.881)	(9.778)	(9.739)
\mathbf{CLF}	1.990	3.575	8.221	30.735	6.828	2.383	16.293	5.175	56.370	13.648	175.787	47.372	641.208	15.529	15.321
	(0.404)	(1.550)	(5.116)	(20.900)	(4.062)	(0.599)	(9.427)	(2.392)	(35.036)	(8.690)	(121.813)	(31.957)	(457.433)	(9.432)	(9.657)
\mathbf{PCL}	1.961	3.057	32.045	114.812	26.331	2.276	13.144	4.295	44.491	47.703	705.650	179.060	2792.766	53.340	52.926
	(0.365)	(1.006)	(22.383)	(88.791)	(17.054)	(0.482)	(6.812)	(1.594)	(26.159)	(33.576)	(530.671)	(132.061)	(2196.379)	(38.326)	(35.561)
\mathbf{PL}	2.311	4.097	8.512	31.172	7.140	2.792	18.638	5.853	63.851	14.094	175.437	47.453	641.179	15.849	15.561
	(0.582)	(2.000)	(5.124)	(21.237)	(4.329)	(0.866)	(11.181)	(3.025)	(40.379)	(9.007)	(123.522)	(32.686)	(452.840)	(9.932)	(9.817)
\mathbf{PF}	2.311	4.097	8.512	31.162	7.140	2.792	18.637	5.853	63.796	14.093	172.226	47.411	617.150	15.848	15.559
	(0.582)	(2.000)	(5.124)	(21.223)	(4.328)	(0.866)	(11.180)	(3.025)	(40.333)	(9.005)	(120.942)	(32.652)	(437.837)	(9.930)	(9.815)

Table 1.22: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 5$ and $\varphi = 0.8$ with aggregation matrix \mathbf{C}_2

The row represents the average and (standard deviation) of the MSEs of HF estimates by a particular model (row-wise) for 1000 artificial samples of a DGP (column-wise).

							INTE	RPOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.505	3.512	6.138	18.783	4.969	2.873	12.068	4.872	35.508	9.125	101.463	29.324	372.292	7.405	81.599
	(0.095)	(0.299)	(1.673)	(7.091)	(1.218)	(0.133)	(2.008)	(0.458)	(7.695)	(2.517)	(45.710)	(11.228)	(192.278)	(2.025)	(35.959)
\mathbf{CLL}	2.499	3.469	4.953	13.404	4.146	2.865	11.623	4.839	32.815	7.312	62.343	20.634	197.971	6.068	50.070
	(0.093)	(0.255)	(0.626)	(2.645)	(0.505)	(0.126)	(1.462)	(0.408)	(5.345)	(1.039)	(14.722)	(4.245)	(54.610)	(0.831)	(12.158)
\mathbf{CLF}	2.498	3.420	4.779	12.661	3.992	2.859	11.225	4.744	31.312	6.998	58.447	19.357	185.285	5.808	46.953
	(0.093)	(0.232)	(0.509)	(2.099)	(0.407)	(0.125)	(1.267)	(0.373)	(4.503)	(0.826)	(11.356)	(3.164)	(43.961)	(0.654)	(9.594)
PCL	2.494	3.368	4.699	12.127	3.937	2.848	10.834	4.667	29.516	6.867	55.227	18.420	171.362	5.667	44.378
	(0.092)	(0.217)	(0.478)	(1.783)	(0.385)	(0.122)	(1.168)	(0.354)	(4.126)	(0.788)	(10.112)	(2.735)	(37.493)	(0.592)	(7.957)
\mathbf{PL}	2.503	3.410	4.624	11.837	3.882	2.866	11.023	4.739	30.001	6.759	54.022	18.013	165.368	5.599	43.000
	(0.093)	(0.228)	(0.462)	(1.749)	(0.362)	(0.125)	(1.223)	(0.368)	(4.261)	(0.747)	(9.640)	(2.637)	(35.247)	(0.592)	(8.116)
\mathbf{PF}	2.503	3.410	4.623	11.783	3.882	2.866	11.022	4.739	29.961	6.754	52.335	17.874	157.888	5.597	41.824
	(0.093)	(0.228)	(0.462)	(1.722)	(0.362)	(0.125)	(1.222)	(0.368)	(4.247)	(0.745)	(8.824)	(2.558)	(31.369)	(0.590)	(7.426)
							EXTR	APOLAT	ION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.582	4.260	52.273	203.764	42.436	3.101	18.607	6.281	65.614	83.436	1279.115	318.096	4837.577	66.791	1046.570
	(0.464)	(1.367)	(38.453)	(153.774)	(31.632)	(0.664)	(8.933)	(2.298)	(36.113)	(60.603)	(951.729)	(248.804)	(3747.381)	(49.508)	(767.398)
CLL	2.704	5.342	12.780	44.867	10.207	3.392	25.807	8.543	88.799	19.340	246.877	71.646	948.718	16.068	200.569
	(0.527)	(2.359)	(8.043)	(28.583)	(6.105)	(0.889)	(15.006)	(4.163)	(58.452)	(12.018)	(169.517)	(47.101)	(631.369)	(9.778)	(143.451)
\mathbf{CLF}	2.694	4.932	12.281	43.959	9.828	3.347	22.835	7.659	79.414	18.878	248.313	69.334	983.171	15.529	203.011
	(0.519)	(1.981)	(7.774)	(28.421)	(5.861)	(0.846)	(12.692)	(3.423)	(50.904)	(11.738)	(173.137)	(46.484)	(667.006)	(9.432)	(142.890)
PCL	2.573	4.172	42.623	159.113	35.659	3.079	17.921	6.184	64.283	65.996	995.771 [´]	247.550	3757.481	53.340	815.243
	(0.456)	(1.347)	(29.788)	(120.223)	(24.760)	(0.656)	(9.110)	(2.289)	(38.351)	(46.774)	(743.189)	(193.020)	(2936.218)	(38.326)	(598.037)
\mathbf{PL}	2.803	5.432	12.688	43.885	10.222	3.556	25.840	8.625	89.073	19.312	248.082	70.834	961.526	15.849	200.794
	(0.606)	(2.467)	(8.115)	(28.077)	(6.208)	(1.045)	(15.335)	(4.306)	(59.952)	(11.938)	(174.318)	(47.020)	(634.799)	(9.932)	(145.311)
\mathbf{PF}	2.803	5.432	12.687	43.850	10.221	3.556	25.837	8.625	88.905	19.308	242.292	70.617	932.282	15.848	197.595
	(0.606)	(2.467)	(8.114)	(28.046)	(6.208)	(1.045)	(15.332)	(4.306)	(59.797)	(11.936)	(168.838)	(46.814)	(627.754)	(9.930)	(141.963)

Table 1.23: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 10$ and $\varphi = 0.2$ with aggregation matrix \mathbf{C}_2

The row represents the average and (standard deviation) of the MSEs of HF estimates by a particular model (row-wise) for 1000 artificial samples of a DGP (column-wise).

							INTE	RPOLAT	TION						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.512	3.486	6.064	18.807	5.052	2.873	12.296	4.889	35.234	9.272	103.285	29.753	373.114	7.429	82.111
	(0.090)	(0.287)	(1.632)	(7.206)	(1.317)	(0.133)	(1.978)	(0.487)	(8.063)	(2.857)	(48.731)	(11.382)	(186.366)	(2.133)	(36.247)
\mathbf{CLL}	2.542	3.477	4.992	13.591	4.198	2.895	11.825	4.859	32.624	7.456	62.854	20.681	199.215	6.046	50.243
	(0.095)	(0.256)	(0.681)	(2.869)	(0.559)	(0.135)	(1.542)	(0.422)	(5.494)	(1.134)	(14.744)	(4.347)	(54.317)	(0.840)	(12.197)
\mathbf{CLF}	2.540	3.432	4.791	12.774	4.040	2.889	11.378	4.762	31.070	7.142	58.977	19.436	186.796	5.805	47.236
	(0.095)	(0.233)	(0.528)	(2.172)	(0.437)	(0.133)	(1.342)	(0.388)	(4.617)	(0.893)	(11.821)	(3.321)	(43.032)	(0.678)	(9.481)
PCL	2.611	3.440	4.747	12.191	4.012	2.968	11.005	4.731	29.218	7.000	55.571	18.632	174.215	5.740	44.389
	(0.095)	(0.215)	(0.458)	(1.838)	(0.369)	(0.131)	(1.276)	(0.372)	(4.244)	(0.769)	(10.121)	(2.864)	(37.053)	(0.596)	(8.145)
\mathbf{PL}	2.619	3.484	4.678	11.906	3.959	2.983	11.196	4.802	29.704	6.913	54.042	18.192	167.440	5.664	42.989
	(0.096)	(0.221)	(0.451)	(1.742)	(0.359)	(0.133)	(1.315)	(0.386)	(4.346)	(0.738)	(9.606)	(2.751)	(34.834)	(0.586)	(7.670)
\mathbf{PF}	2.619	3.484	4.677	11.852	3.958	2.983	11.195	4.802	29.664	6.908	52.397	18.048	159.844	5.661	41.884
	(0.096)	(0.221)	(0.450)	(1.716)	(0.359)	(0.133)	(1.314)	(0.386)	(4.336)	(0.736)	(8.857)	(2.692)	(30.388)	(0.585)	(7.071)
							EXTR	APOLA	ΓΙΟΝ						
	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
CLCL	2.594	4.299	53.208	207.847	41.883	3.042	19.206	6.193	64.232	83.814	1297.589	324.432	5152.244	65.690	1047.137
	(0.454)	(1.468)	(39.396)	(153.094)	(31.614)	(0.653)	(9.402)	(2.332)	(36.409)	(62.020)	(1002.559)	(246.730)	(3732.258)	(46.032)	(773.592)
\mathbf{CLL}	2.779	5.503	12.914	45.764	9.929	3.353	26.266	8.101	87.585	19.585	251.588	66.777	916.470	15.708	203.613
	(0.574)	(2.593)	(8.290)	(31.502)	(6.029)	(0.867)	(16.176)	(4.001)	(55.822)	(12.320)	(172.127)	(46.305)	(657.075)	(9.739)	(143.113)
\mathbf{CLF}	2.764	5.073	12.379	44.701	9.507	3.299	23.454	7.316	78.565	18.961	254.223	65.500	938.457	15.321	208.128
	(0.561)	(2.165)	(7.935)	(30.667)	(5.853)	(0.823)	(13.730)	(3.325)	(47.998)	(11.756)	(176.022)	(45.414)	(681.952)	(9.657)	(148.044)
PCL	2.679	4.244	43.543	162.411	35.036	3.101	18.487	6.064	63.218	66.323	1012.519	252.470	4001.301	52.926	817.163
	(0.482)	(1.470)	(30.357)	(119.525)	(24.594)	(0.660)	(9.433)	(2.332)	(37.501)	(48.142)	(782.869)	(193.301)	(2917.998)	(35.561)	(606.422)
\mathbf{PL}	3.045	5.675	12.989	45.399	9.891	3.692	26.082	8.208	88.727	19.575	252.155	66.611	949.225	15.561	201.520
	(0.754)	(2.737)	(8.365)	(31.198)	(6.058)	(1.125)	(16.394)	(4.183)	(57.309)	(12.358)	(175.987)	(46.254)	(681.914)	(9.817)	(145.461)
\mathbf{PF}	3.045	5.675	12.988	45.336	9.891	3.692	26.079	8.208	88.564	19.571	247.584	66.357	906.342	15.559	199.847
	(0.754)	(2.737)	(8.364)	(31.160)	(6.057)	(1.125)	(16.392)	(4.182)	(57.159)	(12.355)	(171.467)	(46.064)	(662.608)	(9.815)	(143.432)

Table 1.24: The average MSEs of the Monte Carlo HF estimates by the six models under various DGPs with $\sigma^2 = 10$ and $\varphi = 0.8$ with aggregation matrix \mathbf{C}_2

The row represents the average and (standard deviation) of the MSEs of HF estimates by a particular model (row-wise) for 1000 artificial samples of a DGP (column-wise).

Chapter 2

The role of economic and financial indicators in forecasting stock volatility

Abstract

This paper examines the value-added in forecasting volatility by utilising mixed frequency information from market sentiment indicators, economic variables, and activity measures. We consider high-frequency data to generate the volatility measures and low frequency to capture the economic indicators using a mixed sampling frequency approach. The data cover SPY and 100 representative stocks from ten business sectors from 2000 to 2016. Findings are presented on a multi-dimensional scale with three forecasting horizons (daily, weekly, and monthly) and three regimes, including the 2007-08 financial crisis. We find that (extrapolated) macroeconomic and market sentiment indicators exhibit forecasting information for future volatility beyond its lagged values. Further, we find that utilising a representative factor(s) of all potential predictors results in significant forecast gains in predicting long-term financial volatility even during the financial crisis.

Keywords: Realised Volatility; Chow-Lin interpolation; Financial volatility predictors; HARX; Mixed Frequency

2.1 Introduction

Since accurate forecasts can inform action to mitigate financial risk, market traders and regulators take great interest in stock market price volatility. In a seminal approach by Engle (1982), the ARCH model captures the time-varying conditional volatility of financial returns. In an extension to that work, Bollerslev (1986) proposed to use the GARCH model. In a further development by Corsi (2009), the Heterogeneous Autoregressive Realised Volatility (HAR-RV) model presents daily volatility as a sequence of autoregressive components, as realised over daily, weekly and monthly horizons. The general hypothesis is that different categories of market participants have diverse horizons and trading activities, such that they react differently to similar news (Müller et al., 1997). Hence, different market indicators carry distinctive signals. This paper aims to combine financial and economic signals from high and low-frequency variables and test the value-added from each piece of information in forecasting financial volatility.

Considering that stock price volatility, unlike stock returns, is persistent, research papers tend to construct volatility models based on time series information; for example, see Corsi and Reno (2009), Ghysels, Santa-Clara, and Valkanov (2006), and Wang et al. (2016). Researchers explore other factors that can help predict financial volatility; see, for example, Hamilton and Lin (1996), Christiansen, Schmeling, and Schrimpf (2012), and Caporin, Rossi, and Magistris (2015), among others. The overall impact of the business cycle upon financial activity has motivated Officer (1973) followed by Schwert (1989), Hamilton and Susmel (1994), Hamilton and Lin (1996), and So, Lam, and Li (1998). More recently, Bollerslev and Zhou (2006) and Bollerslev and Todorov (2011) finds an association between risk-premium volatility and a number of macroeconomic variables. The general finding is one of increased volatility during recessions. All studies mentioned above have considered economic variables observed on a low-frequency (LF), i.e. monthly or annual. This paper investigates the role of a diverse set of economic and financial indicators in predicting high-frequency (HF), i.e. daily, financial volatility.

Where the record of financial volatility is typically available on (at least) a daily basis, less frequent weekly and monthly values are available for macroeconomic and other relevant variables. Most studies overcome the statistical challenge of modelling mixed frequency data by aggregating the HF variable to match the low-frequency (LF) variables or disaggregating from LF to HF, such as the Chow and Lin (1971) method. Ghysels, Santa-Clara, and Valkanov (2004) suggest a mixed data sampling model, MIDAS, to include HF explanatory variables to predict a LF one. Ghysels (2016) introduced and developed the mixed frequency vector autoregressive, MFVAR, to allow the study of bi-directional causality between LF and HF variables. However, when the frequency ratio of the variables is high, the MFVAR suffers from the curse of dimensionality. For example, if the model includes daily and monthly variables, the vector in the MFVAR consists of the daily and monthly observations within a month. Further, the announcement day of the economic data changes from one calendar month to another, posing a challenge in ordering the variables (according to their occurrence within the calendar month) in the VAR model.

To identify the impact of selected indicators on financial market volatility, we use the HARX framework (i.e., the HAR model with additional variable(s)). The HARX model can be applied directly using explanatory variables that are recorded on HF occurrences (e.g. daily). However, we consider two approaches for variables observed at a lower frequency. The first is the reverse MIDAS by Foroni, Guérin, and Marcellino (2018) where the LF explanatory variables are added to the HARX model without any frequency conversion; the second approach is to use the temporal-extrapolation method of Chow and Lin (1971) to disaggregate the LF series into a HF series. The disaggregated series can then be added as daily explanatory variables in the HARX model. We find the latter approach to generally yield better forecasting results.

A growing body of literature suggests that a subset selection and the derived components of variables from a feasible set improve forecast competence (see Fuentes, Poncela, and Rodríguez (2015) and references therein). Therefore, in addition to studying each indicator's added value, we also examine the benefits of combining all the predictors and adding them to the HARX model. Chun and Keleş (2010) propose the Sparse Partial Least Square (SPLS) approach, which imposes a sparsity constraint, similar to that of LASSO, on the Partial Least Square (PLS) method of Wold (1966). The constraint leads to sparse linear combinations of the original predictors. Since this paper is concerned with forecasting, we employ the SPLS to derive orthogonal unobserved components based on the covariance between the set of predictors and the dependent variable. The derived SPLS factors enter the HARX model as explanatory variables. Overall, we find that all indicators' combined information produces the best forecasts. Information about the economic and market conditions, in addition to the lagged volatility, provide further gains in predicting future stock volatility.

In this paper, we contribute to the literature by empirically investigating which economic and financial indicators exhibit essential data in forecasting HF financial volatility of the market and 100 individual stocks in ten business sectors. We also examine when unique and combined factors are most significant in providing pertinent information to predicting future volatility. We do so by studying the forecasting performance over three regimes: before, during, and post the 2007-08 financial crisis. We find that market sentiments hold valuable information in predicting financial volatility before the financial crisis. The value-added of macroeconomic variables, such as new housing projects and industrial production indices, in forecasting stock volatility increased since the financial crisis.

This paper is organized as follows. Section 2.2 provides background on the various predictors of financial volatility investigated in the literature. Section 2.3 presents the volatility measures and HARX models employed in this paper. Section 2.4 describes the data used in this study. Section 2.5 presents the findings and analysis using forecast evaluation methods. Section 2.6 concludes.

2.2 Predictors of the financial volatility

Due to stocks' persistent financial volatility, various papers often construct volatility models with distributed lags. Hence, the forecasting accuracy can often be improved by incorporating information beyond those included in the volatility's distributed lag. Therefore, exploring factors that contain significant information in forecasting volatility has occupied a number of researchers, see for example Ghysels, Santa-Clara, and Valkanov (2006), Hamilton and Lin (1996), and Christiansen, Schmeling, and Schrimpf (2012) among others. We categorise indicators as macroeconomic, market sentiment and stock-specific activity.

Macroeconomic Indicators. Green (2004) and Andersen et al. (2003b) observe significant intraday effects of macroeconomic announcements. Corradi, Distaso, and Mele (2012) find that macroeconomic factors can explain most of the long-term variation in overall stock volatility. Beltratti and Morana (2006) found that the causality from macroeconomics to financial volatility is stronger than vice versa. In general, any strong contemporaneous relationship between financial volatility and economic conditions would suggest that the use of lagged economic variables might improve volatility forecasts.

We consider macroeconomic indicators with different classifications relative to the business cycle. In particular, the term spread, which is the difference between the three-month LIBOR rate and T-Bill rate, is a forward-looking indicator as it reflects the investors' expectations for future economic activity, demand for credit and monetary policy (see Conrad and Loch (2015) and the references therein). Similarly, the federal fund rates variable is shown to be predictive of volatility jumps (Caporin, Rossi, and Magistris, 2015). Another leading indicator is the housing starts index (HOUST), defined as the total new privately-owned housing units started. It signals the residential investments, so it leads GDP as shown by Kydland, Rupert, and Šustek (2016), among others. The industrial production index (INDPRO) is a coincident

indicator, where an increase in the total industrial output yields a decrease in financial volatility (Officer, 1973; Schwert, 1989).

Another coincident indicator we consider is the Auroba-Diebold-Scotti index (ADS) by Aruoba, Diebold, and Scotti (2009). It reflects the prevailing economic conditions combining jobless claims, growth of payroll enrollment, industrial production growth, real manufacturing, trade sales, real personal income, and real GDP. The unemployment rate is a lagging indicator of the business cycle. For example, due to a recession, people have a more challenging time landing jobs (Conrad and Loch, 2015). Similarly, inflation and inflation uncertainty, Producer Price Index (PPI), are both lagging indicators representing the change in prices (Engle, Ghysels, and Sohn, 2013).

Market Sentiments. Market sentiments signal current and expected business conditions. Andersen, Bollerslev, and Cai (2000) show that indices for consumer confidence (CCI) and business confidence (BCI) can be used to predict financial volatility in the Japanese market. Further, option prices, as reflected in implied volatility, possess information about future stock market volatility (Andersen and Bondarenko, 2007; Chernov, 2007). Becker, Clements, and McClelland (2009) show that the Chicago Board Options Exchange (CBOE) volatility index (VIX) delivers information relevant to future jump activity. Estimated from options prices, squared VIX is interpreted as the conditional return variance using a risk-neutral probability measure. Hence, it is an unbiased predictor of future financial volatility because it incorporates a risk premium (Bekaert and Hoerova, 2014).

Activity Measures. Researchers examined the empirical relationship between stock volatility and trading volume. On the one hand, the mixture of distribution hypothesis (MDH) explains that the volatility-volume relation is a correlation one because any new information yields contemporaneous changes in volume and volatility (Clark, 1973; Harris, 1987). On the other hand, the sequential information arrival hypothesis (SIAH) illustrates that lagged values of volatility may predict current trading volume and vice versa due to the time gap between news and investors' different reactions (Copeland, 1976; Smirlock and Starks, 1988). A positive lead-lag relationship between financial volatility and volume or number of trades has been demonstrated in several empirical studies (Darrat, Zhong, and Cheng, 2007; Girard and Biswas, 2007).

2.3 Measures and Models

Volatility measures

Advances in computing and data technology make it possible to observe HF transaction data. This has led to the so-called realised volatility, RV, which is a non-parametric ex-post estimate of the return variation. The daily RV measure is the sum of intraday squared returns of sub-intervals of the day. Andersen et al. (2003b) show that in the absence of microstructure noise, RV, calculated using HF data, is a consistent estimator of the quadratic variation. Sub-intervals of length 300 sec (e.g. 5 min), which is 78 sub-intervals a day¹, resemble a balance between information gains from high-frequency data and micro-structure effects (Andersen et al., 2001).

Owing to its stylised facts and other desirable statistical properties, RV is preferable over the parametric volatility measures generated from GARCH and Stochastic Volatility (SV) models (Andersen et al., 2003a; Barndorff-Nielsen and Shephard,

¹See Zhang, Mykland, and Aït-Sahalia (2005) for a discussion on optimising the sampling frequency for the estimation of RV.

2002; Liu, Patton, and Sheppard, 2015). The superiority of RV arises from utilising information available at small intervals, which is lost at the lower frequencies of GARCH and SV models.

The price at the start of the *jth* interval is computed as the average of the closing and opening prices of intervals j - 1 and j, respectively. The *jth* return over an interval of length h is:

$$r_j = log(P_{jh}) - log(P_{(j-1)h})$$

Then, dividing the day into M sub-intervals, the intra-day return of the *jth* sub-interval within the *ith* day is:

$$r_{j,i} = log(P_{(i-1)M+j}) - log(P_{(i-1)M+(j-1)}), j = 1, 2, ..., M$$

where M = 78 which is the length of the day with a 5 min sampling frequency.

As a result the daily return is

$$r_t = \sum_{j=1}^M r_{j,t}$$

Hence, the daily realised volatility is

$$RV_t^{(d)} := \Sigma_{j=1}^M r_{j,t}^2$$

 $RV_t^{(w)}$ and $RV_t^{(m)}$ are the weekly and monthly average of RV calculated as follow:

$$RV_t^{(w)} = \frac{1}{5} \Sigma_{i=1}^5 RV_{t-i}^{(d)} \quad ; \qquad RV_t^{(m)} = \frac{1}{22} \Sigma_{i=1}^{22} RV_{t-i}^{(d)}$$

HAR-X models

The Heterogeneous Autoregressive Realised Volatility (HAR) model introduced by Corsi (2009) represents daily volatility as a sequence of autoregressive volatility components realised over daily, weekly and monthly time horizons:

$$RV_t^{(d)} = \beta_0 + \beta_d RV_{t-1}^{(d)} + \beta_w RV_{t-1}^{(w)} + \beta_m RV_{t-1}^{(m)} + \epsilon_t^{(d)}$$
(2.1)

Consider the HARX model:

$$RV_t^{(d)} = \beta_0^{(d)} + \beta_d^{(d)} RV_{t-1}^{(d)} + \beta_w^{(d)} RV_{t-1}^{(w)} + \beta_m^{(d)} RV_{t-1}^{(m)} + \beta_x^{(d)} X_{t-1} + \epsilon_t$$
(2.2)

Similarly, for weekly or monthly realised volatility:

$$RV_t^{(w)} = \beta_0^{(w)} + \beta_d^{(w)} RV_{t-5}^{(d)} + \beta_w^{(w)} RV_{t-5}^{(w)} + \beta_m^{(w)} RV_{t-5}^{(m)} + \beta_x^{(w)} X_{t-5} + \epsilon_t^{(w)}$$
(2.3)

$$RV_t^{(m)} = \beta_0^{(m)} + \beta_d^{(m)} RV_{t-22}^{(d)} + \beta_w^{(m)} RV_{t-22}^{(w)} + \beta_m^{(m)} RV_{t-22}^{(m)} + \beta_x^{(m)} X_{t-22} + \epsilon_t^{(m)}$$
(2.4)

where X is an exogenous variable sampled at a daily level. Some variables sampled at a LF can be incorporated in the model using one of two approaches.

