

1 A geomagnetic polarity stratigraphy for the Middle and Upper Ordovician

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18 Abstract

19 Magnetostratigraphic studies of the Ordovician provide evidence for the nature of core-
20 mantle boundary interactions, and provide means for dating and correlation across differing
21 environmental regimes. We provide new magnetostratigraphic data from the Middle and
22 Upper Ordovician, compiling this into a polarity chronostratigraphic scale for the Dapingian
23 to Hirnantian interval. The new data are derived from the Backside Beck and Cheney
24 Longville sections in Britain, the Mójcza section in Poland and two cores from Poland and
25 Lithuania. The chronology is provided by existing biostratigraphy, principally based on
26 chitinozoans and conodonts for the Ordovician. Correlations between sections are supported
27 by carbon isotope stratigraphy linked to Baltic isotopic zonations, along with lithological and
28 local magnetic susceptibility correlations in Polish cores. The palaeomagnetic signal is
29 carried by both haematite and magnetite, with haematite dominating in red-coloured
30 lithologies (marls and limestones) and magnetite in non-red mudstones and limestones. A
31 positive reversal test (class C) in the Cheney Longville section and positive fold tests in the
32 Backside Beck section provide validation of the isolation of a primary palaeomagnetic signal.
33 Palaeomagnetic directions from cores were re-oriented using Kiaman-age and Brunhes
34 overprints. These new datasets in combination with existing Middle Ordovician data provides
35 a near-complete magnetic polarity chronostratigraphic scale through the Middle and Upper
36 Ordovician. Brief normal-polarity magnetozones extend well into the later parts of what has
37 been considered the Moyero Superchron, which started in the late Tremadocian. Reversal
38 frequencies for the mid and late Ordovician are 1.7 and 1.5 Myr⁻¹ respectively, although that
39 for the late Ordovician may be an underestimate.

40 Keywords: Magnetostratigraphy, Moyero, superchron, palaeomagnetic,
41 chronostratigraphy, reversal frequency

42

43 **1. Introduction**

44 A geomagnetic polarity timescale (GPTS) for the Ordovician has a number of important
45 uses, such as improved understanding of polarity reversal rates and their reflection of core-
46 mantle boundary interactions on geodynamo behaviour (Biggin et al., 2012; Hounslow et al.,
47 2018), as well as providing a means of high resolution stratigraphic correlation to aid
48 chronostratigraphy. Initial attempts to construct an Ordovician GPTS were in the 1960's and
49 1970's, based around sections from the Siberian Platform (Khramov and Rodionov, 1980; see
50 discussion in Pavlov and Gallet 2020). However, the methods used to demagnetise samples
51 were basic and poorly documented, and often in these studies are insufficiently related to
52 other stratigraphic proxies (Pavlov et al., 2008; Trench et al., 1991). In addition, the
53 widespread nature of Siberian re-magnetisation by intrusions or burial events likely
54 compromised results from these early studies (Rodionov and Gurevich, 2010). Studies using
55 modern demagnetisation techniques (Trench et al., 1991; Pavlov and Gallet, 2005, 2020;
56 Grappione et al., 2017; Rodionov, 2016; Kolesov, 2007) have improved our knowledge of
57 Lower and Middle Ordovician geomagnetic polarity. However, the polarity succession in the
58 Upper Ordovician is largely unknown, with only polarity bias-type data, based on a few
59 conventional palaeomagnetic studies (Trench et al., 1991; Algeo, 1996; Pavlov and Gallet,
60 2020). Existing Middle Ordovician magnetostratigraphic data is from sediment sections in
61 Sweden, Estonia, NE Russia and the USA (Farr et al., 1993; Torsvik et al., 1995; Gallet and
62 Pavlov, 1996, Pavlov and Gallet, 2005, Pavlov et al., 2012; Grappione et al., 2017).

63 This work provides the first magnetostratigraphic data from the later parts of the
64 Ordovician, using five sections in the Upper Ordovician. In Britain, this uses sections at
65 Backside Beck in the Howgill Fells (northern England) through the late Katian to the
66 Ordovician-Silurian boundary, along with a Katian section in Shropshire at Cheney Longville
67 (Fig. 1). From Poland, we describe data from the Sandbian to late Katian section at Mójca
68 from the Holy Cross Mountains. We also detail Middle to Upper Ordovician data from the
69 Grabowiec-6 core in Poland, and Core-A from Lithuania (Fig. 1). The data from Core-A
70 provides a template onto which other data can be related. These data are age constrained with
71 existing graptolite, chitinozoan and conodont biostratigraphy and some new carbon isotope
72 stratigraphy. We integrate these new data and existing datasets to yield a detailed polarity
73 timescale from the Middle Ordovician through to the Hirnantian- Rhuddanian boundary,
74 which joins that from the Llandovery of Hounslow et al. (inpress).

75 **2 Methods**

76 Palaeomagnetic samples from the sections were collected using hand samples, oriented
77 with a magnetic compass. Cubic specimens were cut from the hand samples using a circular
78 saw. Samples from Polish and Lithuanian cores were predominantly drill plugs.

79 Outcrops in the Ordovician at Backside Beck are spread discontinuously over ca. 300 m,
80 mostly in the bed of the stream, so the Ordovician samples were located on the ground using
81 a combination of GPS, tape- measurements and map work to produce the composite
82 stratigraphy, based around converting sample locations to stratigraphic position using

83 bedding dips. Bedding dips in the Ordovician section are 12-47° to the north. The precise
84 sample positions are in the supplementary information (SI) Figs. S2 and S3. Only
85 reconnaissance palaeomagnetic sampling of the Upper Ordovician was undertaken at
86 Backside Beck, with most palaeomagnetic effort focussed on the Silurian, described in
87 Hounslow et al. (inpress).

88 Sample measurements from Cheney Longville, Backside Beck, Core-A and Grabowiec-6
89 were performed at Lancaster, and those from the Mójca section in Warsaw. At Lancaster
90 measurements of Natural Remanent Magnetisation (NRM) were made using a CCL 3-axis
91 cryogenic magnetometer (noise level ~2 $\mu\text{A/m}$, four determinations per orthogonal axis),
92 using three specimen positions, from which the magnetisation variance was determined.
93 Some specimens from Core-A were measured on a 2G Enterprises RAPID (noise level ~ 0.5
94 $\mu\text{A/m}$). Specimens were housed in Mu-metal boxes with an ambient magnetic field <10 nT at
95 all times, other than when being measured or demagnetised. In Warsaw the NRM remanence
96 was measured with an AGICO JR6A spinner magnetometer (noise level ~10 $\mu\text{A/m}$).

97 One to three specimens from each sample were treated to stepwise thermal
98 demagnetisation, in 25-50°C steps up to 700°C. Low frequency magnetic susceptibility (K_{lf})
99 was monitored after heating stages, measured using a Bartington MS2B sensor to assess
100 thermal alteration. Certain specimens from all sections, were treated with a combination of
101 thermal and alternating field (AF) demagnetisation; the latter mostly conducted using a
102 Molspin tumbling AF demagnetiser at Lancaster and Warsaw (or for a few static axis-AF on
103 the RAPID, with GRM correction using GM4Edit; Hounslow, 2019). This combined
104 procedure was to better isolate the remanence in the non-red lithologies, some of which were
105 subject to thermal alteration, which started ~250 -350°C. Methods for isolation of
106 characteristic remanent magnetisation (ChRM) and ChRM behaviour classification are
107 further detailed in Hounslow et al. (inpress) and in the Supplementary Information (SI). In
108 addition to fold and reversal tests, we also use the mean VGP A95_{min} and A95_{max} thresholds
109 of Deenen et al. (2011) to express the likely capture of secular variation in the directional
110 datasets.

111 Using a representative sub-set of specimens, progressive isothermal remanent
112 magnetisation (IRM) acquired in fields up to 1.8T, and anhysteretic remanent magnetisation
113 (ARM) were applied to investigate the magnetic mineralogy (methods in Walden 1999). The
114 IRM and ARM was measured using Molspin or JR6 spinner magnetometers. The anisotropy
115 of magnetic susceptibility (AMS) was measured on selected specimens from Backside Beck,
116 and Grabowiec-6 using an Agico KLY3S Kappameter, to assess the impact of tectonic fabric
117 formation (Tarling and Hrouda, 1993).

118 For Core A and Grabowiec-6, surface magnetic susceptibility (K_{surf}) was measured using a
119 Bartington Ltd MS2K surface probe (on flat-sectioned core surface), to assist sub-division of
120 lithostratigraphy and inter-well correlation. Mass specific magnetic susceptibility was
121 measured on a Bartington MS2B probe (at 0.47 kHz) to give χ_{lf} in m^3/kg . At Backside Beck,
122 magnetic susceptibility was also measured (average of 2-4 repeats) using a Bartington MS2F
123 field probe, cross calibrated to 39 hand specimens (measured on the MS2B) to allow the

124 specimen-based dataset from the MS2B probe to be joined to that from the MS2F field probe
 125 in units of m³/kg.

126 Organic carbon isotopes ($\delta^{13}\text{C}_{\text{org}}$) for Backside Beck, Cheney-Longville and Core-A are
 127 given here, along with carbonate carbon isotope data for Core-A and Mójca (see SI Tables
 128 S1, S2). Some of these were measured at Lancaster, but isotope data from Core-A and some
 129 samples from Backside Beck used methods in Sullivan et al. (2018). These additional data
 130 were collected to improve the dating and correlation using carbon isotope stratigraphy
 131 (Ainsaar et al., 2010; Cramer et al., 2011; Bergström et al., 2015).

132 Samples for $\delta^{13}\text{C}_{\text{org}}$ measured at Lancaster were prepared by removing any carbonate
 133 minerals by reacting powdered homogenised material with 6N HCl at 25°C for 24 hours
 134 (Brodie et al., 2011). The residues were washed several times with distilled water to remove
 135 any traces of acids. Residues were then oven-dried at 50°C for 24 hours and subsequently re-
 136 powdered prior to $\delta^{13}\text{C}_{\text{org}}$ isotope analysis. Decarbonated residues were weighed into tin
 137 capsules and loaded into an auto-sampler connected to an Elementar Vario MICROcube,
 138 from where they were dropped into the furnace at 950°C. Produced gases were passed (under
 139 He) through chemical traps to remove sulphur, excess oxygen and water. Large sample
 140 volumes could be used, so we could reliably measure $\delta^{13}\text{C}_{\text{org}}$ down to around 0.02% total
 141 organic carbon (%TOC). Nitrogen was separated from CO₂ by temperature programmed
 142 desorption. The isotopic composition of the resultant purified CO₂ was then measured using
 143 an Isoprime100 Isotope mass spectrometer. Carbon isotope ratios are reported as delta values
 144 ($\delta^{13}\text{C}$) in per mil relative to the international VPDB scale (standards used NBS-18= -5.014
 145 ‰; LSVEC= -46.6 ‰). Analytical precision (1 σ) is estimated to be better than $\pm 0.15\text{‰}$ for
 146 $\delta^{13}\text{C}_{\text{org}}$ based on the replicate analysis of pure, well-mixed, organic compounds used as
 147 laboratory calibration materials. Additional $\delta^{13}\text{C}_{\text{org}}$ data shown for Grabowiec-6 and
 148 Zwierzyniec-1 are from Sullivan et al. (2018).

149 **3.Geology and stratigraphy of sections and cores**

150 **3.1 Cheney-Longville Section, UK**

151 The section sampled is a road cutting (52°28'03.7"N; 2°51'47.4"W; elevation 154 m) on
 152 the A489 in the classic area of the Caradoc Series in the Onny Valley in Shropshire, east of
 153 the Church Stretton Fault (SI Fig. S1b). Conodont alteration indices in the Ordovician east of
 154 the Church Stretton fault are low at around 1.5 (Aldridge, 1986). The Onny Valley strata have
 155 been extensively described for their shelly faunas, conodonts and chitinozoans (Dean, 1958,
 156 Hurst, 1979, Savage and Bassett, 1985; Rushton et al., 2000; Vandenbroucke et al., 2009).
 157 Since the 1970's when the road cuts were made, they have become increasingly vegetated,
 158 and in 2014 only around 20 m of the lower part of the Cheney-Longville Fm (grid reference
 159 SO 4202 8570 to 4208 8566) was suitably exposed for magnetostratigraphic sampling (Fig.
 160 2). The sampled section comprises interbedded brownish sandstones (partly hummocky
 161 cross-bedded) and silty and sandy mudstones, with a limestone bed in the uppermost part of
 162 section. This unit was deposited in an inner shelf environment (Brenchley and Newall, 1982).
 163 Our sampled section starts ca. 41 m above the base of the formation within the Woolstonian
 164 substage (lower part of the British Cheneyan Stage). This is part of the succession between

165 ca. 291-311 m in the Onny Valley succession, as summarised by Hannigan and Brookfield
 166 (2013). The Woolstonian and Cheneyan are some 140 m and 200 m thick in the Onny Valley,
 167 so the sampled section starts ~ 51% and ~36% through the Woolstonian and Cheneyan
 168 respectively. Chitinozoans from the Onny Valley sections indicate the section falls within the
 169 lower part of the *Spinachitina cervicornis* Biozone because of the distribution of the
 170 morphologically similar species *Spinachitina multiradiata* found in the Onny Valley sections
 171 (Vandenbroucke et al., 2009). In Baltoscandian regional stages the *S. cervicornis* Biozone
 172 ranges through the upper Haljala, Keila and lower Oandu stages (Vandenbroucke et al.,
 173 2009), but based on chitinozoan associations, the Woolstonian may equate to the mid parts of
 174 the Baltic Keila Substage (Vandenbroucke, 2008). $\delta^{13}\text{C}_{\text{org}}$ data from the section are broadly
 175 constant throughout the section between -27 to -26 ‰ (apart from two probably spurious
 176 values less than -28 ‰; Fig. 2e), indicating the section does not overlap with the Guttenberg
 177 carbon isotope excursion in the earliest Katian, within the Baltic Keila Stage (Ainsaar et al.,
 178 2010). Therefore in combination with the chitinozoan biostratigraphy, the section probably
 179 occupies the upper part of isotope zone BC5 in the Keila Stage (early Katian). This is
 180 confirmed by the chitinozoan distribution, when compared to graptolites seen in the Whitland
 181 section in South Wales where the *S. cervicornis* Biozone fauna is associated with graptolites
 182 from the *clingani* Biozone (Vandenbroucke, 2008). In addition, the shelly faunas of the
 183 Woolstonian Stage suggest correlation to the lower part of the graptolite *clingani* Biozone
 184 (Fortey et al., 1995).

185 **3.2 Mójcza section, Holy Cross Mountains, Poland**

186 The Mójcza section (50°50'8.8"N; 20°41'40.7"E; 251 m elevation) is located in the Kielce
 187 region of the Holy Cross Mountains and consists of condensed Middle to Upper Ordovician
 188 limestones of the Mójcza Formation, some 10 m thick (Dzik and Pisera, 1994; Trela, 1998,
 189 2000, 2005, 2006a, b; Modliński and Szymański, 2001; Schätz et al., 2006; SI Fig. S1c). This
 190 is overlain by 1 m of ?Hirnantian claystones and siltstones and at least 3.5 m of Silurian
 191 graptolitic shales, assigned to the Prągowiec Beds (Fig. 3). Conodont alteration indices in the
 192 Mójcza Formation are low at 1.0 (Belka, 1990).

193 The Mójcza Formation formed on an isolated offshore carbonate platform showing strong
 194 phosphatisation, particularly in the latest Darriwilian and Sandbian (Trela, 2005). The
 195 limestones lack calcareous algae and reef-building corals, suggesting deposition below the
 196 photic zone, so the limestones are interpreted as deep water deposits, largely swept clear of
 197 major detrital inputs by oceanic currents (Trela, 1998). The section has a rich collection of
 198 conodonts, which allow a detailed zonation to be established (Dzik, 1978; 1994; Dzik and
 199 Pisera, 1994) from the *Eoplacognathus variabilis* zone of the early Darriwilian to the *A.*
 200 *ordovicicus* of the late Katian. A major hiatus occurs in the Darriwilian where the conodont
 201 *E. suecicus* Biozone is missing (Dzik and Pisera, 1994), although a number of hardgrounds
 202 also occur in the younger part of the formation, also indicating possible missing intervals
 203 (Trela, 2005), as might be expected in this highly condensed limestone. A bentonitic marker
 204 bed occurs in the upper part of the *Pygodus anserinus* Biozone near the Darriwilian-Sandbian
 205 boundary. The lower part of the section was previously sampled for magnetostratigraphy by

206 Schätz et al. (2006), which we have extended into the upper part of the section and the newly
207 excavated overlying Silurian (Fig. 3).

208 Above the study interval of Schätz et al. (2006) are medium- to thick-bedded limestones,
209 and thin-bedded brown to yellow argillaceous limestones and overlying marls and calcareous
210 claystones with a sharp, erosional base. This interval forms the topmost part of the Mójcza
211 Formation (Trela, 2006a, b) and belongs to the upper *superbus* and *ordovicicus* conodont
212 biozones (Dzik and Pisera, 1994; Fig. 3). These limestones and marls are represented by
213 ostracod packstones to wackestones with abraded and crushed remains of trilobite fragments,
214 and scarce brachiopod and mollusc shells, deposited in calm-water conditions (Trela, 1998).

215 The topmost part of the Ordovician in the Mójcza section is formed by greenish-grey
216 claystones and siltstones with one thin intercalation of sandstone, corresponding to the
217 Zalesie Formation *sensu* Trela (2006a, b). This can be related to the sea-level drop during the
218 Hirnantian, although there is no biostratigraphic evidence of age. Above this is a succession
219 of Silurian graptolitic shales and siltstones with carbonate concretions and *Odontopleura*
220 trilobites indicating correspondence to the Prągowiec Beds (Tomczyk, 1962; Malec, 2006).
221 The stratigraphic range of these beds in the HCM is from uppermost Sheinwoodian graptolite
222 *rigidus* Biozone to the Ludfordian *leintwardinensis* Biozone (Tomczyk, 1962; Tomczykowa
223 and Tomczyk, 1981; Tomczykowa, 1988; Malec, 2006). As a consequence, the major
224 stratigraphic gap at the Ordovician and Silurian boundary in the Mójcza section includes at
225 least the Llandovery and almost all the Sheinwoodian. The origin of this hiatus, either
226 through no deposition/erosion or tectonic reduction, is uncertain.

227 A $\delta^{13}\text{C}_{\text{carb}}$ stratigraphy is available for the bulk of the Mójcza Formation (Trela, 2000),
228 with overlapping data collected here (Fig. 3e), which broadly agree with the details of Trela
229 (2000). In our sampled interval $\delta^{13}\text{C}_{\text{carb}}$ data show a general upwards trend in the Mójcza Fm
230 from -1‰ to 1‰ (Fig. 3e). This range is similar to that from lower in the section which
231 varies from ca. -2 to 1‰ (Trela, 2000).

232 3.3 Core A, Livonian Tongue, Lithuania

233 Core-A is from the Livonian Tongue, an early Palaeozoic sag basin that extends across
234 western Lithuanian and Latvia and southern Estonia, and was an extension of the Central
235 Baltic Basin to the SW (Modliński et al., 1999; Paškevičius, 2007; Dronov et al., 2011;
236 Dronov, 2017). Core-A ($55^{\circ}25'35.5''\text{N}$; $22^{\circ}18'47.5''\text{E}$; 106 m elevation) covers 80 m of the
237 Ordovician and 15 m of the early Llandovery. A major unconformity separates the
238 Ordovician from the underlying Cambrian (Fig. 4). Bedding dips are near zero and thermal
239 alteration shown by regional conodont alteration indices are around 1.0 (Nehring-Lefeld et
240 al., 1997).

241 The lithostratigraphy of the Ordovician in the Livonian Tongue is well established
242 (Lazauskiene et al., 2003; Hints et al., 2005; Dronov et al., 2011; Meidla et al., 2014), which
243 allows a sub-division of the core into a detailed lithostratigraphy (Fig. 4). This
244 lithostratigraphy can be related to the Baltic regional stages (Männil and Meidla, 1994),
245 which have a detailed conodont, graptolite and chitinozoan biostratigraphy (Nölvak et al.,
246 2006; Meidla et al., 2014). This is supplemented by a detailed carbon isotope sequence

247 stratigraphy and locally an ash bed stratigraphy (Ainsaar et al., 2010; Harris et al., 2004;
 248 Lazauskiene et al., 2003; Kiipli et al., 2009). The surface magnetic susceptibility (K_{surf}) to a
 249 large extent inversely reflects the carbonate content of the cores with lows in carbonate-rich
 250 intervals (e.g. Saldus Fm) and highs in mudstone rich intervals, particularly if they are red-
 251 coloured like in the Kriukai Fm of the Vokhov Stage (Fig. 4). A high in K_{surf} at 41.4 m relates
 252 to the peak of Sandbian organic-rich shale 'kukersite' accumulation in the Kukruse Stage
 253 (Kiipli et al., 2010). A complex of bentonites in the Adza Fm at 38.7 m (including the
 254 Kinnekulle-K bentonite, Kiipli et al., 2009; Fig. 4) is marked by a broad low in K_{surf} , like in
 255 the Vollen section near Oslo (Svensen et al., 2015). The fragmentary carbon isotope
 256 stratigraphy in combination with the lithostratigraphy allows some of the carbon isotope
 257 zones of Ainsaar et al. (2010) from the Jurmala R-1 core (~150 km to the NNE) to be
 258 correlated to Core-A. This assumes the $\delta^{13}\text{C}_{\text{carb}}$ changes are largely synchronous, but which is
 259 not necessarily the case for Baltican lithostratigraphic boundaries, which can be diachronous
 260 (Ainsaar et al., 2010; Hints et al., 2014). The upwards decline in $\delta^{13}\text{C}_{\text{carb}}$ seen through the
 261 Saldus Fm and the low at around depth 1.5 m indicates the equivalent of the terminal part of
 262 the *persculptus* Biozone is present in the core (Gorjan et al., 2012), truncated by the
 263 unconformity at the base of the overlying Remte Fm (Fig. 4), whose basal age is late
 264 Rhuddanian (Hounslow et al., inpress).