The first approach is the reverse MIDAS suggested by Foroni, Guérin, and Marcellino (2018) to use LF data in predicting HF variables:

$$RV_t^{(d)} = \beta_0^{(d)} + \beta_1 B(L,\theta) RV_{t-1}^{(d)} + \beta_x^{(d)} X_{t-1} + \epsilon_t$$
(2.5)

where $B(L, \theta)$ is a weighting function. Note that the HAR model can be thought of as

a particular weighting scheme. While it is interesting to study the optimal restricted AR in modelling RV, it is beyond the scope of this paper. Our approach is to apply Foroni, Guérin, and Marcellino (2018)'s reverse approach to the HARX model. Where this approach holds the LF observation constant within each LF period (month) it allows its coefficient to vary according to the day of the month:

$$\beta_x X_t := \sum_{i=1}^k \gamma_i D_i X_\tau^{LF}$$

where k is the maximum frequency ratio (i.e number of days in a month). γ_i is the coefficient to be estimated, D_i is a dummy variable that takes the value one when t falls on the *i*th day of the month and zero otherwise. X_{τ}^{LF} is the low-frequency observation of variable X in month τ such that t falls in month τ .

The second approach to incorporating a LF (monthly) variable into the HARX model is to disaggregate the LF variable to daily values before adding these as a HF series. Among the numerous extrapolation methods, the Chow and Lin (1971) (CL) is selected because macroeconomic variables' statistical characteristics align with the method's assumptions. Particularly, macroeconomic variables are thought to have low variance within the month. In a nutshell, the CL method uses a statistical relationship between LF data and HF indicator variable(s) to create a HF time series that is consistent with the LF data. By our choice of the Aruoba-Diebold-Scotti Business Conditions Index (ADS) and the daily Term Spread (TS) as daily indicators, we obtain signals of the current state of the business cycle (Berge and Jordà, 2011). Note that, unlike interpolation, no future observation is used in extrapolation, an important feature of the estimated HF series in the out-of-sample window to be used in real-time forecasting. For a more elaborate technical explanation see Dagum and Cholette (2006) and Chow and Lin (1971).

Finally, we add all the explanatory variables to the HARX model. We use the Spare Partial Least Square (SPLS) method of Chun and Keleş (2010) to extract components from the original set of predictors. The advantage of this method is that instead of adding all the predictors to the HARX model, we replace them with a smaller set of factors that capture their joint variation while also considering the in-sample covariation with the dependent variable. The HARX model with SPLS factors is as follows:

$$RV_{t}^{(d)} = \beta_{0}^{(d)} + \beta_{d}^{(d)}RV_{t-1}^{(d)} + \beta_{w}^{(d)}RV_{t-1}^{(w)} + \beta_{m}^{(d)}RV_{t-1}^{(m)} + \beta_{f}^{'(d)}f_{t-1} + \epsilon_{t}$$

$$f_{t} = Wz_{t}$$
(2.6)

where z_t is the $(M \times 1)$ vector of all M predictors listed in table (2.1) at time t. Note that in this model the $X = f_{t-1}$ where f_t is a $(k \times 1)$ vector of latent unobserved factors obtained using SPLS and β_f is its $(k \times 1)$ coefficients vector. W is a $(k \times M)$ matrix of the weights assigned to each one of the predictor variables in each of the SPLS component in f_t .

Firstly, the direction matrix, W in equation (2.6), is found using SPLS. Secondly, once the factor(s), \hat{f}_t , is estimated, it enters the HARX model, equation (2.6), to serve as a reduced set of explanatory variables. We do not include RV lags in the z_t and instead keep the HAR framework.

Estimation is undertaken using OLS. While the error terms suffer from autocorrelation and heteroskedasticity, the GLS estimation shows no significant difference in the coefficients or the forecasting evaluation measures. Therefore, we use OLS for its simplicity and computational convenience. Lastly, we follow Swanson and White (1995) and use an "insanity filter" to ensure that the volatility forecasts are not negative in all the models. In particular, we replace negative forecasted values of any model by the the lagged RV.

2.4 Data

We consider 17 years of high-frequency data for the period 2000 to 2016. The dataset includes the SPY and 100 stocks representing ten sectors. The financial data comprises 4277 trading days observed at the tick level. The use of cleaned data from $TickWrite^2$ makes our results easier to authenticate and replicate. The sample range and sector coverage allow us to examine the sensitivity of the forecasting performance of our models across different market regimes and individual stocks. Table (2.5) provides descriptive statistics of the RV of each stock.

Table (2.1) provides a summary of the explanatory variables, X, in the HARX models (eq. 2.2, 2.3, 2.4) and Table (2.6) presents their descriptive statistics. The variables were downloaded from multiple sources namely: *ALFRED*, *OECD*, and *QUANDL*. Data revisions can be substantial for macroeconomic variables. Thus, employing revised instead of first-release real-time data can be misleading regarding forecasting evaluation (see, for example, Stark (2010)). Hence, for variables that undergo revision, we use *ALFRED*, which allows retrieving vintage versions of the economic data

²TickWrite is a database that provides data on a commercial basis for futures, Index and equity markets. Tick Data is sourced from NYSE's TAQ (Trade and Quote) database. Tick adjusts the TAQ database for ticker mapping, code filtering, price splits, and dividend payments https://www.tickdata.com/.

recorded at the initial release date of each observation. In addition to macroeconomic variables growth rates, we also considered business indices and squared VIX as market sentiments variables and activity measures.

Table 2.1: List of Predicto	rs

Macroeconomic Variables	Frequency	Functional Form	Source
Aruoba-Diebold-Scotti Business Condi-	Daily	Level	QUANDL
tions Index (ADS)			
Effective Federal Funds Rate (FED-	Daily	Logarithmic-square	ALFRED
$\mathrm{FUNDS})$			
Housing Starts Index (HOUST)	Monthly	Growth	ALFRED
Industrial Production Index (INDPRO)	Monthly	Growth	ALFRED
Produce Price Index (PPIACO)	Monthly	Level	ALFRED
Term Spread (TS)	Daily	Logarithmic-square	ALFRED
Unemployment Rate (UNRATE)	Monthly	Growth	ALFRED
Market Sentiments	Frequency	Functional Form	Source
Economic Policy Uncertainty Index	Daily	Level	Policy Uncertainty
(EPU)			
Business Confidence Index (BCI)	Monthly	Level	OECD
Consumer Confidence Index (CCI)	Monthly	Level	OECD
Composite Leading Indicator (CLI)	Monthly	Level	OECD
CBOE Volatility Index (VIX)	Daily	Squared	ALFRED
Activity Measures	Frequency	Functional Form	Source
Number of Trades	Daily	$Level/(10^4)$	TickWrite

2.5 Empirical Results

The explanatory variables under consideration are observed at either daily or monthly frequency. Hence, given that RV's frequency is daily, we have two scenarios based on the frequency of the explanatory variable, X. For explanatory variables with daily frequency, we employ the HARX models (eq. (2.2), (2.3), (2.4)) directly. For explanatory variables with monthly frequency, we use its extrapolated daily data. We also considered the reverse MIDAS approach and found no significant differences with the former in the forecasting results of weekly and monthly horizons. However, for the daily forecast horizon, the CL extrapolation approach, which accounts for daily variations, yields better predictions. Therefore, we only report the results using the CL approach for conciseness.

In what follows, we report the results using the full sample period and sub-samples. Since the time frame of the data includes the 2007-08 financial crisis, we dissect the full sample into three periods: before the financial crisis (Pre-Crisis) from 2000 to 2006, during the financial crisis (Crisis) from 2007 to 2010, and after the financial crisis (Post-Crisis) from 2011 to 2016.

2.5.1 In-sample Estimation

Tables (2.7), (2.8), and (2.9) report the in-sample estimation results of the HAR and HARX models (eq. 2.2, 2.3, & 2.4) with each of the exogenous variables listed in table (2.1) for three sample periods (pre-crisis, crisis, and post-crisis), respectively. The scale of the HAR model coefficients varies across forecast horizons. In particular, $\hat{\beta}_d$ is the highest for equation (2.2) than (2.3) and (2.4). That is to say, more recent daily information, lagged $RV^{(d)}$, is more relevant for the daily forecaster than the longer-term forecaster. Similarly, $\hat{\beta}_w$ and $\hat{\beta}_m$ are higher in long-term predictions models. Further, the magnitude and significance of the HAR and HARX coefficients change over the three sample periods. During the crisis period, $\hat{\beta}_w$ increases while $\hat{\beta}_m$ decreases, which suggests that recent information becomes more valuable during the crisis than the historical average.

Few magnitude and sign changes in the exogenous variables' coefficients are identified across the three periods, but the pattern is consistent across the three forecast horizons. Generally, the magnitude of most of the X variables' coefficients increases during the crisis. Regarding market sentiments, BCI and CLI reflect current conditions and signals about the economy's future, respectively. Hence, an increase in their values yields a decline in the volatility forecasts. Also, CCI measures future developments in households' consumption and saving. Hence, there has been a significant negative relationship between CCI and volatility forecasts since the crisis. Further, the VIX's coefficient is positive and significant in all three sample periods. Lastly, the coefficients of volume and trades become positive and significant since the crisis, which implies that high volume and trades signal high RV of the next day, week, and month.

In terms of the macroeconomic variables, overall, the sign of their coefficients is in line with the counter-cyclical property of long-term volatility observed by Engle, Ghysels, and Sohn (2013) and Conrad and Loch (2015), among others, even at a daily level. We find that the signs of the coefficients of leading indicators such as TS and HOUST are consistently positive and negative, respectively, across the three sample periods. As for lagging indicators, a rise in UE or PPI is linked with an increase in market volatility. Further, coincident indicators such as ADS and the INDPRO signal the current business conditions. Their coefficients are negative before and during the crisis, i.e. an improvement in economic conditions generally reduce market volatility. However, the relationship is positive but insignificant after the crisis. The latter observation suggests that an increase in these particular coincident indicators might not have an instantaneous counter-cyclical effect on HF financial volatility.

The in-sample estimation results over the three periods suggest that the relationship between some exogenous variables and RV has changed over the three periods. Hence, looking at the overall out-of-sample (OOS) performance over the whole sample could be misleading. Therefore, we are interested in examining the OOS forecasting performance over the three subsample periods representing the: pre-crisis, crisis, and post-crisis regimes.

2.5.2 Out-of-Sample Analysis

This paper examines the role of a set of economic and financial variables in predicting financial volatility. To do so, we evaluate the OOS performance of the HARX model against the HAR model as the benchmark. We estimate the models in a rolling window of 1000 daily observations. Specifically, we estimate the model using the first 1000 observations to produce the first out-of-sample forecast value corresponding to the 1001^{st} RV observation. Then, we estimate the model from 2^{nd} to the 1001^{st} observation to have the second OOS forecast corresponding to the 1002^{nd} RV observation. We proceed this way to produce 3277 OOS forecasts corresponding to RV's 1001^{st} to 4277^{th} observation.

Many studies are focused on evaluating and comparing volatility models, see, for

example, Poon and Granger (2003) for an extensive review on the matter. Most apply a loss function, where model-based predictions of the conditional variance are compared to estimates of the conditional variance (e.g. realised volatility). Many researchers have considered the Mean Squared Forecast Error (MSFE) as a forecast evaluation method in volatility models (Andersen, Bollerslev, and Lange, 1999; Corsi and Reno, 2009):

$$MSFE = \frac{1}{N} \sum_{t=1}^{N} [RV_{t+1} - \hat{RV}_{t+1}]^2$$

To ease the comparison between the models, we compute the relative loss function (relative MSFE) defined as:

$$relative \ MSFE_{HARX} = \frac{MSFE_{HARX}}{MSFE_{HAR}}$$
(2.7)

where a value below 1 indicates that the HARX outperforms the HAR model.

Table (2.2) displays the relative MSFE of the various HARX models relative to the HAR model. The combined information using SPLS results in the highest gains over the entire sample for all forecasting horizons. We observe a distinct change in variable predictability gains across the forecast horizons and periods per individual variable. Overall, the variables yield higher forecast gains for the long-term RV, where all variables, except TS and FEDFD, improve the forecast of monthly volatility. Some information is more relevant during particular regimes than others.

For example, we note that market sentiments and macroeconomic variables are more relevant for the weekly and monthly forecasts than daily forecasts. These variables produce the most significant improvements in the RV forecasts during the crisis. These variables flag valuable information beyond the lagged RV during unprecedented times. VIX and activity measures, on the other hand, were better predictors before the financial crisis. Overall, as expected, combined information yields the best forecasts during periods of uncertainty.

To investigate the findings by sectors, we report, for each business sector over each regime using each set of X variables, the OOS median of the daily, weekly, and monthly forecast gains in Tables (2.9) to (2.20). Overall, the financial sector has the highest forecast gains from incorporating market and economic conditions information. There are no significant differences among the other sectors. Hence, we look at the overall impact on forecasting volatility for the 100 individual stocks.

In Table (2.3), we report the mean and the skewness of the relative MSFE. Consistent with the SPY, the findings suggest that information beyond the lagged RV can be more relevant for the long-term horizons of individual stocks' volatility during the financial crisis. We find that, on average, combined information (SPLS factors) yields the highest forecast gains for full-sample, particularly during the crisis. The skewness also enables us to know more about the distribution. If it is negatively skewed, it means that while for some stocks, we get even lower relative MSFE than the average, with the majority of the stocks, the relative MSFE is higher than the mean. We mainly observe negative skewness during the financial crisis.

		Da	aily			We	ekly			Moi	nthly	
	\mathbf{FS}	\mathbf{Pre}	Crisis	\mathbf{Post}	FS	\mathbf{Pre}	Crisis	\mathbf{Post}	FS	\mathbf{Pre}	Crisis	\mathbf{Post}
ADS	1.009	0.987	1.011	0.999	0.940	0.971	0.933	0.996	0.734	0.990	0.717	0.908
BCI	1.023	0.980	1.026	1.003	0.970	0.947	0.966	1.006	0.635	0.890	0.593	1.108
CCI	1.017	1.001	1.020	0.998	0.984	1.011	0.984	0.987	0.711	1.043	0.685	0.980
CLI	1.021	0.990	1.024	1.001	0.970	0.976	0.966	1.002	0.653	0.975	0.621	0.995
\mathbf{EPU}	1.023	0.950	1.026	1.001	1.006	0.983	1.007	0.993	0.950	1.009	0.940	1.060
FEDFD	1.002	1.004	1.002	1.005	1.002	1.048	1.003	0.999	1.005	1.097	1.008	0.958
HOUST	1.006	1.182	1.006	1.002	1.003	1.125	1.003	0.998	0.983	1.075	0.986	0.929
INDPRO	1.008	1.018	1.007	1.012	0.966	1.093	0.955	1.039	0.810	1.232	0.775	1.170
PPI	1.003	0.945	1.000	1.026	0.945	0.935	0.922	1.114	0.728	0.931	0.681	1.268
\mathbf{TRD}	1.024	0.989	1.052	0.815	0.986	0.646	0.998	0.913	0.925	0.379	0.929	0.977
\mathbf{TS}	1.007	1.129	1.008	1.000	1.008	1.233	1.007	1.001	1.009	1.165	1.005	1.034
\mathbf{UE}	1.013	1.013	1.015	1.002	0.989	1.036	0.984	1.022	0.779	1.051	0.737	1.245
VIX	0.859	0.757	0.859	0.861	0.904	0.461	0.908	0.892	0.987	0.471	0.998	0.948
VOL	1.024	1.138	1.043	0.874	1.007	0.893	1.019	0.924	0.972	0.648	0.977	0.961
SPLS	0.842	1.037	0.844	0.826	<u>0.839</u>	0.901	0.829	0.912	0.652	0.578	0.630	0.932

Table 2.2: Relative MSFE of HAR-X in Forecasting Market Volatility

This table presents the MSE values of each HAR-X model relative to the MSFE of the HAR model in forecasting the RV of SPY. The results are listed over the Full-sample period and the sub-samples defined as pre-crisis, crisis, and post-crisis period. A value of less than one in bold indicates forecast gains over the HAR model. The lowest value is underlined, showing the best model in the relevant period.