265 **3.4 Grabowiec-6 Core, Lublin slope, Poland**

266 The Grabowiec-6 well (Fig. 1 50°57'5.2"N; 23°25'56.8"E; 209 m elevation) cored Katian
 267 to late Ludfordian (Sullivan et al., 2018) horizontally bedded units. Only the 23.8 m of Upper
 268 Ordovician carbonates beginning at 3793.1 m are described here (Fig. 5), the overlying
 269 Telchyan is described by Hounslow et al. (inpress). Conodont alteration indices are around
 270 3-4 (Nehring-Lefeld et al., 1997). The Ordovician in the core can be divided into two parts by
 271 a reddened hardground surface at 3797.8 m, which has a substantial K_{surf} spike (Fig. 5c). The
 272 limestone above this surface is barren, but below, it contains chitinozoans and graptolites.
 273 The graptolite *Lasiograptus harknessi* at 3801 m has a wide range through the *gracilis*,
 274 *foliaceus* and *clingani* biozones of Scania (Pålsson, 2001), indicating possible Sandbian-early
 275 Katian strata. The chitinozoan *Lagenochitina baltica* between 3801.1 m and 3815.0 m is
 276 probably the most diagnostic for age (Fig. 5c). In Baltic successions this first appears near the
 277 base of the *Fungochitina fungiformis* Zone (at base of the Baltic Rakvere Stage, mid parts of
 278 the conodont *superbus* Biozone; Kiipli et al., 2014), and disappears near the top of the
 279 *Tanuchitina bergstroemi* Zone (early Pirgu Stage, mid *ordovicicus* Biozone at around the
 280 basal graptolite *complanatus* Biozone; Nölvak and Grahn, 1993). In Avalonian successions,
 281 the first occurrence of *L. baltica* is often used as a proxy for the *Fungochitina spinifera*
 282 Biozone, with an upper range through the *T. bergstroemi* Biozone (through the British
 283 Onnian, Pugsillian and into the Cautleyan stages; Vandenbroucke, 2008, Vandenbroucke et
 284 al., 2013). Additional age constraints can be obtained by correlation to the Zwierzyniec-1
 285 well, 25.7 km to the southwest (Fig. 5a,b). The prominent double peaked in K_{surf} can be seen
 286 in both wells (3801.5- 3795.4 m in Grabowiec-6 and 3011.5-3002 m in Zwierzyniec-1), with
 287 the intervening low in K_{surf} related to an increase in carbonate content, shown by a prominent
 288 low in Al% (Fig. 5b). This interval is also enriched in the Fe/Si ratio and has elevated NRM
 289 intensity (Fig. 5b,c), which is related to the more oxic nature of limestones over this interval.

290 Lower in both cores smaller amplitude modulation of the K_{surf} signal along with sharp low
 291 troughs (high carbonate contents) in K_{surf} tentatively allow a more detailed correlation that is
 292 consistent with the distribution of the chitinozoan *Belonechitina hirsuta* complex. In Baltic
 293 successions the chitinozoan *Spinachitina multiradiata* is a component of the *Lagenochitina*
 294 *dalbyensis* and *Sp. cervicornis* biozones (Nölvak and Grahn, 1993; Grahn and Nölvak, 2007),
 295 having a range through upper *foliaceus* and lower *clingani* biozones, suggesting a level near
 296 the Sandbian-Katian boundary in the bottom of the Zwierzyniec-1 core (Fig. 5a). This and the
 297 overlying biostratigraphy (Sullivan et al., 2018) is consistent with the $\delta^{13}\text{C}_{\text{org}}$ excursions in
 298 Zwierzyniec-1, which suggest the Rakvere, Saunja, Moe and Paroveja excursions of Ainsaar
 299 et al. (2010) are present. These correlation relationships suggest the barren unit above 3797.8
 300 m in Grabowiec-6 represents the early part of the Hirnantian (Fig. 5c).

301 These ages are supported by the nearby (15 km to NW) Łopiennik IG-1 well, where the
 302 17.5m of the uppermost Ordovician, Tyśmienica Formation overlies the Kodeniec Formation
 303 (Modliński and Szymański, 2008). The Kodeniec Formation is 4–5 m of grey and red-
 304 brownish marly and organodetritic limestones (Modliński and Szymański, 2008, 2012;
 305 Drygant et al., 2006) which is lithologically similar to the 3799.5 -3796 m interval in
 306 Grabowiec-6. The Tyśmienica Formation contains trilobites and brachiopods with
 307 *Mucronaspis* sp., *?Proteus* sp., *?Platymenta* cf. *polonica*, *Eostropheodonta* sp., *Plactatrypa*
 308 sp., and *Orbiculoidea* sp. indicating the Hirnantian (op. cit.). Therefore the interval 3799.5 m
 309 to top most Ordovician in Grabowiec-6 likely correlates to the Kodeniec and overlying
 310 Tyśmienica formations in Łopiennik IG-1. Underlying the Kodeniec Formation in wells
 311 Białopole IG-1 (20 km to NE) and Łopiennik IG-1 is 53–57 m of dark claystones with
 312 intercalations of mudstones, marls and marly limestones of the Udal Formation. The middle
 313 part of which has a fauna of *D. clingani*, *Amplexograptus vasae*, *Climacograptus bicornis*,
 314 *Orthograptus calcaratus tenuicornis* (Modliński and Szymański, 2008, 2012) indicating the
 315 *clingani* Biozone, and is lithologically much like the cored interval in Grabowiec-6 below
 316 3799.5 m. These data indicate the Ordovician in Grabowiec-6 is younger than the Cheney
 317 Longville section, but overlaps in age with the mid part of the Mójcza section. Consequently,
 318 the hiatus at 3793.18 m in Grabowiec-6 has most of the Llandovery and probably late
 319 Hirnantian missing (Sullivan et al., 2018).

320 **3.5 Backside Beck section, Lake District, UK**

321 The Backside Beck section (54°23'33.9"N; 2°27'51.0"E; 383 m elevation), in the
 322 Westerdale inlier of the Howgill Fells, provides one of the most continuous sections in
 323 England through the Upper Ordovician and Llandovery (SI Fig. S1a), and is also in the type
 324 area of the British Ashgill Series (Ingham and Wright, 1970). In the section a significant N-S
 325 striking fault cuts the Upper Ordovician Cautley Volcanic Member, just below the base of the
 326 studied section (Fig. 5 and SI Figs. S2, S3). Conodont alteration indices in the Upper
 327 Ordovician of the central Lake District to the west are around 5 (Bergström, 1980), although
 328 in the Howgill Fells may be less thermally mature than this (Oliver, 1988). Although thermal
 329 alteration is high, this region contains one of the few good quality palaeomagnetic datasets
 330 from European Silurian sediments (Channell et al., 1993), so was investigated in more detail.

331 The sampled part of the Backside Beck section (Fig. 6) begins in bedded, pink to buff,
 332 rhyolitic, vitric tuffs of the Cautley Volcanic Member. The water lain tuff units grade into,
 333 and are interbedded with, blue-grey, calcareous mudstones containing light grey calcareous
 334 nodules typical of the enclosing Cautley Mudstone Formation. This upper part of the Cautley
 335 Mudstone Formation above the main volcanic interval (–60 to –35 m; Fig. 5), and below the
 336 Cystoid Limestone Member was assigned to trilobite zone 7 (late Rawtheyan) by Ingham
 337 (1966). Rickards (2002, 2004) re-assessed the graptolite faunas from Ingham’s zones and
 338 suggested that Ingham’s Zone 6 fauna (mid Rawtheyan) belongs to the *linearis* Biozone,
 339 hinting that the slightly different graptolite assemblage in trilobite Zone 7 may be indicative
 340 of a transition into the *complanatus* Biozone (earliest Rawtheyan; late Katian; Zalasiewicz et
 341 al., 2009). This implied that the main body of the *complanatus* and *anceps* biozones, and any
 342 Hirnantian must be restricted to the overlying beds (i.e. uppermost part of Cautley Mudstone
 343 through Ashgill formations). The chitinozoan *Bursachitina umbilicata* Biozone occurs in the
 344 younger part of the Cautley Mudstone (Vandenbroucke et al., 2005; Vandenbroucke, 2008).
 345 Within this unit chitinozoan *Ancyrochitina merga* allows correlation to the northern
 346 Gondwana *A. merga* Biozone, the terminal chitinozoan zone of the Katian, which conflicts
 347 with the older age suggested by the ‘transitional’ *complanatus* Biozone interpretation of
 348 Rickards (2002). This younger age for the upper part of the Cautley Mudstone Fm is
 349 consistent, with beds assigned to Ingham’s zone 7 at Girvan which contain *Paraorthograptus*
 350 *pacificus*, the index species of the upper subzone of the *anceps* Biozone of the latest Katian
 351 (Floyd et al., 1999; Zalasiewicz et al., 2009).

352 A 2 m thick muddy limestone; the Cystoid Limestone Member, marks the base of the
 353 Ashgill Formation (Fig. 6). In the nearby Taythes Inlier, the volcanic member and two shelly
 354 faunal zones are missing below this limestone (Ingham, 1966), but this hiatus is not evident at
 355 Backside Beck. The main Ashgill Formation comprises about 70 m of blue-grey mudstones
 356 lacking the calcareous nodules of the Cautley Mudstone Formation (Fig. 6). Apart from the
 357 benthic shelly fauna in the Cystoid Limestone, the Ashgill Formation is sparsely fossiliferous
 358 and restricted to one trilobite species (*Mu. mucronata*) and three brachiopod species (e.g.
 359 *Eostropheodonta*, *Dalmanella*, *Plectatrypa*), and *Hirnantia cf. sagittifera* (Ingham, 1966).
 360 The bulk of this ‘Hirnantian-type’ shelly fauna comes from the upper part of the Ashgill Fm,
 361 but over a wide area of the Howgill Fells (Ingham, 1966). The Ashgill Fm is barren of
 362 graptolites and chitinozoans (Rickards 2002; Vandenbroucke et al., 2005). A sandy mudstone
 363 12.2 m below the top of the formation, may correlate with conglomerates elsewhere in the
 364 Cautley inliers and with the Wharfe Conglomerate Member from the Craven inliers (Rickards
 365 and Woodcock 2005) probably representing a sea level low-stand.

366 The base of the overlying Skelgill Formation is marked by a 1.0 metre thick limestone (at
 367 Backside Beck), the Spengill Member (Kneller *et al.*, 1994), which is the ‘basal beds’ of
 368 Rickards (1970, 1988) and earlier authors (Fig. 6). In Lake District sections to the west, this
 369 member contains *persculptus* Biozone graptolites (Hutt, 1974) and the distinctive low-
 370 diversity *Hirnantia* shelly fauna including *Hirnantia sagittifera* (Scott and Kneller, 1990).
 371 The Spengill Member is therefore Hirnantian rather than Llandovery in age and is overlain by
 372 black shales starting in the *acuminatus* Biozone.

373 The $\delta^{13}\text{C}_{\text{org}}$ shows a decline through the Ashgill Fm to more negative values from a high
 374 in $\delta^{13}\text{C}_{\text{org}}$ of ca -28‰ at around the level of the Cystoid Limestone Mbr, with a low in
 375 $\delta^{13}\text{C}_{\text{org}}$ in the Ashgill Fm at ca. 60 m (Fig. 6e). Above this there is some fluctuation to a
 376 second high of ca. -28.3‰ , in the uppermost two samples from the Spengill Mbr. The
 377 $\delta^{13}\text{C}_{\text{org}}$ changes beginning around 60 m and into the Spengill Mbr are inferred to be the initial
 378 part of the rising limb of the Hirnantian isotope excursion (HICE), with the corresponding
 379 underlying low in $\delta^{13}\text{C}$ seen in Baltic sections, close to the base of the Porkuni Stage
 380 (Jurmala R-1; Fig. 4) in the latest Katian (Ainsaar et al., 2010; Bergström et al., 2015). The
 381 increasingly more positive values downwards through the Ashgill Fm are probably part of
 382 isotope zone BC15 in the underlying upper part of the Baltic Pirgu Stage (Fig. 4) of latest
 383 Katian age. This correlation is compatible with the Hirnantian faunas from the upper Ashgill
 384 Fm and Spengill Mbr.

385 **4 Magnetic results.**

386 **4.1 Magnetic mineralogy**

387 The magnetic mineralogy of samples from the sections and cores range from hard to soft
 388 coercivity behaviour, with higher coercivity minerals (haematite or goethite) seen by non-
 389 saturation in an IRM field of 300 mT (Fig. 7), and remanent coercivities (H_{cr}) >200 mT (SI
 390 Fig. S4). Reddish-and brownish coloured samples at Mójcza and Core-A have $>80\%$ of IRM
 391 acquisition above 200 mT and $H_{\text{cr}} >280$ mT (Figs. 7a,b; SI Fig. S4). The hard ferrimagnetic
 392 mineral in most of these samples is haematite as shown by the resistance of the NRM to
 393 thermal demagnetisation (see later).

394 Grey-coloured samples from Backside Beck and Grabowiec-6 have some of the softest
 395 behaviour ($H_{\text{cr}} <50$ mT; SI Fig. S4), but with most samples tested having 10-20% of IRM
 396 acquisition above 200 mT (Figs. 7c,d). Behaviours intermediate between these hard and soft
 397 end members are found in most sections in the non-red lithologies. The relative consistency
 398 of the H_{cr} and $\text{SIRM}_{\text{IT}}/\chi_{\text{ARM}}$ values in those samples with softer coercivity behaviour, suggest
 399 a similar mineralogy in most of the sections and cores (SI Fig. S4). Using data from Peters
 400 and Dekkers (2003) suggests this is most likely magnetite, with magnetic particle sizes <0.1
 401 μm in size (See Fig. S4).

402 **4.2 Palaeomagnetic and magnetostratigraphic results**

403 **4.2.1 Cheney-Longville section**

404 The palaeomagnetic specimens responded best to thermal demagnetisation to round 350°C
 405 to 400°C , followed by AF demagnetisation (Fig. 8b,c). There are no systematic changes in
 406 NRM intensity (mean = 0.48 mA/m) or magnetic susceptibility (mean = 23×10^{-6} SI) in the
 407 section (Fig. 2). A low stability component typically dominates the NRM, up to around
 408 200°C to 250°C with some 80% of the NRM intensity composed of this component (Fig.
 409 8b,c). This well-defined component (mean of 359 , $+75$, $\alpha_{95}=3.1$; SI Fig. S6) is interpreted as
 410 a primarily a Brunhes component (geocentric axial dipole inclination at site = 69°).
 411 Intermediate stability components are dominated by a weak Kiaman-like component, along
 412 with fairly random composite directions that may be mixtures of Brunhes, Kiaman or ChRM

413 directions (SI Fig. S6). The dual polarity ChRM (Fig. 9) is usually the highest stability
 414 component (often through the origin), but in nine specimens is in the intermediate stability
 415 range. In these specimens this may relate to the Kiaman-like component sometimes residing
 416 in haematite. 82% of specimens contain evidence of ChRM polarity, with 59% of these being
 417 line-fits (Fig. 2). Mean directions using either Fisher mean or the combined great circle
 418 method pass reversal tests (Table 1; Fig. 9). Minor bedding dip divergence of 15-25°
 419 precluded use of a fold test. The mean VGP, A95 is within the thresholds of Deenan et al.
 420 (2011) indicating dispersion is within the secular variation range (Table 1). The section
 421 polarity is dominantly reverse with three thin normal-polarity magnetozones and two further
 422 tentative normal magnetozones (Fig. 2)

423 *4.2.2. Mójcza section*

424 Low stability (LT) components (100 - 250°C or to early AF range) are mostly Brunhes or
 425 composites of Brunhes and Kiaman, since some of the LT components are skewed to the
 426 south in geographic coordinates, due to component overlap with the Kiaman component (SI
 427 Fig. S6). In 9% of specimens this component comprised the entire NRM. We did not find the
 428 randomly oriented low stability component identified by Schätz et al. (2006) that they
 429 assigned to a goethite remanence (Fig. 8d, e).

430 Mid-stability (MT) components are interpreted as mostly Kiaman, or a composite of
 431 ChRM and Kiaman components (steps from around 200°C into mid or high AF
 432 demagnetisation range). This is a post-tilting component (SI Fig. S6) that is the same as
 433 component-B identified by Schätz et al. (2006). Other MT components seem to be composite
 434 directions between the Kiaman component and the Brunhes or the Kiaman and the high
 435 temperature (HT) components (SI Fig. S6). The Kiaman overprints are strongly concentrated
 436 in the Ordovician part of the section, with the Silurian part of the section mostly having the
 437 intermediate stability Kiaman-Brunhes composite components. In 20% of specimens the MT
 438 component persisted until complete demagnetisation.

439 High stability components are interpreted as Ordovician and Silurian ChRMs (Fig. 9).
 440 Line fits through these components are few, since mostly this component is seen as, great-
 441 circle trends toward the expected polarity (Fig. 3b). Schätz et al. (2006) found similar
 442 behaviour in the lower part of the section. The Wenlock age part of the section is dominated
 443 by great circle behaviours. The Upper Ordovician and the Prągowiec Beds parts of the
 444 section have quite different mean directions (Table 1).

445 Our Ordovician data fails the reversal test with reverse and normal 25° apart, with a
 446 critical angle of 14° (McFadden and McElhinney, 1990). However, the data of Schätz et al.
 447 (2006) from the lower part of the section passed a reversal test (class C). Bedding dips are
 448 not sufficiently different to enable a meaningful fold test. 65% of specimens contain evidence
 449 of ChRM polarity, with 31% of these being line-fits (Table 1; Fig. 3b). The mean VGP, A95
 450 of both the Silurian and Ordovician parts of the section are within the thresholds of Deenan et
 451 al. (2011) indicating dispersion is consistent with capture of secular variation (Table 1) These
 452 data define five normal and three reverse-polarity magnetozones, with one and two tentative
 453 normal and reverse submagnetozones respectively (Fig. 3d).

454 **4.2.3 Core-A, Lithuania**

455 Thermal demagnetisation to 250-340°C followed by AF demagnetisation, best suited
 456 samples from the Saldus Fm and some of the remaining Ordovician grey-coloured
 457 lithologies, whereas thermal demagnetisation to around 500-550°C worked best for
 458 Ordovician red-lithologies, and most of the paler limestones. Low temperature (LT)
 459 components extracted between the NRM and around 250 or 300°C are steep down-directed,
 460 and interpreted as largely a Brunhes age component (Fig. 8h,i,j). In a small proportion of
 461 samples, particularly from the upper parts of the core, the LT component is very steep and
 462 may be in part drilling-induced origin (De Wall and Worm, 2001). In most specimens, the LT
 463 component tends to dominate the magnetisation intensity, but with evidence of an additional
 464 dual polarity ChRM component remaining to the highest demagnetisation stages. The mean
 465 inclination of the LT component is 73° ($\alpha_{95}=2.4^\circ$, n=155 method of McFadden and Reid,
 466 (1982); expected geocentric axial dipole field inclination of 71°; Fig. 10e). The LT
 467 component was used to re-orient the core runs (e.g. Hailwood and Ding, 1995), and recover
 468 mean ChRM directions (Figs. 10f, g). The LT data from specimens in contiguous core runs
 469 were averaged to determine the mean azimuth, where possible. However, some 24% of
 470 ChRM specimen data could not be oriented using these methods, so both VGP latitude and
 471 inclination data is shown (Fig. 11d). No evidence for the presence of a Kiaman component
 472 was found in these samples, a feature also inferred by Grappone et al. (2017) in Middle
 473 Ordovician limestones from northern Estonia, although contrastingly Plado et al. (2010)
 474 identified Permo-Triassic re-magnetisations carried by haematite in Middle Ordovician
 475 limestones in northern Estonia.

476 95% of specimens contained evidence of the ChRM, often in the mid to late stages of AF
 477 demagnetisation, or from around 300-400°C to the maximum thermal demagnetisation step,
 478 for thermally demagnetised specimens (Fig. 8h, i, j). 98% of specimens yielded a polarity
 479 interpretation, and 88% of these were s-class line fits (Fig. 11b), with the remaining 12%
 480 interpreted as T-class great-circle trends, based on the re-orientation of the core (Figs. 10,
 481 11). Mean inclination (and re-oriented mean directions) shows a systematic decline up though
 482 the core (Figs. 10g, 11d). Reversal tests (McFadden and McElhinney, 1990) are varied
 483 ranging from indeterminate to fail due to the larger than usual dispersion (Table 1), due to
 484 additional declination dispersion from re-orienting the core using the LT component (Fig.
 485 10f). This additional dispersion has lead to larger A95 (Table 1) outside the usual secular
 486 variation thresholds of Deenan et al. (2011). The VGP latitude was determined using the
 487 mean direction from depths 80 - 30 m and 30 - 16 m (Silurian data above 0 m in Hounslow et
 488 al. inpress), due to the systematic difference in inclinations. These data define 10 normal and
 489 11 reverse-polarity magnetozone, with 5 tentative single-specimen magnetozone (Fig. 11).
 490 The palaeomagnetic data from Core-A is the best quality of those we have studied here, in
 491 spite of the apparently larger ChRM dispersion due to the imperfect core re-oriented.

492 **4.2.4 Grabowiec-6 core**

493 Samples responded best to thermal demagnetisation to around 430°C followed by AF
 494 demagnetisation, or sometimes entirely thermal magnetisation (Fig. 8f,g). A low stability
 495 (component LT) component was isolated by thermal demagnetisation between 100 and 210-

496 250°C. The mean of this component has a slightly steeper inclination ($73.5^\circ \alpha_{95}=2.0$;
 497 McFadden and Reid, 1982) than the expected modern field at the core site (of 68°). The LT
 498 component is interpreted as predominantly a Brunhes age component. The intermediate, and
 499 often the high stability components are dominated by an often negative inclination
 500 component, which is often stable until the last stages of demagnetisation (Fig. 8f,g). Re-
 501 orientation of the core runs using the LT component shows the largely SSW-directed nature
 502 of this component (Fig. 10a). This is interpreted as a Kiaman partial re-magnetisation, which
 503 is widespread in early Palaeozoic sediments in Poland and the East European Craton margins
 504 (Smethurst and Khramov, 1992; Jelenska et al., 2005; Nawrocki, 2000). This Kiaman
 505 component in the specimens is very well defined, and when used to re-orient the core,
 506 indicates the LT component is scattered between the Brunhes field direction and the Kiaman
 507 component (Fig. 10b). We therefore used the better-defined Kiaman component to re-orient
 508 the specimens to assist in interpreting the behaviour of the Ordovician component. We used a
 509 mean Kiaman palaeopole (palaeopole at latitude = -42.2° , longitude = 346° ; giving an
 510 expected Kiaman field direction at the well site of 207° , $/-17^\circ$), based on 8 published studies
 511 from eastern Europe (SI Fig. S6). Data from specimens in contiguous core runs (Hailwood
 512 and Ding, 1995) were averaged to determine the mean azimuth for runs, where available.

513 74% of specimens contained evidence of an additional magnetisation component at the
 514 highest demagnetisation stages, at AF demagnetisation $>60\text{mT}$ or thermal demagnetisation
 515 $>400^\circ\text{C}$ (Fig. 12b). In 30% of these specimens (6 specimens), a ChRM line-fit direction could
 516 be extracted (Fig. 10c). In the remaining 14 specimens this component is shown as great
 517 circle trajectory trends (Figs. 10d; 12b) towards expected Ordovician normal and reverse-
 518 polarity directions, defining four normal and three reverse-polarity magnetozones. (Fig. 12d).
 519 Great-circle trends were combined with the ChRM directions to define a mean direction and a
 520 palaeopole (Table 1). A95 larger (and close to upper range) than the secular variation
 521 thresholds of Deenen et al. (2011) is probably due to some additional declination dispersion
 522 from the less than perfect core re-orientation (Table 1). The AMS shows exclusively
 523 sedimentary type fabrics (See SI Fig. S5).