 $\frac{98}{100}$

		Da	aily			We	ekly			Mor	nthly	
	FS	\mathbf{Pre}	\mathbf{Crisis}	\mathbf{Post}	FS	Pre	\mathbf{Crisis}	\mathbf{Post}	FS	Pre	\mathbf{Crisis}	\mathbf{Post}
ADS	0.991	0.994	0.991	0.998	0.930	0.989	0.920	<u>0.980</u>	0.762	0.988	0.734	0.823
ADS	-3.214	-6.739	-2.562	0.466	-3.667	-0.579	-3.963	-2.993	-0.524	1.804	-1.018	-1.250
BCI	1.002	1.003	1.002	1.043	0.957	0.999	0.946	1.099	0.718	0.973	0.657	1.291
Der	-3.271	6.362	-2.906	5.855	-4.164	2.224	-4.028	4.536	-0.695	0.568	-0.125	2.254
CCI	1.002	1.015	1.003	1.029	0.972	1.042	0.964	1.050	0.765	1.074	0.712	1.157
001	-3.621	3.213	-3.115	6.023	-4.146	1.663	-4.098	4.796	-0.777	0.373	-0.418	2.635
CLI	1.000	1.005	1.000	1.018	0.953	1.014	0.943	1.041	0.711	1.031	0.655	1.075
011	-3.152	2.696	-2.812	6.341	-4.471	2.935	-4.244	3.679	-0.811	3.869	-0.576	2.108
\mathbf{EPU}	1.001	0.979	1.001	1.044	0.986	0.964	0.985	1.049	0.953	0.995	0.940	1.138
EIU	-0.430	-2.552	-0.180	4.856	-2.676	-2.165	-2.664	3.212	-1.195	-0.339	-1.073	2.363
FEDFD	1.005	1.038	1.003	<u>0.996</u>	1.011	1.128	1.004	0.971	1.027	1.324	1.013	0.874
FEDFD	5.616	5.442	1.286	-6.047	5.835	2.620	0.625	-3.489	5.153	1.905	-0.056	-1.014
HOUST	1.003	1.048	1.001	0.997	1.001	1.080	0.995	0.976	0.985	1.116	0.976	0.887
110051	3.360	5.149	-1.710	-4.517	4.856	3.720	-2.031	-3.474	3.755	3.021	-2.265	-1.764
INDPRO	0.998	1.015	0.997	1.042	0.969	1.045	0.956	1.116	0.864	1.132	0.813	1.278
mol no	-4.043	6.178	-3.612	5.160	-3.372	2.311	-4.060	3.553	0.095	2.252	-0.199	1.827
PPI	0.997	1.010	0.993	1.070	0.956	1.009	0.933	1.217	0.789	1.010	0.725	1.396
FFI	2.967	9.705	0.139	4.576	1.248	8.331	-0.195	4.198	-0.004	2.920	0.604	2.710
TRD	0.987	1.175	0.996	1.030	0.969	1.019	0.967	1.050	0.933	0.887	0.928	1.059
IKD	-3.531	4.615	-3.451	2.894	-3.993	2.855	-4.589	4.389	-3.001	1.752	-3.376	4.233
\mathbf{TS}	1.005	1.062	1.003	1.005	1.011	1.155	1.003	1.042	1.033	1.243	1.016	1.162
15	5.272	8.101	-3.524	4.086	4.824	3.497	-2.288	6.206	2.118	1.068	-0.624	3.926
UE	1.004	1.009	1.004	1.054	0.983	1.028	0.975	1.153	0.834	1.068	0.770	1.535
0E	-3.076	1.488	-2.626	5.195	-3.893	1.377	-3.266	4.687	-0.330	2.003	-0.117	2.398
VIX	0.935	0.940	0.929	1.076	<u>0.892</u>	0.802	0.887	1.095	0.913	0.748	0.926	1.051
VIX	-1.980	-2.685	-1.552	5.009	-2.126	-0.177	-1.773	5.633	-1.817	0.440	-1.980	7.019
VOL	1.006	1.174	1.008	1.036	1.009	1.073	1.002	1.023	1.002	1.008	0.991	1.023
VOL	-2.201	6.444	-2.148	5.844	-1.521	5.653	-3.512	2.914	1.395	3.080	-5.630	6.441
~~~~	<u>0.933</u>	1.062	0.923	1.077	0.898	0.830	0.888	1.087	0.674	0.824	0.629	1.071
SPLS	-1.666	4.943	-1.152	5.091	-1.138	-0.185	-0.934	5.612	-0.345	0.907	-0.399	2.446

Table 2.3: Mean and Skewness of Relative MSE of HAR-X in Forecasting Individual Stocks Volatility

This table presents the mean and skewness of MSE values of each HAR-X model relative to the MSFE of the HAR model in forecasting the RV of the 100 individual stocks. The results are listed over the Full-sample period, and the sub-samples are defined as pre-crisis, crisis, and post-crisis periods. A mean of less than one in bold indicates forecast gains over the HAR model. The lowest value is underlined, showing the best model in the relevant period.

We note that macroeconomic variables, particularly the unemployment rate, became more valuable during the crisis. Similarly, on average, economic conditions indicators are more valuable during unprecedented times, except for a few individual stocks. In the case of SPY, activity measures are more valuable during the financial crisis than pre-and post-crisis regimes. Hence, we only find supporting evidence for the volatility-volume and the trades-volatility lead-lag relationship among the individual stocks during the crisis period. The finding suggests that, on average, the stock activity measures do not necessarily add valuable information beyond those included in the lagged volatility for all individual stocks.

#### 2.5.3 Forecasting Significance

To test for forecasting significance, we consider the Clark and West (2007) (CW) statistics on MSFE which has been widely adopted in the literature for comparing nested models (see, for example, Audrino and Hu (2016) and Audrino, Sigrist, and Ballinari (2019) for forecasting RV). The test adds an adjustment term to the OOS difference in MSFE that accounts for parameter estimation noise. The null hypothesis involves the population difference in MSFE between the two nested models for a given forecasting horizon. The test statistic is defined as:

$$\hat{f}_{t+1} = (RV_{t+1} - \hat{RV}_{t+1}^{HAR})^2 - [(RV_{t+1} - \hat{RV}_{t+1}^{HARX})^2 - (\hat{RV}_{t+1}^{HAR} - \hat{RV}_{t+1}^{HARX})^2] \quad (2.8)$$

Using the resulting  $\hat{f}_{t+1}$  (2.8), we estimate the regression  $\hat{f}_{t+1} = \mu + \epsilon_t$ , then test the null hypothesis:  $\mu \leq 0$  which implies that the HARX model is not better than the HAR model. We reject the null if the t-statistic is greater than +1.282 for 10% significance or +1.645 for 5% significance. The results are reported in table (2.4) for all HARX models across the three forecasting horizons.

According to the CW test, lagging and coincident indicators such as UE and INDPRO show significant gains during the crisis. In contrast, leading indicators such as HOUST show significant gains after the crisis. HOUST's long-term forecasting gains are also significant during the crisis. Further, most market sentiments are significant for longterm forecasting horizons during and after the crisis. However, VIX is the only variable that yields significant gains during the crisis period for the daily forecasts. Volume and number of trades variables are found to predict the market RV before and after the crisis. Lastly, SPLS is the only one that significantly improves the prediction of the market volatility consistently in all three regimes and three forecasting horizons.

The previously mentioned forecast evaluation methods examine the global forecasting performance over a chosen interval. However, our entire sample period includes unprecedented times, which caused instability in the economy. While dividing the sample into three sub-samples relative to the financial crisis period, namely the precrisis, crisis, and post-crisis period, mitigates part of the regime-change problem, it still suffers from other challenges. For example, it is nearly impossible to depict the exact day of the beginning and end of a regime.

		$\mathbf{D}_{i}$	aily			We	ekly			Mont	hly	
	FS	$\mathbf{Pre}$	Ċrisis	Post	FS	$\mathbf{Pre}$	Crisis	Post	$\mathbf{FS}$	$\mathbf{Pre}$	Ċrisis	Post
ADS	0.016	0.001*	0.032	0.013**	0.218*	0.002	$0.657^{*}$	0.034***	0.994*	0.002	3.015*	0.132***
BCI	-0.053	$0.001^{***}$	-0.175	0.001	0.100**	$0.004^{***}$	$0.308^{*}$	0.009	$1.016^{*}$	$0.019^{***}$	$3.236^{*}$	0.020
CCI	-0.040	0.000	-0.140	0.007	0.059**	0.000	0.160	0.021	$0.765^{*}$	0.001	$2.368^{*}$	$0.069^{**}$
CLI	-0.049	$0.001^{***}$	-0.162	0.002	0.091*	$0.002^{***}$	$0.283^{*}$	0.008	$0.948^{*}$	$0.008^{***}$	$3.007^{*}$	$0.034^{*}$
$\mathbf{EPU}$	-0.050	$0.004^{***}$	-0.177	0.007	0.017	$0.001^{***}$	0.031	0.015	0.214	0.000	0.656	0.024
FEDFD	0.004	0.002	0.003	0.005	0.008	0.003	-0.002	$0.017^{**}$	0.062	0.006	0.128	$0.047^{*}$
HOUST	-0.013	-0.004	-0.053	$0.010^{***}$	0.008	-0.003	-0.008	$0.025^{***}$	0.067***	-0.003	$0.087^{*}$	$0.089^{**}$
INDPRO	-0.001	0.000	0.004	-0.006	0.121*	0.000	$0.400^{*}$	-0.006	$0.549^{*}$	-0.006	$1.800^{*}$	-0.015
PPI	0.012	$0.003^{***}$	0.032	0.002	0.149*	$0.005^{***}$	$0.464^{*}$	0.010	$0.694^{*}$	$0.018^{***}$	$2.122^{*}$	$0.071^{**}$
$\mathbf{TRD}$	0.097	$0.003^{***}$	-0.090	$0.270^{*}$	0.140**	$0.028^{***}$	$0.298^{*}$	$0.091^{*}$	0.340	$0.128^{***}$	0.973	$0.020^{*}$
$\mathbf{TS}$	-0.013	-0.004	-0.053	0.010	0.008	-0.003	-0.008	0.025	0.067	-0.003	0.087	0.089
$\mathbf{UE}$	-0.035	0.000	-0.115	0.002	0.040**	-0.001	$0.124^{**}$	0.004	$0.563^{*}$	-0.002	$1.769^{*}$	0.034
VIX	1.513*	$0.035^{***}$	$4.489^{**}$	$0.269^{**}$	0.563**	$0.043^{***}$	$1.645^{**}$	$0.100^{**}$	0.165	$0.070^{***}$	0.452	0.020***
VOL	0.009	-0.003	-0.231	$0.176^{**}$	0.056	0.010***	0.040	$0.089^{**}$	$0.135^{*}$	$0.069^{***}$	0.326	$0.039^{***}$
SPLS	1.467*	0.021***	$4.420^{**}$	$0.220^{*}$	0.592**	$0.043^{***}$	$1.641^{**}$	$0.166^{*}$	$1.063^{*}$	$0.076^{***}$	$3.278^{*}$	0.067***

Table 2.4: Clark-West MSFE significance test of forecasting RV of SPY using HARX model compared to the HAR model

*** 1% significance, ** 5% significance, and * 10% significance.

Therefore, we examine local superiority using the fluctuation test developed by Giacomini and Rossi (2010) (GR). The methods are designed to examine the relative forecasting performance of two competing models in the presence of instabilities. Thus, we can more accurately detect any reversal in the relative forecasting ability of HAR and HARX models. The general formula of the local relative performance statistic is defined as:

$$F_{t,m}^{OOS} = \hat{\sigma}^{-1} m^{-1} \sum_{j=t-m/2}^{t+m/2-1} \Delta L_j(\hat{\theta}_{j-h,R}^{HAR}, \hat{\theta}_{j-h,R}^{HARX}), \quad t = R + h + m/2, ..., T - m/2 + 1 \quad (2.9)$$

Where m is the length of the of the moving window, set to 250 (i.e. approximately one year of data).  $\Delta L_j$  is the loss function difference between the two models, HAR and HARX. With the CW statistic for nested models, the local relative performance MSFE becomes:

$$F_{t,m}^{OOS} = \hat{\sigma}^{-1} m^{-1} \Sigma_{j=t-m/2}^{t+m/2-1} \hat{f}_{t+\tau}, \quad t = R+h+m/2, ..., T-m/2+1$$
(2.10)

where  $\hat{f}_{t+\tau}$  is defined in equation (2.8) and R = 1000 is the number of observations in the in-sample. For every t, the parameters of the models are re-estimated over the in-sample period in a rolling window scheme: t - h - R + 1, ..., t - h. The null hypothesis for this test is that the two competing models, HAR and HARX, have equal OOS performance at each point in time. Wherever the local relative MSFE,  $F_{t,m}^{OOS}$ , exceeds the critical value we reject the null hypothesis and conclude that the HARX produced better forecasts than the HAR model over the period of length mcentered at t.

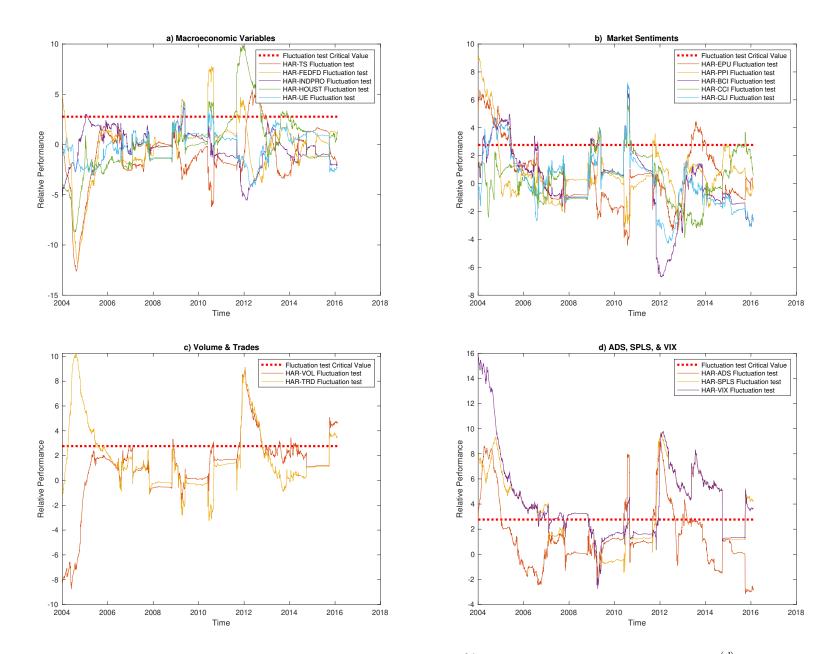


Figure 2.1: GR statistic tests plot for HARX models with 5% confidence level for forecasting  $RV^{(d)}$  of SPY

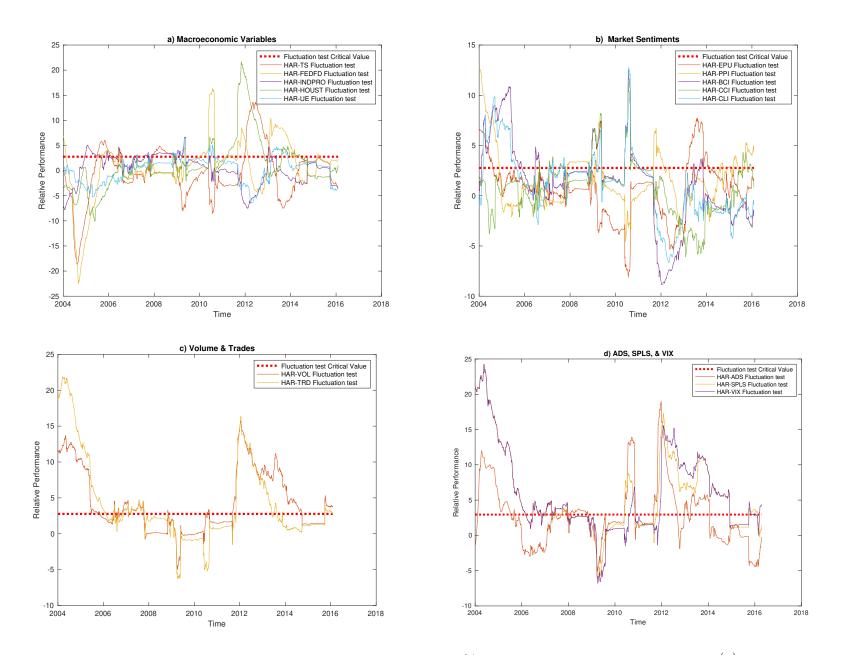


Figure 2.2: GR statistic tests plot for HARX models with 5% confidence level for forecasting  $RV^{(w)}$  of SPY

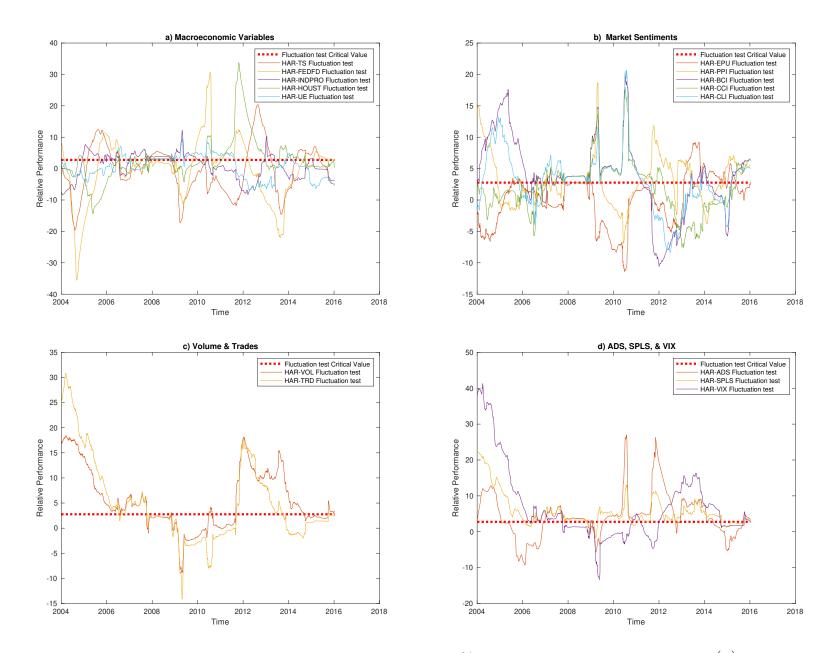


Figure 2.3: GR statistic tests plot for HARX models with 5% confidence level for forecasting  $RV^{(m)}$  of SPY

Figures (2.1), (2.2), and (2.3) show the significance of the forecast gains of HARX, with each explanatory variable (X), for the daily, weekly, and monthly forecasts, respectively. The figures reveal when the HARX models outperform the HAR model. It is evident that the overall level of significance falls during the crisis period and increases afterwards. It falls around 2014 but starts to rise again by the end of 2015. The decline in significance around 2014 can be explained by economic trends not reflected explicitly in any of the HARX models, such as the plunge in oil prices and the dollar's appreciation.

Most macroeconomic variables yield significant long-term forecast gains over the covered years, although it generally falls during the crisis. The window over which the HARX has significant outperformance increases with the forecast horizon, including the crisis period for the monthly forecast. In particular, as leading indicators, both TS and HOUST show predictability for market volatility, especially directly after the crisis. Similarly, INDPRO, a coincident indicator, and UE, a lagging indicator, yield significant forecast gains for market volatility even during the crisis for the weekly and monthly forecast. Another lagging indicator, the PPI, shows significant forecast gains before and towards the end of the crisis period. ADS, a daily indicator of the current business environment, shows significant weekly and monthly forecast improvement mainly after the crisis.

The significant-high gains observed in the weekly and monthly forecasts using macroeconomic factors confirm that the latter can explain long-term variation in the stock market. We have shown that this observation holds not only for calendar months volatility, as used in previous literature, but also at a higher frequency level, i.e. for one week,  $RV_{t+5}^{(w)}$ , or one month,  $RV_{t+22}^{(m)}$ , ahead from any given day, t.

Furthermore, market sentiments yield significant forecast gains for most of the sample period, with lower gains during the crisis. BCI and CLI yield the highest significant gains over the majority of the sample period for the weekly and monthly forecast, while VIX shows higher significance for the daily forecast. Moreover, we find evidence of the SIAH using SPY before and after the crisis period only and more significant for weekly and monthly forecast horizons. The number of trades shows higher significant forecast gains than the volume before the crisis, while the latter yields higher significant gains after the crisis. Lastly, the forecast gains of using combined information (SPLS factors) are significant for most of the sample period, particularly for the long-term forecasting horizon.

### 2.6 Conclusion

This study examines the value-added from utilising information for various financialeconomic indicators observed at different frequencies. We do so by adding "X" variable(s) to the HAR model, hence, HARX. We consider macroeconomic indicators, market sentiments, and activity measures that reflect the economic and market conditions. The findings reveal that economic variables include essential information beyond the lagged and historical average of volatility that significantly improves the RV's forecast. The market and economic conditions are particularly valuable for individual stocks during the crisis. The forecast gains are higher when forecasting long-term volatility than short-term ones.

One challenge of studying the relationship between financial markets and macroeconomy is the frequency of the variables. While the former's data are available at a high-frequency level, the latter's are recorded at a lower frequency. Hence, researchers suggest different methods to handle the frequency mismatch. One approach that can be used to study a LF variable's effect on a HF variable is employing the long-known Chow-Lin extrapolation method (Chow and Lin, 1971). A recent approach is the reverse MIDAS by Foroni, Guérin, and Marcellino (2018). The former extrapolates the LF into the HF series, allowing the monthly variable's value to fluctuate within the month while its coefficient in the HARX model is constant in a given estimation window. However, the latter approach keeps the value of the LF variable constant during the month but allows its coefficient to vary based on the day of the month. There is no significant difference in forecasting gains in the long-term forecasts while favouring the extrapolation method in the daily forecasts as it incorporates daily variations.

The forecast gains of including macroeconomic variables in the HARX model are more pronounced for long-term forecasting (weekly and monthly) than one day ahead. While unemployment and industrial production growth rates yield significant forecasting gains for most of the sample period, the housing starts growth rate, a leading indicator, showed more persistent significance after the financial crisis. Similarly, most market sentiments yield significant forecasting gains, particularly leading indicators. The CLI (VIX) produces significant forecasting gains for most of the OOS period for the weekly and monthly (daily) forecast. Also, the predictive power of volume and number of trades for RV is significant at all forecast horizons but not during a recession.

This study shows that incorporating financial and macroeconomic indicators in forecasting the RV produces significant forecasting gains. We note that combining the information using SPLS components in the HARX framework has the lowest average forecast errors over the sample period and is particularly beneficial for long-term investors during an unprecedented time, such as the financial crisis. Hence, a study of the economic gains from incorporating this information in portfolio risk management would be interesting to examine for future research. Furthermore, speaking of blended information, it is of interest to develop a resilient economic-financial conditions index robust to periods of high uncertainty such as financial crises or pandemics. Ideally, such an index would reflect the overall environment by including new information and adjusting the measurements accordingly.

# Appendix

#### 2.A Tables

	Ticker	Mean	Median	Minimum	Maximum	St. Dev.
Amazon.com Inc	AMZN	8.284	3.344	0.225	229.244	14.846
Best Buy Co. Inc.	BBY	5.760	3.110	0.244	1439.571	23.491
Gap (The)	GPS	2.966	1.452	0.143	145.706	5.015
Interpublic Group	IPG	2.974	1.478	0.060	170.100	6.221
Marriott International	MAR	5.088	2.830	0.146	178.723	7.586
McDonald's	MCD	3.121	1.573	0.156	103.477	4.938
News Corporation NWS	FOX	5.524	2.632	0.235	615.148	13.956
The Home Depot	HD	3.537	1.782	0.154	104.578	5.364
Time Warner Inc.	TWX	2.164	1.090	0.087	161.156	4.333
Walt Disney	DIS	4.099	1.722	0.163	148.875	7.289
Consumer Discretionary Sector	Avg.	4.352	2.101	0.161	329.658	9.304
Avon Products	AVP	5.098	2.159	0.140	342.665	11.720
Brown-Forman Corp.	BFb	1.920	1.152	0.074	240.414	4.861
Coca-Cola	KO	3.087	1.497	0.126	83.955	4.728
Costco	COST	2.679	1.505	0.158	62.689	3.863
Estee Lauder Cos.	EL	1.566	0.856	0.061	100.855	2.745
Kimberly-Clark	KMB	1.561	0.836	0.046	58.808	2.535
PepsiCo Inc.	PEP	1.693	0.828	0.083	132.093	3.638
Procter & Gamble	$\mathbf{PG}$	1.492	0.766	0.101	79.549	2.907
Unilever	UL	1.302	0.646	0.078	25.319	2.054
Wal-Mart	WMT	2.045	0.976	0.090	71.485	3.277
Consumer Staples Sector Avg.		2.244	1.122	0.096	119.783	4.233

Table 2.5: Descriptive Statistics of Stocks

This table reports the descriptive statistics of daily Realised Variance (RV) for each stock over the period 01/2000 to 12/2016.

	Ticker	Mean	Median	Minimum	Maximum	St. Dev.
Baker Huges	BHI	4.312	2.699	0.318	155.091	6.361
Chesapeaks Energy	CHK	5.609	3.584	0.321	218.401	7.742
Chevron Corporation	CVX	13.008	4.537	0.418	2216.263	41.965
Devon Energy Corp	DVN	2.113	1.263	0.112	137.535	4.181
Exxon Mobil	XOM	4.595	2.747	0.274	136.649	7.012
Halliburton Co.	HAL	6.305	3.579	0.229	265.432	10.443
Hess Corporation	HES	4.474	2.570	0.211	271.511	8.675
Occidental Petroleum	OXY	3.743	2.179	0.208	161.023	6.349
Sunoco Inc.	APA	10.216	3.338	0.267	1633.848	42.946
Williams Cos.	WMB	1.987	1.141	0.107	141.130	3.955
Energy Sector Avg.		5.636	2.764	0.246	533.688	13.963
Allstate Corp	ALL	3.188	1.173	0.098	255.375	8.091
American Express	AXP	4.020	1.465	0.088	299.968	9.195
Bank of America	BAC	5.387	1.772	0.085	406.683	16.736
Citigroup Inc.	С	5.086	1.771	0.156	1059.901	23.703
Goldman Sachs Group	GS	6.742	2.110	0.137	975.858	27.240
JPMorgan Chase	JPM	4.186	1.757	0.153	400.346	11.978
Morgan Stanley	MS	4.615	1.770	0.114	252.877	10.848
The Bank of New York Mellon	BK	8.191	2.797	0.236	1732.782	46.295
Travelers Group Inc	TRV	2.968	1.186	0.102	263.929	7.866
Wells Fargo	WFC	4.302	1.330	0.104	226.609	12.139
Financials Sector Avg.		4.869	1.713	0.127	587.433	17.409

Table 2.5: Descriptive Statistics of Stocks (cont.)

Stock	Ticker	Mean	Median	Minimum	Maximum	St. Deviation
Abbott Laboratories	ABT	2.168	1.226	0.119	65.347	3.045
Amgen Inc	AMGN	3.542	1.835	0.203	94.858	5.259
Boston Scientific	BSX	5.010	2.908	0.200	151.610	7.026
Gilead Sciences	GILD	6.308	2.839	0.198	187.286	10.469
Humana Inc.	HUM	6.679	2.609	0.240	157.529	11.366
Johnson & Johnson	JNJ	1.385	0.692	0.076	179.016	3.482
Medtronic Inc	MDT	2.311	1.236	0.109	181.758	4.348
Merck	MRK	2.431	1.359	0.135	223.255	5.227
Pzer	PFE	2.332	1.382	0.150	62.697	3.224
United Health Group	UNH	3.407	1.745	0.129	225.956	6.883
Health Care Sector Avg.		3.557	1.783	0.156	152.931	6.033
Boeing	BA	2.812	1.602	0.086	55.570	3.900
Caterpillar	CAT	3.239	1.873	0.185	105.908	4.889
Cummins Inc.	CMI	4.905	2.566	0.157	199.866	9.070
General Dynamics	$\operatorname{GD}$	2.237	1.281	0.081	63.282	3.259
General Electric	GE	3.020	1.303	0.108	180.389	6.982
Honeywell Int'l Inc.	HON	3.253	1.609	0.104	268.331	6.513
Minnesota Mining & Mfg Co	MMM	4.844	3.030	0.285	150.858	6.413
Southwest Airlines	LUV	1.853	1.008	0.082	91.955	3.216
United Parcel Service	UPS	1.649	0.851	0.081	216.939	4.140
United Technologies	UTX	2.300	1.211	0.104	75.915	3.793
Industrials Sector Avg.		3.011	1.633	0.127	140.901	5.217

Table 2.5: Descriptive Statistics of Stocks (cont.)

Stock	Ticker	Mean	Median	Minimum	Maximum	St. Deviation
Apple Inc.	AAPL	5.292	2.578	0.079	126.172	7.806
Cisco Systems	CSCO	4.275	1.986	0.184	87.833	6.707
Dell Inc.	EBAY	6.511	2.782	0.202	236.419	12.762
EMC Corp.	YAHOO	4.302	2.267	0.122	156.740	7.066
Hewlett-Packard	HPQ	2.026	0.986	0.102	71.293	3.527
IBM	IBM	4.075	2.038	0.154	89.885	5.754
Intel Corp.	INTC	2.679	1.416	0.083	62.386	3.854
Microsoft	MSFT	5.021	2.174	0.129	123.804	8.381
Oracle Corp.	ORCL	6.548	2.864	0.299	276.588	13.732
Xerox Corp.	XRX	7.419	3.100	0.240	220.822	12.760
Information Technolog	gy Sector	4.815	2.219	0.159	145.194	8.235
Avg.						
AK Steel Holding Corp	AKS	16.137	10.585	0.872	559.612	21.802
Alcoa	ARNC	5.288	3.070	0.339	291.089	9.601
Dow Chemical	DOW	2.761	1.491	0.100	83.487	4.076
DuPont	DD	3.976	2.039	0.146	216.937	7.353
Freeport-McMoran	FCX	8.034	4.327	0.317	188.580	12.279
International Paper	IP	4.290	2.103	0.162	171.959	7.500
Newmont Mining	NEM	5.098	3.435	0.470	109.879	6.395
Nucor Corp.	NUE	4.910	2.754	0.334	266.824	10.582
United States Steel Corp.	Х	3.699	2.043	0.252	131.603	5.916
Weyerhauser Co	WY	8.696	5.557	0.736	344.652	12.579
Materials Sector Avg.		6.289	3.740	0.373	236.462	9.808

Table 2.5: Descriptive Statistics of Stocks (cont.)

Stock	Ticker	Mean	Median	Minimum	Maximum	St. Deviation
Ameren Corp	AEE	1.828	1.061	0.109	113.488	3.453
American Electric Power	AEP	2.444	1.208	0.151	207.378	6.592
Constellation Energy Group	CEG	4.497	1.825	0.168	772.838	20.605
Duke Energy	DUK	2.433	1.182	0.051	189.935	6.070
Entergy Corp.	ETR	2.230	1.177	0.085	118.560	4.028
Exelon Corp.	EXC	2.635	1.429	0.158	130.875	4.791
PG&E Corp.	PCG	3.402	1.668	0.160	411.055	8.590
Progress Energy, Inc.	CNP	4.162	1.265	0.149	1532.118	27.721
Public Serv. Enterprise Inc.	PEG	2.461	1.376	0.148	122.428	4.827
The Southern Company	SO	1.744	0.937	0.092	97.041	2.773
Utilities Sector Avg.		2.784	1.313	0.127	369.572	8.945
American Tower Corp A	AMT	9.628	1.957	0.165	1048.656	32.874
AT&T	Т	2.311	1.162	0.100	59.568	3.228
BT Group plc (ADR)	BT	2.052	1.063	0.082	65.965	3.608
CenturyTel Inc	CTL	2.720	1.346	0.152	279.942	6.856
Frontier Communications	FTR	5.426	2.933	0.232	1738.725	27.431
Qwest Communication Int	CHL	20.499	10.917	0.242	1159.384	35.569
Sprint Nextel Corp	LVLT	2.655	1.184	0.108	141.846	4.767
Telefonica S.A. (ADR)	TEF	1.820	1.021	0.117	162.204	3.601
Verizon Communications	VZ	1.986	0.926	0.110	70.936	3.063
Vodafone Group Plc (ADR)	VOD	2.336	1.162	0.122	102.221	3.860
Telecom Sector Avg.		5.143	2.367	0.143	482.945	12.486
SPDR ETF	SPY	1.037	0.485	0.013	59.863	3.453

Table 2.5: Descriptive Statistics of Stocks (cont.)

Variable	Functional form	Mean	Median	Minimum	Maximum	Std. dev	Skewness	Kurtosis
ADS	Level	-0.331	-0.192	-4.234	0.958	0.759	-2.418	10.907
BCI	Level	99.807	99.852	96.022	101.997	1.069	-0.919	4.893
CCI	Level	99.657	99.809	96.707	102.702	1.376	-0.221	2.671
CLI	Level	99.751	100.005	94.614	101.892	1.409	-1.351	5.339
FEDFD	Log of Squared	-0.796	-0.020	-6.438	3.900	3.186	0.013	1.394
HOUST	% Growth	0.199	0.109	-18.830	24.922	8.294	0.193	2.979
INDPRO	% Growth	0.046	0.098	-4.337	1.459	0.676	-2.029	12.843
PPI	Level	171.332	174.850	128.100	208.300	26.372	-0.251	1.644
TS	Log of Squared	0.791	1.493	-9.210	2.696	1.998	-1.945	6.945
UE	% Growth	0.116	0.000	-7.463	8.000	2.721	0.499	3.325
VIX	Squared	491.688	335.989	97.812	6538.340	539.523	4.562	33.363

Table 2.6: Descriptive Statistics

This table reports the descriptive statistics of the variables in the specified functional form over the period 01/2000 to 12/2016.

	$\hat{eta}_0^{(d)}$		$\hat{eta}_{d}^{(d)}$		$\hat{eta}^{(d)}_w$		$\hat{eta}_m^{(d)}$		$\hat{eta}_x^{(d)}$		Adj. $R^2$
HAR	0.131	***	0.236	**	0.241	**	0.393	***			36.598%
	(0.042)		(0.117)		(0.119)		(0.115)				
HAR-ue	0.148	***	0.236	**	0.234	**	0.379	***	0.055		36.753%
	(0.053)		(0.116)		(0.117)		(0.106)		(0.064)		
HAR-ind	0.171	**	0.235	**	0.240	**	0.375	***	-3.226		36.673%
	(0.068)		(0.117)		(0.118)		(0.109)		(3.157)		
HAR-hst	0.132	***	0.236	**	0.234	**	0.401	***	-0.015		36.607%
	(0.042)		(0.116)		(0.119)		(0.115)		(0.011)		
HAR-ts	0.131	***	0.236	**	0.241	**	0.393	***	0.002		36.563%
	(0.042)		(0.116)		(0.119)		(0.117)		(0.012)		
HAR-cci	-2.445		0.236	*	0.240	**	0.391	***	0.026		36.578%
	(4.633)		(0.121)		(0.119)		(0.097)		(0.046)		
HAR-cli	6.508	*	0.235	*	0.242	**	0.360	***	-0.064		36.714%
	(3.936)		(0.121)		(0.118)		(0.099)		(0.039)		
HAR-bci	8.055		0.234	*	0.241	**	0.352	***	-0.079		36.766%
	(5.055)		(0.122)		(0.119)		(0.089)		(0.050)		
HAR-ppi	1.524	***	0.234	**	0.239	**	0.339	***	-0.009	***	36.864%
	(0.382)		(0.118)		(0.115)		(0.102)		(0.002)		
HAR-ads	0.124	***	0.236	**	0.241	**	0.379	***	-0.076		36.608%
	(0.038)		(0.117)		(0.118)		(0.116)		(0.106)		
HAR-fd	0.092	*	0.236	**	0.240	**	0.388	***	0.023		36.596%
	(0.051)		(0.117)		(0.119)		(0.113)		(0.026)		
HAR-vol	0.185	*	0.238	**	0.242	**	0.380	***	-0.130		36.591%
	(0.106)		(0.112)		(0.118)		(0.131)		(0.169)		
HAR-trd	0.251	**	0.237	**	0.242	**	0.354	***	-0.212	**	36.727%
	(0.099)		(0.111)		(0.116)		(0.127)		(0.097)		
HAR-vix	-0.120	**	0.203	**	0.172	*	0.143		0.503	***	39.368%
	(0.060)		(0.103)		(0.089)		(0.096)		(0.112)		

Table 2.7: a) In-sample estimation of HARX (eq. (2.2)) using SPY data for  $y = RV^{(d)}$ , Pre-Crisis Period

	$\hat{oldsymbol{eta}}_0^{(d)}$		$\hat{eta}_{d}^{(d)}$	$\hat{eta}^{(d)}_w$		$\hat{eta}_m^{(d)}$	$\hat{eta}^{(d)}_x$		Adj. $R^2$
HAR	0.165	*	0.221	0.493	**	0.194			54.048%
	(0.097)		(0.138)	(0.199)		(0.135)			
HAR-ue	0.121		0.218	0.499	**	0.153	0.140		54.173%
	(0.078)		(0.138)	(0.203)		(0.151)	(0.099)		
HAR-ind	0.184	*	0.217	0.495	**	0.157	-5.142	*	54.224%
	(0.101)		(0.137)	(0.203)		(0.146)	(3.078)		
HAR-hst	0.080		0.219	0.502	**	0.161	-0.089		54.072%
	(0.082)		(0.138)	(0.203)		(0.159)	(0.114)		
HAR-ts	0.157	*	0.221	0.493	**	0.193	0.012		54.006%
	(0.091)		(0.138)	(0.199)		(0.137)	(0.019)		
HAR-cci	11.414		0.219	0.498	***	0.173	-0.114		54.079%
	(7.998)		(0.155)	(0.188)		(0.153)	(0.080)		
HAR-cli	5.195		0.220	0.499	***	0.174	-0.050		54.053%
	(4.711)		(0.155)	(0.188)		(0.156)	(0.046)		
HAR-bci	20.314	*	0.216	0.506	***	0.127	-0.202	*	54.246%
	(11.836)		(0.155)	(0.189)		(0.164)	(0.118)		
HAR-ppi	-1.279		0.221	0.489	***	0.201	0.008		54.039%
	(1.146)		(0.155)	(0.188)		(0.151)	(0.006)		
HAR-ads	0.086		0.218	0.498	**	0.149	-0.177		54.212%
	(0.074)		(0.138)	(0.203)		(0.151)	(0.114)		
HAR-fd	0.184	*	0.220	0.486	**	0.208	0.036	*	54.082%
	(0.097)		(0.138)	(0.197)		(0.136)	(0.022)		
HAR-vol	-0.813	***	0.122	0.476	**	0.160	0.639	***	55.274%
	(0.265)		(0.146)	(0.190)		(0.143)	(0.203)		
HAR-trd	-0.376	**	0.161	0.493	***	0.131	0.168	**	54.728%
	(0.188)		(0.147)	(0.190)		(0.147)	(0.069)		
HAR-vix	-0.589	***	0.093	0.443	**	-0.357	0.907	***	58.407%
	(0.227)		(0.147)	(0.210)		(0.228)	(0.257)		

Table 2.7: b) In-sample estimation of HARX (eq. (2.2)) using SPY data for  $y = RV^{(d)}$ , Crisis Period

	$\hat{eta}_0^{(d)}$		$\hat{eta}_{d}^{(d)}$		$\hat{eta}^{(d)}_w$	$\hat{eta}_m^{(d)}$		$\hat{eta}_x^{(d)}$		Adj. $R^2$
HAR	0.103	***	0.422	**	0.169	0.204	**			35.336%
	(0.036)		(0.197)		(0.153)	(0.092)				
HAR-ue	0.127	***	0.422	**	0.168	0.202	**	0.058		35.329%
	(0.030)		(0.196)		(0.152)	(0.092)		(0.059)		
HAR-ind	0.087	*	0.422	**	0.168	0.202	**	1.685		35.315%
	(0.046)		(0.196)		(0.152)	(0.092)		(1.863)		
HAR-hst	0.117	***	0.422	**	0.167	0.202	**	-0.030		35.338%
	(0.031)		(0.195)		(0.152)	(0.092)		(0.029)		
HAR-ts	0.083		0.422	**	0.168	0.205	**	0.014		35.298%
	(0.055)		(0.197)		(0.153)	(0.091)		(0.022)		
HAR-cci	5.128	*	0.420	**	0.168	0.175	*	-0.050	*	35.536%
	(2.865)		(0.193)		(0.153)	(0.096)		(0.029)		
HAR-cli	8.241	**	0.420	**	0.170	0.177	*	-0.081	**	35.465%
	(3.290)		(0.197)		(0.156)	(0.093)		(0.033)		
HAR-bci	4.557		0.422	**	0.169	0.195	**	-0.045		35.336%
	(3.819)	***	(0.198)		(0.157)	(0.091)		(0.038)		
HAR-ppi	0.110		0.422	**	0.169	0.204	**	0.000		35.293%
	(0.499)		(0.198)		(0.157)	(0.090)		(0.002)		
HAR-ads	0.105	***	0.422	**	0.168	0.205	**	0.015		35.295%
	(0.033)		(0.196)		(0.153)	(0.092)		(0.042)		
HAR-fd	0.062		0.422	**	0.169	0.201	**	-0.011		35.307%
	(0.056)		(0.197)		(0.154)	(0.092)		(0.010)		
HAR-vol	-0.382	**	0.241		0.082	0.072		0.619	**	40.999%
	(0.188)		(0.157)		(0.087)	(0.106)		(0.250)		
HAR-trd	-0.458	*	0.181		0.050	0.097		0.223	*	40.556%
	(0.277)		(0.247)		(0.084)	(0.108)		(0.119)		
HAR-vix	-0.138	**	0.284		0.080	-0.333		0.690	***	41.753%
	(0.065)		(0.184)		(0.098)	(0.204)		(0.218)		

Table 2.7: c) In-sample estimation of HARX (eq. (2.2)) using SPY data for  $y = RV^{(d)}$ , Post-Crisis Period

	$\hat{eta}_0^{(w)}$		$\hat{eta}_{d}^{(w)}$		$\hat{eta}^{(w)}_w$		$\hat{eta}_m^{(w)}$		$\hat{eta}_x^{(w)}$		Adj. $R^2$
HAR	0.189	***	0.096	*	0.324	***	0.392	***			52.638%
	(0.052)		(0.050)		(0.082)		(0.111)				
HAR-ue	0.207	***	0.095	*	0.317	***	0.378	***	0.060		53.001%
	(0.065)		(0.050)		(0.086)		(0.104)		(0.074)		
HAR-ind	0.247	***	0.094	*	0.323	***	0.366	***	-4.726		53.027%
	(0.076)		(0.050)		(0.081)		(0.108)		(3.305)		
HAR-hst	0.190	***	0.095	*	0.315	***	0.403	***	-0.021		52.766%
	(0.052)		(0.050)		(0.078)		(0.104)		(0.015)		
HAR-ts	0.189	***	0.096	*	0.324	***	0.392	***	0.000		52.611%
	(0.052)		(0.050)		(0.082)		(0.111)		(0.017)		
HAR-cci	-3.623		0.096	*	0.324	***	0.389	***	0.038		52.676%
	(6.415)		(0.050)		(0.081)		(0.106)		(0.064)		
HAR-cli	9.287		0.094	*	0.326	***	0.346	***	-0.091		53.152%
	(5.749)		(0.049)		(0.079)		(0.121)		(0.057)		
HAR-bci	11.388	**	0.094	*	0.325	***	0.334	***	-0.111	**	53.323%
	(5.214)		(0.049)		(0.080)		(0.104)		(0.052)		
HAR-ppi	2.188	***	0.093	*	0.322	***	0.315	***	-0.013	***	53.694%
	(0.629)		(0.048)		(0.078)		(0.108)		(0.004)		
HAR-ads	0.180	***	0.095	*	0.325	***	0.375	***	-0.100		52.750%
	(0.048)		(0.050)		(0.081)		(0.118)		(0.125)		
HAR-fd	0.133	**	0.095	*	0.324	***	0.386	***	0.033		52.735%
	(0.060)		(0.051)		(0.083)		(0.107)		(0.032)		
HAR-vol	0.345	***	0.102	**	0.329	***	0.355	***	-0.378	**	53.038%
	(0.105)		(0.052)		(0.082)		(0.104)		(0.175)		
HAR-trd	0.390	***	0.097	*	0.326	***	0.328	***	-0.036	***	53.429%
	(0.104)	***	(0.051)		(0.080)		(0.100)		(0.011)		
HAR-vix	-0.040		0.066	*	0.262	***	0.166		0.456	***	56.641%
	(0.065)		(0.035)		(0.081)		(0.123)		(0.135)		

Table 2.8: a) In-sample estimation of HARX (eq. (2.3)) using SPY data for  $y = RV^{(w)}$ , Pre-Crisis Period

	$\hat{oldsymbol{eta}}_0^{(w)}$		$\hat{eta}_{d}^{(w)}$		$\hat{eta}^{(w)}_w$		$\hat{eta}_m^{(w)}$		$\hat{eta}^{(w)}_x$		Adj. $R^2$
HAR	0.263	***	0.212	***	0.344	**	0.298	**			64.863%
	(0.095)		(0.048)		(0.147)		(0.118)				
HAR-ue	0.192	**	0.208	***	0.353	**	0.230	*	0.229	*	65.484%
	(0.082)		(0.047)		(0.157)		(0.136)		(0.125)		
HAR-ind	0.296	***	0.206	***	0.348	**	0.233	*	-9.017	**	65.798%
	(0.084)		(0.047)		(0.155)		(0.134)		(3.809)		
HAR-hst	0.134		0.210	***	0.357	**	0.248	**	-0.135		65.057%
	(0.129)		(0.047)		(0.152)		(0.123)		(0.127)		
HAR-ts	0.253	***	0.212	***	0.344	**	0.296	**	0.015		64.837%
	(0.087)		(0.048)		(0.147)		(0.118)		(0.028)		
HAR-cci	18.453	*	0.210	***	0.352	**	0.263	**	-0.184	*	65.118%
	(10.584)		(0.043)		(0.143)		(0.115)		(0.107)		
HAR-cli	7.679		0.210	***	0.353	**	0.267	**	-0.074		64.985%
	(5.381)		(0.043)		(0.143)		(0.116)		(0.053)		
HAR-bci	31.012	*	0.205	***	0.364	**	0.196		-0.308	*	65.632%
	(16.481)		(0.042)		(0.147)		(0.127)		(0.165)		
HAR-ppi	-2.596		0.212	***	0.335	**	0.311	***	0.016		65.034%
	(2.090)		(0.044)		(0.136)		(0.106)		(0.012)		
HAR-ads	0.132		0.207	***	0.352	**	0.223		-0.292	**	65.642%
	(0.094)		(0.047)		(0.157)		(0.137)		(0.147)		
HAR-fd	0.283	***	0.211	***	0.336	**	0.314	***	0.041		64.980%
	(0.103)		(0.048)		(0.143)		(0.113)		(0.031)		
HAR-vol	-0.485	*	0.136	***	0.331	**	0.271	**	0.489	**	65.888%
	(0.281)	***	(0.044)		(0.144)		(0.120)		(0.203)		
HAR-trd	-0.100		0.172	***	0.344	**	0.255	**	0.011	*	65.295%
	(0.165)		(0.043)		(0.148)		(0.125)	***	(0.006)		
HAR-vix	-0.138		0.144	***	0.317	*	0.004		0.484	***	66.615%
	(0.115)		(0.042)		(0.167)		(0.152)		(0.128)		

Table 2.8: b) In-sample estimation of HARX (eq. (2.3)) using SPY data for  $y = RV^{(w)}$ , Crisis Period

	$\hat{eta}_0^{(w)}$		$\hat{eta}_{d}^{(w)}$		$\hat{eta}^{(w)}_w$		$\hat{eta}_m^{(w)}$		$\hat{eta}_{x}^{(w)}$		Adj. $R^2$
HAR	0.167	***	0.232	**	0.133	*	0.302	***			32.982%
	(0.042)		(0.112)		(0.069)		(0.076)				
HAR-ue	0.206	***	0.231	**	0.133	*	0.299	***	0.093		33.101%
	(0.050)		(0.111)		(0.070)		(0.076)		(0.094)		
HAR-ind	0.143	***	0.232	**	0.133	*	0.299	***	2.521		33.024%
	(0.052)		(0.111)		(0.070)		(0.076)		(3.212)		
HAR-hst	0.190	***	0.231	**	0.131	*	0.298	***	-0.049		33.149%
	(0.046)		(0.110)		(0.070)		(0.076)		(0.048)		
HAR-ts	0.091		0.232	**	0.132	**	0.305	***	0.051		33.060%
	(0.075)		(0.109)		(0.053)		(0.050)		(0.048)		
HAR-cci	8.105		0.228	**	0.132	**	0.256	***	-0.080		33.984%
	(5.838)		(0.104)		(0.054)		(0.065)		(0.058)		
HAR-cli	12.929	*	0.229	**	0.135	***	0.259	***	-0.127		33.667%
	(7.834)		(0.107)		(0.051)		(0.059)		(0.078)		
HAR-bci	7.409		0.231	**	0.135	***	0.287	***	-0.072		33.137%
	(5.530)		(0.109)		(0.051)		(0.052)		(0.055)		
HAR-ppi	0.119		0.232	**	0.133	**	0.302	***	0.000		32.938%
	(0.863)		(0.109)		(0.052)		(0.050)		(0.004)		
HAR-ads	0.170	***	0.232	**	0.133	*	0.303	***	0.022		32.945%
	(0.042)		(0.112)		(0.070)		(0.076)		(0.064)		
HAR-fd	0.124		0.232	**	0.134	***	0.299	***	-0.011		32.963%
	(0.086)		(0.109)		(0.051)		(0.051)		(0.019)		
HAR-vol	-0.268		0.070		0.056		0.184	**	0.554	**	40.813%
	(0.179)		(0.051)		(0.058)		(0.078)		(0.252)		
HAR-trd	-0.275		0.042		0.039		0.218	***	0.018	**	38.587%
	(0.168)	***	(0.095)		(0.065)		(0.073)		(0.008)		
HAR-vix	-0.009		0.131		0.069	*	-0.088		0.501	***	38.839%
	(0.037)		(0.099)		(0.040)		(0.114)		(0.129)		

Table 2.8: c) In-sample estimation of HARX (eq. (2.3)) using SPY data for  $y = RV^{(w)}$ , Post-Crisis Period

	$\hat{eta}_0^{(m)}$		$\hat{eta}_d^{(m)}$		$\hat{eta}_w^{(m)}$		$\hat{eta}_m^{(m)}$		$\hat{eta}_x^{(m)}$		Adj. $R^2$
HAR	0.313	***	0.070	**	0.253	***	0.366	***			50.834%
	(0.083)		(0.034)		(0.064)		(0.093)				
HAR-ue	0.320	***	0.070	**	0.251	***	0.361	***	0.001		50.875%
	(0.090)		(0.034)		(0.067)		(0.092)		(0.001)		
HAR-ind	0.391	***	0.069	**	0.251	***	0.331	***	-0.063		51.832%
	(0.114)		(0.033)		(0.063)		(0.100)		(0.049)		
HAR-hst	0.316	***	0.070	**	0.238	***	0.385	***	-0.001		51.429%
	(0.082)		(0.033)		(0.060)		(0.082)		(0.001)		
HAR-ts	0.314	***	0.070	**	0.253	***	0.366	***	-0.004		50.815%
	(0.088)		(0.034)		(0.062)		(0.093)		(0.025)		
HAR-cci	-7.644		0.070	**	0.252	***	0.360	***	0.079		51.191%
	(8.838)		(0.034)		(0.058)		(0.083)		(0.088)		
HAR-cli	14.266		0.068	**	0.256	***	0.295	**	-0.139		52.533%
	(9.723)		(0.033)		(0.057)		(0.129)		(0.097)		
HAR-bci	17.631	**	0.067	**	0.254	***	0.275	***	-0.172	**	53.125%
	(6.883)		(0.034)		(0.058)		(0.095)		(0.068)		
HAR-ppi	3.655	***	0.066	**	0.249	***	0.238	**	-0.022	***	54.870%
	(1.150)		(0.031)		(0.054)		(0.096)		(0.007)		
HAR-ads	0.302	***	0.070	**	0.254	***	0.345	***	-0.122		51.085%
	(0.077)		(0.034)		(0.063)		(0.104)		(0.169)		
HAR-fd	0.217	**	0.070	**	0.253	***	0.354	***	0.058		51.329%
	(0.100)		(0.034)		(0.063)		(0.085)		(0.047)		
HAR-vol	0.666	***	0.084	**	0.263	***	0.282	***	-0.857	***	53.801%
	(0.138)		(0.039)		(0.066)		(0.080)		(0.230)		
HAR-trd	0.677	***	0.073	**	0.256	***	0.249	***	-0.065	***	54.472%
	(0.144)		(0.035)		(0.064)		(0.079)		(0.016)		
HAR-vix	0.104		0.043	**	0.197	***	0.162		0.414	***	55.351%
	(0.078)		(0.021)		(0.064)		(0.142)		(0.160)		

Table 2.9: a) In-sample estimation of HARX (eq. (2.4)) using SPY data for  $y = RV^{(m)}$ , Pre-Crisis Period

	$\hat{eta}_0^{(m)}$		$\hat{eta}_d^{(m)}$		$\hat{eta}_w^{(m)}$		$\hat{eta}_m^{(m)}$		$\hat{eta}_x^{(m)}$		Adj. $R^2$
HAR	0.509	***	0.122	***	0.348	*	0.249				57.038%
	(0.173)		(0.033)		(0.180)		(0.179)	**			
HAR-ue	0.353	***	0.111	***	0.369	**	0.090		0.005	*	61.496%
	(0.122)		(0.028)		(0.183)		(0.247)	***	(0.003)		
HAR-ind	0.596	***	0.104	***	0.359	**	0.070		-0.247	**	66.151%
	(0.155)		(0.025)		(0.155)		(0.221)	*	(0.118)		
HAR-hst	0.180		0.115	***	0.383	*	0.117		-0.004		58.952%
	(0.229)		(0.031)		(0.201)		(0.261)		(0.003)		
HAR-ts	0.481	***	0.122	***	0.348	*	0.241		0.048		57.110%
	(0.153)		(0.032)		(0.178)		(0.180)		(0.054)		
HAR-cci	35.711		0.117	***	0.363	***	0.179		-0.357		58.464%
	(23.154)	**	(0.020)		(0.124)		(0.119)		(0.233)	**	
HAR-cli	11.680		0.119	***	0.361	***	0.202	*	-0.112		57.446%
	(9.511)		(0.020)		(0.125)		(0.117)		(0.094)		
HAR-bci	48.044		0.111	***	0.380	***	0.093		-0.476		59.347%
	(29.977)		(0.020)		(0.128)		(0.153)		(0.299)		
HAR-ppi	-8.676		0.120	***	0.318	***	0.288	***	0.051		59.753%
	(6.545)		(0.020)		(0.107)		(0.085)		(0.037)		
HAR-ads	0.218		0.110	***	0.366	***	0.080		-0.652	*	62.126%
	(0.163)		(0.022)		(0.136)		(0.175)		(0.374)		
HAR-fd	0.529	***	0.121	***	0.337	*	0.271		0.054		57.328%
	(0.187)		(0.033)		(0.173)		(0.169)		(0.046)		
HAR-vol	-0.163		0.053		0.337	**	0.224		0.442		58.112%
	(0.365)		(0.044)		(0.162)		(0.170)		(0.310)		
HAR-trd	0.232		0.091	***	0.348	**	0.215		0.009		57.356%
	(0.192)		(0.034)		(0.160)		(0.166)		(0.008)		
HAR-vix	0.481	**	0.117	***	0.346	*	0.228		0.034		57.006%
	(0.236)		(0.042)		(0.179)		(0.195)		(0.142)		

Table 2.9: b) In-sample estimation of HARX (eq. (2.4)) using SPY data for  $y = RV^{(m)}$ , Crisis Period

	$\hat{eta}_0^{(m)}$		$\hat{eta}_d^{(m)}$		$\hat{eta}_w^{(m)}$		$\hat{eta}_m^{(m)}$		$\hat{eta}_x^{(m)}$		Adj. $R^2$
HAR	0.235	***	0.100	*	0.068	**	0.360	***			29.782%
	(0.058)		(0.052)		(0.031)		(0.073)				
HAR-ue	0.295	***	0.100	*	0.067	**	0.355	***	0.001		30.433%
	(0.079)		(0.051)		(0.029)		(0.065)		(0.001)		
HAR-ind	0.213	***	0.100	*	0.068	**	0.357	***	0.024		29.872%
	(0.064)		(0.052)		(0.031)		(0.073)		(0.053)		
HAR-hst	0.269	***	0.100	**	0.066	**	0.354	***	-0.001		30.579%
	(0.071)		(0.051)		(0.030)		(0.067)		(0.001)		
HAR-ts	0.052		0.101	*	0.067	**	0.367	***	0.123		31.014%
	(0.115)		(0.053)		(0.030)		(0.067)		(0.087)		
HAR-cci	10.471		0.096	**	0.067	**	0.302	***	-0.103		32.756%
	(8.794)		(0.046)		(0.031)		(0.046)		(0.088)		
HAR-cli	16.468		0.097	*	0.071	**	0.305	***	-0.162		31.808%
	(12.267)		(0.050)		(0.029)		(0.070)		(0.122)		
HAR-bci	9.341	*	0.099	*	0.070	**	0.342	***	-0.091		30.320%
	(5.599)		(0.053)		(0.029)		(0.077)		(0.056)		
HAR-ppi	-0.197		0.101	*	0.067	**	0.363	***	0.002		29.800%
	(1.294)		(0.053)		(0.031)		(0.071)		(0.007)		
HAR-ads	0.241	***	0.100	*	0.068	**	0.361	***	0.035		29.771%
	(0.064)		(0.052)		(0.031)		(0.072)		(0.102)		
HAR-fd	0.080		0.100	*	0.069	**	0.350	***	-0.040		30.296%
	(0.166)	***	(0.052)	***	(0.028)	*	(0.058)		(0.046)		
HAR-vol	-0.061		-0.009		0.016		0.280	***	0.376	**	36.154%
	(0.120)	***	(0.034)	***	(0.037)	***	(0.065)		(0.173)		
HAR-trd	-0.029		-0.013		0.012		0.310	***	0.010	**	33.297%
	(0.092)		(0.040)		(0.044)	*	(0.073)		(0.004)		
HAR-vix	0.118	**	0.034		0.026		0.102		0.332	***	34.307%
	(0.052)		(0.040)		(0.028)		(0.084)		(0.101)		

Table 2.8: c) In-sample estimation of HARX (eq. (2.4)) using SPY data for  $y = RV^{(m)}$ , Post-Crisis Period

_	ADS	BCI	CCI	CLI	EPU	FEDFD	HOUST	INDPRO	PPI	TRD	$\mathbf{TS}$	UE	VIX	VOL	SPLS
CD	1.001	1.005	0.997	0.937	0.992	1.005	1.001	1.003	0.998	1.003	1.006	1.003	1.004	0.994	0.950
$\mathbf{CS}$	1.002	0.999	0.999	0.899	0.981	1.000	0.991	0.997	0.995	0.989	0.991	0.985	0.991	0.965	0.921