524 **4.2.5 Backside Beck section, UK**

525 The χ_{lf} in the section broadly follows the formational units, with lows in the Cautley
 526 Mudstone Fm, which may reflect increased ash content in the mudstones from the upper part
 527 of the Cautley Mudstone Fm. The Ashgill Fm has largely consistent χ_{lf} values, with perhaps
 528 some variation, which is partly defined with the field probe data (Fig. 6a), but a decline in the
 529 uppermost Ashgill Fm, which presumably may be a reflection of the environmental changes
 530 heralding the Hirnantian. The magnetic susceptibility is primarily carried by paramagnetic
 531 minerals in the section (Hounslow et al., submitted), so the χ_{lf} changes must be primarily
 532 reflecting Fe-silicate content. The largest χ_{lf} are in the lower part of the Spengill Mbr (Fig.
 533 6a).

534 Our reconnaissance palaeomagnetic sampling of the Ordovician at Backside Beck
 535 indicates some 62% of samples showed evidence of Ordovician polarity (Fig. 6b), the
 536 remainder were overprinted. Some 77% of these show great circle behaviour (T-class data;

537 Fig. 6b). Ordovician samples show a dual polarity ChRM often seen by thermal
 538 demagnetisation steps above 400°C, or by AF demagnetisation above 40 mT (Fig. 8a). The
 539 palaeomagnetic behaviour of these is similar to the larger number of samples measured from
 540 the overlying Silurian in the section by Hounslow et al. (inpress), and the interpretation of
 541 overprint components does not differ (SI Fig. S7). Their data shows a low stability
 542 components (often up to 200-400 °C), which is a Brunhes-like component. Intermediate
 543 stability components are either dual polarity Lower Devonian in age, or southerly directed
 544 and shallow inclinations, with both up and downwards dipping directions, likely of Kiaman
 545 (late Carboniferous) age, or composite in nature. Further palaeomagnetic details are in
 546 Hounslow et al. (inpress) and the Ordovician data is compared to their data in SI Fig. S7. The
 547 combined great circle mean suffers from too few samples to provide a sufficiently precise
 548 directional mean (Table 1). However, the VGP A95 is within the thresholds of Deenen et al.
 549 (2011) for the expected range of secular variation (Table 1). Fold tests are possible, but only
 550 using the Ordovician and Silurian data combined (See SI Figs. S8, S10 to S13). The
 551 proportional and DC-fold tests indicate a positive fold test, with the 95% confidence intervals
 552 on unfolding including 100% unfolding (Table 2). The McFadden fold test indicates that
 553 100% unfolding is the most likely option for ChRM acquisition ($P_f > 5\%$; Table 2). Channell
 554 et al. (1993) also obtained a positive regional fold test using data from the late Llandovery of
 555 this area.

556 The data tentatively define two normal and one reverse-polarity magnetozones, with one
 557 additional very tentative reverse submagnetozone at the section base (Fig. 6). The AMS data
 558 from the Ordovician shows a similar style to the overlying Silurian described in detail by
 559 Hounslow et al. (inpress), which shows evidence of some tectonic modification, shown by the
 560 K_1 axes corresponding to the bedding-cleavage intersection, but also by more complex
 561 inverse-related fabrics, described in more detail by Hounslow et al. (inpress), with the
 562 Ordovician sample data in SI Fig. S9.

563 **5 Middle and Upper Ordovician geomagnetic polarity**

564 The higher quality of the magnetostratigraphic data from Core-A, along with the detailed
 565 lithostratigraphy of this core, linked to the Baltic regional stages provides a suitable template
 566 for the other Middle and Upper Ordovician magnetostratigraphic data. Other existing
 567 magnetostratigraphic datasets are from Sweden, Estonia and USA (Arkansas) in the Middle
 568 Ordovician (Farr et al., 1993; Torsvik et al., 1995; Grappione et al., 2017), more extensive
 569 data from Siberia (Pavlov and Gallet, 1998, 2005; Pavlov et al., 2012) and some data from
 570 Poland (Schätz et al., 2006). Gallet and Pavlov (1996) dismissed the data from Farr et al.
 571 (1993) due to an ‘inversion’ in the component stability, with a Kiaman component more
 572 stable than the inferred Ordovician component. Whilst this is unusual, similar behaviour from
 573 the Ukrainian Silurian is known (Jelenska et al., 2005), and the magnetostratigraphy seems
 574 repeatable from multiple sections, so the composite-section data from Farr et al. (1993) is
 575 used here in a tentative fashion.

576 Although the Siberian dataset is the largest, the Siberia regional stages are problematic to
 577 link to Baltican and international stages due to the endemic faunas (Dronov, 2013, 2017;
 578 Sennikov et al., 2015). The best means to do this independently of the magnetostratigraphy is

579 carbon isotope stratigraphy, supported by limited cosmopolitan biostratigraphic data and
 580 sequence stratigraphy (Kanygin et al., 2010; Ainsaar et al., 2015). Two biostratigraphic tie
 581 points in the Volginian and basal Chertovskian stages allows linkage of the Siberian sections
 582 into the late Darriwilian- Sandbian interval (Dronov, 2017). The carbon isotope zones (BC1
 583 to BC17) of Ainsaar et al. (2010) and their proposed equivalent isotope zones in the Siberia
 584 stages (Ainsaar et al., 2015) also allows a framework onto which the existing
 585 magnetostratigraphic data for the Middle to Upper Ordovician can be assembled (Fig. 13).

586 Magnetochrons have been labelled (LO, MO, UO for Lower, Middle and Upper
 587 Ordovician) in the polarity chronostratigraphic composite to allow ease of description. This
 588 labelling differs from Hounslow (2016); for example magnetochron MO1n is now labelled
 589 MO3n, since normal magnetozones now appear to extend to older intervals in the Middle
 590 Ordovician. Pavlov and Gallet (2020) have described the Lower Ordovician data in detail.

591 The earliest Middle Ordovician substantive normal magnetozones (MO1n) occurs on the
 592 rising limb of the MDICE isotope excursion (BC3) in both Core-A and the Gullhögen Quarry
 593 section (Meidla et al., 2004; Torsvik et al., 1995). Differences in the relationship of MO1n to
 594 the Baltic Kunda-Volkov stage boundary reflects the lithostratigraphic nature of this
 595 distinction in Core-A, which is instead better defined to be within the *Lenodus antivariabilis*
 596 conodont zone near Gullhögen (Lindskog et al., 2014).

597 Magnetozones MO1n is likely the equivalent of the major normal polarity magnetozones in
 598 the early to mid part of Member B of the Everton Fm (Farr et al., 1993; Fig. 13), which is an
 599 unconformity bounded package of carbonates (Etherington et al., 2012). This age relationship
 600 is inferred, since the lower part of the Everton Fm contains mid to late Dapingian conodonts
 601 (*Histiodelia altifrons* Biozone) as supported by regional correlations (Etherington et al.,
 602 2012; Cooper et al., 2012). Normal polarity magnetozones from the underlying Sneys Mbr
 603 (of the Everton Fm), and the older Lower Ordovician, Powell Dolomite, may represent
 604 additional normal magnetozones. However, these are not as well validated as those from the
 605 overlying parts (Member B) of the Everton Fm, which are based on data from more than one
 606 section (Farr et al., 1993). The submagnetochrons MO1n.1n and MO1n.2n are missing in
 607 Core-A probably due to the regional disconformity (Meidla et al., 2014) at the Volkhov -
 608 Kunda boundary. MO1n at Gullhogen quarry (base of Holen Limestone, magnetozones N1 of
 609 Torsvik et al. 1995) contains 12 normal-polarity specimens. This combined with data from
 610 Core A and the Everton Mbr indicates this normal magnetochron is well substantiated. A
 611 normal magnetozones at this level is not known in the Siberian sections (around the base of
 612 the Siberian Vikhorevian Stage, Dronov 2017), an interval there of exclusively reverse
 613 polarity (Gallet & Pavlov, 1996; Pavlov et al., 2012).

614 Magnetozones MO2n, seen in Core A (Fig. 13), is probably equivalent to the latest part of
 615 the Kunda Stage seen at Gullhögen, where a single-sample normal-polarity interval (Torsvik
 616 et al., 1995) occurs in the top of the Holen Limestone, near the top of the MDICE excursion
 617 (Meidla et al., 2004). Magnetozones MO2n was also detected by 3 normal-polarity specimens
 618 from nearby Hallekis quarry in the upper Holen Limestone (Torsvik & Trench, 1991; their
 619 'N1' magnetozones). Smethurst et al. (1998) have also detected a normal polarity

620 magnetozone in the late Kunda Stage. This is from 4 specimens from two levels in the top ~ 1
 621 m of the Tosna section. Tolmacheva (2005) indicates that the upper part of the Kunda Stage
 622 in the St Petersburg area is within the *Eoplacognathus pseudoplanus* conodont zone of the
 623 mid Darriwilian.

624 Tentative normal magnetozones in the Siberian Volginian Stage from the Polovinka (Fig.
 625 13) and Rozhkova sections may be equivalent to MO2n, considering the provisional nature of
 626 correlation using the Siberian isotope data (Ainsaar et al., 2015). However, Dronov (2017)
 627 has correlated the sequence boundary at the base of the Volginian to that at the base of the
 628 Baltic Aseri Stage, suggesting that MO2n may be in the Siberian Mukteian Stage underlying
 629 the Volginian. In Arkansas, the Jasper Mbr (of the Everton Fm) is early to mid Darriwilian in
 630 age (*Histiodella sinuosa* to *H. holodentata* conodont zones; Etherington et al., 2012; Cooper
 631 et al., 2012) suggesting that magnetochron MO2n is the probable equivalent of the normal
 632 polarity magnetozone in the top part of Member C and into the overlying Jasper Mbr of the
 633 Everton Fm (Fig. 13). This suggests the geomagnetic polarity during the Darriwilian may be
 634 best represented by the Everton Fm data, since other sections do not show such complexity in
 635 reversal pattern through MO1 and MO2n. Greater uncertainties remain about the dating and
 636 duration of hiatus in the Siberian Middle Ordovician (Ainsaar et al., 2015; Dronov, 2017), but
 637 possibly MO2n has been eroded at the Siberian Mukteian-Volginian sequence boundary. The
 638 overlying reverse-polarity dominated MO2r contains at least two brief normal-polarity
 639 submagnetozones, represented by those in the Volginian-Kirinian stages in Siberian sections,
 640 and in the uppermost part of the Everton Fm (Fig. 13).

641 Magnetozone MO3n is seen in the Baltic Uhaku Stage in Core-A, in the Pakri Cape and
 642 Vao Quarry sections and in a single-sample level at Gullhögen, within the BC4 isotope zone
 643 (Fig. 13). MO3n is dated to the upper part of the *P. serra* conodont zone (ca. base of
 644 *Eoplacognathus lindstroemi* Subzone) in the Gullhögen Limestone (Holmer, 1989), and
 645 therefore its base likely correlates to the lowest normal magnetozone (MO3n.1n) at Mójcza
 646 (Fig. 13). The Siberian Rozhkova section at around the upper parts of BC4 has two normal
 647 magnetozones (MO3n.1n and MO3n.2n) compared to the Baltic sections. The Mójcza section
 648 similarly has two normal magnetozones in the upper part of the Baltic Uhaku Stage (which
 649 overlaps the *serra* and *anserinus* biozones; Meidla et al., 2014; of mid to late Darriwilian
 650 age). This suggests the polarity data from Baltic sections is incomplete (in Core-A strata
 651 missing across the Taurupe- Dreimani boundary?), and so we use the Siberian Kudrino
 652 section to construct the composite polarity scale through MO3n and MO3r. Using these
 653 relationships suggests magnetozone MO3n ranges into the early part of the *P. anserinus*
 654 Biozone (late Darriwilian; Fig. 13).

655 A substantive normal-polarity interval occurs in the Dreimani Fm in Core A, the Dalby
 656 Limestone at Gullhögen and the later parts of the Chertovskian in Siberian sections, which
 657 we call UO1n. This magnetozone is what Trench et al. (1991) referred to as bias-
 658 magnetozone 'C(N)', characterised by normal-polarity palaeopole-type data from a number
 659 of volcanic-units. At Gullhögen, the base of UO1n is not detected due to a large sampling gap,
 660 so UO1n may extend into the underlying Ryd limestone (Torsvik et al., 1995; Fig. 13). The
 661 Siberian Rozhkova section appears to provide the most detailed record through the interval

662 MO3r.1n to UO1n, and is used for the composite polarity. At Mójcza, the base of UO1n is
 663 within the *Prionlodus gerdæ* Subzone of the *tvaerensis* Biozone, but the youngest part of the
 664 *tvaerensis* Biozone may be condensed because of phosphatisation at this level (Trela, 2005).
 665 In Core-A, UO1n is within the upper part of the Baltic Kukruse Stage, which is within the
 666 mid and lower parts of the *tvaerensis* Biozone (Meidla et al., 2014).

667 Magnetozone UO1r is detected through the Baltic Haljala-Keila-Nabala-Vormsi stages in
 668 Core-A, with a probable higher resolution record of the lower part of UO1r (containing
 669 UO1r.1n and UO1r.2n) in the Cheney-Longville section (Fig. 13); a correlation supported by
 670 chitinozoans (Vandenbroucke, 2008). The Mójcza magnetostratigraphy appears incomplete in
 671 the interval covering UO1r (also a sampling gap), likely because of three hiatus in the
 672 *superbus* Biozone (Trela, 2005), suggesting the normal magnetozone beginning at 3 m is
 673 UO2n.1n (Fig. 13). The zonal interpretation of carbon isotope data at Mójcza is tentative, but
 674 may range through the BC10-BC11 isotope zones from ~3.0- 3.5 m (Fig. 13). The Guttenberg
 675 carbon isotope excursion (GICE) is located in magnetochron UO1r.3r probably above the
 676 sampled Cheney-Longville section and below the oldest part of the Grabowiec-6 core.

677 Magnetozone interval UO2n-UO2r is found in Core-A, and appears in the lower part of
 678 the Grabowiec-6 core, based on the $\delta^{13}\text{C}_{\text{org}}$ excursions in this core (Figs. 4, 5). This interval is
 679 also seen at Mójcza at 3- 3.8 m, since the *ordovicicus* Biozone ranges through the Upper
 680 Nabala –Vormsi - Pirgu and Porkuni stages (Meidla et al., 2014). At Mójcza more negative
 681 $\delta^{13}\text{C}_{\text{carb}}$ values between 3.1 – 3.5 m may be BC11, with overlying BC12 (more positive
 682 isotope values) in the *ordovicicus* Biozone (Figs. 4, 13). In all sections isotope zone BC12
 683 spans the mid part of magnetochron UO2r (containing UO2r.1n).

684 Magnetochron UO3 is seen in isotope zones BC13 – BC14 to lower part of BC15 in Core-
 685 A and Grabowiec-6. The top of the Mójcza Fm is likely within magnetozone UO3r within the
 686 BC14 isotope zone, like that in Core-A. Kolesov (2007) has also probably detected the
 687 UO3n-UO3r magnetochrons in two Siberian sections on the Lena and Nuya rivers equated to
 688 the *Pleurograptus linearis* Zone. Study of the Dolbor Fm from Siberia (of mid Katian age;
 689 Dronov, 2017) by Pavlov and Gallet (2020) has also found only normal polarity, probably
 690 equivalent to UO3n. Several other studies of rather poorly dated rocks around this age
 691 (Bachtadse et al., 2000; Bazhenov et al., 2003) has lead Pavlov and Gallet (2020) to infer
 692 entirely normal polarity for much of the Upper Ordovician, which conflicts with our data,
 693 which clearly shows a more complex reversal pattern (Fig. 13).

694 In Core-A, the transition from the Pirgu Stage to the overlying lower part of the Porkuni
 695 Stage is interrupted by two hiatus, which may have removed part of the early Hirnantian, a
 696 feature commonly inferred in Latvia and Lithuanian (Brenchley et al., 2003; Meidla et al.,
 697 2014). In Core-A, magnetozone UO4r is likely truncated by the lowest of these hiatus at the
 698 base of the Kuldiga Fm, and the top of UO5n is likely truncated by hiatus at the base of the
 699 Saldus Fm (Figs. 11, 13). More expanded intervals of the Kuldiga Fm show the progressive
 700 increase in $\delta^{13}\text{C}$, younging through to the peak of the HICE excursion in the late Kuldiga Fm
 701 or lower parts of the Saldus Fm (Hints et al., 2010). For this reason we have expanded the
 702 polarity pattern from the Kuldiga Fm, covering the upper part of UO3r and lower part of

703 UO4n. This expansion is supported by the reverse-polarity seen in the lower part of isotope
 704 zone BC16 in the Grabowiec-6 core (Figs. 12, 13). The sparse Ordovician data through the
 705 Ashgill Fm at Backside Beck may represent the interval from upper UO3r into the early part
 706 UO5n, with the normal polarity in the Cautley Mudstone Fm being the equivalent of UO4n.
 707 This is broadly the correlation suggested by the chitinozoans from the upper part of the
 708 Cautley Mudstone Fm, and the Hirnantian fauna in the upper Ashgill Fm and Spengill Mbr.
 709 This places the base of the Porkuni Stage in the upper part of the Ashgill Fm at Backside
 710 Beck (at ~ 60 m level), consistent with the sedimentological evidence of sea-level lowstand
 711 in the formation, and the latest Katian low in $\delta^{13}\text{C}$. This would make the base Hirnantian
 712 within UO5n, rather than in the top of UO4r. The Katian-Hirnantian boundary needs more
 713 work to better place the base Hirnantian with respect to the magnetostratigraphy.

714 The upper part of the Saldus Fm is the only representation of magnetozone UO6, which is
 715 in the youngest part of the decline from the $\delta^{13}\text{C}$ peak in the HICE (Figs 4, 11), with the
 716 likely minimum in $\delta^{13}\text{C}_{\text{carb}}$ at 1.4m below the top of the formation, within UO6r. The top-
 717 most sample in the Saldus Fm is of normal polarity, and likely represents the first normal
 718 magnetochron, LL1n of the Llandovery (Hounslow et al., inpress).

719 In the original definition of the Moyero Superchron, Pavlov and Gallet (2005) placed the
 720 upper boundary of the superchron equivalent to the top of MO2r, based on data from the
 721 Siberian Kudrinian Stage at Kudrino and Moyero (Fig. 13). Retaining this definition,
 722 suggests that the Moyero Superchron does not have 100% bias to a single polarity, like the
 723 late Cretaceous Normal-Polarity Superchron (Ogg, 2012), but instead is more like the
 724 Carboniferous-Permian, Kiaman Superchron in containing brief normal-polarity intervals
 725 (Hounslow and Balabanov, 2016; Hounslow, inpress). In this sense the Paleozoic superchrons
 726 seem physically distinct from the Cretaceous Normal-Polarity Superchron. The exclusively
 727 reverse-polarity part of the Moyero Superchron appears to be at maximum ca. 10 Myr
 728 duration, substantially short than the 39 Myr duration for the Cretaceous Superchron
 729 (Olierook et al., 2019). The absence of substantive normal polarity in the Middle Ordovician
 730 Siberian sections below MO3n may relate to the issue of potentially missing intervals at the
 731 sequence boundaries in these successions (Dronov, 2017). The number of magnetozones
 732 through the Darriwilian-Sandbian is larger than that through the Katian-Hirnantian, a
 733 probable expression of the larger dataset from the Middle Ordovician, rather than actual
 734 larger reversal frequency.

735 **6 Conclusions**

736 An Ordovician geomagnetic polarity scale is defined using polarity data from five new
 737 successions covering the Middle to Upper Ordovician through the Dapingian to Hirnantian
 738 stages. These new datasets are supported by some fold and reversal tests, suggesting the
 739 primary nature of the magnetisations, which is carried by variable mixtures of both haematite
 740 and magnetite. Using existing biostratigraphy, supplemented by correlations based around
 741 new carbon isotope and magnetic susceptibility data, we construct a correlation and age
 742 framework onto which these new polarity datasets are tied.

743 In combination with existing magnetostratigraphic data from the Dapingian to early
 744 Sandbian, we construct the first geomagnetic polarity chronostratigraphic scale through the
 745 entire ca. 23.5 Myr of the mid to latest Ordovician. Our new datasets demonstrate a
 746 substantial revision in the polarity pattern for the Upper Ordovician. This makes the
 747 Ordovician System the first in the early Paleozoic to have a nearly complete record of
 748 geomagnetic polarity. The interval which may prove to have a more complex pattern of
 749 polarity reversals is around the Katian - Hirnantian boundary.

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1136

1137 **Figure Captions**

1138 Fig. 1. Location of sampled sections (block dots) and core (red dots) and other locations
 1139 discussed in text (black circle).

1140 Fig. 2. Detailed magnetic and chemostratigraphic data for the Cheney Longville section. a)
 1141 NRM intensity and magnetic susceptibility (K). b) Specimen demagnetisation behaviour
 1142 showing categorisation into good (S1) and poor (S3) ChRM line-fits; great circle fit quality
 1143 range from good (T1) to noisy (T3), and specimens with uninterpretable magnetisation, or
 1144 those entirely remagnetised are indicated in the X column (see text for details). c) Interpreted
 1145 specimen polarity quality, with those specimens in the middle column (X) not assigned a
 1146 polarity. Best quality specimens indicated as R or N and poorest quality in column headed ??.
 1147 d) Specimen VGP latitude and section magnetic polarity. Filled symbols = those specimens
 1148 possessing an S-class ChRM, and unfilled symbols for specimens with T-class, great-circle
 1149 behaviour. Polarity bar widths in the section polarity column correspond to interpreted
 1150 quality of the polarity interpretation, with full-bar width corresponding to good quality and
 1151 1/4 bar width to lowest confidence. e) Carbon isotope and total organic carbon (%TOC) data.

1152 Fig. 3. Detailed magneto and chemo-stratigraphic data for the Mójcza section. See Fig. 2 for
 1153 details of columns. Surface gamma counts are also shown in a).

1154 Fig. 4. Lithostratigraphy, surface magnetic susceptibility (K_{surf}) and carbon isotope data for
 1155 Core-A. The carbon isotope stratigraphy and lithostratigraphy for well Jurmala R-1 (Latvia,
 1156 Fig. 1) from Ainsaar et al. (2010), allows interpretation of the fragmented isotope stratigraphy
 1157 from Core A. The BC isotope zones and positive isotope peak names from Jurmala R-1 are
 1158 those of Ainsaar et al. (2010). Isotope zones on Core-A are interpreted. MDICE, GICE, HICE
 1159 are the mid Darriwilan, Guttenburg and Hirnantian isotope excursions respectively.

1160 Fig. 5. Correlation between the Ordovician in Grabowiec-6 and Zwierzyniec-1 wells, based
 1161 primarily on the surface magnetic susceptibility (K_{surf}) data between the two cores, but
 1162 assisted by the biostratigraphy. Geochemical and biostratigraphic data from Sullivan et al.
 1163 (2018). Isotope zone labels and names as in Fig. 4. Coloured bands and dotted lines represent
 1164 correlated levels and intervals.

1165 Fig. 6. Detailed magnetostratigraphic data for the Backside Beck section. See Fig. 2 for
 1166 column details. A more detailed log of sampling spots and lithological key is in SI Fig. S3.