Energy	1.002	1.002	1.002	0.951	0.993	1.000	0.999	1.002	0.995	1.003	1.003	1.003	1.017	0.993	0.911
FIN	1.004	1.001	0.998	0.974	0.979	1.004	0.994	0.996	0.992	0.987	0.989	0.985	0.964	0.948	0.958
HC	1.007	1.002	0.999	0.912	0.992	1.002	0.996	1.002	0.998	1.001	1.002	0.998	0.998	0.895	0.917
IND	1.003	1.007	1.012	0.883	1.006	1.003	1.005	1.013	1.004	1.018	1.014	1.019	1.044	1.030	0.886
IT	1.007	1.008	1.001	0.965	0.992	1.006	0.999	1.002	0.994	0.998	0.999	0.997	1.008	1.003	0.974
MAT	1.003	1.006	1.003	0.960	0.996	1.003	1.000	1.007	0.998	1.007	1.007	1.006	1.014	1.015	0.953
TELE	1.008	1.006	1.001	0.927	0.997	1.003	1.003	1.007	0.997	1.008	1.008	1.006	1.014	1.009	0.919
UTI	1.010	1.017	0.996	0.943	0.982	1.005	0.996	1.008	1.002	1.004	1.005	1.000	1.008	1.022	0.943

Table 2.9: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(d)}$  of ten individual stocks in each sector over the full-sample period

	ADS	BCI	CCI	CLI	EPU	FEDFD	HOUST	INDPRO	PPI	TRD	$\mathbf{TS}$	UE	VIX	VOL	SPLS
CD	1.004	1.011	0.978	0.888	0.940	1.004	0.981	0.990	0.966	0.970	0.989	0.968	1.009	0.976	0.889
$\mathbf{CS}$	1.004	1.001	0.976	0.819	0.912	0.995	0.953	0.973	0.959	0.935	0.949	0.926	0.994	0.934	0.864
Energy	1.002	1.001	0.985	0.943	0.932	0.997	0.964	0.973	0.945	0.949	0.964	0.949	1.046	0.994	0.962
FIN	1.004	1.003	0.992	0.933	0.879	0.995	0.946	0.950	0.936	0.901	0.925	0.895	0.938	0.888	0.910
HC	1.011	1.003	0.991	0.880	0.947	1.000	0.973	0.993	0.970	0.976	0.985	0.970	1.045	0.929	0.893
IND	1.005	1.010	0.994	0.842	0.937	1.000	0.974	0.997	0.955	0.979	0.989	0.978	1.038	1.001	0.838
IT	1.029	1.026	0.991	0.953	0.941	1.009	0.977	0.981	0.955	0.950	0.967	0.949	1.014	1.002	0.956
MAT	1.003	1.010	0.992	0.957	0.940	1.000	0.973	0.987	0.952	0.967	0.981	0.964	1.011	1.006	0.950
TELE	1.029	1.015	0.985	0.845	0.950	1.000	0.985	0.994	0.962	0.971	0.990	0.970	1.001	0.963	0.859
UTI	1.017	1.033	0.975	0.863	0.918	1.012	0.963	0.995	0.962	0.972	0.981	0.962	0.997	0.992	0.855

Table 2.10: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(w)}$  of ten individual stocks in each sector over the full-sample period

	ADS	BCI	CCI	CLI	EPU	FEDFD	HOUST	INDPRO	PPI	TRD	$\mathbf{TS}$	UE	VIX	VOL	SPLS
CD	1.022	1.050	0.954	0.903	0.809	0.985	0.906	0.875	0.835	0.781	0.824	0.767	1.000	0.931	0.660
$\mathbf{CS}$	1.013	1.022	0.943	0.855	0.731	0.975	0.831	0.820	0.791	0.689	0.742	0.675	0.989	0.881	0.638
Energy	1.006	1.008	0.947	0.943	0.761	0.982	0.827	0.782	0.744	0.656	0.704	0.667	1.009	0.920	0.656
FIN	1.008	1.027	0.981	0.936	0.723	0.973	0.851	0.805	0.771	0.669	0.715	0.653	0.910	0.875	0.656
HC	1.021	1.019	0.957	0.923	0.766	0.973	0.873	0.870	0.811	0.770	0.805	0.758	1.087	0.959	0.698
IND	1.014	1.018	0.943	0.908	0.730	0.980	0.852	0.827	0.764	0.721	0.768	0.710	1.021	0.935	0.669
IT	1.092	1.079	0.972	0.983	0.794	1.016	0.901	0.838	0.800	0.713	0.768	0.709	1.016	0.981	0.757
MAT	1.009	1.032	0.961	0.981	0.756	0.981	0.857	0.817	0.769	0.701	0.752	0.690	0.998	0.968	0.700
TELE	1.064	1.021	0.948	0.839	0.813	0.984	0.908	0.876	0.826	0.765	0.820	0.769	0.992	0.906	0.672
UTI	1.023	1.049	0.921	0.858	0.734	0.999	0.831	0.830	0.775	0.720	0.751	0.709	0.998	0.971	0.633

Table 2.11: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(m)}$  of ten individual stocks in each sector over the full-sample period

Table 2.12: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(d)}$  of ten individual stocks in each sector during the Pre-Crisis period

Pre-Crisis	ADS	BCI	CCI	CLI	EPU	FEDFD	HOUST	INDPRO	PPI	TRD	$\mathbf{TS}$	UE	VIX	VOL	SPLS
CD	1.020	1.039	0.985	0.893	0.989	1.060	1.010	1.008	0.998	1.015	1.024	1.009	1.033	0.998	1.018
$\mathbf{CS}$	1.011	1.028	0.987	0.927	0.999	1.028	1.006	1.005	1.000	0.995	1.009	1.001	1.017	1.000	0.972
Energy	1.022	1.002	0.977	1.008	1.007	1.010	1.021	1.001	1.002	0.994	1.004	1.005	1.401	1.225	1.151
FIN	1.046	1.132	0.951	0.881	0.980	1.089	1.051	1.015	0.995	1.001	1.025	1.005	1.998	1.557	1.210
$\mathbf{HC}$	1.005	1.007	0.996	0.970	1.005	1.019	1.006	1.003	1.001	0.999	1.003	1.002	1.019	1.013	1.000
IND	1.006	1.034	0.987	0.917	0.991	1.028	1.007	1.006	0.994	1.003	1.010	1.003	1.030	1.035	0.969
IT	1.084	1.095	0.974	0.908	1.002	1.094	1.022	1.018	1.003	0.996	1.016	0.999	1.033	1.049	0.991
$\mathbf{MAT}$	1.009	1.023	0.995	1.032	1.001	1.026	1.002	1.003	1.000	1.003	1.008	1.002	1.012	1.014	1.028
TELE	1.064	1.041	0.968	0.916	1.029	1.043	1.032	1.023	1.002	0.999	1.036	1.015	1.039	1.146	0.989
$\mathbf{UTI}$	1.111	1.215	0.967	0.947	0.938	1.088	0.995	1.005	1.107	1.029	1.017	1.009	1.161	1.711	1.289

Table 2.13: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(w)}$  of ten individual stocks in each sector during the Pre-Crisis period

	ADS	BCI	CCI	$\mathbf{CLI}$	$\mathbf{EPU}$	FEDFD	HOUST	INDPRO	$\mathbf{PPI}$	$\mathbf{TRD}$	$\mathbf{TS}$	$\mathbf{UE}$	VIX	VOL	SPLS
CD	1.128	1.161	0.946	0.707	0.959	1.099	1.033	1.028	0.995	1.020	1.078	1.025	1.045	0.922	0.795
$\mathbf{CS}$	1.066	1.108	0.977	0.780	0.994	1.051	1.028	1.018	0.999	0.979	1.026	1.004	1.023	0.901	0.811
Energy	1.050	0.998	0.963	0.962	0.995	1.017	1.045	0.993	1.007	0.995	1.001	1.010	1.186	1.153	1.016
FIN	1.192	1.350	0.937	0.655	0.944	1.120	1.033	1.032	0.992	0.997	1.054	1.015	1.325	1.010	0.690
HC	1.026	1.021	0.996	0.890	1.017	1.040	1.028	1.011	0.993	0.994	1.014	1.012	1.004	0.977	0.904
IND	1.061	1.143	0.973	0.732	0.958	1.030	1.019	1.026	0.983	1.011	1.033	1.011	1.009	0.919	0.770
IT	1.429	1.347	0.986	0.780	0.999	1.186	1.123	1.069	1.014	0.980	1.062	0.999	1.053	0.979	0.787
MAT	1.035	1.090	0.987	1.006	0.996	1.051	1.012	1.012	0.993	1.010	1.026	1.005	1.012	0.994	0.962
TELE	1.145	1.052	0.971	0.655	1.116	1.052	1.125	1.073	1.020	0.991	1.089	1.047	1.004	1.073	0.787
UTI	1.145	1.284	0.907	0.858	0.916	1.156	1.010	1.016	1.089	1.014	1.035	1.016	1.067	1.263	0.782

Table 2.14: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(m)}$  of ten individual stocks in each sector during the Pre-Crisis period

	ADS	BCI	CCI	$\mathbf{CLI}$	$\mathbf{EPU}$	FEDFD	HOUST	INDPRO	$\mathbf{PPI}$	TRD	$\mathbf{TS}$	$\mathbf{UE}$	VIX	VOL	SPLS
CD	1.435	1.323	0.970	0.632	0.916	1.107	1.111	1.072	0.986	0.968	1.102	1.024	1.027	0.785	0.805
$\mathbf{CS}$	1.277	1.242	1.020	0.683	0.996	1.098	1.109	1.058	0.976	0.921	1.070	1.013	0.993	0.750	0.807
Energy	1.096	0.976	0.984	0.933	0.989	1.023	1.064	1.015	1.050	1.012	0.969	1.038	1.000	1.013	0.813
FIN	1.476	1.519	0.980	0.586	0.920	1.141	1.096	1.061	0.994	0.956	1.095	1.019	1.018	0.858	0.629
HC	1.172	1.060	1.027	0.849	1.051	1.071	1.139	1.047	0.999	0.975	1.064	1.038	0.999	0.882	0.863
IND	1.274	1.295	1.000	0.691	0.892	1.044	1.062	1.065	0.975	0.984	1.065	1.022	0.951	0.776	0.749
IT	1.920	1.495	1.039	0.965	0.953	1.306	1.318	1.108	1.053	0.937	1.122	0.997	1.090	0.864	1.050
MAT	1.175	1.242	1.005	0.924	0.995	1.118	1.054	1.070	0.961	1.006	1.075	1.019	1.026	0.973	1.052
TELE	1.214	0.955	1.014	0.581	1.243	1.066	1.276	1.126	1.057	0.982	1.134	1.102	0.926	0.887	0.825
UTI	1.199	1.319	0.911	0.632	0.922	1.185	1.093	1.062	1.044	0.989	1.046	1.033	1.046	1.084	0.647

Table 2.15: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(d)}$  of ten individual stocks in each sector during the Crisis period

Crisis	ADS	BCI	CCI	CLI	$\mathbf{EPU}$	FEDFD	HOUST	INDPRO	$\mathbf{PPI}$	TRD	$\mathbf{TS}$	UE	VIX	VOL	$\mathbf{SPLS}$
CD	1.001	1.005	0.997	0.940	0.992	1.000	0.998	1.003	0.995	1.002	1.005	1.002	1.005	1.003	0.943
$\mathbf{CS}$	1.002	0.999	1.001	0.881	0.980	0.999	0.989	0.999	0.991	0.990	0.991	0.986	0.997	0.973	0.905
Energy	1.002	1.003	1.002	0.946	0.992	1.000	0.997	1.004	0.989	1.004	1.004	1.004	1.013	0.989	0.897
FIN	1.004	1.001	0.997	0.971	0.978	1.004	0.992	0.994	0.990	0.985	0.988	0.984	0.963	0.947	0.955
$\mathbf{HC}$	1.004	1.004	1.000	0.920	0.996	1.000	0.998	1.009	0.996	1.006	1.006	1.003	1.023	0.997	0.926
IND	1.003	1.007	1.013	0.876	1.005	1.003	1.003	1.013	1.001	1.018	1.014	1.019	1.044	1.030	0.879
IT	1.004	1.005	1.001	0.961	0.990	1.002	0.995	1.001	0.989	0.997	0.998	0.996	1.009	1.003	0.969
MAT	1.003	1.006	1.002	0.948	0.995	1.002	0.998	1.007	0.992	1.007	1.007	1.006	1.015	1.017	0.937
TELE	1.002	1.003	1.002	0.911	0.996	1.001	1.000	1.006	0.994	1.008	1.006	1.006	1.013	1.007	0.904
UTI	1.001	0.999	0.996	0.934	0.986	0.998	0.995	1.009	0.988	1.000	1.005	1.000	0.999	0.990	0.918

Table 2.16: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(w)}$  of ten individual stocks in each sector during the Crisis period

	ADS	BCI	CCI	$\mathbf{CLI}$	$\mathbf{EPU}$	FEDFD	HOUST	INDPRO	$\mathbf{PPI}$	$\mathbf{TRD}$	$\mathbf{TS}$	$\mathbf{UE}$	VIX	VOL	SPLS
CD	1.001	1.007	0.982	0.895	0.938	0.990	0.969	0.982	0.948	0.960	0.977	0.957	1.007	0.986	0.893
$\mathbf{CS}$	1.003	0.998	0.975	0.801	0.899	0.992	0.940	0.966	0.937	0.925	0.943	0.915	0.991	0.931	0.844
Energy	1.003	1.001	0.984	0.948	0.916	0.996	0.949	0.963	0.911	0.937	0.955	0.937	1.031	0.974	0.962
FIN	1.005	0.999	0.989	0.929	0.876	0.996	0.938	0.940	0.927	0.894	0.921	0.890	0.938	0.888	0.906
HC	1.005	1.005	0.984	0.855	0.927	0.992	0.960	0.992	0.942	0.962	0.976	0.956	1.005	0.968	0.858
IND	1.005	1.008	0.993	0.834	0.930	1.000	0.965	0.990	0.939	0.972	0.986	0.971	1.037	0.998	0.827
IT	1.007	1.008	0.990	0.961	0.926	0.997	0.956	0.964	0.927	0.935	0.954	0.935	1.013	1.001	0.961
MAT	1.005	1.010	0.990	0.950	0.928	0.998	0.958	0.976	0.925	0.955	0.974	0.951	1.014	1.008	0.942
TELE	1.003	1.004	0.984	0.835	0.938	0.995	0.968	0.985	0.944	0.963	0.978	0.959	1.001	0.961	0.829
UTI	1.001	0.994	0.978	0.861	0.923	0.994	0.956	0.992	0.926	0.959	0.977	0.955	0.985	0.953	0.852

Table 2.17: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(m)}$  of ten individual stocks in each sector during the Crisis period

	ADS	BCI	CCI	CLI	$\mathbf{EPU}$	FEDFD	HOUST	INDPRO	$\mathbf{PPI}$	$\mathbf{TRD}$	$\mathbf{TS}$	$\mathbf{UE}$	VIX	VOL	SPLS
CD	1.011	1.045	0.951	0.935	0.799	0.968	0.860	0.819	0.783	0.727	0.776	0.715	0.995	0.945	0.620
$\mathbf{CS}$	1.010	1.010	0.930	0.864	0.706	0.970	0.783	0.763	0.735	0.636	0.697	0.626	0.986	0.882	0.592
Energy	1.009	1.007	0.938	0.957	0.703	0.981	0.768	0.702	0.654	0.574	0.628	0.591	1.010	0.909	0.614
FIN	1.012	1.007	0.974	0.935	0.724	0.980	0.827	0.766	0.749	0.640	0.694	0.635	0.910	0.874	0.644
$\mathbf{HC}$	1.014	1.015	0.922	0.901	0.702	0.971	0.804	0.790	0.724	0.677	0.727	0.668	0.994	0.943	0.610
IND	1.012	1.008	0.930	0.917	0.715	0.982	0.818	0.778	0.720	0.678	0.735	0.674	1.022	0.933	0.640
$\mathbf{IT}$	1.022	1.033	0.958	0.989	0.763	0.981	0.821	0.755	0.720	0.641	0.698	0.642	1.009	0.988	0.683
MAT	1.015	1.027	0.947	0.980	0.730	0.980	0.815	0.750	0.706	0.636	0.700	0.637	0.999	0.965	0.652
TELE	1.013	1.016	0.937	0.885	0.771	0.974	0.841	0.803	0.759	0.703	0.764	0.703	0.997	0.926	0.609
$\mathbf{UTI}$	1.009	0.992	0.910	0.898	0.723	0.976	0.788	0.774	0.700	0.661	0.707	0.662	0.985	0.913	0.627

Table 2.18: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(d)}$  of ten individual stocks in each sector during the Post-Crisis period

	ADS	BCI	CCI	CLI	EPU	FEDFD	HOUST	INDPRO	PPI	TRD	$\mathbf{TS}$	UE	VIX	VOL	SPLS
CD	0.992	0.995	1.032	1.200	0.980	0.990	1.026	1.016	1.040	1.018	1.008	1.009	0.994	0.989	1.227
$\mathbf{CS}$	0.981	0.996	1.017	0.985	0.975	0.982	1.044	1.104	1.051	1.069	1.048	1.018	0.966	0.982	0.974
Energy	1.001	1.002	1.007	1.004	0.999	1.003	1.011	1.005	1.038	1.004	1.000	1.002	0.996	0.996	1.009
FIN	0.989	1.056	1.279	1.249	1.073	0.980	1.240	1.370	1.308	1.286	1.219	1.123	1.331	1.213	1.379
HC	1.002	1.000	1.008	0.972	0.991	1.003	1.005	1.003	1.011	1.005	1.003	1.001	0.979	0.892	0.973
IND	1.000	1.001	1.020	1.026	0.999	1.006	1.027	1.009	1.068	1.013	1.004	1.008	0.989	1.015	0.988
IT	1.001	1.003	1.008	0.988	0.997	1.001	1.012	1.008	1.021	1.004	1.000	1.002	0.993	0.986	0.990
MAT	0.999	1.003	1.017	1.049	0.998	1.006	1.020	1.008	1.051	1.011	1.004	1.008	0.986	0.991	1.036
TELE	1.000	1.001	1.011	1.114	0.995	1.005	1.012	1.006	1.031	1.008	1.002	1.004	1.019	1.039	1.063
UTI	0.994	0.994	1.046	1.171	0.973	0.989	1.027	1.012	1.080	1.016	1.001	1.008	1.101	1.198	1.134

Table 2.19: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(w)}$  of ten individual stocks in each sector during the POst-Crisis period

	ADS	BCI	CCI	$\mathbf{CLI}$	$\mathbf{EPU}$	FEDFD	HOUST	INDPRO	PPI	TRD	$\mathbf{TS}$	UE	VIX	VOL	SPLS
CD	0.971	0.991	1.071	1.303	0.940	0.966	1.071	1.067	1.125	1.048	1.012	1.022	1.039	1.044	1.292
$\mathbf{CS}$	0.960	1.018	1.022	0.981	0.950	0.967	1.076	1.169	1.108	1.100	1.062	1.027	0.969	0.998	0.969
Energy	0.990	1.013	1.008	0.985	0.997	0.996	1.031	1.031	1.132	1.011	1.000	1.000	1.001	1.013	0.985
FIN	0.906	1.360	1.246	1.256	1.065	0.864	1.609	1.978	1.795	1.600	1.391	1.252	1.000	1.002	1.259
$\mathbf{HC}$	0.995	1.000	1.013	1.029	0.973	1.001	1.020	1.017	1.063	1.021	1.007	1.008	1.074	0.955	0.980
IND	0.984	1.008	1.025	1.039	0.999	0.994	1.088	1.056	1.204	1.047	1.006	1.027	0.999	1.074	1.062
IT	0.987	1.013	1.002	0.985	0.979	0.993	1.040	1.044	1.083	1.012	0.993	1.001	0.998	1.006	0.998
MAT	0.976	1.016	1.028	1.053	0.998	1.004	1.079	1.071	1.203	1.050	1.020	1.031	0.986	0.990	1.050
TELE	0.983	1.004	1.012	1.160	0.977	1.009	1.055	1.044	1.153	1.040	1.008	1.021	1.033	1.097	1.124
UTI	0.959	0.993	1.064	1.160	0.926	0.963	1.093	1.056	1.305	1.056	1.003	1.023	1.127	1.319	1.150

Table 2.20: OOS median MSE % gains of HARX relative to the HAR model in forecasting  $RV^{(m)}$  of ten individual stocks in each sector during the POst-Crisis period

	ADS	BCI	CCI	CLI	EPU	FEDFD	HOUST	INDPRO	PPI	TRD	$\mathbf{TS}$	UE	VIX	VOL	SPLS
CD	0.886	1.035	1.159	1.250	0.795	0.929	1.216	1.388	1.255	1.211	1.099	1.048	1.062	1.101	1.014
$\mathbf{CS}$	0.889	1.118	1.060	0.973	0.819	0.912	1.199	1.369	1.303	1.196	1.086	1.026	0.979	1.027	0.969
Energy	0.943	1.077	1.068	1.015	0.952	0.952	1.134	1.311	1.289	1.138	1.076	1.011	0.977	1.008	0.954
FIN	0.743	1.959	1.269	1.136	0.662	0.629	1.866	2.700	2.001	2.023	1.699	1.381	0.973	1.001	1.291
$\mathbf{HC}$	0.944	1.029	1.085	1.040	0.888	0.937	1.131	1.217	1.155	1.127	1.051	1.040	1.219	1.031	0.987
IND	0.909	1.080	1.138	0.983	0.890	0.904	1.272	1.439	1.342	1.236	1.082	1.069	0.992	1.087	1.051
$\mathbf{IT}$	0.919	1.093	1.051	0.989	0.853	0.944	1.188	1.340	1.273	1.148	1.051	1.005	0.998	1.015	1.084
MAT	0.887	1.094	1.150	1.052	0.898	0.942	1.258	1.504	1.378	1.265	1.161	1.079	0.964	0.973	1.086
TELE	0.870	1.057	1.108	1.051	0.807	0.939	1.216	1.398	1.346	1.237	1.101	1.056	1.011	1.104	1.088
$\mathbf{UTI}$	0.747	1.074	1.296	1.020	0.665	0.781	1.299	1.690	1.613	1.334	1.162	1.032	1.058	1.244	1.181

## Chapter 3

# SHARP: A State-Space HAR model using Particle Gibbs Sampling

#### Abstract

We propose a general state-space autoregressive (AR) model with time-varying coefficients that follow an AR process with stochastic volatility. We implement these new specifications in the HAR framework to capture the time-varying salient feature of volatility using a two-state representation via a) allowing the time-varying coefficients to follow an AR(1) specification. b) introducing stochastic volatility for the innovations of the coefficients. Using high-frequency data of the SPY-ETF and representative NYSE stocks from 2000 to 2016, we show that the proposed model estimated using particle Gibbs sampling consistently outperforms different HAR

model specifications in forecasting financial volatility.

**Key words:** Forecasting; Heterogeneous Autoregressive Realised Volatility model; Particle GIBBS sampling; Sequential Monte Carlo; State Space models.

#### 3.1 Introduction

There is considerable evidence of time-varying parameters and volatility in time series macroeconomic and financial data. For out-of-sample analyses, models that fail to anticipate changes in the data generating process produce inaccurate predictions. Hence, analysing and accounting for parameters' behaviour is essential in the statistical modelling of such data. Much attention has been given to forecast evaluation (Giacomini and Rossi, 2009), estimation approaches (Inoue, Jin, and Rossi, 2017; Pesaran and Timmermann, 2005), and dynamic models (Bekierman and Manner (2018), Buccheri and Corsi (2021), Chen et al. (2018), Clark and Ravazzolo (2015), and D'Agostino, Gambetti, and Giannone (2013), among others).

This work presents a novel state-space model where the coefficients are assumed to be latent factors that follow an autoregressive (AR) process with Stochastic Volatility (SV). Such features can be utilised for any model specification, including exogenous regressors, trends, seasonal dummies, or autoregressive structures. We use Bayesian inference to estimate the model with Particle Gibbs sampling (Andrieu, Doucet, and Holenstein, 2010) following the Creal and Tsay (2015) procedure. This method permits an efficient computation of the latent variables. We implement the new specifications in modelling and forecasting the volatility of financial assets returns. Financial volatility is a preeminent topic in theoretical and empirical finance due to its significance in risk management, portfolio selection, and asset pricing. In a seminal work, to capture the time-varying conditional volatility of financial returns, Engle (1982) introduced the ARCH model. Extending that work, Bollerslev (1986) proposed the GARCH model, which became widely accepted. In further development, the non-parametric realised volatility (RV) - defined as the sum of intraday returns is preferred to parametric volatility measures generated by GARCH and SV models. The improvement derives from the delivery of information over much smaller intervals of time (Andersen and Bollerslev, 1998; Barndorff-Nielsen and Shephard, 2002; Liu, Patton, and Sheppard, 2015).

Since then, significant work has been done on linear reduced-form modelling of the RV series, where the heterogeneous autoregressive (HAR) model by Corsi (2009) has become the preferred specification for that purpose. The model is distinctive by its heterogeneous market hypothesis, i.e., that, with diverse horizons and trading activities, market participants may react differently to identical news. It models daily volatility as a sequence of RV's daily, weekly, and monthly historical average. By its design, it parsimoniously captures RV's persistence. Further, Corsi (2009) notes that modelling the RV's log-transformation in the HAR model (henceforth, HARL) ensures the partial volatilities' positiveness, improves the dynamic specification for the RV and delivers more promising forecasts.

Despite its empirical success, the linear HARL model does not account for measurement errors and nonlinear dependencies, which may lead to biased OLS estimates, autocorrelated and highly heteroskedastic residuals, and time-varying OLS coefficients. Several proposals have emerged to address these miss-specifications in the HARL model. Inspired by the work of Bollerslev, Patton, and Quaedvlieg (2016), but instead of explicitly using the realise quarticity (RQ), a proxy of the measurement error variance, Bekierman and Manner (2018) presents a state-space HARL model that specifies the first-order AR coefficient as a latent Gaussian AR(1) process. This specification also captures additional sources of temporal variations in the coefficient of interest. In parallel work, Chen et al. (2018) propose, for the HARL model, time-varying coefficients of an unknown functional form. Further, in an attempt to model both time-varying coefficients and measurement errors, Buccheri and Corsi (2021) propose score-driven coefficients and heteroskedastic innovations. Several other papers address nonlinear dependency in different approaches, such as regime-changing (McAleer and Medeiros, 2008) and dynamic model averaging (Wang et al., 2016), among others.

Our work is mostly related to that of Bekierman and Manner (2018), Buccheri and Corsi (2021), and Chen et al. (2018) in the HARL framework. We model all the coefficients (including the intercept) as an AR(1) process either with or without SV. Hence, we account for the model's nonlinearity and relax the assumption of homoskedastic errors in the coefficients' process. We also account for autocorrelation and heteroskedasticity of the residuals in the main equation. We estimate the statespace model in the HARL framework by using Particle Gibbs sampling. We refer to the proposed model i) without SV by SHARP and ii) with SV by SHARP-sv.