1167 Fig. 7. Isothermal remanent magnetisation (IRM) acquisition curves for representative
 1168 specimens from the sections and cores. Specimen names (e.g. CL33 from Cheney Longville
 1169 section in a)) and sample depths marked on Figs. 2, 3, 6, 11 and 12.

1170 Fig. 8. Example demagnetisation data for specimens. In each case a) to j) consists of a
 1171 Zijderveld diagram, stereonet and intensity decay (J/J_0) plot (sometimes aligned vertically).
 1172 All in stratigraphic coordinates, except a) which is in geographic coordinates. Points plotted
 1173 in black are thermal demagnetisation and in blue are AF demagnetisation steps. Key to
 1174 Zijderveld and stereonets in b). Where can be clearly shown, the Zijderveld plot has the
 1175 ranges over which components extracted are shown with coloured arrows, and on the
 1176 stereonet the direction of the associated components (key under a)). Measurement confidence
 1177 cones around some steps shown when γ_{95} (Briden and Arthur, 1981) is $>20^\circ$. A) Backside
 1178 Beck, BB54.2 (66.8 m, demagnetisation type=T1, polarity rating=N) shows a Kiaman
 1179 component from 100°C-340 (40 mT) and normal-polarity ChRM from 340 (40 mT) to
 1180 origin. The Kiaman component in geographic coordinates is shown and the ChRM in
 1181 stratigraphic coordinates are shown on the stereonet. B) Cheney Longville (CL1.2, S3, N). A
 1182 three component magnetisation with a dominant Brunhes-component NRM-250°C, Kiaman-
 1183 like component 415-440 (15 to 40 mT) and a normal-polarity ChRM from 450 (50 mT) to the
 1184 origin. C) Cheney Longville (CL20.1, S3, R), height= 10.23 m. A two component
 1185 magnetisation with a dominant Brunhes component NRM-200°C, and a reverse-polarity
 1186 ChRM from 430 (30 mT) to the origin. D) Mójcza (MOZ-4y), height = 0.75 m. Three
 1187 component magnetisation, with a Brunhes-component NRM-250°C, Devonian-like
 1188 component 250-400°C, and a much weaker reverse-polarity ChRM, extracted during AF
 1189 steps 405-425 (5 to 25 mT). The remaining undemagnetised direction is Brunhes-like. E)
 1190 Mójcza (MOZ-1y, T1, N), height = 0.15 m. Three component magnetisation, with a Brunhes
 1191 component 150-225°C, a Kiaman component 225-350°C, with the last steps showing a partial
 1192 great circle trend to normal-polarity ChRM (demagnetisation class T1). F) Grabowiece-6
 1193 (3809.67 m, S3, R?). Three component magnetisation with a strong Brunhes component 100-
 1194 210 °C, Kiaman component 400-430°C, and a reverse-polarity ChRM from 430°C to origin.
 1195 G) Grabowiece-6 (3806.77 m, S3, N). Three component magnetisation with a strong Brunhes
 1196 component, NRM-300°C; Kiaman component 300-530 (30 mT) and a normal-polarity ChRM
 1197 from 560 (60 mT) to the origin. H) Core-A (69.15 m, S3, N). A two component
 1198 magnetisation with a 'Brunhes-like' component, NRM-150°C (not used to orient core) and a
 1199 normal-polarity ChRM component from 440°C to the origin. Intermediate steps show strong
 1200 component overlap. Thermal steps above 460°C have been removed due to large thermal
 1201 alteration. I) Core-A (4.13 m, S2, R). A two component magnetisation with a dominant
 1202 Brunhes component, NRM-200°C, and reverse-polarity ChRM from 415 (15 mT) to the
 1203 origin. J) Core-A (62.38 m, S1, N). A two component magnetisation with a Brunhes
 1204 component, NRM-150°C and a normal-polarity ChRM from 250-525°C.

1205 Fig. 9. Characteristic remanence (ChRM) directions for the Cheney Longville (top panel)
 1206 and Mójcza (bottom panel) sections. In situ directions shown on the left, bedding corrected in

1207 the middle and bedding corrected poles to remagnetisation circles on the right. The planes
1208 through the ChRM great circle poles have a pole which is near the mean of the ChRM line-fit
1209 directions in each site (or age interval for Mójcza). The great circle line-fit ChRM plane
1210 (derived from the S-class data) is shown on the right most stereonets. S1 to S3 indicate the
1211 demagnetisation behaviour explained in the text. Symbol keys for stereonets in box.

1212 Fig. 10. Topmost panel data for Grabowiec-6 and bottom panel for Core-A. a) Kiaman
1213 component in Grabowiec-6 re-oriented using the LT component (i.e. rotated to 0°
1214 declination). b), e) Low stability (LT) components for Grabowiec-6 and Core-A, with in b),
1215 the re-oriented LT component using the Kiaman component (for Grabowiec-6). c), f) ChRM
1216 directions re-oriented using the Kiaman and LT component respectively. d) Poles to great
1217 circles for the T-class ChRM from Grabowiec-6 (see Fig. 9 caption). g) Mean ChRM
1218 inclination, 95% confidence cone (α_{95}), Fisher k, and n number of specimens of Baltican
1219 stage intervals for Core-A.

1220 Fig. 11. Detailed magnetostratigraphic data for Core-A. See Fig. 2 for column details. The
1221 intervals used for the mean inclinations are those in Fig. 10g. Information about core-
1222 reorientation of directions in Fig. 10f. Width of bar for mean inclination is $2*\alpha_{95}$.

1223 Fig. 12. Detailed magnetostratigraphic data for the Grabowiec-6 core. See Fig. 2 for column
1224 details. Biostratigraphic data from Sullivan et al. (2018). Stages, $\delta^{13}\text{C}$ zones of Ainsaar et al.
1225 (2010) and chitinozoans from Fig. 5.

1226 Fig. 13. Summary polarity chronoscale for the Dapingian to Hirnantian (rightmost) and the
1227 source polarity data. Each section or core has its own meter scale. Magnetic polarity data
1228 from: lower part of Mójcza section from Schätz et al. (2006); Gullhögen from Torsvik et al.
1229 (1995); Pakri Cape and Vao Quarry from Grappone et al. (2017); Popovka, Tosna, Lava from
1230 Smethhurst et al. (1998); Polovinka from Pavlov et al. (1999); Moyera from Gallet & Pavlov
1231 (1996); Kudrino from Pavlov et al. (2008); Rozhkova from Pavlov et al. (2012). Buffalo and
1232 White River from Farr et al. (1993). The exclusively reverse polarity of the Polovinka and
1233 Moyero sections below the Siberian Volginian Stage is not shown. See section 5 for details.

1234

Section/core [code]	Age/ mean type	Mean Dec/Inc (°)	k/ α_{95} (°)	N _s /N _T /N _l /N _p	Reversal Test [γ_o/γ_c] (°)	VGP Pole Long./Lat. (°)	D _p /D _m (°)	A95 (min, max), %VGP ₄₅
Backside Beck [BB]	Ordovician/ GC mean	31.7/-17.1	12.1/6.9	7/21/3/9	Ro [22.1, 20.3]	323.6/-21.5	3.7/7.1	8.1 (5.5, 24.1), 0
Cheney	Fisher mean	338.7/-57.7	17.4/6.2	13/68/33/0	Rc [9.5/17.8]	13.7/2.7	6.7/9.1	11.9(4.2, 15.6)
Longville [CL]	GC mean	341.5/-58.2	12.2/4.3	34/68/33/19	Rc [7.0/14.2]	11.5/2.8	4.7/6.4	7.1(2.9, 8.9), 1.5
Möjca [M]	Silurian/ GC mean	35.3/-19.9	8.8/15.2	8/28/4/9	n.p	343/-21.7	8.3/15.9	13.2 (5.2, 22)
“	Ordovician/ GC mean	3.8/-53.3	10.7/5.6	10/26/10/9	R-[24.7/14.1]	17.5/-5.2	5.4/7.8	10.5 (4.8, 19.2), 9
Grabowiec-6	Fisher mean	27/-41.2	6.6/21.7	6/46/9/0	n.p	-	-	30 (5.6, 26.5)
“	GC mean	29.6/-37.0	5.8/9.0	20/46/9/11	R-[25.2/22.6]	355.0/-13.8	6.2/10.6	12.0 (3.6, 12.4), 10
Core-A	Pirgu-Porkuni/Fisher Mean	14.6/-33.6	7.3/9.4	37/59/37/0	R-[22.4/17.7]	7.8/-16.5	6.1/10.7	9.9 (2.8, 8.4), 17
“	Uhaku-Vormsi/ Fisher Mean	359.5/-59.5	8.7/9.6	29/36/29/0	R _l [13.9/20.7]	22.6/-4.2	8.5/12.8	11.6 (3.1, 9.8)
“	Billingen-Kunda/Fisher Mean	344.0/-51.6	6.6/13.4	21/29/21/0	Ro[40/42.5]	35.7/-1.2	12.4/18.3	13.7 (3.6, 12.1)

Table 1. Mean palaeomagnetic directions for the sections and cores. Code=sample code. k/α_{95} = Fisher concentration parameter and 95% cone of confidence. N_s= number of levels (sites) used, N_T=Total specimens measured, N_l=Number of specimens with line-fits, N_p=Number of specimens with great circle (GC) fits. GC means determined with method of McFadden and McElhinney (1988) for N_l+N_p data and Fisher mean for N_l data. For the reversal test (McFadden and McElhinney, 1990), γ_o =observed angle; γ_c =critical angle. Virtual geomagnetic pole (VGP) is the normal pole. N.p= not possible. A95 (min, max) = Fisher 95% confidence interval for VGP-based site mean (N_s sites), and A95_{min} and A95_{max} threshold values of Deenen et al. (2011). %VGP₄₅= percent of samples yielding VGP latitude < |45|, as a reflection of the match to modern geomagnetic field models and palaeomagnetic data in which %VGP₄₅ is a 3-4% (Cromwell et al., 2018). %VGP₄₅ applies to all the section or core. Statistics determined with Pmagtool v.5 (Hounslow, 2006).

Section	Proportional %uf [lower, upper]	Direction- correction %uf [confidence interval]	McFadden %0 [f_d, %P_f], 100% [f_d, %P_f], Ng
Backside Beck (N=37)	100 [66, 147]	132 [± 34]	0 [5.7, 0.5], 100[1.5, 23],2

Table 2. Fold test data for the Backside Beck section (N=number of data used). The three right hand columns indicate results from three types of fold tests, the proportional (Tauxe & Watson, 1994), the direction correction (DC) fold test (Enkin, 2003) and the ‘means’ fold test of McFadden (1998). The proportional and DC fold tests display the 95% confidence interval on the degree of unfolding. %uf= best degree of unfolding, values in [...] indicate the unfolding % of the confidence interval. Bootstrap confidence intervals use 2000 simulations. N= number of groups defined in McFadden fold test. Uses combined Silurian and Ordovician data (Silurian data from Hounslow et al. submitted-a). In the McFadden fold test the value of P_f indicates the probability the magnetisation was acquired at that percent folding, so that when P_f exceeds 5% it suggests it could have been acquired at that state (only 0% and 100% shown). Ng=number of bedding attitude groups in McFadden fold test. F_d is the f-statistic in the McFadden fold test. Tests determined with the Pmagtool software (Hounslow, 2006), and further details on these tests are in SI Figs. S10 to S13 and Table S3.

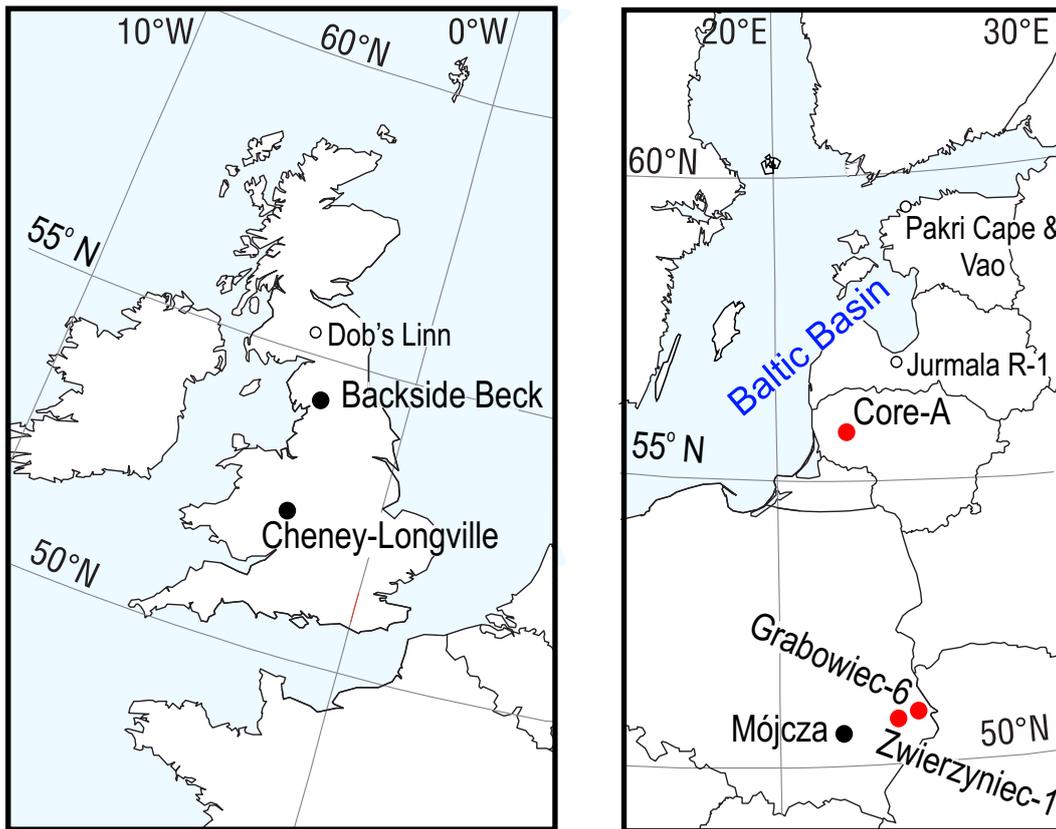


Fig. 1.

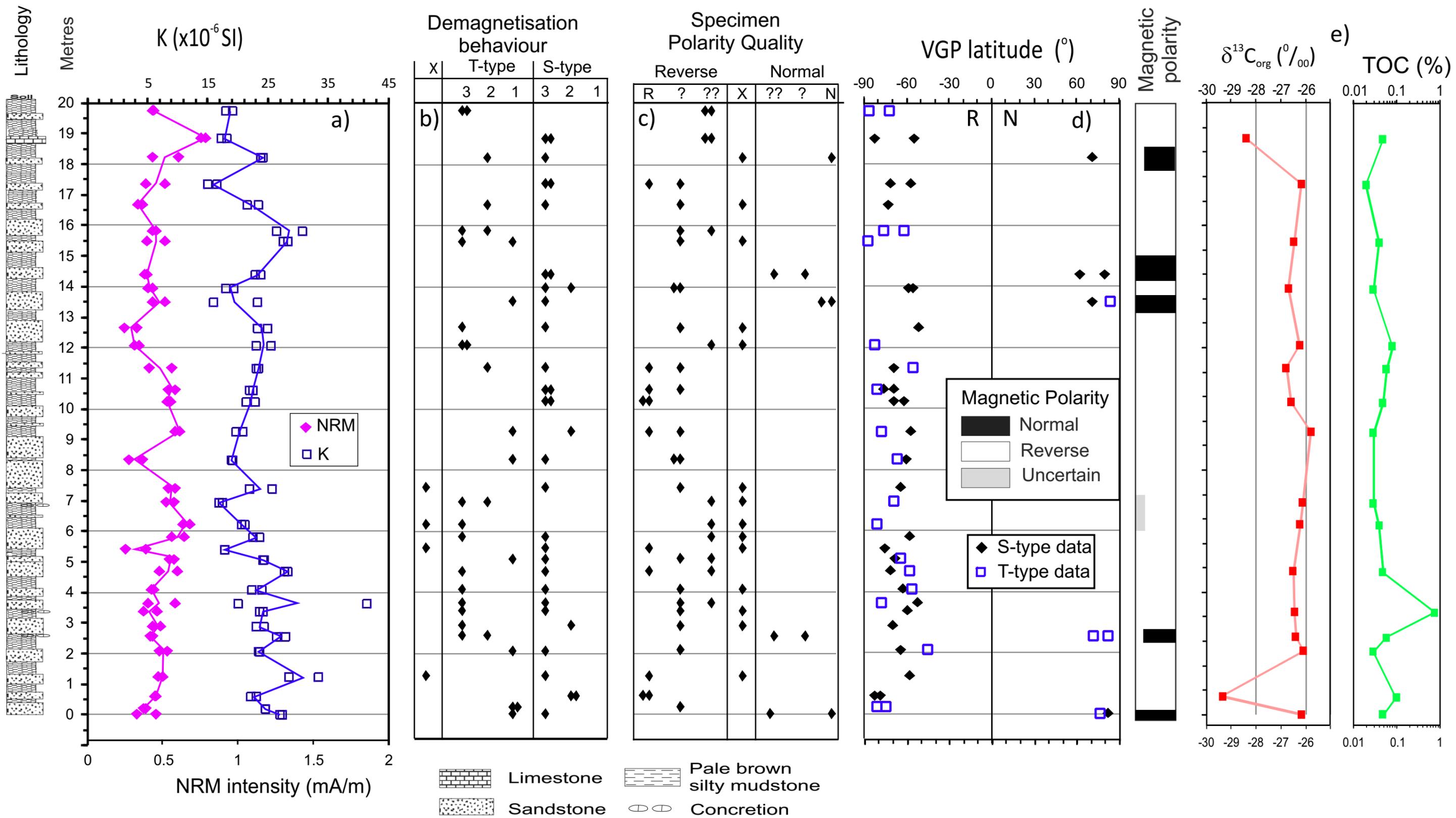
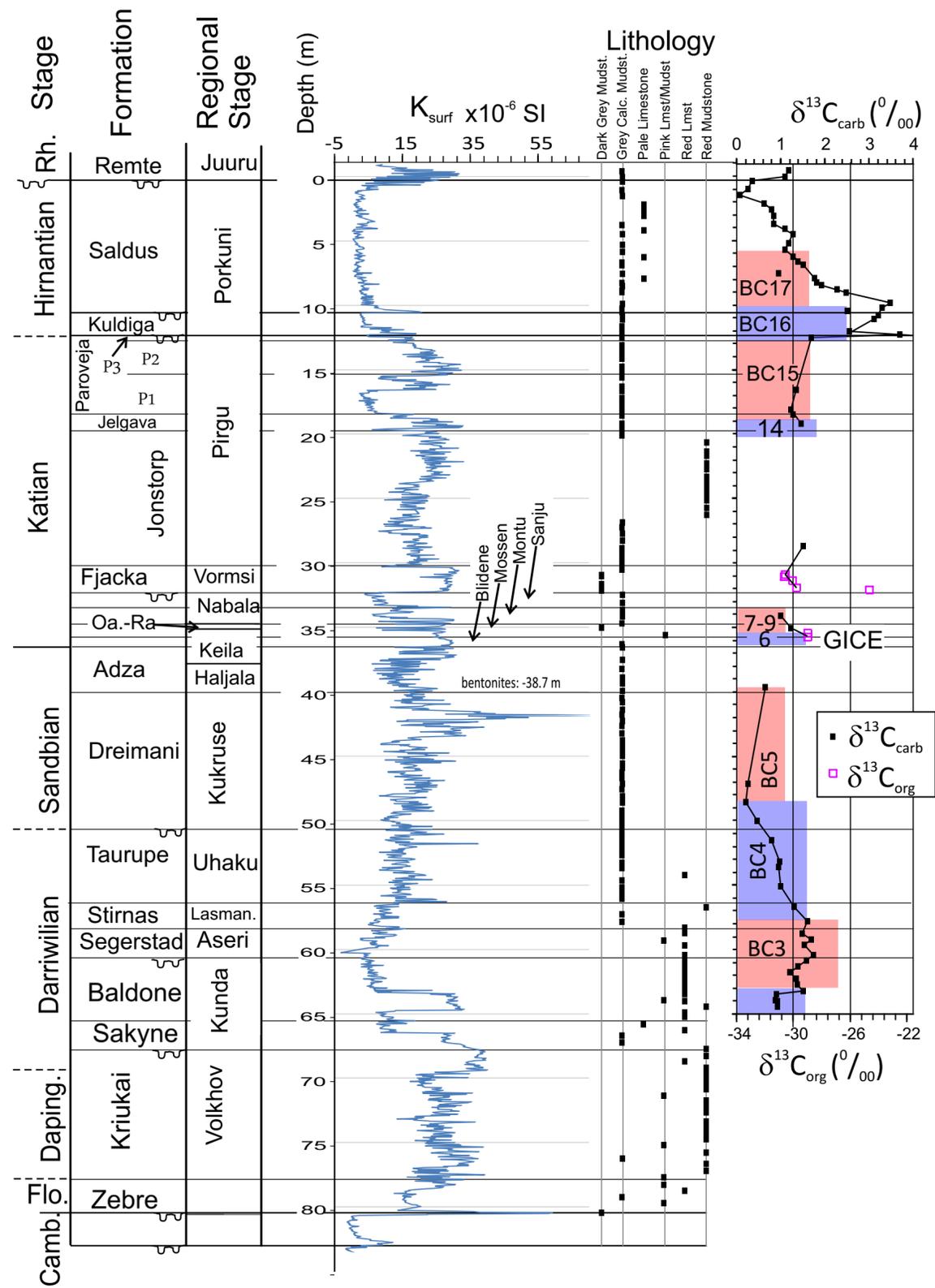


Fig. 2.



JURMALA R-1

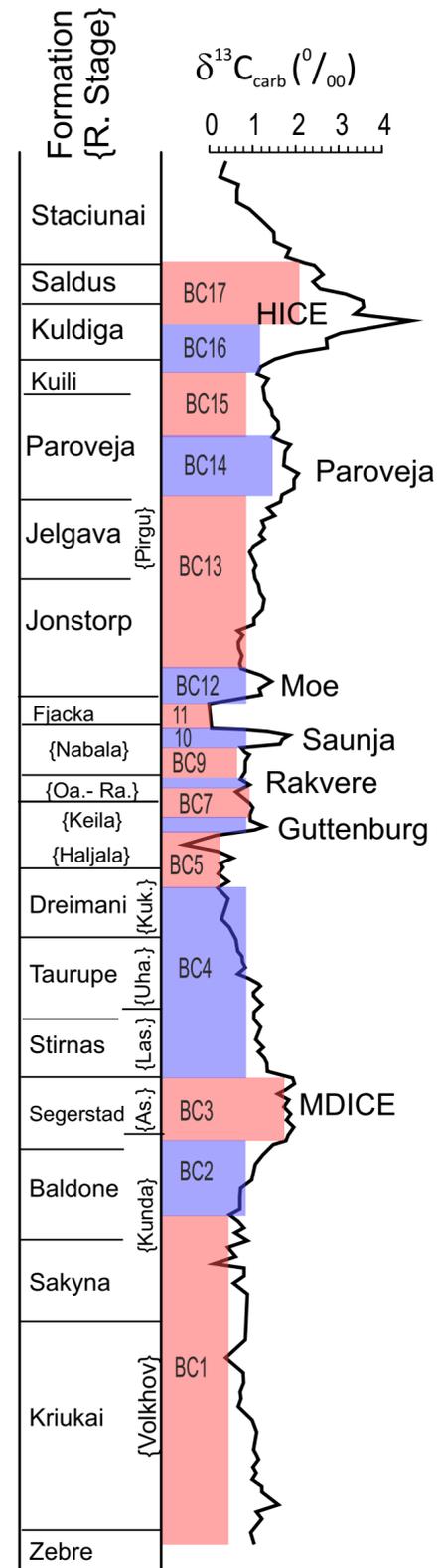


Fig. 4.