We compare the models' forecasting performance with a recent (parallel) work within the HARL framework. In an empirical study using daily RV of SPY-ETF, as a tradable US market index, and twenty representative NYSE individual stocks over seventeen years from 2000 to 2016, both the SHARP and SHARP-sv models significantly outperform the HARL model and its existing time-varying extensions in predicting RV. While the SHARP-sv model offers a moderate forecasting advantage over the SHARP model, both: (i) are consistently included in the model confidence set of Hansen, Lunde, and Nason (2011); (ii) significantly outperform the HARL model in forecasting RV according to the predictive ability test by Giacomini and White (2006).

Our contribution to the literature can be summarised as follows: i) we introduce a general state-space model with relaxed features (time-varying coefficients that follow an AR(1) process with SV) and suggest a feasible estimation method. ii) We utilise these new features in the HARL framework (SHARP and SHARP-sv) to improve the forecast of RV. iii) We compare the proposed models against recent proposals in modelling the logRV series with time-varying coefficients.

The remainder of this paper is organised as follows. Firstly, in Section 3.2, we describe the realised volatility measure and provide an overview of existing dynamic and time-varying parameters extensions to the HARL model. We introduce the proposed models in Section 3.3 (refer to the Appendix for details on the estimation method). Section 3.4 analyses and evaluates the proposed extensions under the HARL framework in forecasting stock volatility using real data. Section 3.5 concludes.

# 3.2 Volatility Measure and HARL Family of Models

Consider an asset whose log-price,  $\log(P_s)$ , process is given by the stochastic differential equation:

$$d\log(P_s) = \mu_s + \sigma_s dW_s \tag{3.1}$$

where  $\mu_s$  denotes the drift,  $\sigma_s$  is the instantaneous volatility and  $W_t$  a standard Brownian motion. The latent integrated variance for day t is defined as:

$$IV_t = \int_{t-1}^t \sigma_s^2 ds. \tag{3.2}$$

RV is a nonparametric ex-post estimate of the return variation, where the daily RV measure is the sum of high-frequency squared returns of sub-intervals of the day. Specifically, the daily RV is calculated by averaging intra-daily squared returns over a one-day horizon, t, using M sub-intervals.

$$RV_t^{(d)} := \sum_{j=1}^M r_{j,t}^2, \tag{3.3}$$

where  $r_{j,t} = log(P_{(t-1)M+j}) - log(P_{(t-1)M+(j-1)})$  is the intra-day return of the *jth* sub-interval within the  $t^{th}$  day. P is the asset price at the start of the  $j^{th}$  interval computed as the average of the closing and opening prices of intervals j - 1 and j, respectively. Zhang, Mykland, and Aït-Sahalia (2005) provide a discussion on optimising the sampling frequency for the estimation of RV. However, Buccheri and Corsi (2021) show that the relative forecast performance of models with time-varying coefficients, such as the ones discussed in this paper, is independent of the sampling

frequency. Therefore, for conciseness, we use sub-intervals of length 300 seconds in constructing the daily RV series. The latter defines 78 intraday sub-intervals and combines balanced information from high-frequency data and microstructure effects (Andersen et al., 2001).

In the aim of forecasting the daily RV, Corsi (2009) introduced the Heterogeneous Autoregressive Realised Volatility (HARL) model defined as:

$$logRV_t^{(d)} = \beta_1 + \beta_2 logRV_{t-1}^{(d)} + \beta_3 logRV_{t-1}^{(w)} + \beta_4 logRV_{t-1}^{(m)} + v_t ; \quad v_t \sim N(0, \sigma_v^2)$$
(3.4)

where  $log RV^{(d)}$  denotes the logarithmic transformation of daily RV. To account for the bias generated by the logarithmic transformation, one-step-ahead forecasts of RVare estimated using the moment generating function of the normal distribution.

Respectively,  $logRV^{(w)}$  and  $logRV^{(m)}$  are the weekly and monthly logRV realised at time t. These are computed over a recursive rolling window of fixed length (week or month) as follows:

$$log RV_t^{(w)} = \frac{1}{5} \Sigma_{i=1}^5 log RV_{t-i}^{(d)} \quad ; \qquad log RV_t^{(m)} = \frac{1}{22} \Sigma_{i=1}^{22} log RV_{t-i}^{(d)}$$

Henceforth, we refer to  $\beta_1$ ,  $\beta_2$ ,  $\beta_3$ , and  $\beta_4$  as the intercept, daily, weekly, and monthly coefficient, respectively.

Inspired by Barndorff-Nielsen and Shephard (2002) and Bollerslev, Patton, and Quaedvlieg (2016), Bekierman and Manner (2018) explain that the measurement error, though less severe, also exists in logRV series, where:

$$logRV_t = logIV_t + \epsilon_t, \quad \epsilon_t \sim MN(0, 2\Delta \frac{IQ_t}{IV_t^2}).$$
 (3.5)

and  $V_t = \frac{\sum_{j=1}^{M} r_{j,t}^4}{(\sum_{j=1}^{M} r_{j,t}^2)^2}$  is a consistent estimate of the variance of  $\epsilon_t$  (Buccheri and Corsi, 2021).

Hence, Bekierman and Manner (2018) propose the HARSL model, a state-space HARL model for the logRV series that specifies the daily coefficient as a latent Gaussian AR(1) process. The model is estimated using maximum likelihood with a standard Kalman filter. Their model consistently outperforms the HAR model in forecasting the RV. Its empirical success lies in the specification of capturing other sources of temporal variation in addition to the variance of the measurement error. However, they note that the model's maximum likelihood estimator is inefficient. Further, allowing all the coefficients to follow an AR process is computationally challenging to perform using their employed estimation method.

In parallel work, Chen et al. (2018) suggest a HARL model with time-varying coefficients of an unknown functional form (TVCHAR). They apply a local linear smoothing method to estimate the model¹. While their proposed model is flexible by allowing all the coefficients to be time-varying, they show that it only outperforms the HARL model for long forecasting horizons. Within the same line of research, Buccheri and Corsi (2021) propose the SHARK model that features time-varying coefficients and heteroskedastic error terms and handles measurement errors. They show that their SHARK model yields moderate improvements in the forecasts for the one day

¹We refer the reader to their paper for more details about model estimation and to Casas and Fernandez-Casal (2019) for their R code.

ahead but is more prominent for long-term forecasting.

The above studies are primarily concerned with estimation techniques with specific underlying assumptions that better reflect the underlying process of logRV series. They suggest remedies for particular miss-specifications and successfully improve longterm forecasting of RV. In comparison, the employed estimation method presented in the next section makes it feasible to relax some of the assumptions. Consequently, our proposed model of time-varying coefficients with SV is general and, by nature, accounts for all forms of the HARL miss-specifications. It captures the dynamic statistical characteristics of logRV and its relationship with the lagged observations.

#### 3.3 Methodology

In this section we describe the two variations of the introduced model. The general framework is:

$$y_t = x'_t \beta_t + v_t, \quad v_t \sim \mathcal{N}(0, \sigma_v^2), \quad t = 1, \dots, n,$$
(3.6)

$$\beta_{t,j} = \alpha_j + \rho_j \beta_{t-1,j} + \varepsilon_{tj}, \quad j = 1, \dots, k, \quad t = 1, \dots, n,$$
(3.7)

In the first specification, we assume:

$$\varepsilon_{tj} \sim \mathcal{N}\left(0, \sigma_{\varepsilon_j}^2\right), \quad j = 1, \dots, k, \quad t = 1, \dots, n,$$
(3.8)

In the second one, we assume stochastic volatility (SV) :

$$\varepsilon_{tj}|h_{tj} \sim \mathcal{N}(0, h_{tj}), \quad j = 1, \dots, k, \quad t = 1, \dots, n,$$

$$(3.9)$$

where

$$\log h_{tj} = \gamma_j + \delta_j \log h_{t-1,j} + u_{tj}, \quad j = 1, \dots, k, \quad t = 1, \dots, n,$$
(3.10)

$$u_{tj} \sim \mathcal{N}(0, \sigma_{uj}^2), \quad j = 1, \dots, k, \quad t = 1, \dots, n.$$
 (3.11)

We treat  $h_0$  and  $\beta_{0j}$  as unknown parameters. Note that the model can be extended to include autoregressive structure, deterministics such as a constant, (linear or nonlinear) trends, or seasonal dummies. Also, other specifications for time-varying heteroskedasticity are possible.

Utilising the introduced models in forecasting volatility, we set  $y_t = log(RV_t^{(d)})$  and  $x_t = (log(RV_{t-1}^{(d)}), log(RV_{t-1}^{(w)}), log(RV_{t-1}^{(m)}))'$ . Hence, the first variation, denoted by SHARP, is a state space HARL model where the coefficients are time-varying that follow an AR(1) process. The second variation, denoted by SHARP-sv, is the same as the SHARP model with the additional feature of SV in the AR process of the coefficients.

The measurement equation (3.6) describes the vector of observations,  $log(RV_t^{(d)})$ , in terms of the independent variables, the state vector,  $\beta_t$ , and the disturbances,  $v_t$ . The "coefficients" transition equation (3.7) describes the evolution of the coefficients over time. Indeed, the AR process of the intercept also captures the autocorrelation of the residuals in the measurement equation. Under stationarity, while the unconditional first and second moments of  $\beta_{jt}$  are constant, its conditional second moment,  $var(\beta_{jt}|h_{jt})$ , can change over time. The "variance" transition equation (3.10) describes the change over time of the logarithmic transformation of the conditional variance of the innovations,  $\varepsilon_{tj}$ , in the "coefficients" transition equation. The latter feature allows not only for heteroskedasticity in the main equation (through the intercept) but for heteroskedasticity in the coefficients' process as well. Note that, under stationarity, the first and second moments of  $h_{jt}$  are constant (see Appendix 3.A).

To estimate the model, we use a modification of the sequential Monte Carlo method known as the particle Gibbs (PG) sampler, see Andrieu, Doucet, and Holenstein (2010). The latent variables in our model are  $\lambda_t = [\beta'_t, h'_t]'$  where  $\beta_t = [\beta_{t1}, \ldots, \beta_{tk}]$  and  $h_t = [h_{t1}, \ldots, h_{tk}]$  whose prior can be described by  $p(\lambda_t | \lambda_{t-1}, \theta)$ . The joint posterior is  $p(\theta, \lambda_{1:T} | \mathbf{y}_{1:T})$ . In the PG sampler, we can draw the structural parameters as usual, from their posterior conditional distributions  $\theta | \lambda_{1:T}, \mathbf{y}_{1:T}$  (see Appendix 3.B). This is important because, in this way, we can avoid mixture approximations or other Monte Carlo procedures that need considerable tuning and may not have good convergence properties. The latent variables can be integrated out of the joint posterior using the Creal and Tsay (2015) procedure (see Appendix 3.C).

Our choice of conjugate prior for this particular empirical exercise is as follows:

$$\alpha_j, \gamma_j \sim \mathcal{N}(0, 1), \quad j = 1, \dots, k,$$

$$\rho_j, \delta_j \sim \mathcal{N}(0.5, 1) \mathbb{I}_{\rho_j \in (0, 1)} \mathbb{I}_{\delta_j \in (0, 1)}, \quad j = 1, \dots, k,$$

$$\sigma_{uj}^2 \sim \Gamma(6.5, 0.005), \quad j = 1, \dots, k$$

$$\sigma_v^2 \sim \Gamma(6.5, 0.005)$$

Using the above prior, we restrict the estimation range of  $\rho$  and  $\delta$  between (0,1) to guarantee stationarity and reflect the belief that coefficients are positively autocorrelated. The prior of  $\gamma$  and  $\delta$  can be left more flexible in this exercise with a prior mean of 0. The conjugate prior specification of the variances,  $\sigma_u^2$  and  $\sigma_v^2$ , has a low mean.

#### Algorithm of Forecasting RV using SHARP or SHARP-sv

Let n be the total number of observations, is be the in-sample estimation window size (approximately four years), nsim = 15000 be the total number of MCMC iterations, and nburn = 5000 be the number burned iterations.

- For T = is, ..., n 1
  - A. For i = 1, ..., nsim
    - i Draw  $\lambda_t$  for t = is T + 1, ..., T as illustrated in Appendix (3.C)
    - ii Sample the parameters of (3.7) and (3.10) using their posterior distributions in Appendix (3.B)
    - iii Forecast  $\widehat{logRV}_{T+1,i}$  for i = nburn + 1, ..., nsim
  - B. Estimate  $\widehat{logRV}_{T+1} = \sum_{i=nburn+1}^{isim} \frac{\widehat{logRV}_{T+1,i}}{isim-nburn}$

Forecasts of the realised volatility are then computed based on the expectation of a log-normal distribution, as follows:

$$\hat{RV}_{t+1} = \exp(\widehat{log(RV_{t+1})} + \frac{\hat{\omega}_{t+1|t}^2}{2})$$
(3.12)

where:

$$\hat{\omega}_{t+1|t}^{2} = \hat{\sigma}_{v}^{2} + \widehat{Var}(\beta_{1,t+1|t}) + (\log(RV_{t}^{d}))^{2} \times \widehat{Var}(\beta_{2,t+1|t}) \\ + (\log(RV_{t}^{w}))^{2} \times \widehat{Var}(\beta_{3,t+1|t}) + (\log(RV_{t}^{m}))^{2} \times \widehat{Var}(\beta_{4,t+1|t})$$

Here, the first term in the expression of  $\hat{\omega}_{t+1|t}^2$  is the variance of the measurement equation whereas the subsequent terms represent the variance of each of the state

equations entering through the coefficients in the measurement equation. In the SHARP model,  $\widehat{Var}(\beta_{j,t+1|t}) = \widehat{\sigma}_{\epsilon,j}^2$  whereas in the SHARP-sv model,  $\widehat{Var}(\beta_{j,t+1|t}) = exp(\widehat{log(h_{j,t+1})} + \widehat{\sigma}_{u_j}^2/2)$ , for j = 1, ..., 4.

#### 3.4 Empirical Study

#### 3.4.1 Data

The data relates to 4277 trading days of SPY-ETF and twenty NYSE individual stocks RV from 03/01/2000 to 31/12/2016 computed from tick level price observations obtained from TickWrite². The cleaned data makes our results easier to authenticate and replicate. The selected assets allow examining the model performance to predict the price volatility of the market and diverse individual stocks from multiple economic sectors. Also, the sample range enables us to explore the forecasting performance of the models across different market regimes, including the 2008 global financial crisis (GFC). Table (3.1) presents the descriptive statistics of RV of SPY-ETF and the twenty NYSE individual stocks.

²TickWrite is a database that provides data on a commercial basis for futures, Index, and equity markets. Tick Data is sourced from NYSE's TAQ (Trade and Quote) database. Tick adjusts the TAQ database for ticker mapping, code filtering, price splits, and dividend payments https://www.tickdata.com/.

Stock names	Ticker	Sector	Mean	Median	St. Dev.	Skewness	Kurtosis	Min	Max
SPDR S&P 500 ETF Trust	SPY		1.0372	0.4853	2.2589	10.3174	172.6643	0.0128	59.8630
Apple Inc.	AAPL	Information Technology	5.2923	2.5776	7.8061	4.4878	37.8812	0.0791	126.1716
Constellation Energy Group	AEE	Utilities	1.8275	1.0614	3.4532	14.5795	362.0346	0.1089	113.4878
Brown-Forman Corp.	BFB	Consumer Staples	1.9203	1.1517	4.8608	32.2490	1449.2850	0.0742	240.4141
BT Group plc (ADR)	BT	Communications Services	2.3113	1.1621	3.2277	4.7116	47.6503	0.1004	59.5677
Exelon Corp.	EXC	Utilities	2.6354	1.4288	4.7911	9.4223	161.6486	0.1585	130.8746
Freeport-McMoran	FCX	Materials	8.0344	4.3266	12.2786	5.6708	50.4463	0.3168	188.5795
General Dynamics	GD	Industrials	2.2370	1.2810	3.2592	6.2650	67.3871	0.0807	63.2822
General Electric	GE	Industrials	3.0201	1.3030	6.9818	10.3833	172.3605	0.1077	180.3886
The Home Depot	HD	Consumer Discretionary	3.1214	1.5733	4.9381	6.7444	83.0200	0.1557	103.4768
TECO Energy	HES	Energy	4.4738	2.5702	8.6746	12.7967	280.6983	0.2109	271.5113
Humana Inc.	HUM	Health Care	6.6787	2.6090	11.3665	4.4673	33.5169	0.2404	157.5287
IBM	IBM	Information Technology	2.0255	0.9862	3.5274	7.4383	92.6942	0.1019	71.2926
Coca-Cola	KO	Consumer Staples	1.5608	0.8355	2.5353	8.6219	138.4686	0.0456	58.8085
Marriott Int'l.	MAR	Consumer Discretionary	3.5370	1.7819	5.3635	5.4536	55.7000	0.1543	104.5781
Nucor Corp.	NUE	Materials	4.9098	2.7544	10.5817	13.8696	279.0002	0.3337	266.8244
Pfizer	PFE	Health Care	2.3324	1.3819	3.2242	6.4660	77.5016	0.1498	62.6970
AT&T	Т	Communications Services	2.6549	1.1840	4.7673	9.4248	195.7955	0.1082	141.8456
Travelers - Travelers Group Inc	TRV	Financials	2.9683	1.1863	7.8664	15.2588	379.9201	0.1020	263.9287
Wells Fargo	WFC	Financials	4.3023	1.3299	12.1389	8.2667	94.9689	0.1036	226.6092
ExxonMobil	XOM	Energy	1.9866	1.1409	3.9555	15.8314	430.8596	0.1067	141.1297

Table 3.1: Descriptive statistics for RV of SPY and selected twenty NYSE individual stocks.

### 3.4.2 Empirical Estimation of the SHARP and SHARP-sv models

This section reports the estimation results of the SHARP and SHARP-sv models. While the flexible specification of the coefficients as an AR process (eq. 3.7) is possible, we report results using  $\rho = 0.99$ ,  $\delta = 0.99$ ,  $\gamma = 0$ , and  $\delta = 0$ . The general AR process does not yield significantly better out-of-sample results in this exercise. The time-varying estimates of the coefficients using the SHARP and SHARP-sv models are plotted in Figure (3.1) along with horizontal lines representing the fullsample OLS estimates of the simple HARL model (eq. 3.4). Further, the correlation matrix in Table (3.2) depicts the relationship between the estimated time-varying coefficients. The SHARP and SHARP-sv models result in a comparative estimation of the coefficients.

Firstly, the plot of the intercept coefficient,  $\hat{\beta}_{1,t}$ , has a wide range, implying that there may be time-varying factors beyond the lagged logRV terms that explain the variation of daily logRV. Also,  $\hat{\beta}_{1,t}$  is remarkably below the corresponding constant-coefficient during the tranquil periods. Secondly, the daily and weekly coefficients are negatively correlated. For example, the daily coefficient is more prominent during periods of uncertainty, such as the GFC period, while the opposite effects emerge in the weekly coefficient, which declines during such periods. On the other hand, the monthly coefficient is (weekly) positively correlated with the daily coefficient, increasing during the GFC period. During times of uncertainty in the financial market, the observed phenomenon can be attributed to the primacy and recency effects. The current longterm conditions, reflected in the monthly average of  $logRV^d$ , can be regarded as the primary information. The recent information is depicted in the daily  $logRV^d$ . When predicting short-term (daily) volatility during times of uncertainty, primary and recent information become more pertinent than information in the middle, such as the weakly average volatility.

Figure 3.1: Smoothed posterior means of  $\beta_{tj}$  parameters of SHARP and SHARP-sv models along with the corresponding constant-coefficient by the HARL model estimated on SPY-ETF realised variance in the period January 1st, 2000 to December 31st, 2016.

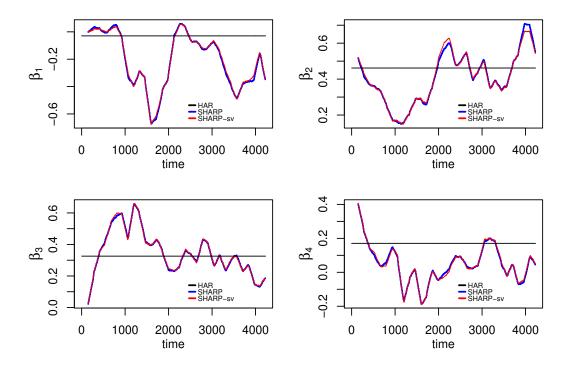


Table 3.2: Correlation matrix of  $\beta_{tj}$  parameters obtained by estimating the SHARP and SHARPsv models on SPY-ETF realised variance in the period January 1st, 2000 to December 31st, 2016.

	SHARP						SHARP	-sv	
	$\hat{eta}_1$	$\hat{\beta}_2$	$\hat{eta}_3$	$\hat{eta}_4$		$\hat{\beta}_1$	$\hat{eta}_2$	$\hat{\beta}_3$	$\hat{eta}_4$
$\hat{\beta}_1$	1	0.1924	-0.0780	0.5014	$\hat{\beta}_1$	1	0.2144	-0.0951	0.5088
$\hat{eta}_2 \\ \hat{eta}_3$	$0.1924 \\ -0.0780$	1 -0.6284	-0.6284 1	$0.2382 \\ -0.2734$	$\begin{vmatrix} \hat{\beta}_2 \\ \hat{\beta}_3 \end{vmatrix}$	0.2144 -0.0951	1 -0.6353	-0.6353 1	$0.2563 \\ -0.2909$
$\hat{eta}_4$	0.5014	0.2382	-0.2734	1	$\hat{\beta}_4$	0.5088	0.2563	-0.2909	1

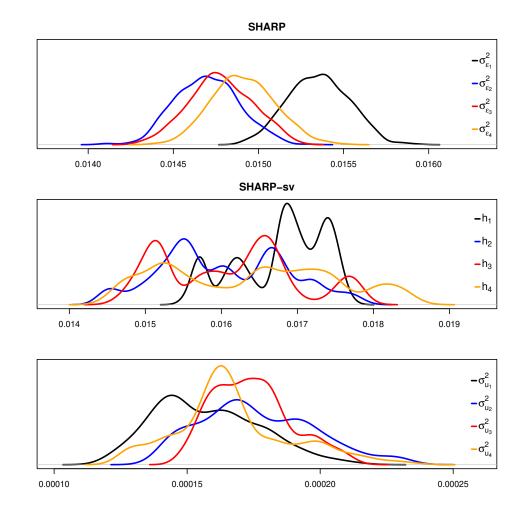


Figure 3.2: Marginal posterior densities of  $\hat{\sigma}_{\epsilon_j}^2$  in SHARP as well as  $\hat{h}_{t,j}$  and  $\hat{\sigma}_{u_j}^2$  in SHARP-sv obtained by estimating the models on SPY-ETF realised variance in the period January 1st, 2000 to December 31st, 2016.