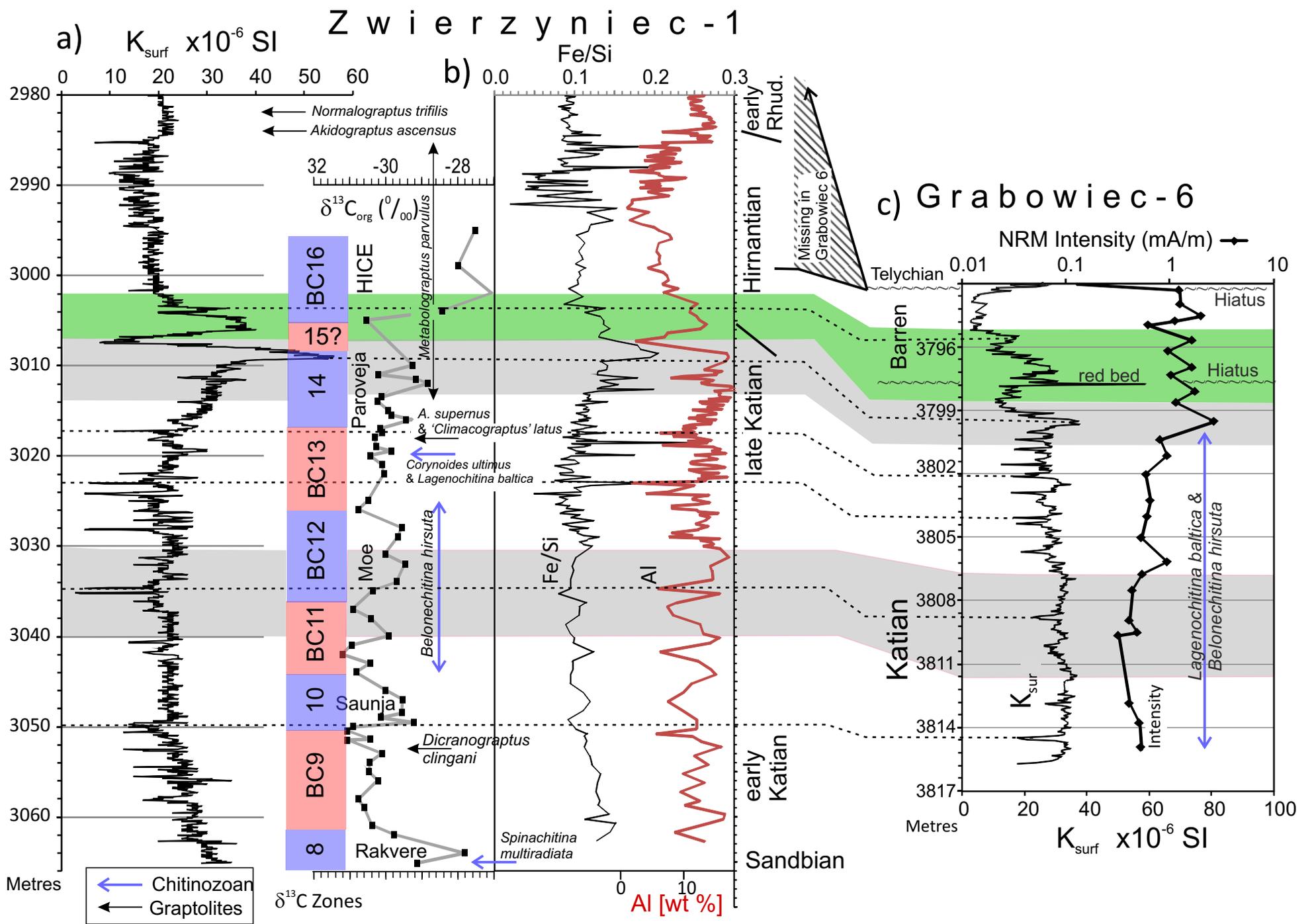


Fig. 5.

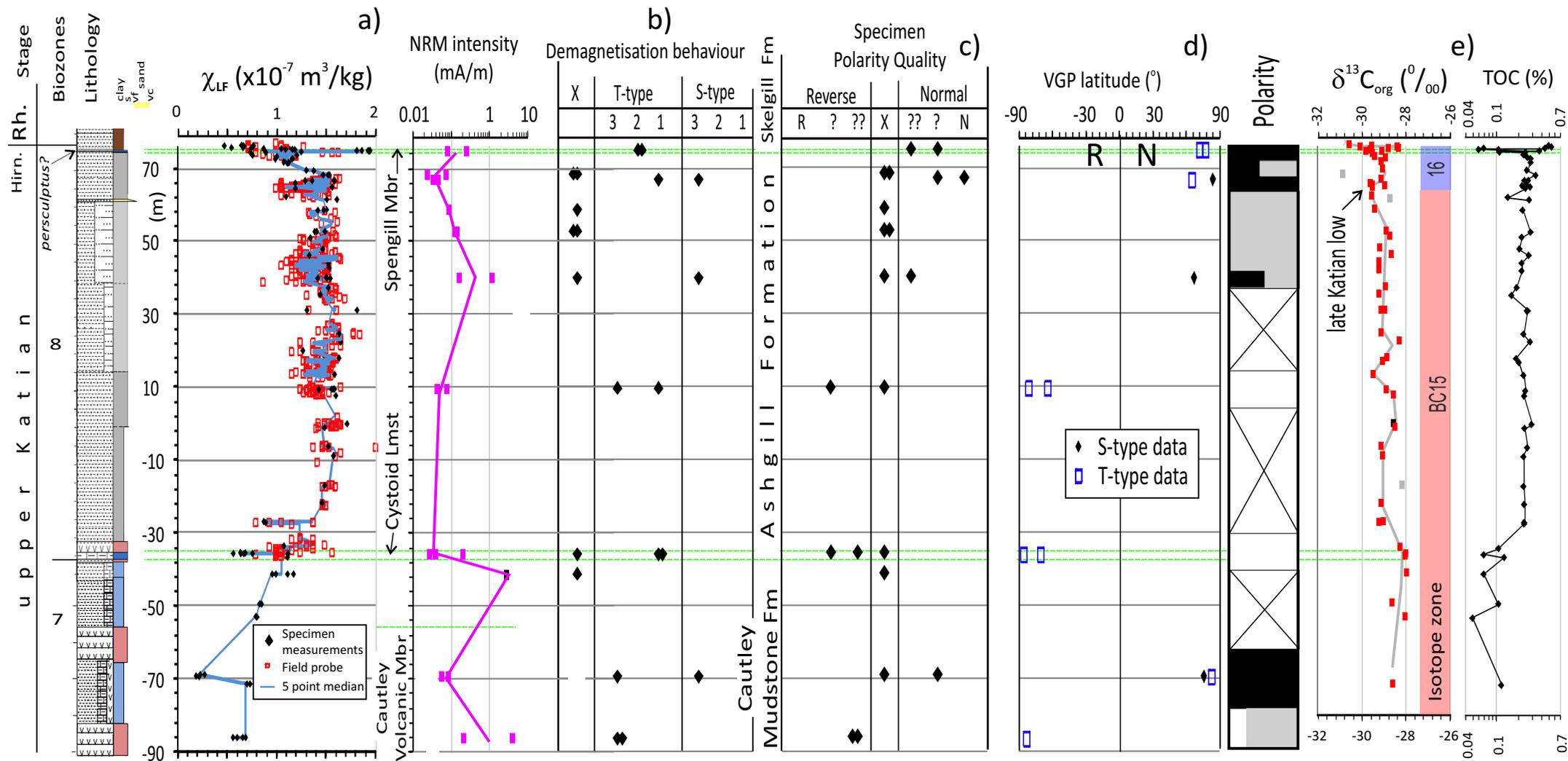


Fig. 6.

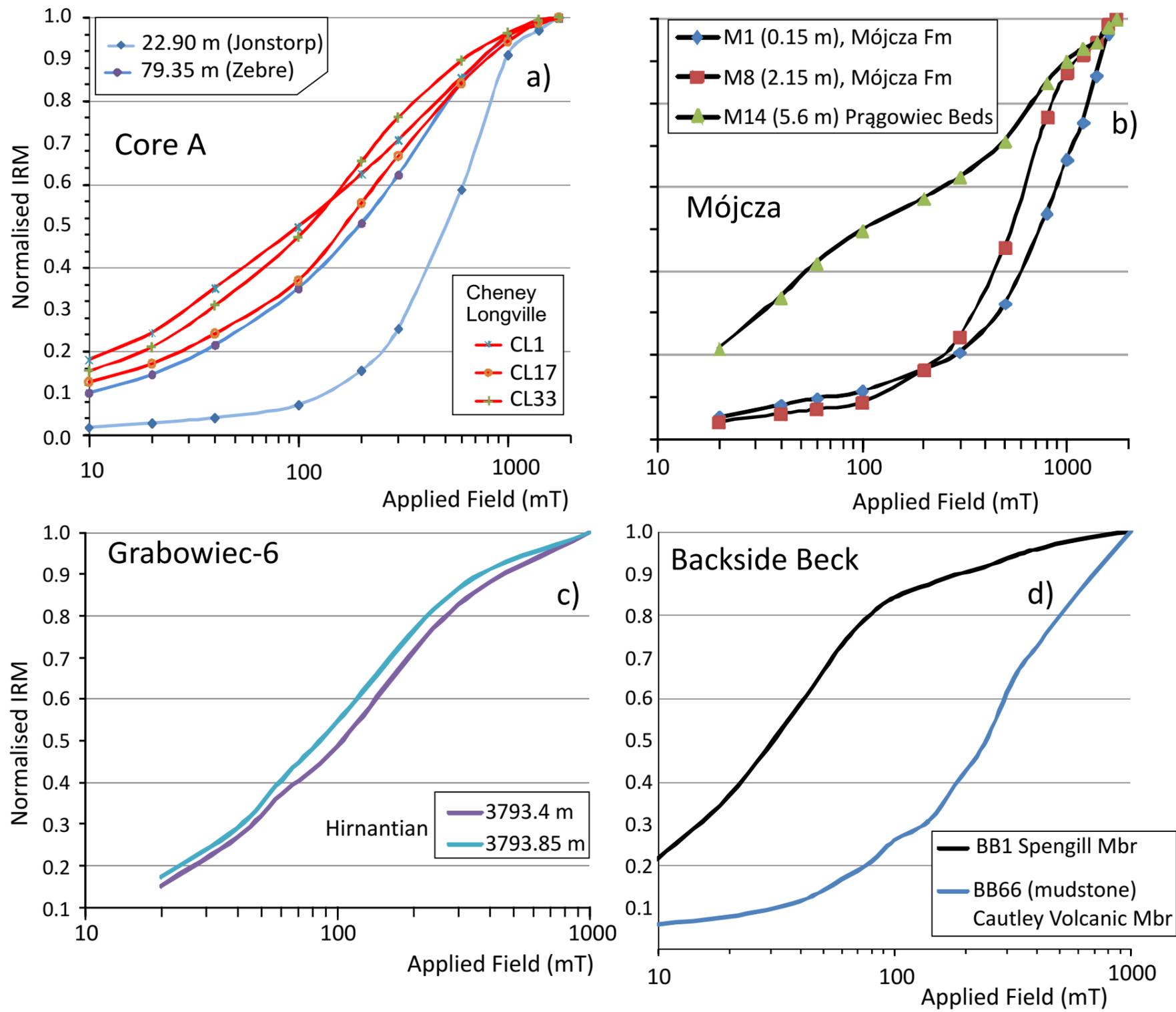


Fig. 7.

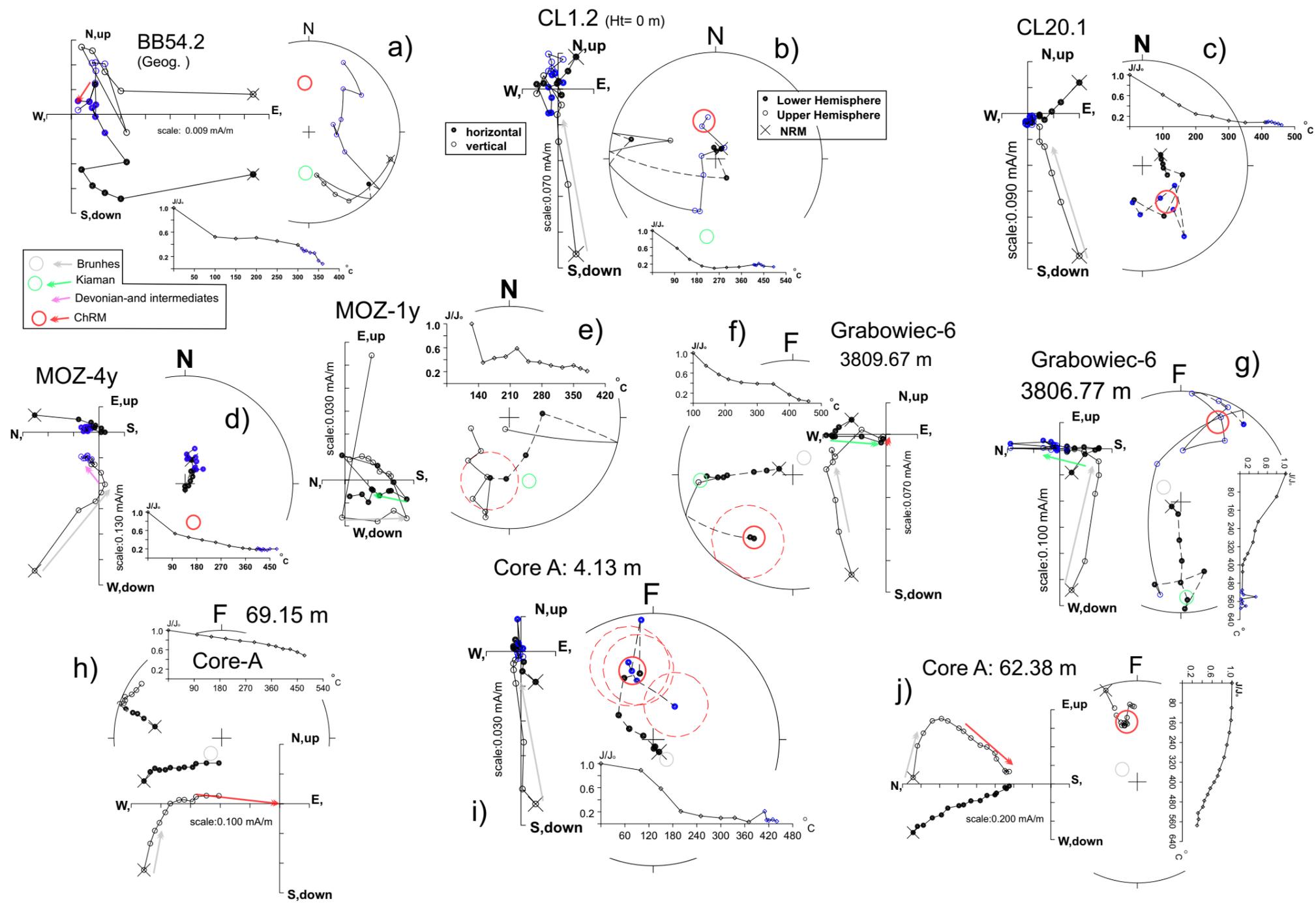


Fig. 8.

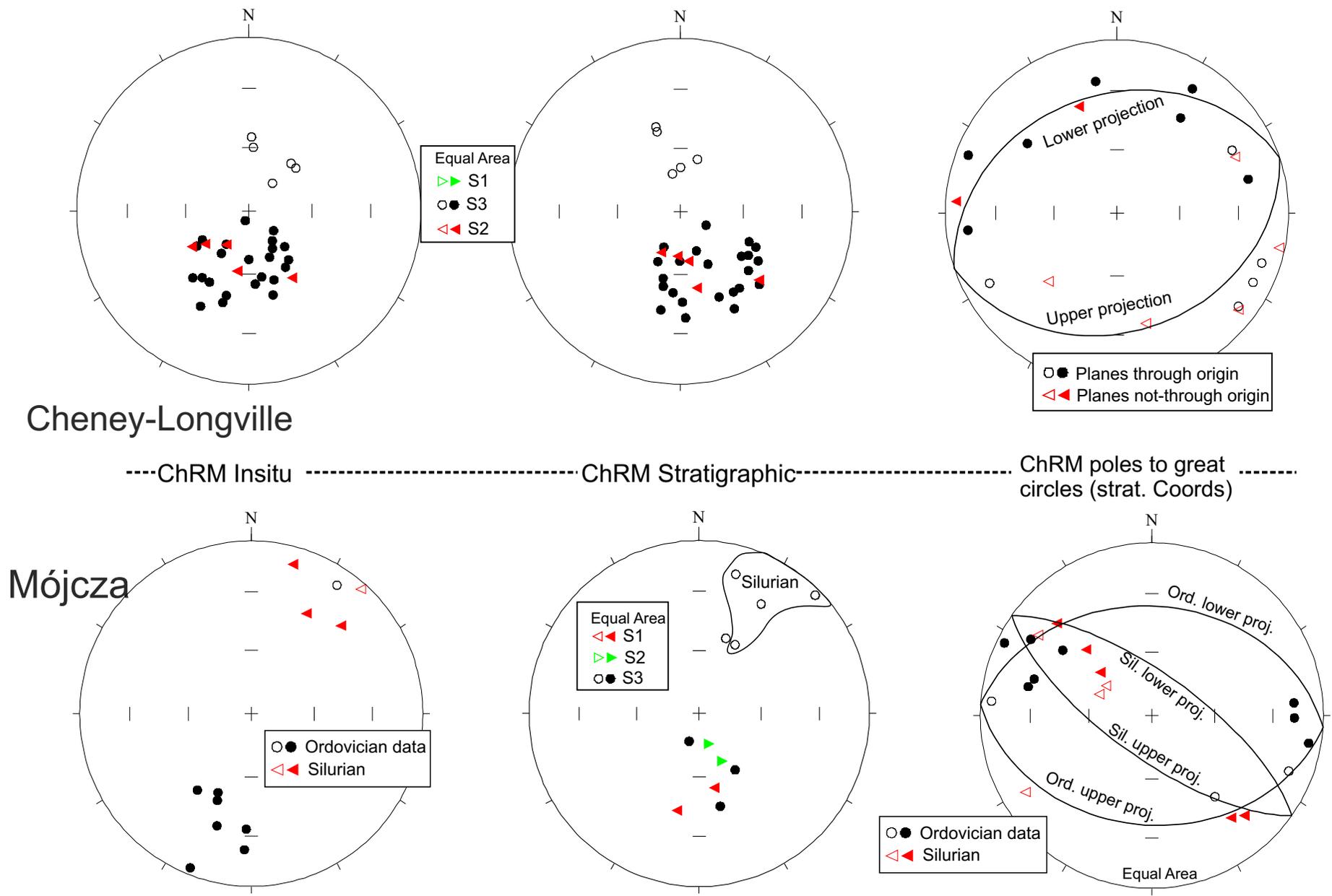
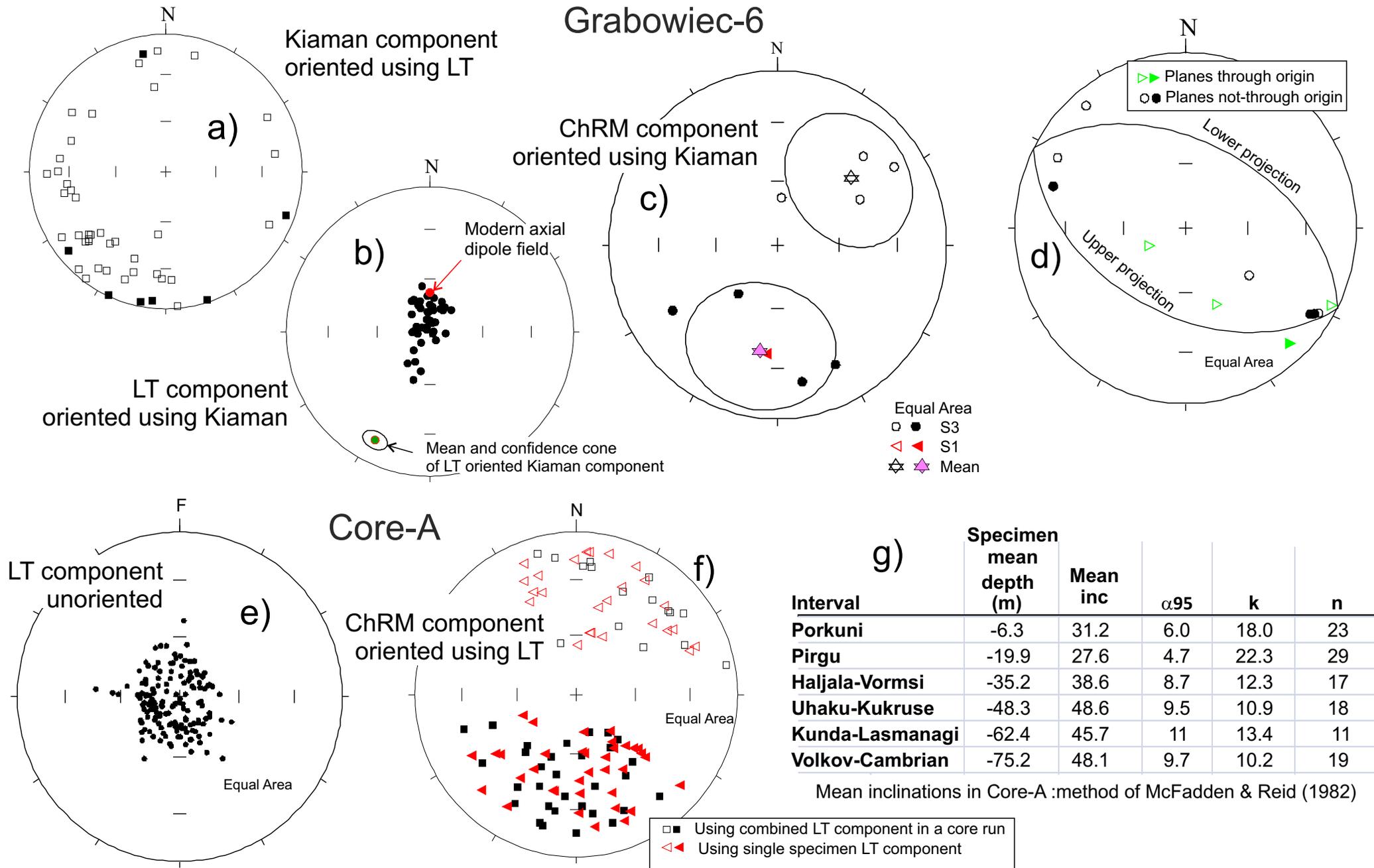


Fig. 9. .