We also plot the marginal posterior densities of the variance of the residuals in each state equation (3.7) and (3.10) in Figure (3.2). The density plots of  $\hat{h}_{t,j}$  have slightly higher means and variances than that of the  $\hat{\sigma}_{\epsilon_j}^2$ . Also, it is evident from the plots that the SHARP-sv model depicts the presence of heteroskedasticity in the variances of the coefficients' state equations. In both models, we observe that the mean of the variance of the state equation of the intercept coefficient ( $\beta_1$ ) is higher than that of

the other coefficients  $(\beta_2, \beta_3, \beta_4)$ .

Lastly, we use the Bayes factor to assess if sufficient evidence in the SPY-ETF logRV data series supports the introduced features represented by the statistical models: SHARP and SHARP-sv. Firstly, we compute the Bayes factor to compare the two hypotheses " $H_1$ : model's coefficients ( $\beta$ 's) are constant", represented by the HARL model, and " $H_2$ : model's coefficients ( $\beta$ 's) are time-varying", represented by the SHARP model. Secondly, we compare the hypothesis " $H_3$ : model's coefficients are time-varying with SV", represented by the SHARP-sv model, and " $H_2$ ". Assigning equal priors to the two models under consideration, the Bayes factor can then be computed as  $BF_{i,j} = \frac{P(Data|H_i)}{P(Data|H_j)}$ , where  $P(Data|H_k)$  is the probability of observing the data under the model representing hypothesis  $H_k$ . A higher value of the Bayes factor means greater evidence for  $H_i$  against  $H_j$ . We follow Berger and Pericchi (1996)'s interpretation of twice the natural logarithm of  $BF_{i,j}$  in Table (3.3).

According to the reported results, we conclude that there is evidence favouring timevarying parameters with SV. However, the evidence of  $H_2$  against  $H_1$  is stronger than  $H_3$  against  $H_2$ . Also, the evidence of each of the time-varying and SV features is stronger in the second sub-sample period, which coincides with the GFC period. Hence, all the preliminary findings are promising for both the SHARP and SHARP-sv models, where, in line with Chen et al. (2018) and Buccheri and Corsi (2021), it shows supporting evidence that the coefficients of the HARL model are time-varying with SV.

$2 \log(\mathbf{B_{2,1}})$	Decision	$2\log(\mathbf{B_{3,2}})$	Decision
	January 2004 to D	ecember 2006	
8.1258 > 6	Strong evidence for H2 against H1	2.1789 > 2	Positive evidence for H3 against H2 $$
	January 2007 to D	ecember 2010	
12.7844 > 10	Very Strong evidence for H2 against H1	2.9059 > 2	Positive evidence for H3 against H2
	January 2011 to D	ecember 2016	
6.5801 > 6	Strong evidence for H2 against H1	2.7126 > 2	Positive evidence for H3 against H2 $$

Table 3.3: Bayes Factor obtained by estimating the HARL, SHARP, and SHARP-sv models on SPY-ETF realised variance in the period January 1st, 2000 to December 31st, 2016

#### 3.4.3 Out-of-sample Comparative analyses

This section presents an out-of-sample analysis of the introduced SHARP and SHARP-sv models and a set of comparable models outlined below in forecasting the daily *RV* of SPY-ETF and twenty NYSE individual stocks. We report the outof-sample forecasts using the rolling window approach with an estimation window of 1000 daily observations (approximately four years). Hence, the in-sample period is 1000 observations coinciding roughly from January 2000 to December 2003. The out-of-sample period starts from January 2004 to December 2016. We perform the out-of-sample (OOS) analyses over three sub-periods relative to the GFC period. The first sub-sample spans from January 2004 to December 2006; the second sub-sample covers the GFC period, the most volatile period in our data sample, starting from January 2007 to December 2010; the last sub-sample is from January 2011 until December 2016.

There is extensive evidence in the literature showing the HARL, i.e. modelling the logRV, yields better RV forecasts than the HAR model estimated using the RV series. The HARL and all the other models estimated on the log series are less

affected by the huge peaks of RV. Therefore, we consider the HARL as our benchmark model. This paper's set of competing methods consists of 1) the HARL model (Corsi, 2009), eq. (3.4), is estimated by OLS. 2) HARLQ (inspired from Bollerslev, Patton, and Quaedvlieg (2016) and Buccheri and Corsi (2021)) is a HARL model with timevarying daily coefficient where  $\beta_{2,t} = \beta_2 + \gamma \frac{\sqrt{RQ}_{t-1}}{RV_{t-1}}$ , estimated by OLS. 3) HARSL (Bekierman and Manner, 2018), which is a HARL model with a time-varying daily coefficient that follows an AR(1) process, estimated by maximum likelihood using the Kalman filter. 4) TVCHAR (Chen et al., 2018) is a HARL model with time-varying coefficients of unspecified functional form, estimated by the local linear method. 5) SHARK (Buccheri and Corsi, 2021), a HARL model with heteroskedastic disturbances and score-driven parameters, estimated by maximum likelihood using the Kalman filter. 6) SHARP model, proposed in this paper, is a HARL model with time-varying coefficients that follow an AR(1) process, estimated by particle Gibbs sampling. 7) SHARP-sv model is similar to the SHARP model with the additional feature of SV in the process of the coefficients.

The selected competing set allows us to evaluate the proposed SHARP and SHARP-sv models and their recent rivals to highlight new comparisons across these models. For example, i) HARL vs HARLQ: knowing that the measurement error of *logRV* series still exists, we explore whether accounting for this measurement error would improve the forecasts by HARL as evident in the case of HARQ vs HAR in Bollerslev, Patton, and Quaedvlieg (2016). ii) HARSL vs HARLQ: outlines whether modelling the daily coefficient as an AR(1) process is better in forecasting than the more restrictive yet straightforward approach of the HARLQ model. iii) TVCHAR vs SHARK vs SHARP: demonstrates which coefficients' time-varying specification yields better forecasts. Finally, iv) SHARP vs SHARP-sv: reveals whether the additional SV feature in the AR process of the coefficients improves the forecasts.

In Table 3.4, we report the out-of-sample average of the estimated coefficients used at each forecasting point by each model. The logRV series has lower measurement errors, implying a minor impact on the magnitude of the coefficients than when modelling the RV series. The average values of the coefficients across the models are within a close range. The variances of the coefficients are the lowest for the HARL model and the highest for the SHARK model. As we pointed out in the previous subsection 3.4.2, the daily and monthly coefficients' estimates increase during the GFC period. In contrast, the weekly coefficient estimate decreases. Also, the coefficients' variances increase during that period.

Next, we compare the models' accuracy in forecasting the one-day-ahead RV. We use two popular unit-free loss functions (see, Wang, Wu, and Xu (2015) and Zhang, Ma, and Liao (2020), among others): (i) the heteroskedasticity-adjusted version of the mean squared error,  $HMSE = \frac{1}{n-k}\sum_{t=k+1}^{n}(1-\frac{\hat{y}_t}{y_t})^2$ , (iii) heteroskedasticity-adjusted version of the mean absolute error,  $HMAE = \frac{1}{n-k}\sum_{t=k+1}^{n}|1-\frac{\hat{y}_t}{y_t}|$ . Both HMSE and HMAE quantify the relative forecast error with respect to the true objective value, RV.

			January 20	004 to Decemb	oer 2016		
	HARL	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-s
$\bar{\hat{\beta}}_1$	-0.0674	-0.0672	-0.0518	-0.0699	-0.0808	-0.0695	-0.0693
_	(0.0691)	(0.0701)	(0.0393)	(0.0722)	(0.0858)	(0.0721)	(0.0733)
$\bar{\hat{\beta}}_2$	0.4225	0.4257	0.4332	0.4225	0.5404	0.4222	0.4220
_	(0.1010)	(0.0988)	(0.1114)	(0.1021)	(0.1191)	(0.1024)	(0.1033)
$\bar{\hat{\beta}}_3$	0.3633	0.3579	0.3500	0.3632	0.3043	0.3724	0.3731
	(0.0886)	(0.0852)	(0.0830)	(0.0896)	(0.0877)	(0.0938)	(0.0948)
$\bar{\hat{\beta}}_4$	0.1393	0.1410	0.1572	0.1377	0.0572	0.1308	0.1311
	(0.0386)	(0.0394)	(0.0427)	(0.0391)	(0.0731)	(0.0346)	(0.0380)
			January 20	004 to Decemb	per 2006		
	HARL	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-s
$\overline{\hat{\beta}}_1$	-0.0255	-0.0247	-0.0278	-0.0271	-0.1085	-0.0278	-0.0279
	(0.0205)	(0.0217)	(0.0222)	(0.0224)	(0.0680)	(0.0227)	(0.0270)
$\bar{\hat{\beta}}_2$	0.2703	0.2849	0.2891	0.2678	0.4064	0.2668	0.2670
	(0.0260)	(0.0450)	(0.0929)	(0.0246)	(0.0527)	(0.0246)	(0.0296)
$\bar{\hat{\beta}}_3$	0.4951	0.4794	0.4708	0.4978	0.3942	0.5171	0.5173
, 0	(0.0389)	(0.0539)	(0.0537)	(0.0364)	(0.0503)	(0.0342)	(0.0379)
$\bar{\hat{\beta}}_4$	0.1871	0.1887	0.1990	0.1858	0.1075	0.1678	0.1683
/ 1	(0.0136)	(0.0139)	(0.0326)	(0.0124)	(0.0364)	(0.0117)	(0.0200)
			January 20	007 to Decemi	oer 2010		
	HARL	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-s
$\bar{\hat{\beta}}_1$	-0.0668	-0.0675	-0.0486	-0.0697	-0.0277	-0.0680	-0.0673
_	(0.0708)	(0.0722)	(0.0416)	(0.0759)	(0.0698)	(0.0754)	(0.0769)
$\bar{\hat{\beta}}_2$	0.4343	0.4354	0.4735	0.4367	0.5447	0.4370	0.4363
	(0.0835)	(0.0840)	(0.0933)	(0.0847)	(0.1360)	(0.0845)	(0.0852)
$\overline{\hat{\beta}}_3$	0.3669	0.3641	0.3394	0.3650	0.3252	0.3697	0.3700
	(0.0446)	(0.0402)	(0.0516)	(0.0459)	(0.0754)	(0.0471)	(0.0494)
$\bar{\hat{\beta}}_4$	0.1241	0.1238	0.1329	0.1211	0.0128	0.1194	0.1199
	(0.0286)	(0.0288)	(0.0221)	(0.0281)	(0.0979)	(0.0275)	(0.0322)
			January 20	11 to Decemb	per 2016		
	HARL	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-s
$\overline{\hat{\beta}}_1$	-0.0882	-0.0878	-0.0656	-0.0909	-0.1027	-0.0911	-0.0911
	(0.0739)	(0.0746)	(0.0382)	(0.0762)	(0.0874)	(0.0764)	(0.0774)
$\bar{\hat{\beta}}_2$	0.4894	0.4882	0.4769	0.4888	0.6033	0.4888	0.4885
1- 4	(0.0332)	(0.0408)	(0.0597)	(0.0336)	(0.0641)	(0.0338)	(0.0377)
$\bar{\hat{\beta}}_3$	0.2963	0.2942	0.2978	0.2961	0.2462	0.3030	0.3044
~3	(0.0428)	(0.0428)	(0.0412)	(0.0432)	(0.0628)	(0.0432)	(0.0460)
-		(0.0420) 0.1292	(0.0412) 0.1529	0.1251	0.0623	0.1202	0.1204
$\bar{\hat{\beta}}_4$	0.1260						

Table 3.4: Average and (standard deviation) of coefficients' out-of-sample estimates used at each forecasting point obtained by estimating the HAR, HARL, HARQ, HARLQ, HARSL, TVCHAR, SHARK, SHARP, and SHARP-sv models on SPY-ETF realised variance over a rolling window of 1000 observations.

The relative loss measure of a model, M, over the benchmark model, HARL, using a loss function, LF (i.e. HMSE or HMAE), is defined as:

$$RL_M := \frac{LF_M}{LF_{HAR}} \tag{3.13}$$

 $RL_M < 1$  indicates that Model M outperforms the benchmark model. We report the  $RL_M$  for each model, M, using SPY-ETF data in Table (3.5) and their average over the twenty NYSE individual stocks in Table (3.6). We also present the box plot of the  $RL_M$  of each model, M, using the twenty NYSE individual stocks data in Figure (3.3).

To test the significance of the forecasting gains of a model, we employ the Model Confidence Set (MCS) developed by Hansen, Lunde, and Nason (2011). We test the null hypothesis that all the models are equally good against the alternative that there is a smaller subset of superior models. We follow Hansen, Lunde, and Nason (2011) and choose 0.1 as the critical p-value based on the range statistics. Hence, models with p-values below the significance level,  $\alpha = 0.1$ , are excluded from the superior subset, denoted by  $\hat{M}_{90\%}$ . Using the SPY-ETF RV data, in Table (3.5), an asterisk indicates the model is included in the  $\hat{M}_{90\%}$ . Table (3.6) reports the number of individual financial stocks (out of 20) where a model is included in the  $\hat{M}_{90\%}$ .

		J	anuary 200	4 to Decembe	er 2006		
	HARL	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv
HMSE	1	1.0134	$0.8994^{*}$	0.9547	0.8918*	0.8608*	$0.8520^{*}$
HMAE	1	1.0044	$0.9468^{*}$	0.9777	$0.9399^{*}$	$0.9403^{*}$	$0.9349^{*}$
		J	anuary 200	7 to Decembe	er 2010		
	HARL	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv
HMSE	1	1.0060	0.9867	0.8832	1.1296	$0.7454^{*}$	$0.7062^{*}$
HMAE	1	1.0005	0.9905	0.9428	1.0444	$0.8672^{*}$	$0.8638^{*}$
		J	anuary 201	1 to Decembe	er 2016		
	HARL	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv
HMSE	1	1.0072	0.9525	0.9140	0.9176	0.8827	$0.8236^{*}$
HMAE	1	1.0004	0.9768	0.9552	0.9636	0.9397	$0.9187^{*}$

Table 3.5: Out-of-sample relative loss measure (3.13) of HARL, HARLQ, HARSL, TVCHAR, SHARK, SHARP, and SHARP-sv models obtained by estimating the models on SPY-ETF realised variance over a rolling window of 1000 observations in the period January 1st, 2000 to December 31st, 2016.

* indicates that the model is included in  $\hat{M}_{90\%}$ .

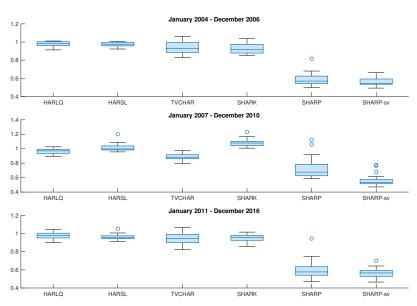
Table 3.6: Average of the out-of-sample relative loss measure (3.13) of HARL, HARLQ, HARSL, TVCHAR, SHARK, SHARP, and SHARP-sv models obtained by estimating the models on twenty individual NYSE stocks' realised variance over a rolling window of 1000 observations in the period January 1st, 2000 to December 31st, 2016.

January 2004 to December 2006							
	HARL	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv
HMSE HMAE	$1^{(0)}$ $1^{(0)}$	$\begin{array}{c} 0.9779^{(0)} \\ 0.9941^{(0)} \end{array}$	$\begin{array}{c} 0.9732^{(0)} \\ 0.9871^{(0)} \end{array}$	$0.9351^{(0)}$ $0.9673^{(0)}$	$\begin{array}{c} 0.9252^{(0)} \\ 0.9640^{(0)} \end{array}$	$\begin{array}{c} 0.5923^{(20)} \\ 0.7935^{(20)} \end{array}$	$\begin{array}{c} 0.5631^{(20)} \\ 0.7812^{(19)} \end{array}$
January 2007 to December 2010							
	HARL	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv
HMSE HMAE	$     1^{(0)}     1^{(0)} $	$\begin{array}{c} 0.9648^{(0)} \\ 0.9877^{(0)} \end{array}$	$\frac{1.0167^{(1)}}{0.9952^{(0)}}$	$\begin{array}{c} 0.8873^{(5)} \\ 0.9459^{(3)} \end{array}$	$\frac{1.0832^{(0)}}{1.0356^{(0)}}$	$\begin{array}{c} 0.7302^{(18)} \\ 0.8478^{(19)} \end{array}$	$\begin{array}{c} 0.5622^{(20)} \\ 0.7915^{(20)} \end{array}$
			January 201	11 to Decemb	er 2016		
	HARL	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv
HMSE HMAE	$1^{(0)}$ $1^{(0)}$	$\begin{array}{c} 0.9748^{(0)} \\ 0.9946^{(0)} \end{array}$	$\begin{array}{c} 0.9623^{(1)} \\ 0.9807^{(1)} \end{array}$	$\begin{array}{c} 0.9452^{(0)} \\ 0.9748^{(0)} \end{array}$	$\begin{array}{c} 0.9523^{(1)} \\ 0.9807^{(1)} \end{array}$	$\begin{array}{c} 0.6046^{(20)} \\ 0.7978^{(20)} \end{array}$	$\begin{array}{c} 0.5652^{(20)} \\ 0.7788^{(20)} \end{array}$

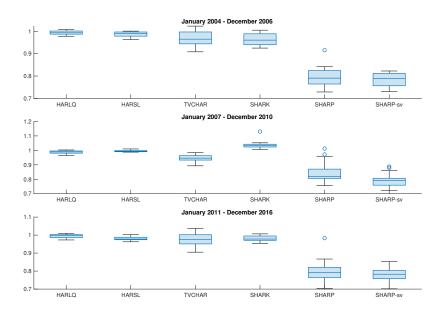
We show in parenthesis the number of times each model is included in  $\hat{M}_{90\%}$ 

Generally, the results obtained using the *RV* of SPY-ETF and individual stocks are consistent. Correcting for the measurement error in the HARL model relatively improves the forecasts, on average, since the average RL of the HARLQ is less than one. However, the HARLQ does not consistently outperform the HARL model across the twenty individual stocks and the SPY-ETF. The HARSL model, on the other hand, yields moderate forecast improvements compared to the HARLQ, meaning a general time-varying specification such as the AR process for the daily coefficient is better than the restrictive correction in the daily coefficient based on the measurement error variance. It indicates that other factors beyond the measurement error lead to changes in the daily coefficient. Hence, an AR process allows for such flexibility.

Further, the TVCHAR model is more promising than the HARSL model. The former has all its coefficients time-varying, not just the daily coefficient. However, as pointed out by their respective authors, both the TVCHAR and HARSL models do not consistently outperform the HARL model for the one-day ahead forecasts. The SHARK model is better than the TVCHAR model in the first and second subsample periods. Still, the forecasting accuracy of the SHARK model deteriorates while that of TVCHAR improves during the GFC period. Figure 3.3: Box plot of the out-of-sample relative loss measure (3.13) of HARL, HARLQ, HARSL, TVCHAR, SHARK, SHARP, and SHARP-sv models obtained by estimating the models on twenty individual NYSE stocks' realised variance over a rolling window of 1000 observations in the period January 1st, 2000 to December 31st, 2016.



## (a) HMSE



#### (b) HMAE

For most individual stocks, some form of time-varying specification of the coefficients generally improves the model's forecasting. However, only the SHARP-sv model's forecasting performance is in all cases superior to the HARL model. The SHARP and SHARP-sv models have the lowest RL using the SPY-ETF data and average RL using the twenty individual stocks data. Adding the SV feature to the AR process of the coefficients yields a further moderate improvement compared to the forecasts by the SHARP model. Hence, we find that the general specification of the time-varying coefficients leads to the most significant improvement in the forecasts of RV.

The findings based on the model confidence set reveal that our new extensions, SHARP and SHARP-sv, are always included in the confidence set. It excludes the remaining models with few exceptions. For example, between January 2004 and December 2006, the HARSL and SHARK models are included in the confidence set using the SPY-ETF data but not the individual stocks data. Also, the TVCHAR model is in the superior set during the GFC period for a few individual stocks.

Lastly, we use the unconditional predictive ability (uCPA) test by Giacomini and White (2006) to evaluate the out-of-sample predictions produced by the models. It builds on Diebold and Mariano (1995) test and provides a framework for an unconditional forecast evaluation criterion robust to misspecified forecasting models. Our forecast evaluation involves seven models with various specifications and estimation methods. Giacomini and White (2006) test does not impose restrictions on the estimation methods used and allows the models to be nested or non-nested. In Tables (3.7) and (3.9), we report the results for pairwise tests of uncodnitional predictive ability over three OOS periods with significance level of  $\alpha = 0.05$ . Based on the results, we rank the models from best to worse in Tables (3.8) and (3.10).

Table 3.7: Results of the uCPA test for out-of-sample forecasting performance between the models with the selected loss function (HMSE) obtained by estimating the models on SPY-ETF realised variance over a rolling window of 1000 observations in the period January 1st, 2000 to December 31st, 2016. The model that is significantly better than the other is stated inside the table unless both models are equally good.

	January 2004 to December 2006						
vs	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv	
HARL HARLQ HARSL TVCHAR SHARK SHARP	Equally good	HARSL HARSL	TVCHAR TVCHAR HARSL	SHARK SHARK Equally good SHARK	SHARP SHARP Equally good SHARP Equally good	SHARP-sv SHARP-sv Equally good SHARP-sv Equally good Equally good	
		January	2007 to Decen	nber 2010			
vs	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv	
HARL HARLQ HARSL TVCHAR SHARK SHARP	Equally good	Equally good Equally good	TVCHAR TVCHAR TVCHAR	HARL HARLQ HARSL TVCHAR	SHARP SHARP SHARP SHARP SHARP	SHARP-sv SHARP-sv SHARP-sv SHARP-sv SHARP-sv SHARP-sv	
		January	2011 to Decen	nber 2016			
vs	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv	
HARL HARLQ HARSL TVCHAR SHARK SHARP	Equally good	HARSL HARSL	TVCHAR TVCHAR TVCHAR	SHARK SHARK Equally good Equally good	SHARP SHARP SHARP Equally good Equally good	SHARP-sv SHARP-sv SHARP-sv SHARP-sv SHARP-sv SHARP-sv	

Table 3.8: Ranking of models from best to worse in every subsample period according to the uCPA results in Table (3.7). Models appearing on the same row in a given period of time are equally good.

January 2004 to December 2006	January 2007 to December 2010	January 2011 to December 2016