Interval	Specimen mean depth (m)	Mean inc	α_{95}	k	n
Porkuni	-6.3	31.2	6.0	18.0	23
Pirgu	-19.9	27.6	4.7	22.3	29
Haljala-Vormsi	-35.2	38.6	8.7	12.3	17
Uhaku-Kukruse	-48.3	48.6	9.5	10.9	18
Kunda-Lasmanagi	-62.4	45.7	11	13.4	11
Volkov-Cambrian	-75.2	48.1	9.7	10.2	19

Mean inclinations in Core-A :method of McFadden & Reid (1982)

Fig. 10.

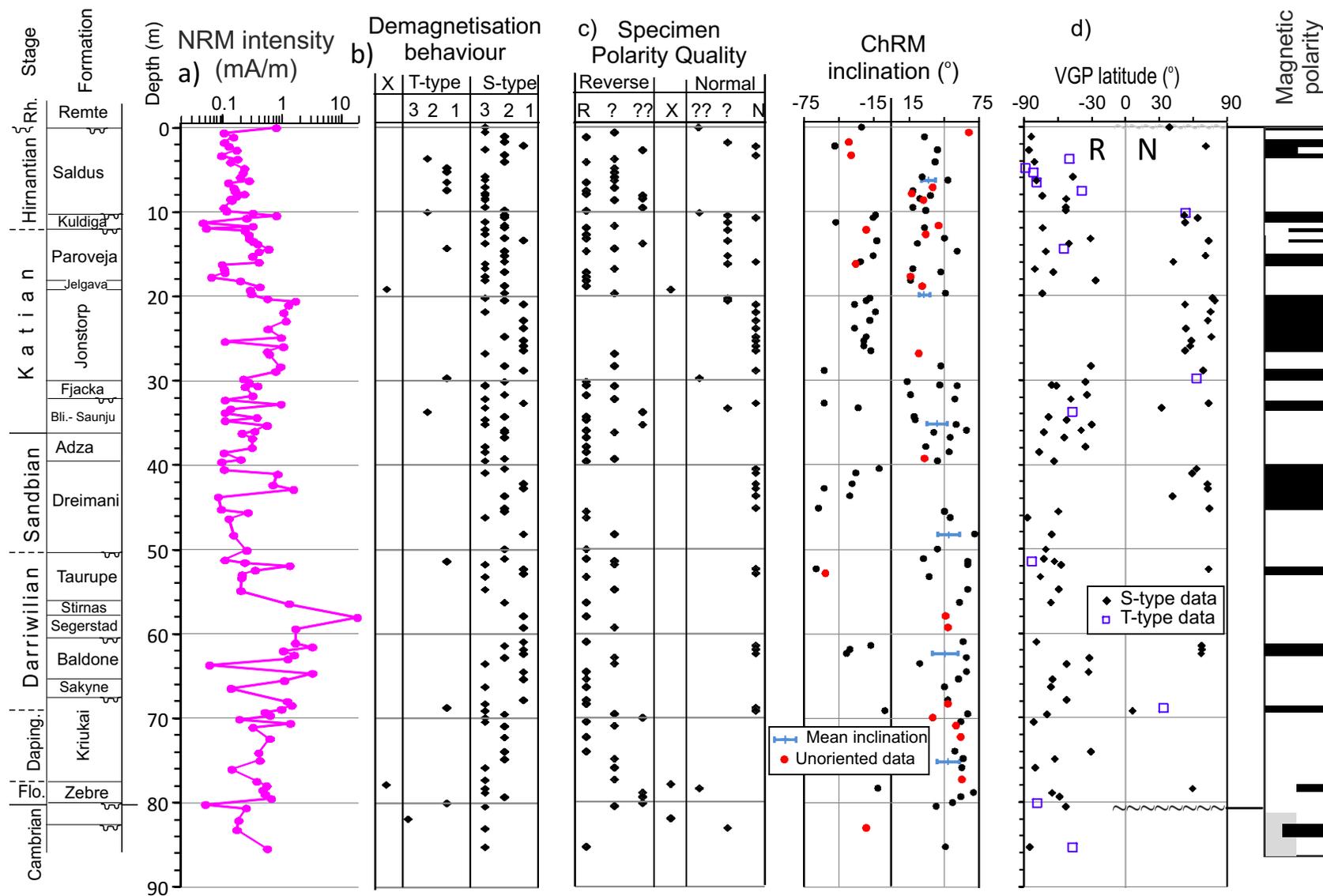


Fig. 11.

Supplementary Information for: A geomagnetic polarity stratigraphy for the Middle and Upper Ordovician

By: Mark W. Hounslow, Samuel, E. Harris, Krystian Wójcik, Jerzy Nawrocki, Kenneth T. Ratcliffe, Nigel H. Woodcock, Paul Montgomery

The supplementary information contains the following:

- a) Details of sampling locations both on maps and in more detailed sedimentary logs (Figs. S1 to S3).
- b) Additional rock magnetic data (Fig. S4).
- c) Carbon isotope and organic matter concentration data for the Cheney Longville and Mójcza sections and Core-A (Tables S1, S2).
- d) Anisotropy of magnetic susceptibility data for Grabowiec-6 (Fig. S5)
- e) Palaeomagnetic overprint components from sections and cores (Figs. S6, S7).
- f) Palaeomagnetic Ordovician ChRM components from the Backside Beck section show with the Silurian data from the section (Fig. S8).
- g) Anisotropy of magnetic susceptibility (AMS) data for the Ordovician at Backside Beck (Fig. S9).
- h) Additional data pertinent to the fold tests at Backside Beck (Figs. S10 to S13, Table S3).
- i) Additional information about how the demagnetisation data was analysed with the LINEFIND software of Kent et al. (1983) (Fig. S14 to S16), and information on the statistics of the fitted lines and planes in Table S3.
- j) Compilation of the palaeomagnetic data for each specimen from all sections (in the associated excel file).

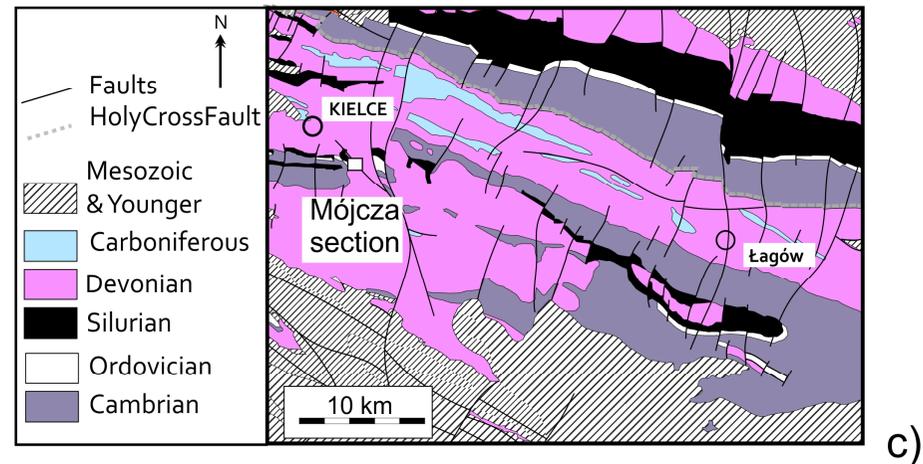
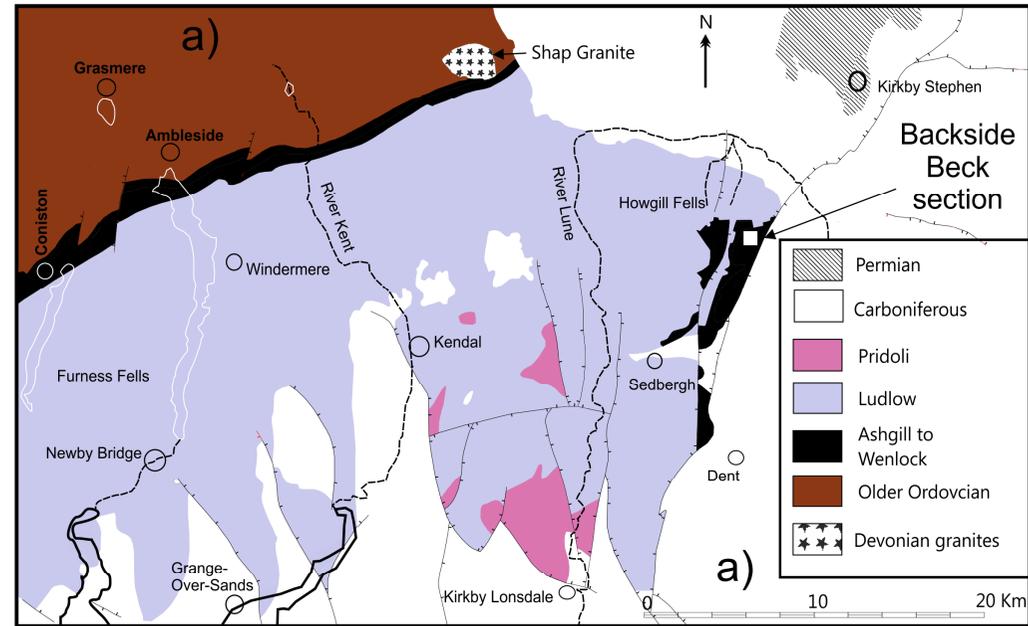
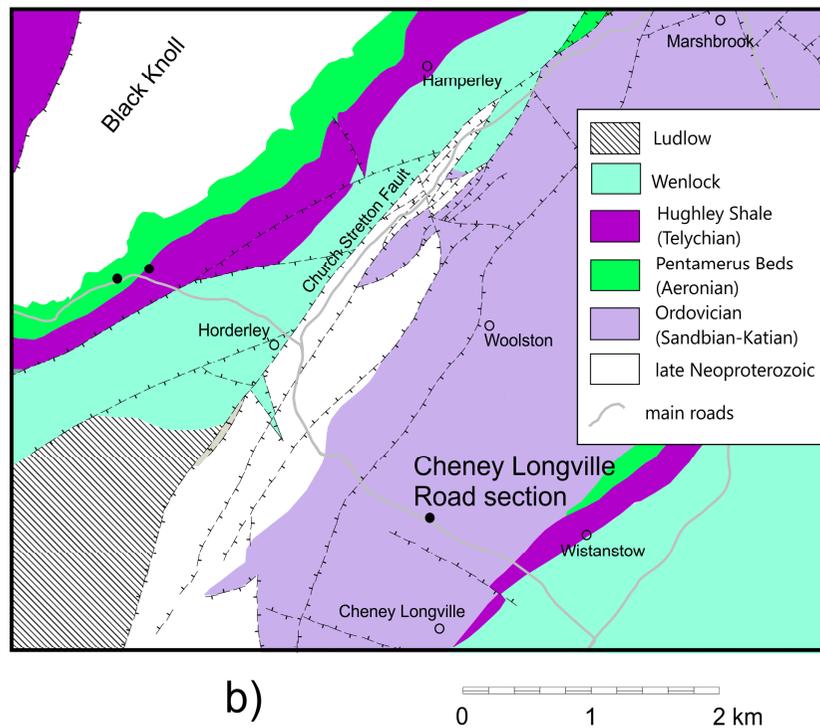


Fig. S1. a), b) and c) Simplified geological maps of the sampling sites (unfilled squares in a and c). a) the Backside Beck section, Howgill Fells, UK, b) Cheney Longville road section, Shropshire, UK and c) Mójcza sections, Holy Cross Mountains, Poland; Holy Cross Mountain map after Rühle (1977), Kowalczewski et al. (1990).

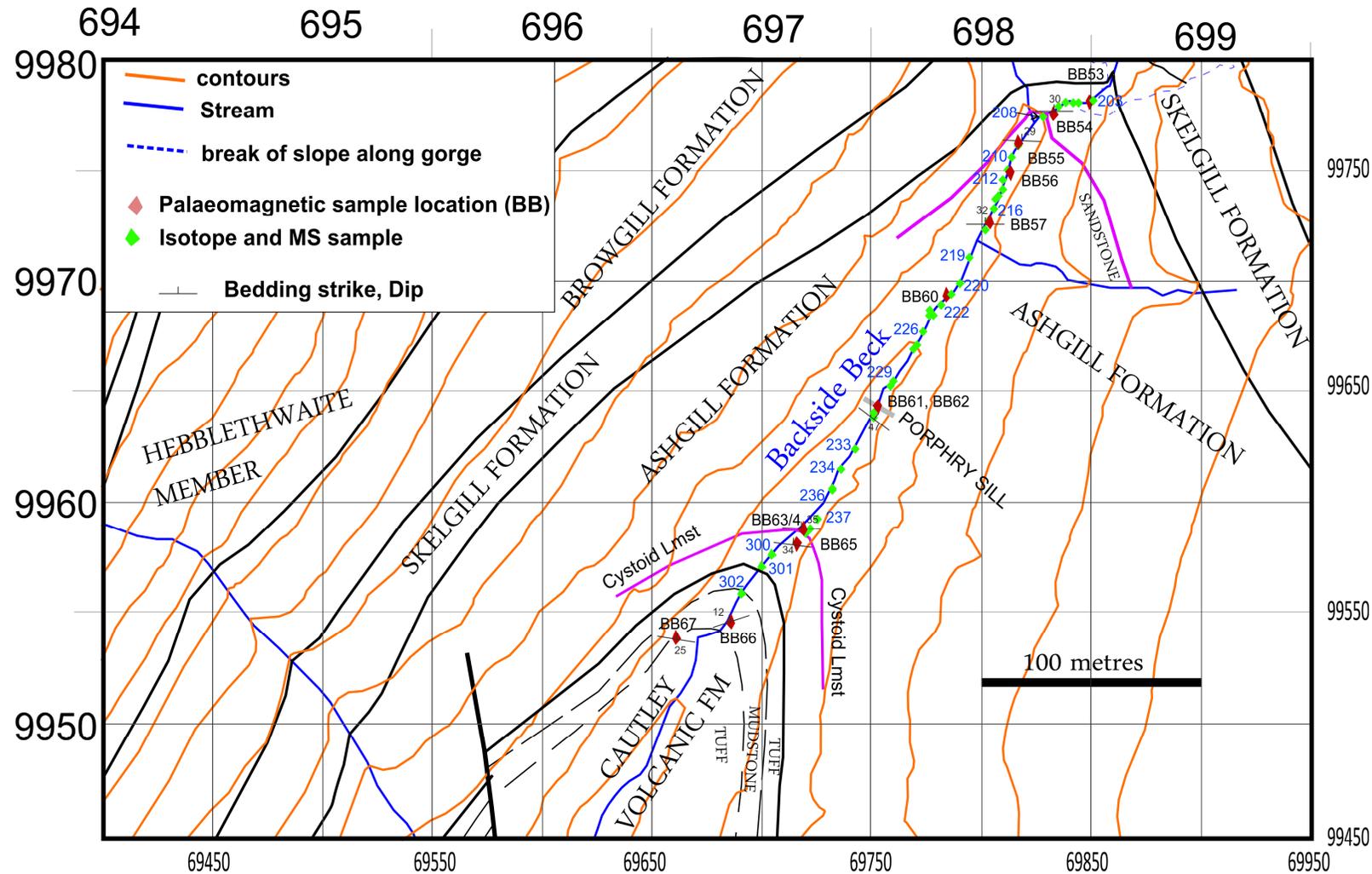


Fig. S2. Detail of the sampling locations of the palaeomagnetic and carbon isotope plus magnetic susceptibility (MS) samples in the Ordovician of the Backside Beck Section, UK. The coordinates are the British National Grid. The bedding strike displayed assumes the grid is true north-directed. Base geology map based on Woodcock and Rickards (2006).

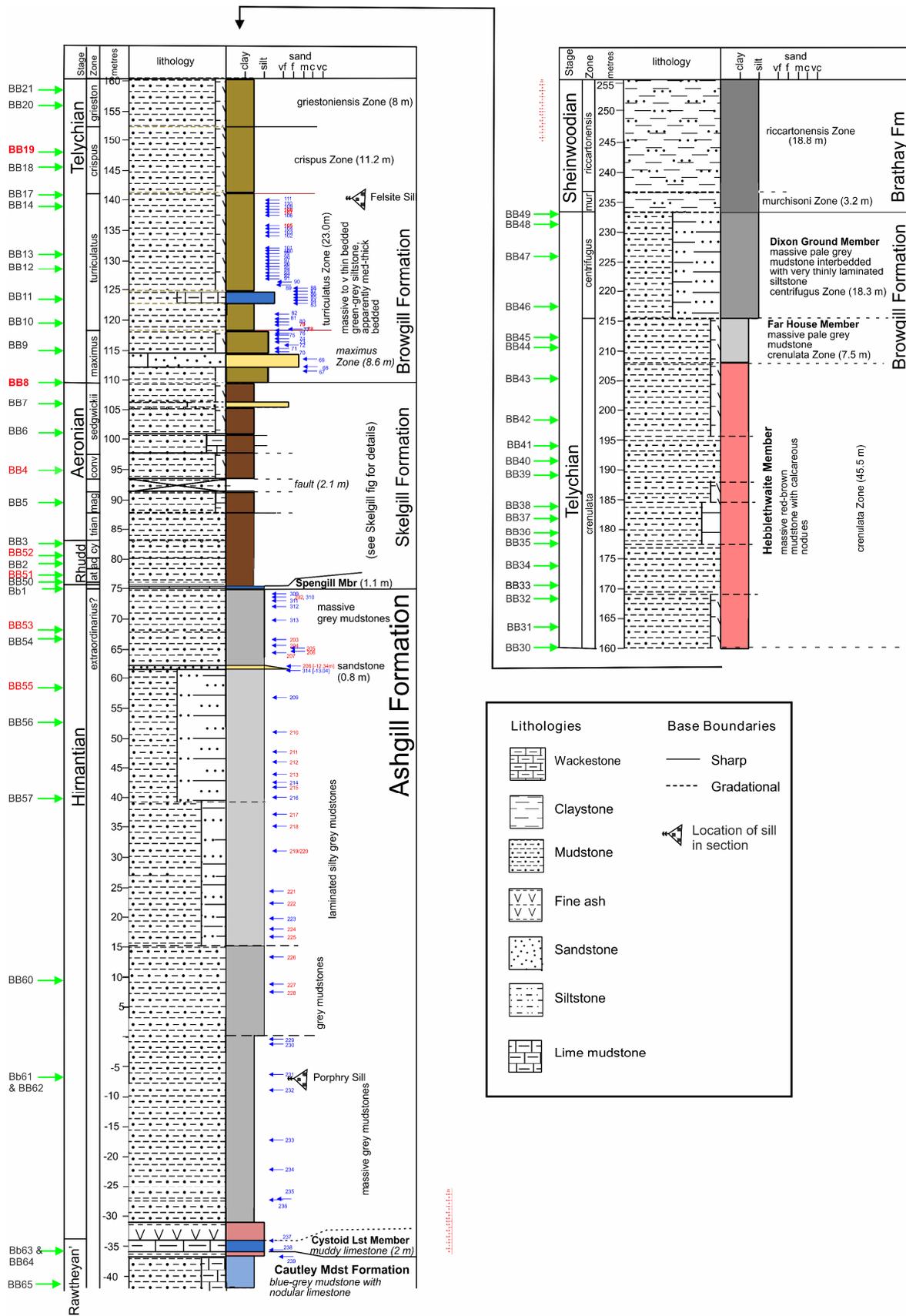


Fig. S3. Sampling positions of the palaeomagnetic (green arrows) and carbon isotope (blue arrows) samples in the log of the Backside Beck Section, UK. The carbon isotope and MS data from the Silurian part of the section and the Silurian palaeomagnetic data in Hounslow et al. (2021). Graptolite zones based on Rickards (1970, 1989), Woodcock & Rickards (2006).

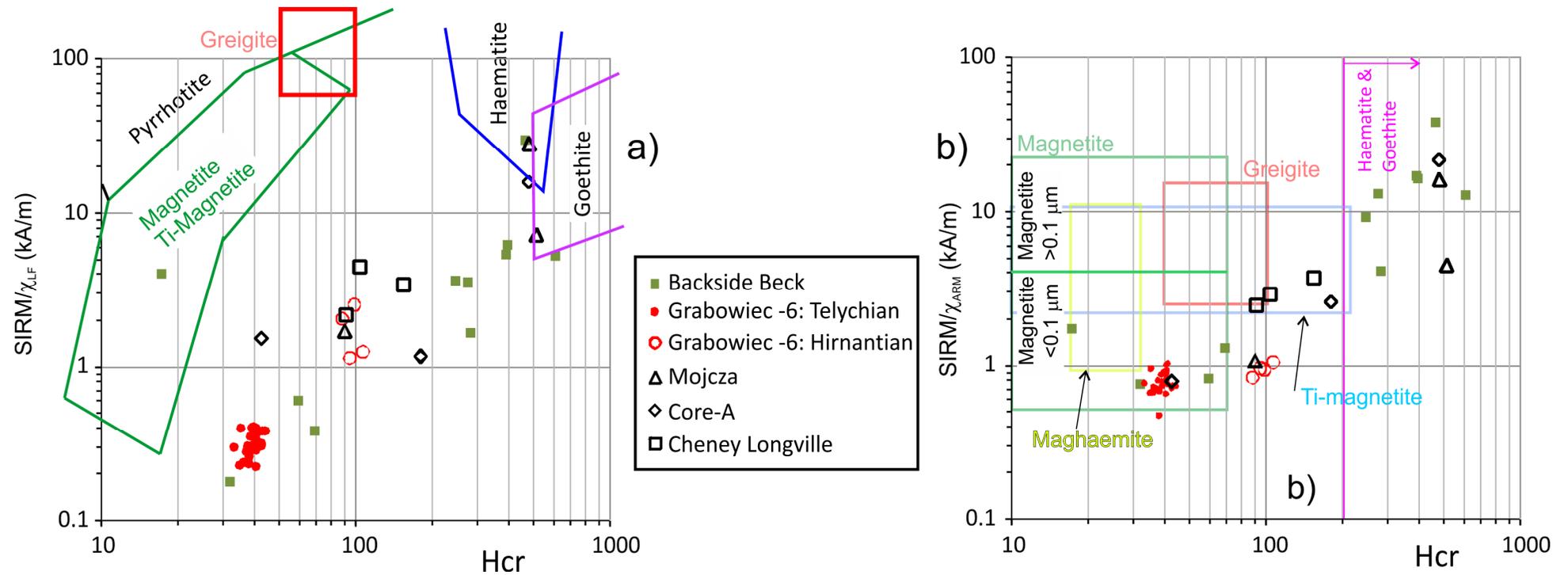


Fig. S4. Additional rock magnetic data. a) Data showing H_{cr} and $SIRM/\chi_{LF}$, along with the discrimination fields of Peters and Thompson (1998; based on natural minerals). The large range in H_{cr} suggests haematite dominates and makes a substantial contribution in some samples (with $H_{cr} > 0.2$ T, mostly red or reddish samples), and lower H_{cr} (< 0.1 T) in other drab coloured lithologies. Intermediate H_{cr} likely indicates mixtures between these two extremes. Our H_{cr} data < 100 mT falls largely outside magnetite or magnetic sulphide envelopes in a), due to a substantial paramagnetic contribution to χ_{LF} , which lowers the $SIRM/\chi_{LF}$ by a factor of $1/(1-p\%)$, where $p\%$ is the paramagnetic contribution to χ_{LF} . B) H_{cr} and $SIRM/\chi_{ARM}$ data with the coloured boxes demarcating the limits of these parameters for different minerals (natural and synthetics) from Peters and Dekkers (2003). Using this data it seems most likely that magnetite is the low H_{cr} mineral with probable particle sizes $< 0.1\mu\text{m}$.

Sample code	%Nitrogen (wt%)	%Carbon (wt%)	$\delta^{13}\text{C}_{\text{org}}$ VPDB (‰)	Height (m)
Cheney Longville				
CL-33	0.02	0.05	-28.38	18.83
CL-31	0.01	0.02	-26.18	17.35
CL-28	0.02	0.04	-26.46	15.46
CL-26	0.02	0.03	-26.69	13.93
CL-23	0.04	0.08	-26.24	12.05
CL-22	0.03	0.06	-26.78	11.32
CL-20	0.03	0.05	-26.59	10.23
CL-19	0.02	0.03	-25.80	9.24
CL-16	0.02	0.03	-26.12	6.94
CL-15	0.02	0.04	-26.22	6.21
CL-11	0.02	0.05	-26.50	4.67
CL-8	0.08	0.76	-26.45	3.36
CL-6	0.02	0.06	-26.39	2.55
CL-5	0.02	0.03	-26.08	2.07
CL-3	0.03	0.10	-29.34	0.59
CL-1	0.02	0.05	-26.17	0.00
Backside Beck Ordovician: Cautley Mudstone Fm, Ashgill Fm, Spengill Mbr.				
304	0.08	0.07	-28.37	75.45
305	0.08	0.06	-28.33	75.40
306	0.08	0.11	-28.82	75.20
307	0.09	0.38	-29.59	74.75
308	0.10	0.25	-29.09	74.45
311	0.12	0.29	-28.94	72.72
312	0.11	0.29	-29.15	71.74
313	0.11	0.26	-29.04	69.43
314	0.10	0.28	-28.74	61.40
BB56	0.09	0.29	-28.90	52.60
229	0.09	0.30	-28.56	-0.27
237	0.05	0.11	-28.25	-34.08
BB64	0.04	0.07	-28.00	-35.80
239	0.05	0.13	-28.03	-36.63
BB65	0.04	0.07	-27.94	-41.30
300	0.04	0.11	-28.64	-49.36
301	0.03	0.05	-28.05	-53.11
302	0.05	0.12	-28.58	-71.52
3	-	0.11	-29.83	74.70
2	-	0.24	-29.54	73.70
202	-	0.26	-29.45	73.27
BB 53	-	0.34	-30.89	68.30
BB 54	-	0.24	-33.55	66.76
203	-	0.27	-29.13	66.65
204	-	0.25	-29.61	65.66
205	-	0.23	-29.53	65.28
206	-	0.28	-28.97	64.90
207	-	0.25	-29.54	64.60
208	-	0.15	-29.57	62.16
BB 55	-	0.23	-29.45	58.52
210	-	0.22	-28.73	51.10
211	-	0.21	-29.17	47.88

212	-	0.28	-28.68	46.23
213	-	0.22	-29.22	44.12
215	-	0.22	-29.22	41.94
217	-	0.19	-28.91	37.35
218	-	0.16	-29.24	35.33
219	-	0.27	-28.99	31.13
220	-	0.26	-29.13	31.13
221	-	0.23	-29.14	24.50
222	-	0.29	-28.30	22.42
224	-	0.19	-28.83	18.09
225	-	0.21	-29.05	16.78
226	-	0.24	-29.49	13.41
227	-	0.25	-28.89	9.01
228	-	0.24	-28.56	7.70
230	-	0.24	-28.50	-1.21
231	-	0.26	-29.13	-6.48
232	-	0.24	-29.04	-8.92
233	-	0.24	-28.18	-17.19
234	-	0.24	-29.13	-22.09
235	-	0.24	-29.00	-27.10
236	-	0.24	-29.22	-27.31
Sample code		%TOC (wt%)	$\delta^{13}\text{C}_{\text{org}}$ VPDB (‰)	Height (m)
Core-A				
CA-1	-	2.28	-28.12	14.41
CA-2	-	6.30	-30.00	14.24
CA-3	-	3.92	-28.86	13.37
CA-4	-	6.63	-29.03	12.77
CA-5	-	0.53	-28.22	12.43
CA-6	-	1.87	-28.60	11.65
CA-7	-	1.70	-28.73	11.59
CA-8	-	1.28	-28.67	11.35
CA-9	-	3.00	-29.31	10.60
CA-10	-	9.73	-29.90	6.54
CA-11	-	6.64	-29.92	5.76
CA-12	-	5.98	-29.84	5.45
CA-13	-	2.18	-29.28	5.28
CA-14	-	1.23	-30.13	4.85
CA-15	-	3.21	-30.61	4.50
CA-16	-	1.30	-29.96	2.44
CA-17	-	2.72	-30.59	-30.49
CA-18	-	2.37	-30.66	-30.72
CA-19	-	6.78	-30.06	-31.00
CA-20	-	3.43	-29.81	-31.55
CA-21	-	0.59	-24.65	-31.69
CA-22	-	0.44	-29.01	-35.10
CA-23	-	0.81	-28.99	-35.35

Table S1. Organic carbon isotope data. At Backside Beck, the top of the Cystoid Lmst Mbr=-33.9 m; base of the Ashgill Fm sandstone = 61.5 m; base of the Spengill Mbr=74.5 m; base of the Silurian=75.5 m. Sample codes in red measured using methods in Sullivan et al. (2018). Sample codes for Backside Beck samples shown in Figs. S2 and S3. Others relate to depths/heights shown in main text figures.

Sample code	$\delta^{13}\text{C}_{\text{carb}}$ VPDB (‰)	$\delta^{18}\text{O}_{\text{carb}}$ VPDB (‰)	Height
Mójcza			
MZ-9b	0.84	-4.54	2.50
MZ-9a	0.91	-4.98	2.45
MZ-8b	1.04	-3.97	2.34
MZ-7g	0.81	-4.01	2.18
MZ-7d	0.36	-4.005	2.00
MZ-7b	0.575	-4.455	1.86
MZ-7	0.62	-4.58	1.65
MZ-6b	0.05	-3.93	1.37
MZ-5c	-0.56	-4.71	1.15
MZ-4d	-0.04	-4.205	0.92
MZ-4b	-1.19	-4.78	0.75
MZ-3a	-0.84	-4.69	0.58
MZ-2	-0.23	-3.98	0.44
MZ-1a	-0.92	-3.89	0.17
MZ-0	-0.29	-3.85	0.07
Core-A			
CA_C-13	1.12	-6.61	2.74
CA_C-14	1.22	-5.09	1.38
CA_C-15	1.20	-6.23	0.77
CA_C-16	1.11	-6.47	0.28
CA_C-17	0.36	-4.66	-0.03
CA_C-18	0.26	-5.87	-0.65
CA_C-19	0.09	-7.74	-1.15
CA_C-20	0.63	-5.37	-1.79
CA_C-21	0.80	-5.62	-2.24
CA_C-22	0.84	-5.73	-2.70
CA_C-23	0.84	-7.87	-3.35
CA_C-24	1.09	-6.58	-3.76
CA_C-25	1.28	-5.06	-4.13
CA_C-26	1.20	-5.91	-4.85
CA_C-27	1.11	-7.51	-5.35
CA_C-28	1.29	-6.49	-5.90
CA_C-29	1.40	-5.86	-6.30
CA_C-30	1.53	-5.56	-6.55
CA_C-31	0.96	-9.53	-7.15
CA_C-32	1.77	-4.84	-7.55
CA_C-33	1.81	-4.53	-7.88
CA_C-34	1.94	-6.00	-8.12
CA_C-35	2.29	-5.27	-8.48
CA_C-36	2.49	-5.10	-8.65
CA_C-37	3.50	-4.92	-9.52
CA_C-38	3.30	-6.16	-9.88
CA_C-39	2.51	-5.32	-10.13
CA_C-40	3.21	-7.46	-10.43
CA_C-41	3.12	-7.88	-10.72
CA_C-42	2.56	-6.04	-11.68
CA_C-43	3.69	-6.39	-11.92
CA_C-44	1.70	-4.70	-12.18

CA_C-45	1.36	-6.11	-16.22
CA_C-46	1.24	-6.75	-17.72
CA_C-47	1.29	-5.52	-18.16
CA_C-48	1.47	-5.09	-18.85
CA_C-49	1.52	-5.48	-28.30
CA_C-50	1.01	-4.81	-33.74
CA_C-51	1.24	-3.48	-34.68
CA_C-52	0.67	-5.84	-39.28
CA_C-53	0.26	-6.85	-46.80
CA_C-54	0.22	-6.14	-48.22
CA_C-55	0.47	-6.24	-49.65
CA_C-56	0.81	-6.76	-51.14
CA_C-57	0.98	-6.33	-52.88
CA_C-58	0.96	-6.72	-53.27
CA_C-59	1.00	-6.49	-54.78
CA_C-60	1.30	-5.77	-56.32
CA_C-61	1.60	-6.11	-57.44
CA_C-62	1.49	-5.79	-58.39
CA_C-63	1.70	-5.51	-58.89
CA_C-64	1.53	-6.28	-59.30
CA_C-65	1.76	-5.34	-60.07
CA_C-66	1.59	-6.82	-60.52
CA_C-67	1.40	-5.66	-60.98
CA_C-68	1.22	-6.12	-61.43
CA_C-69	1.36	-5.32	-61.90
CA_C-70	1.38	-6.07	-62.38
CA_C-71	1.51	-5.54	-62.85
CA_C-72	0.91	-13.43	-63.13
CA_C-73	0.89	-13.45	-63.57
CA_C-74	0.95	-13.49	-63.66
CA_C-75	0.93	-13.47	-64.06

Table S2. Carbon and oxygen isotope data for bulk carbonate for the Mójcza section and Core-A. Sample codes in red measured using methods in Sullivan et al. (2018).

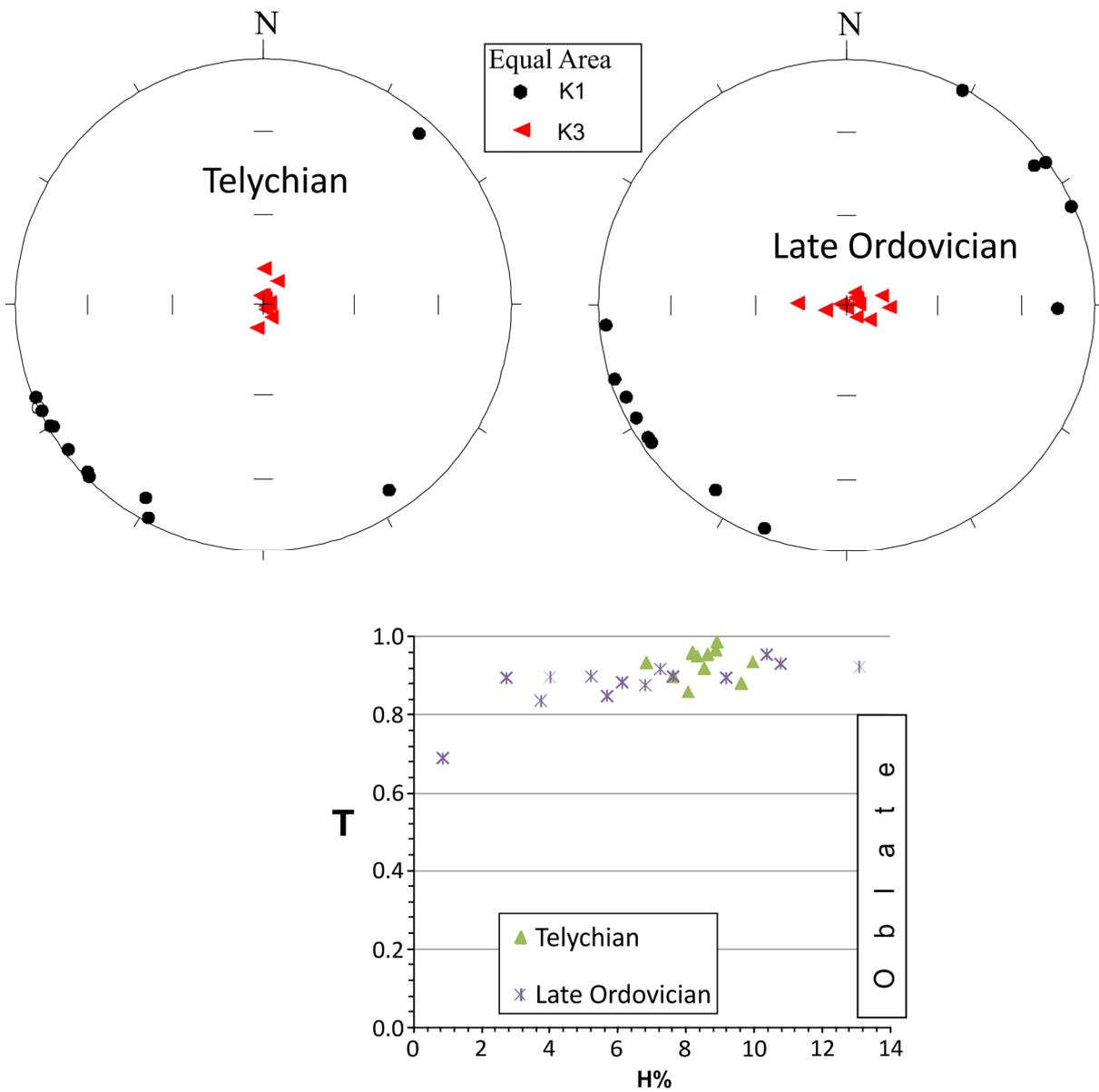


Fig. S5. AMS data for the Grabowiec-6 core. This also shows the data for the Telychian described by Hounslow et al. (2021). The lower hemisphere AMS maximum (K₁) and minimum (K₃) axes are shown in the stereonets and the ellipsoid shape and intensity (T versus h%). Data is oriented using the Kiaman partial remagnetisations. All samples have a sedimentary type fabric, shown by oblate shape (T > 0). The mean K₁ axis is at bearing 067° and indicates probable distal deposition on a palaeoslope directed to the SW, away from the East European platform coastal zone to the east (Teller, 1997).

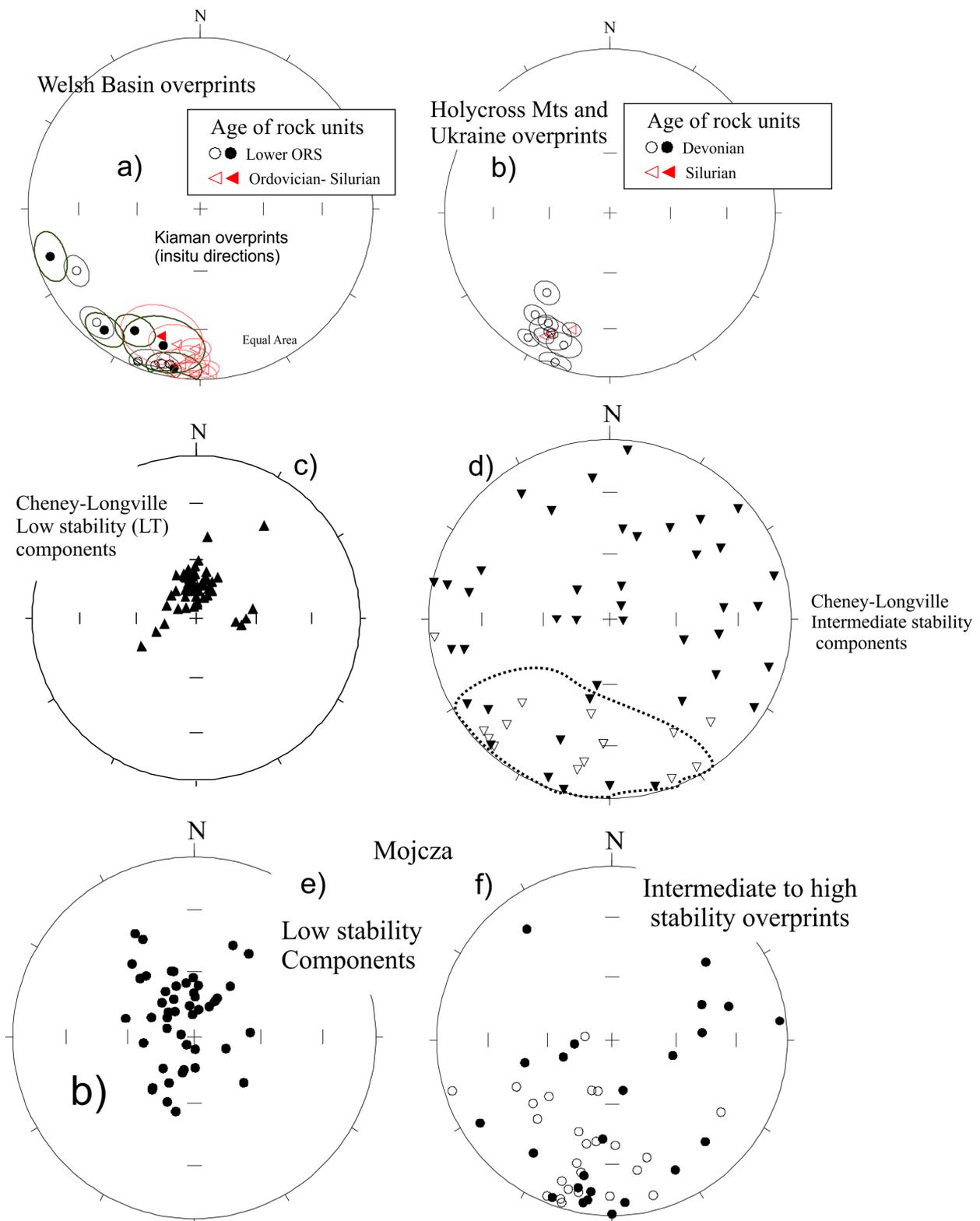


Fig. S6. a), b) Overprint and remagnetisation directions (igneous and sedimentary rocks) in other published studies from pre-Carboniferous units (Old Red Sandstone, [ORS], Ordovician and Silurian) from: a) the Welsh Basin and b) Poland and Ukraine. Directions are shown with α_{95} confidence cones. Filled symbols lower hemisphere, unfilled upper hemisphere. Equal area projection. Welsh Basin data from: Chamaluan and Creer (1964), McClelland Brown (1983), Smith and Piper (1984), Piper (1995), Setiabudidaya et al. (1994); Stearns and Van der Voo (1987), Channell et al. (1992a,b) and McCabe and Channell (1990), Poland and Ukraine data from: Smethurst and Khramov (1992), Grabowski and Nawrocki (1996, 2001), Zwing, (2003), Jeleńska et al (2005), Szaniawski (2008), Szaniawski & Lewandowski (2010), Szaniawski et al. (2011).

c), d), e), f) Low and intermediate stability overprint components from section at: c), d) Cheney Longville and e), f) Mójcza. Low stability components are inferred as Brunhes overprints, with varying degrees of scatter due to unblocking spectra overlap with intermediate stability components. Intermediate stability components at Mójcza are clearly Kiaman-like. The dotted line around the SSW directed intermediate stability components in d), loosely define a Kiaman-like component like seen in other studies, shown in a). Filled symbols lower hemisphere, unfilled upper hemisphere. Equal area projection. All data in insitu coordinates.

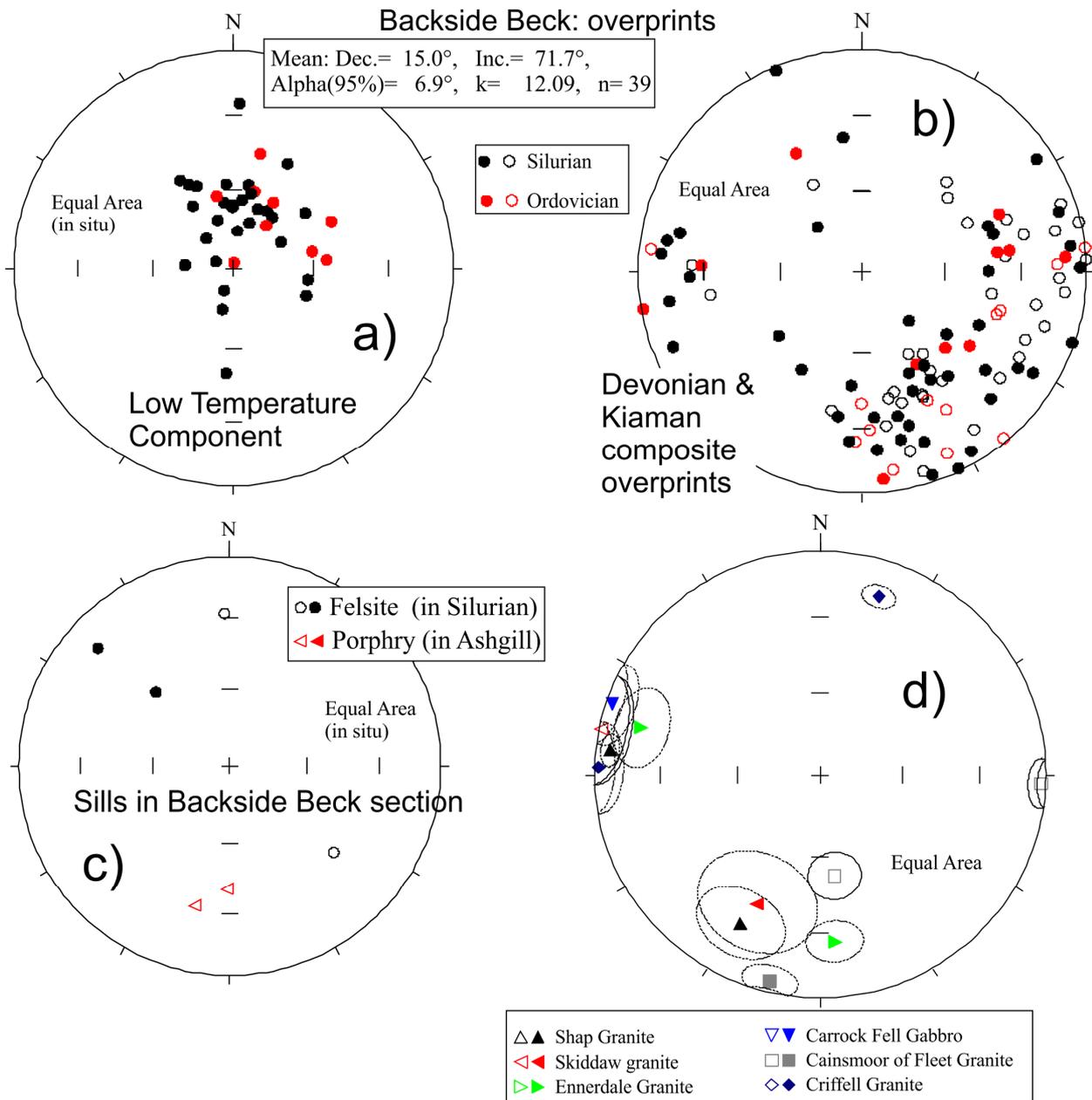


Fig. S7. a), b) Ordovician Backside Beck overprint magnetisation directions from this study. These are shown with the Silurian data (in black in a) and b) from Hounslow et al. (2021). c) Shows the ChRM directions determined from the felsite and porphyry sills in the Backside Beck section. d) Post Silurian magnetization directions (and α_{95} confidence cones) for igneous bodies in the Lake District and southern Scotland detailed in Piper (1997). These are both primary Devonian directions (approx. E-W dual polarity group) and later overprints (Kiaman SSW group) identified in the various units detailed in Piper (1997). All data in geographic coordinates. Filled symbols lower hemisphere, unfilled upper hemisphere.