SHARP-sv, SHARP, SHARK, HARSL TVCHAR HARL, HARLQ	SHARP-sv SHARP TVCHAR HARL, HARLQ, HARSL SHARK	SHARP-sv SHARP, TVCHAR, SHARK SHARK, HARSL HARL, HARLQ

Table 3.9: Results of the uCPA test for out-of-sample forecasting performance between the models with the selected loss function (HMAE) obtained by estimating the models on SPY-ETF realised variance over a rolling window of 1000 observations in the period January 1st, 2000 to December 31st, 2016. The model that is significantly better than the other is stated inside the table unless both models are equally good.

	January 2004 to December 2006						
vs	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv	
HARL HARLQ HARSL TVCHAR SHARK SHARP	Equally good	HARSL HARSL	TVCHAR TVCHAR HARSL	SHARK SHARK Equally good SHARK	SHARP SHARP Equally good SHARP Equally good	SHARP-sv SHARP-sv Equally good SHARP-sv Equally good Equally good	
		January	2007 to Decen	nber 2010			
vs	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv	
HARL HARLQ HARSL TVCHAR SHARK SHARP	Equally good	Equally good Equally good	TVCHAR TVCHAR TVCHAR	HARL HARLQ HARSL TVCHAR	SHARP SHARP SHARP SHARP SHARP	SHARP-sv SHARP-sv SHARP-sv SHARP-sv SHARP-sv Equally good	
		January	2011 to Decen	nber 2016			
vs	HARLQ	HARSL	TVCHAR	SHARK	SHARP	SHARP-sv	
HARL HARLQ HARSL TVCHAR SHARK SHARP	Equally good	HARSL HARSL	TVCHAR TVCHAR TVCHAR	SHARK SHARK SHARK Equally good	SHARP SHARP SHARP SHARP SHARP	SHARP-sv SHARP-sv SHARP-sv SHARP-sv SHARP-sv SHARP-sv	

Table 3.10: Ranking of models from best to worse in every subsample period according to the CPA results in Table 3.9. Models appearing on the same row in a given period of time are equally good.

January 2004 to December 2006	January 2007 to December 2010	January 2011 to December 2016
SHARP-sv, SHARP, SHARK, HARSL	SHARP-sv, SHARP	SHARP-sv
TVCHAR	TVCHAR	SHARP
HARL, HARLQ	HARSL, HARLQ	TVCHAR, SHARK
	HARLQ, HARL	HARSL
	SHARK	HARLQ, HARL

The SHARP-sv consistently ranks among the best models. The SHARP model ranks as good as the SHARP-sv or the second best. During the second and third subsample periods, the SHARP-sv is first exclusively. While the SHARK model is among the top in the first and third sub-sample periods, its performance deteriorates during the GFC period. The TVCHAR ranks better than the SHARK except for the first subsample period and is always superior to the HARL and HARLQ models. Also, the HARSL ranks better than the HARL and HARLQ models except during the GFC period. Lastly, the HARLQ and HARL models are equally good.

### 3.5 Conclusion

We propose a dynamic state-space model that defines time-varying parameters as an AR process with heteroskedastic innovations: the log-transformation of the conditional variance of those innovations is an AR process. Thereby, we allow for time-varying heteroscedasticity in the measurement equation and the state process of coefficients. We follow Creal and Tsay (2015) in using particle filtering as a computationally efficient approach (Andrieu, Doucet, and Holenstein, 2010). In building upon three recent analyses of the dynamic features of the HARL (HAR with the log of RV) model coefficients (Bekierman and Manner, 2018; Buccheri and Corsi, 2021; Chen et al., 2018), the SHARP and SHARP-sv models incorporate additional innovations with a feasible estimation method.

With our empirical study, we apply the SHARP and SHARP-sv models to forecast the volatility of financial stock returns. Comparisons are made between the out-of-sample forecasting performance of the models, across three sub-sample periods, including the financial crisis of 2008. For a robustness check, we evaluate forecasts using two loss functions. We also use the Model Confidence Set by Hansen, Lunde, and Nason (2011) and conditional predictive ability test by Giacomini and White (2006) to assess the

significance of our results.

First: While the literature shows that accounting for the measurement error variance improves the forecast by the HAR model, its failure to do likewise by the HARL may be attributed to reduced measurement errors when using the logarithmic form. Second: Permitting the coefficients to be time-varying, with time-varying conditional variance, improves the forecasts of RV. By their overall improved performance, the proposed SHARP and SHARP-sv models significantly outperform other extensions of the HARL model in forecasting the realised volatility of financial stocks.

## Appendices

### **3.A** Derivations of the first and second moments

For conciseness we will assume for what follows that k = 1 so the transition equations are defined as:

$$\beta_t = \alpha + \rho \beta_{t-1} + \varepsilon_t, \quad \varepsilon_t | h_t \sim \mathcal{N}(0, h_t) \tag{3.14}$$

and

$$\ln h_t = \gamma + \delta \ln h_{t-1} + u_t, \quad u_t \sim \mathcal{N}(0, \sigma_u^2)$$
(3.15)

With  $\delta < 1$ , one can show that:

$$\mu := E(\ln h_t) = \frac{\gamma}{1 - \delta^2} \quad \text{and} \quad \omega^2 := var(\ln h_t) = \frac{\sigma_u^2}{1 - \delta^2} \tag{3.16}$$

$$E(h_t) = \exp(\mu) \exp(\omega^2/2) \quad \text{and} \quad var(h_t) = \exp(2\mu) \left(\exp(2\omega^2) - \exp(\omega^2)\right) \quad (3.17)$$

With  $\rho < 1$ , one can show that:

$$E(\beta_t) = \frac{\alpha}{1 - \rho^2} \tag{3.18}$$

The conditional second moment (variance) of  $\beta_t$  could change over time:

$$var(\beta_t|h_t) = \sum_{i=0}^{\infty} \rho^{2i} h_{t-i}$$
$$= \exp(\mu) \sum_{i=0}^{\infty} \rho^{2i} \exp(S_{t-i})$$

Where  $S_t = \sum_{i=0}^{\infty} \delta^i u_{t-i} \sim N(0, \omega^2)$ . Note that, since  $|\delta| < 1$ ,  $S_t$  converges in mean square to some finite number as  $T \to \infty$  provided that  $\sum_{i=0}^{\infty} (\delta^i)^2 < \infty$ . Similarly, since  $|\rho^2| < 1$ ,  $\sum_{i=0}^{T} \rho^{2i} \exp(S_{t-i})$  converges in mean square to some finite number as  $T \to \infty$  provided that  $\sum_{i=0}^{\infty} (\rho^{2i})^2 < \infty$ .

Finally, one can show that the unconditional variance of  $\beta_t$  is constant:

$$var(\beta_t) = E(var(\beta_t|h_t)) + var(E(\beta_t|h_t))$$
$$= \frac{\exp(\mu + \omega^2/2)}{1 - \rho^2}$$

# 3.B Posterior Derivation

Model:

$$y_t = x'_t \beta_t + v_t; \quad v_t \sim \mathcal{N}(0, \sigma_v^2), \quad t = 1, \dots, n$$
$$\beta_{tj} = \alpha_j + \rho_j \beta_{t-1,j} + \varepsilon_{tj}; \quad \varepsilon_{tj} | h_{tj} \sim \mathcal{N}(0, h_{tj}), \quad j = 1, \dots, k, \ t = 1, \dots, n$$
$$\log h_{tj} = \gamma_j + \delta_j \log h_{t-1,j} + u_{tj}; \quad u_{tj} \sim \mathcal{N}(0, \sigma_{uj}^2), \quad j = 1, \dots, k, \ t = 1, \dots, n$$

$$\begin{array}{lll} \mathbf{Let} \quad \lambda_t = [\beta'_t, logh'_t]' \\ \mathbf{where} \quad \beta_t = [\beta_{t1}, \dots, \beta_{tk}] \quad \mathbf{and} \quad logh_t = [logh_{t1}, \dots, logh_{tk}] \\ \mathbf{Let} \quad \theta = \{\theta_1, \dots, \theta_k\}, \quad \mathbf{where} \quad \theta_j = \{\alpha_j, \rho_j, \gamma_j, \delta_j, \sigma_{uj}\}, \quad j = 1, \dots, k, \end{array}$$

Prior

$$\alpha_j, \gamma_j \sim \mathcal{N}(0, 1), \quad j = 1, \dots, k,$$
  

$$\rho_j, \delta_j \sim \mathcal{N}(0.5, 1) \mathbb{I}_{\rho_j \in (0, 1)} \mathbb{I}_{\delta_j \in (0, 1)}, \quad j = 1, \dots, k,$$
  

$$\sigma_{uj}^2 \sim \Gamma(6.5, 0.005), \quad j = 1, \dots, k$$
  

$$\sigma_v^2 \sim \Gamma(6.5, 0.005)$$

Equivalently:

$$p(\alpha_j) \propto exp(-\frac{n}{2}\alpha_j^2), \quad j = 1, \dots, k$$

$$p(\gamma_j) \propto exp(-\frac{n}{2}\gamma_j^2), \quad j = 1, \dots, k$$

$$p(\rho_j) \propto exp(-\frac{n}{2}(\rho_j - 0.5)^2) \mathbb{I}_{\rho_j \in (0,1)}, \quad j = 1, \dots, k$$

$$p(\delta_j) \propto exp(-\frac{n}{2}(\delta_j - 0.5)^2) \mathbb{I}_{\delta_j \in (0,1)}, \quad j = 1, \dots, k$$

$$p(\sigma_{uj}) \propto \sigma_{uj}^{-(\underline{n}+1)} exp\left(-\frac{q}{2\sigma_{uj}^2}\right), \quad j = 1, \dots, k, \quad \underline{n} = 10 \quad \text{and} \quad \underline{q} = 0.01$$

$$p(\sigma_v) \propto \sigma_{uj}^{-(\underline{n}+1)} exp\left(-\frac{q}{2\sigma_{uj}^2}\right), \quad \underline{n} = 10 \quad \text{and} \quad \underline{q} = 0.01$$

### Posterior

$$P\left(\theta, \sigma_{v}, \{\lambda\}_{t=1}^{T} | Y\right) \propto (\sigma_{v}^{-n}) exp\left(-\frac{1}{2\sigma_{v}^{2}} \sum_{t=1}^{n} (y_{t} - x_{t}'\beta_{t})^{2}\right)$$

$$\left(\prod_{i=1}^{k} \prod_{t=1}^{n} h_{it}^{-1}\right) exp\left(-\sum_{i=1}^{k} \sum_{t=1}^{n} \frac{1}{2h_{it}^{2}} (\beta_{it} - \alpha_{i} - \rho_{i}\beta_{i,t-1})^{2}\right)$$

$$\left(\prod_{i=1}^{k} \sigma_{u_{i}}^{-n}\right) exp\left(-\sum_{i=1}^{k} \frac{1}{2\sigma_{u_{i}}^{2}} \sum_{t=1}^{n} (logh_{it} - \gamma_{it} - \delta_{it}logh_{i,t-1})^{2}\right)$$

$$\prod_{j=1}^{k} p(\sigma_{u_{j}}) \prod_{i=1}^{k} p(\alpha_{i}) \prod_{i=1}^{k} p(\gamma_{j}) \prod_{i=1}^{k} p(\rho_{i}) \prod_{i=1}^{k} p(\delta_{i})$$

1. For  $P(\theta|\{\lambda\}_{t=1}^{T}, Y)$ , for i = 1, ..., k: 2. For  $\gamma_i \left| \{\{\theta \setminus \{\gamma_i\}\}, \sigma_v, \{\lambda\}_{t=1}^{T}, Y\}$  we use mixed estimator:

$$\begin{aligned} \det Z &= \begin{bmatrix} logh_{i,2} - \delta_i logh_{i,1} \\ \vdots \\ logh_{i,T} - \delta_i logh_{i,T-1} \end{bmatrix} \text{ and } u_i = \begin{bmatrix} u_{i,2} \\ \vdots \\ u_{i,T} \end{bmatrix} \\ \\ \text{then, } \begin{bmatrix} Z \\ 0 \end{bmatrix} &= \begin{bmatrix} 1_{(n\times 1)} \\ 1 \end{bmatrix} \gamma_i + \begin{bmatrix} u_i \\ \eta \end{bmatrix}; \begin{bmatrix} u_i \\ \eta \end{bmatrix} \sim N(0, \begin{bmatrix} \sigma_i I_n \\ 1 \end{bmatrix}) \\ \\ \text{where } n &= T-1 \\ \\ \implies \gamma_i \middle| \{\{\theta \setminus \{\gamma_i\}\}, \{\lambda\}_{t=1}^T, Y\} \sim N(\frac{(logh_{i,T} - logh_{i,1})}{1 + \sigma_i^2}, \frac{\sigma_i^2}{1 + \sigma_i^2}) \\ \\ 3. \text{ For } \delta_i \middle| \{\{\theta \setminus \{\delta_i\}\}, \sigma_v, \{\lambda\}_{t=1}^T, Y\}, \text{ we use mixed estimator:} \end{aligned}$$

$$\det Z = \begin{bmatrix} logh_{i,2} - \gamma_i \\ \vdots \\ logh_{i,T} - \gamma_i \end{bmatrix}, \ \kappa = \begin{bmatrix} logh_{i,1} \\ \vdots \\ logh_{i,T} \end{bmatrix} \text{ and } u_i = \begin{bmatrix} u_{i,2} \\ \vdots \\ u_{i,T} \end{bmatrix}$$
$$then, \ \begin{bmatrix} Z \\ 0.5 \end{bmatrix} = \begin{bmatrix} \kappa \\ 1 \end{bmatrix} \delta_i + \begin{bmatrix} u_i \\ \eta \end{bmatrix}; \ \begin{bmatrix} u_i \\ \eta \end{bmatrix} \sim N(0, \begin{bmatrix} \sigma_i I_n \\ 1 \end{bmatrix})$$
$$where \ n = T - 1$$

 $\xrightarrow{\sigma_{u_i}^2} \left| \{ \{ \theta \setminus \{ \sigma_i \} \}, \{ \lambda \}_{t=1}^t, Y \} \right|$ 5. For  $\alpha_i \left| \{ \{ \theta \setminus \{ \alpha_i \} \}, \{ \lambda \}_{t=1}^T, Y \} \right|$  we use mixed estimator:

$$\det Z = \begin{bmatrix} \beta_{i,2} - \rho_i \beta_{i,1} \\ \vdots \\ \beta_{i,T} - \rho_i \beta_{i,T-1} \end{bmatrix}, H_i = \begin{bmatrix} h_{i,1} \\ \vdots \\ h_{i,T} \end{bmatrix}, \text{ and } \epsilon_i = \begin{bmatrix} \epsilon_{i,2} \\ \vdots \\ \epsilon_{i,T} \end{bmatrix}$$

then, 
$$\begin{bmatrix} Z\\ 0 \end{bmatrix} = \begin{bmatrix} 1_{(n \times 1)}\\ 1 \end{bmatrix} \alpha_i + \begin{bmatrix} \epsilon_i\\ \eta \end{bmatrix}$$
;  $\begin{bmatrix} u_i\\ \eta \end{bmatrix} \sim N(0, \begin{bmatrix} H_i\\ & 1 \end{bmatrix})$   
where  $n = T - 1$ 

$$\Longrightarrow \alpha_i \bigg| \big\{ \{\theta \setminus \{\alpha_i\}\}, \{\lambda\}_{t=1}^T, Y \big\} \sim N\big(\frac{\sum_{t=2}^T \frac{Z_{i,t}}{h_{i,t}}}{\sum_{t=2}^T \frac{1}{h_{i,t}} + 1}, \frac{1}{\sum_{t=2}^T \frac{1}{h_{i,t}} + 1}\big)$$

6. For  $\rho_i \left| \left\{ \{\theta \setminus \{\rho_i\}\}, \sigma_v, \{\lambda\}_{t=1}^T, Y \right\} \text{ we use mixed estimator:} \right.$ 

$$\begin{aligned} &\text{let } Z = \begin{bmatrix} \beta_{i,2} - \alpha_i \\ \vdots \\ \beta_{i,T} - \alpha_i \end{bmatrix}, \, \kappa = \begin{bmatrix} \beta_{i,1} \\ \vdots \\ \beta_{i,T-1} \end{bmatrix}, \, H_i = \begin{bmatrix} h_{i,1} \\ \vdots \\ \vdots \\ h_{i,T} \end{bmatrix}, \\ &\text{h}_{i,T} \end{bmatrix}, \end{aligned}$$

$$\begin{aligned} &\text{and } \epsilon_i = \begin{bmatrix} \epsilon_{i,2} \\ \vdots \\ \epsilon_{i,T} \end{bmatrix} \\ &\text{then, } \begin{bmatrix} Z \\ 0.5 \end{bmatrix} = \begin{bmatrix} \kappa \\ 1 \end{bmatrix} \rho_i + \begin{bmatrix} \epsilon_i \\ \eta \end{bmatrix}; \begin{bmatrix} u_i \\ \eta \end{bmatrix} \sim N(0, \begin{bmatrix} H_i \\ 1 \end{bmatrix}) \\ &\text{where } n = T - 1 \end{aligned}$$

$$\implies \rho_i \left| \left\{ \{\theta \setminus \{\rho_i\}\}, \{\lambda\}_{t=1}^T, Y \right\} \sim N(\frac{\sum_{t=2}^T \frac{Z_{i,t}}{h_{i,t}} + 0.5}{\sum_{t=2}^T \frac{1}{h_{i,t}} + 1}, \frac{1}{\sum_{t=2}^T \frac{1}{h_{i,t}} + 1}) \right|$$

7.

$$P\left(\sigma_{v} \middle| \{\theta\}, \{\lambda\}_{t=1}^{T}, Y\right)$$

$$\propto$$

$$\sigma_{v}^{-n} exp\left(-\frac{1}{2\sigma_{v}^{2}} \sum_{t=1}^{n} (y_{t} - x_{t}'\beta_{t})^{2}\right) \quad \sigma_{v}^{-(\underline{n}+1)} exp\left(-\frac{q}{2\sigma_{v}^{2}}\right)$$

$$\propto$$

$$\sigma_{v}^{-(n+\underline{n}+1)} \quad exp\left(-\frac{1}{2\sigma_{i}^{2}} \left(\sum_{t=1}^{n} (y_{t} - x_{t}'\beta_{t})^{2} + \underline{q}\right)\right)$$

$$\Longrightarrow \frac{\sum_{t=1}^{n} (y_{t} - x_{t}'\beta_{t})^{2} + \underline{q}}{\sigma_{v}^{2}} \middle| \{\{\theta\}, \{\lambda\}_{t=1}^{T}, Y\} \sim \chi^{2}(n + \underline{n} + 3)$$

### **3.C** Particle filtering within MCMC

Particle filtering is a simulation-based algorithm that sequentially approximates continuous marginal distributions using discrete distributions. This is performed by using a set of support points called "particles" and probability masses; see Creal (2012) for a review. The PG sampler draws a single path of the latent or state variables from this discrete approximation. As the number of particles M goes to infinity, the PG sampler draws from the exact full conditional distribution. The advantage of the algorithm is that it allows for drawing paths of the state variables in large blocks.

As mentioned in Creal and Tsay (2015), the PG sampler is a standard Gibbs sampler but defined on an extended probability space where a particle filter generates all the random variables. Chopin and Singh (2015) analysed the theoretical properties of the PG sampler, and showed that the sampler is uniformly ergodic. Unlike the standard particle filter, the PG sampler involves a "conditional" resampling algorithm in the last step. Namely, for draws from the particle filter to be a valid Markov transition kernel on the extended probability space, the state variables drawn at the previous iteration must have a positive sampling probability (Andrieu, Doucet, and Holenstein, 2010). The conditional resampling step within the PG forces the pre-existing path to survive the particle filter's resampling steps. We use the conditional multinomial resampling algorithm from Andrieu, Doucet, and Holenstein (2010), although other resampling algorithms exist (see, for example, Fearnhead et al. (2010) and Chopin and Singh (2015)).

Suppose the posterior is  $p(\theta, \lambda_{1:T} | \mathbf{y}_{1:T})$  where  $\lambda_{1:T}$  denotes the latent variables whose prior can be described by  $p(\lambda_t | \lambda_{t-1}, \theta)$ . In the PG sampler, we can draw the structural parameters  $\theta | \lambda_{1:T}, \mathbf{y}_{1:T}$  as usual, from their posterior conditional distributions. This is important because, in this way, we can avoid mixture approximations or other Monte Carlo procedures that need considerable tuning and may not have good convergence properties. Suppose we have  $\lambda_{1:T}^{(1)}$  from the previous iteration. The particle filtering procedure consists of two phases, forward and backward filtering.

Phase I: Forward filtering (Andrieu, Doucet, and Holenstein, 2010).

For  $t = 1, \ldots, T$ 

- Draw a proposal  $\lambda_t^{(m)}$  from an importance density  $q(\lambda_t | \lambda_{t-1}^{(m)}, \theta), \quad m = 2, ..., M.$
- Compute the importance weights:

$$w_t^{(m)} = \frac{p(y_t | \lambda_t^{(m)}, \theta) p(\lambda_t^{(m)} | \lambda_{t-1}^{(m)}, \theta)}{q(\lambda_{it} | \lambda_{t-1}^{(m)}, \theta)}, \quad m = 1, ..., M.$$
(3.19)

• Normalise the weights:  $\tilde{w}_t^{(m)} = \frac{w_t^{(m)}}{\sum_{m=1}^M w_t^{(m)}}, \quad m = 1, ..., M.$ 

• Re-sample, conditionally, the particles  $\{\lambda_t^{(m)}, m = 1, ..., M\}$  with probabilities

 $\{\tilde{w}_t^{(m)}, m = 1, ..., M\}.$ 

In the original PG sampler, the particles are stored for t = 1, ..., T and a single trajectory is sampled using the probabilities from the last iteration. An improvement upon the original PG by drawing the path of the latent variables from the particle approximation is using the backwards sampling algorithm of Godsill, Doucet, and West (2004). In the forward pass, we store the normalised weights and particles then we draw a path of the latent variables as we detail below (the draws are from a discrete distribution).

Phase II: Backward filtering (Chopin and Singh, 2015; Godsill, Doucet, and West, 2004).

- At time t = T draw a particle  $\lambda_T^* = \lambda_T^{(m)}$ .
- Compute the backward weights:  $w_{t|T}^{(m)} \propto \tilde{w}_t^{(m)} p(\lambda_{t+1}^* | \lambda_t^{(m)}, \theta).$
- Normalise the weights:  $\tilde{w}_{t|T}^{(m)} = \frac{w_{t|T}^{(m)}}{\sum_{m=1}^{M} w_{t|T}^{(m)}}, m = 1, ..., M.$
- Draw a particle  $\lambda_{it}^* = \lambda_t^{(m)}$  with probability  $\tilde{w}_{t|T}^{(m)}$ .

Therefore,  $\lambda_{1:T}^* = \{\lambda_1^*, ..., \lambda_T^*\}$  is a draw from the full conditional distribution. When the state vector dimension is large, we can draw  $\lambda_{i,1:T}$ , conditional on all other paths  $\lambda_{-i,1:T}$  that are not path *i*. Therefore, we can draw from the full conditional distribution  $p(\lambda_{i,1:T}|\lambda_{-i,1:T}, \mathbf{y}_{1:T}, \theta)$ . The backwards step often results in dramatic improvements in computational efficiency and strictly dominates the original PG (Chopin and Singh, 2015). For example, Creal and Tsay (2015) find that M = 100particles is enough.

There remains the problem of selecting an importance density  $q(\lambda_t|\lambda_{t-1},\theta)$ . We use

an importance density implicitly defined by  $\lambda_{it} = a_i + \sum_{p=1}^{P} b_{i,p} \lambda_{i,t-1}^p + \Omega_i \xi_{it}$  where  $\xi_{it}$  follows a standard (zero location and unit scale) Student-*t* distribution with  $\nu = 5$  degrees of freedom. That is, we use polynomials in  $\lambda_{i,t-1}$  of order *P*. We select the parameters  $a_i, b_i$  and  $\Omega_i$  during the burn-in phase (using P = 1 and P = 2) so that the weights  $\{\tilde{w}_{it}^{(m)}, m = 1, ..., M\}$  and  $\{\tilde{w}_{t|T}^{(m)}, m = 1, ..., M\}$  are approximately not too far from a uniform distribution.

### 3.D HARSL estimation using Kalman Filter

The HARSL equation is:

$$lnRV_{t} = \beta_{1} + (\beta_{2} + \lambda_{t})lnRV_{t-1}^{(d)} + \beta_{3}lnRV_{t-1}^{(w)} + \beta_{4}lnRV_{t-1}^{(m)} + \epsilon_{t}$$

$$\lambda_{t+1} = \phi\lambda_{t} + \eta_{t}$$
(3.20)

It can be rewritten as:

$$lnRV_{t} = \beta_{1} + \beta_{2}lnRV_{t-1}^{(d)} + \beta_{3}lnRV_{t-1}^{(w)} + \beta_{4}lnRV_{t-1}^{(m)} + \lambda_{t}lnRV_{t-1}^{(d)} + \epsilon_{t}$$

$$\lambda_{t+1} = \phi\lambda_{t} + \eta_{t}$$
(3.21)

We use the fkf package on R to estimate the model. The above model can be rewritten as the general form of the transition and measurement equation for Kalman filter is given by:

$$y_t = c_t + Z_t \alpha_t + G_t . \epsilon_t$$

$$\alpha_{t+1} = d_t + T_t \alpha_t + H_t . \eta_t$$
(3.22)

where:

$$y_{t} = lnRV_{t}$$

$$Z_{t} = (1, lnRV_{t-1}^{(d)}, lnRV_{t-1}^{(w)}, lnRV_{t-1}^{(m)}, lnRV_{t-1}^{(d)})$$

$$c_{t} = 0$$

$$\alpha_{t} = (\beta_{1}, \beta_{2}, \beta_{3}, \beta_{4}, \lambda_{t})'$$

$$G_{t} = G = \sigma_{\epsilon}^{2}$$

$$d_{t} = 0$$

$$T_{t} = (1, 1, 1, 1, \eta)'$$

$$H_{t} = \begin{bmatrix} 0_{(4 \times 4)} \\ & \sigma_{\eta_{t}}^{2} \end{bmatrix}$$

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