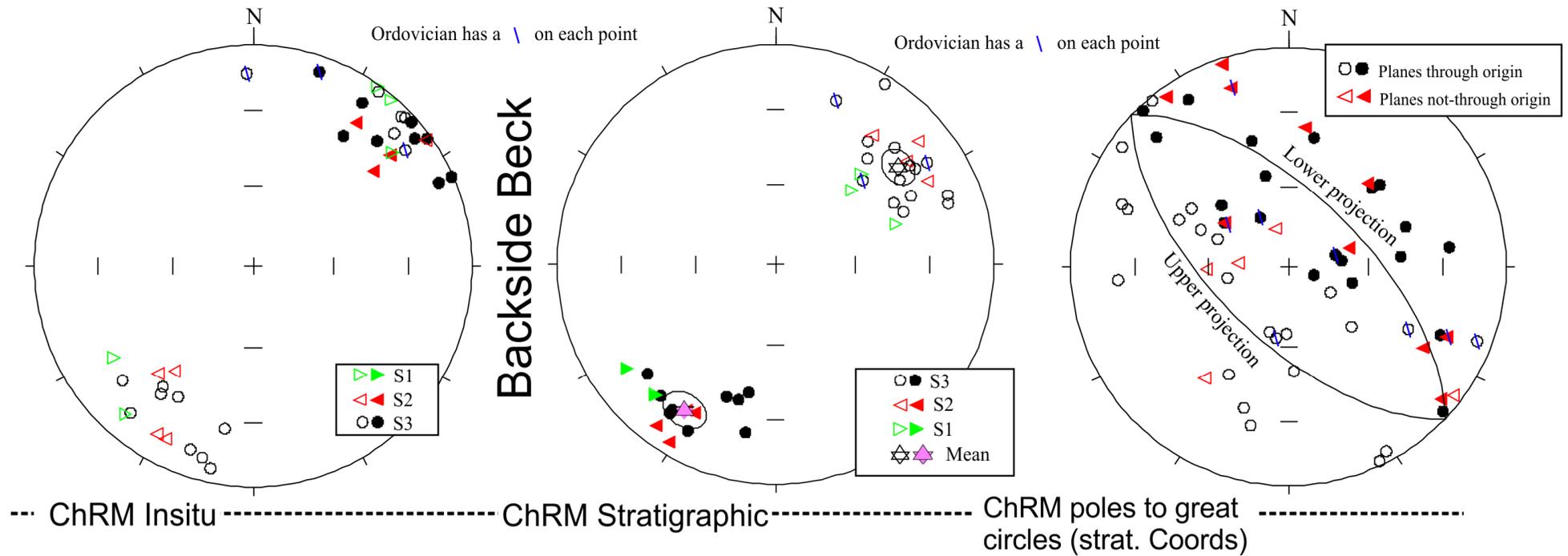


Fig. S8. Characteristic remanence (ChRM) directions for the Backside Beck section. The Ordovician data (blue ticks through symbols) are shown with the Silurian data from Hounslow et al. (2021). Insitu directions shown on the left, bedding corrected in the middle and bedding corrected poles to great circles on the right. The planes through the ChRM great circle poles have a pole which is near the mean of the ChRM line-fit direction in middle panel. The great circle line-fit ChRM plane, which is orthogonal to the S-class mean (derived from S-class data in mid stereonet) is that shown on the right most stereonet. S1 to S3 indicate the demagnetisation behaviour explained in the text.

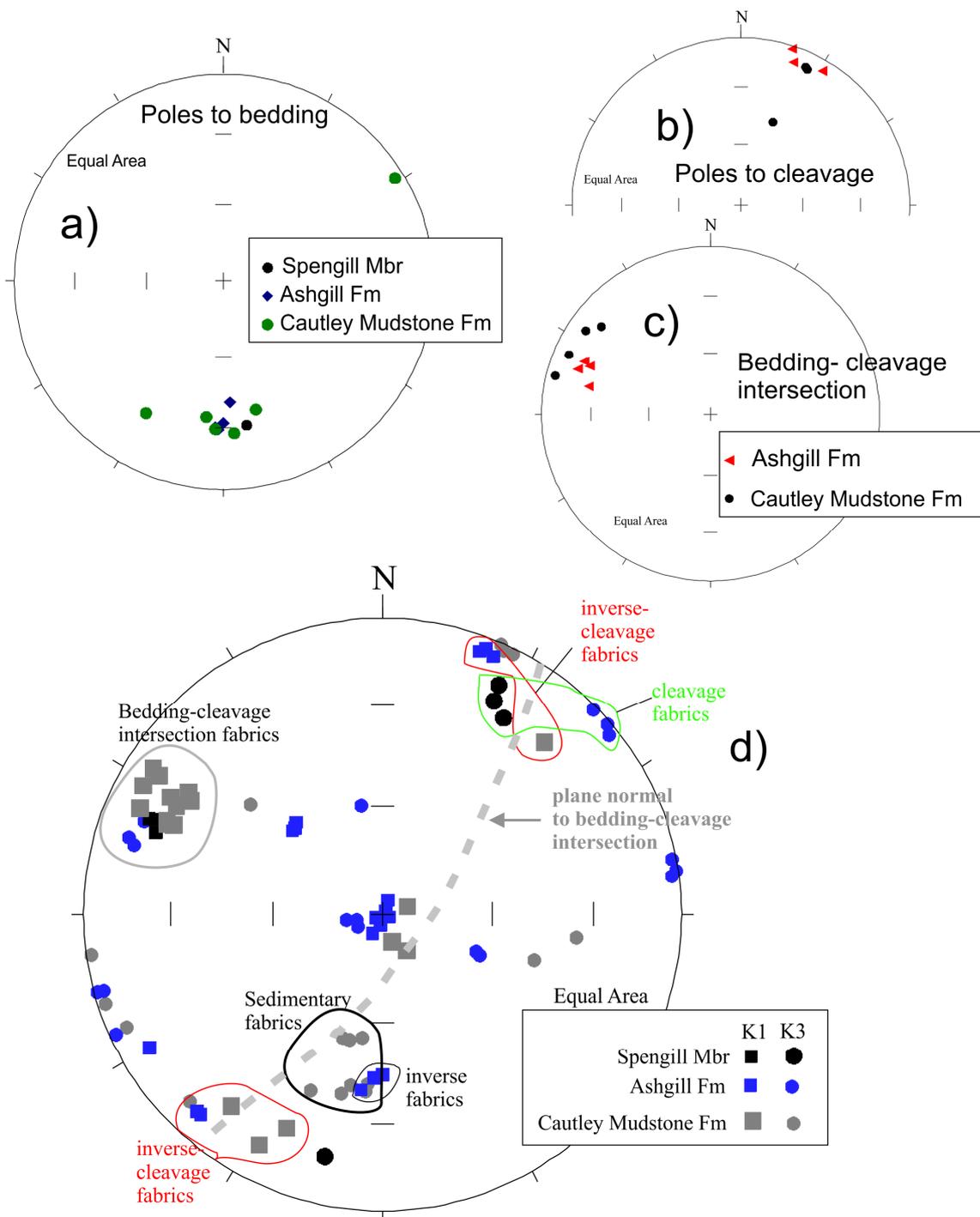


Fig. S9. AMS and structural data for the Ordovician at Backside Beck. a) poles to bedding; b) poles to the incipient cleavage in the section. c) Bedding- cleavage intersection lineations. d) The AMS directional data (in geographic coordinates) with the interpretation overlaid. Some samples from the Cautley Mudstone show primary-like fabrics with K_3 near normal to the bedding. A small group from the Ashgill Fm show an inverse-type fabric (Rochette, 1988; Ihmlé et al., 1989; Hounslow, 2001) with K_1 near normal to the bedding. Most of the sample set show evidence of a tectonic fabric overprint, with K_1 close to the bedding-cleavage intersection, typical for the initial indicators of strain. Both K_1 and K_3 are spread out in a girdle normal to the bedding-cleavage intersection, with some from the Ashgill Fm and Spengill Mbr showing evidence of K_3 normal to the cleavage ('cleavage fabrics'), or an inverse fabric to the cleavage ('inverse cleavage fabrics') expressed by the K_1 axes normal to the cleavage. The inverse fabrics are likely the results of Fe-Mn carbonates in the section (see Hounslow et al., 2021, for more details, and the comparison to the larger set of Silurian data).

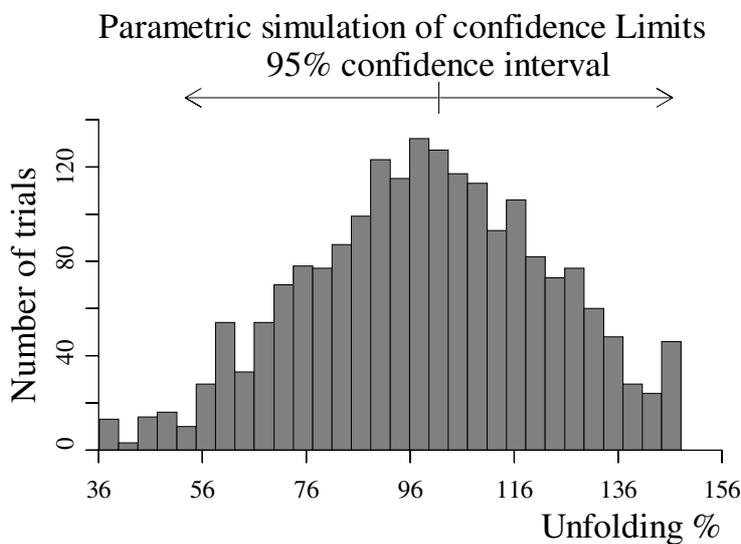
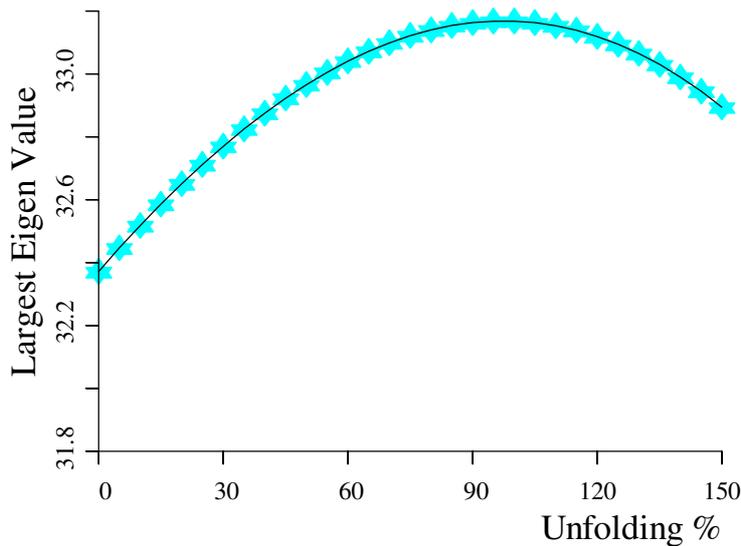


Fig. S10. Backside Beck section fold test data. Upper panel-the progressive unfolding test using the eigen-based test of Tauxe and Watson (1994). Lower panel-the re-sampling (simple) bootstrap simulation of confidence limits, using 2000 simulations. This simulation has median (best) Unfolding = 102.0%, 95% confidence limits on unfolding% of 53% and 146%. Ordovician and Silurian data combined (confidence limits can vary by a few%, since it's a simulation).

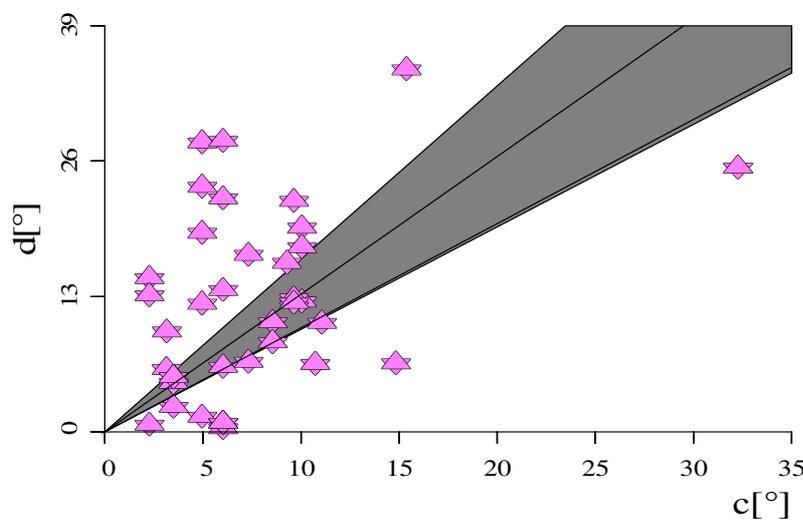


Fig. S11. Backside Beck section showing the DC fold test data of Enkin (2003). Ordovician and Silurian data combined. This is a single sample-test, rather than a site-mean test. Best fitting slope (unfolding) = 132.3% with 95% confidence interval: $\pm 33.9\%$, standard error (sigma) = 16.7. The DC fold test is positive, indicating the magnetisation was probably acquired PRIOR to folding. See Enkin (2003) for definition of the d_i and c_i parameters plotted here.

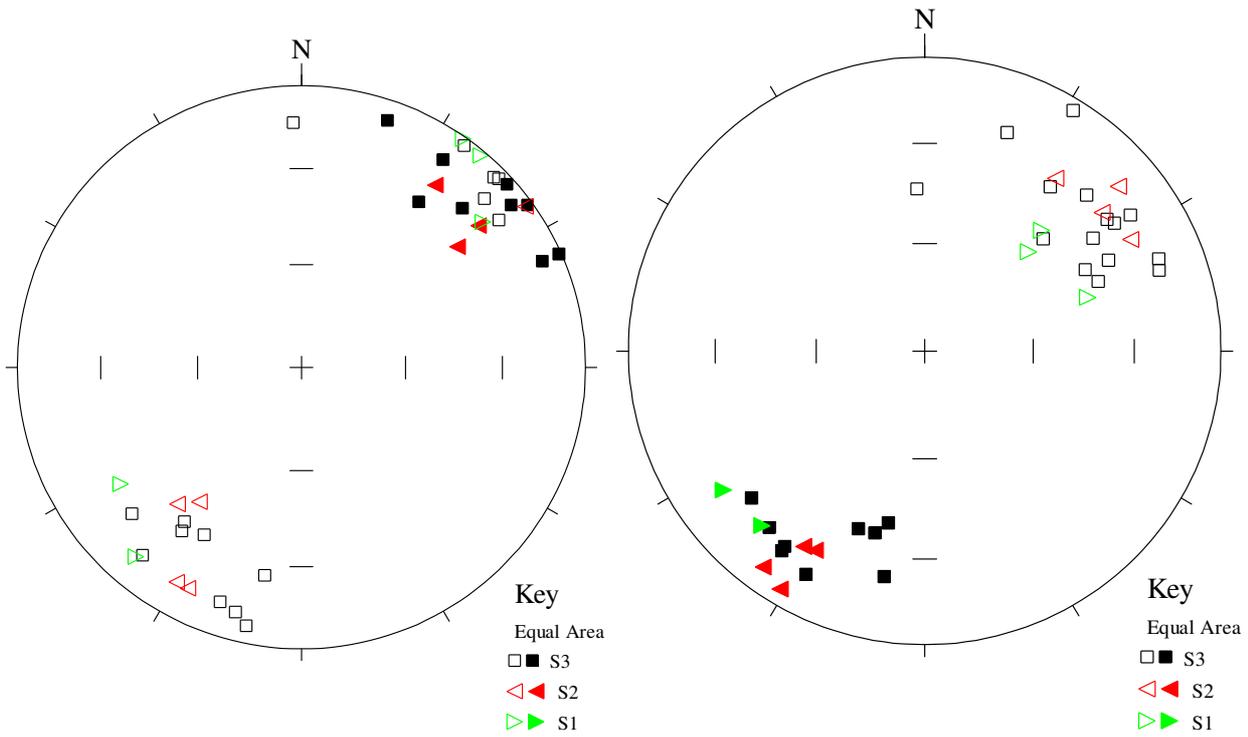


Fig. S12. Backside Beck data used for fold test. Left in geographic, on right in stratigraphic coordinates. geographic coordinate mean = 039.2°, 11.5°, α_{95} = 6.4°, k= 14.6, n= 37; stratigraphic coordinate mean 222.6°, 27.1°, α_{95} = 5.7°, k= 17.9, n= 37

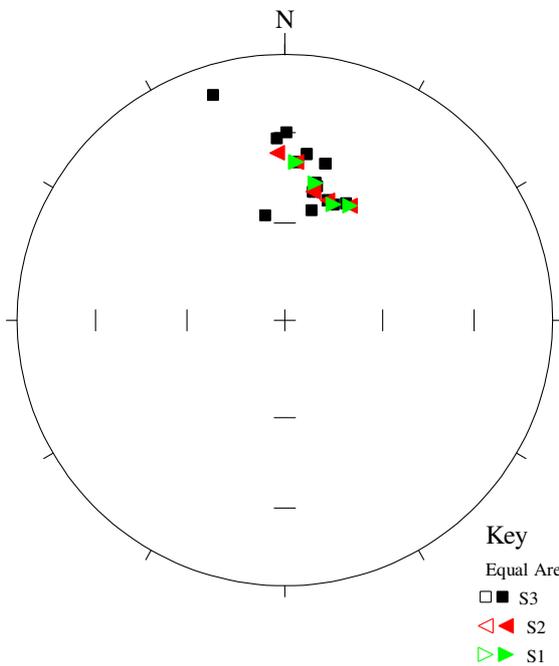


Fig. S13. Bedding dip directions used in the fold test at Backside Beck.

Cluster	n	r	k	Dec.	Inc.
For 0% unfolding					
1	16	14.95	14.3	42.4	2.7
2	21	19.93	18.7	36.7	18.1
For 100% unfolding					
1	16	15.05	15.8	47.4	-25.8
2	21	20.02	20.4	38.9	-28.1

Table S3. The data used for the McFadden (1998) fold test with the statistics for the two data clusters used at 0 and 100 percent unfolding.

Explanation of component fitting procedures using Linefind.

LINEFIND is explained in detail in Kent et al. (1983) and in outline in McFadden & Schmidt (1986), and has a number of features which make it substantially different in use to 'standard' manual-PCA fitting as widely used in much software. Heslop & Roberts (2016) have also recently re-examined some of the fundamental problems with standard PCA fits. One of the conundrums as they put it is: "It is not apparent why the

- 3) These variance estimates (one of the **two** 'noise' models) to use are added to by an 'excess noise' proxy using the parameter rho (the so called 'excess standard deviation'). That is when rho=1 the total variance is like in the error model chosen. The additional 'excess-noise' comes from things like between-step orientation errors, short-term viscous remanence, GRM, thermal alteration, magnetometer/ambient noise, changes in holder magnetisation. Generally a rho between ~1 to ~5 is required to give sensible looking fits to the data (but sometimes outside this range). **The choice of what rho to use is the major fundamental choice the user needs to make.**

Fig. S15. The change of rho and the partial AIC for various fitting models

RHO = 0.20	Partial AIC = 0.12458	NDF= 12	MDF= 36
RHO = 0.40	Partial AIC = 3.2475	NDF= 22	MDF= 26
RHO = 0.60	Partial AIC = 3.2475	NDF= 22	MDF= 26
RHO = 0.80	Partial AIC = 3.2984	NDF= 22	MDF= 26
RHO = 1.00	Partial AIC = 3.7467	NDF= 26	MDF= 22
RHO = 1.20	Partial AIC = 5.5767	NDF= 26	MDF= 22
RHO = 1.40	Partial AIC = 5.5767	NDF= 26	MDF= 22
RHO = 1.70	Partial AIC = 5.5625	NDF= 26	MDF= 22
RHO = 2.00	Partial AIC = 5.5625	NDF= 26	MDF= 22
RHO = 2.50	Partial AIC = 5.5625	NDF= 26	MDF= 22
RHO = 3.00	Partial AIC = 5.5625	NDF= 26	MDF= 22
RHO = 4.00	Partial AIC = 4.8158	NDF= 24	MDF= 18
RHO = 5.00	Partial AIC = 4.8158	NDF= 24	MDF= 18
RHO = 7.00	Partial AIC = 5.4435	NDF= 28	MDF= 20
RHO = 10.00	Partial AIC = 5.4435	NDF= 28	MDF= 20
Best model: RHO= 0.40000		AICP = 3.2475	

- 4) Selecting a rho~1 will often over specify the number of linear segments in the data, and massively overestimating rho will give only one linear segment (or none). The selection of the best rho is guided by two things. Firstly the Akaike's Information criteria (AIC, actually the partial AIC; Kent et al., 1983), which should plateau (or ideally drop slightly from the previous rho-step) to give the best rho-model-estimate. E.g. in Fig. S15 example, there is rho=4 (pAIC=4.8158) which is lower than the previous step at 5.5625. However, there may be several such steps-changes in pAIC, and several choices may be possible (e.g. 0.4 to 0.6; 1.7 to 3.0 in Fig. S15). Secondly, and obviously most importantly, rho needs to be selected based on LINEFINDS interpretation of the Zijdeveld and steronet-plots of the demagnetisation data, to guide which parts of the data may have suitable linear segments. In this case pAIC selection worked ok (i.e rho=4, but actually closer to 3.5), giving 3 components and a GC plane (Fig. S16). Component A is a possible Brunhes (but rather steep, so may be ?drilling-induced- this is an example from Core-A). The C component is the inferred ChRM, in this case through the origin (the origin step is 999). The B component is inferred as a composite intermediate component and is largely seen during AF demagnetisation (the 'Range' in Fig. S16 is a composite Temp-AF scale). The fitted plane in this case is one, which includes the entire dataset, implying there is some unblocking-spectra overlap of all these 3 'linear' components.

```

Number of acceptable linear Segments= 3
Number of planar chunks = 1
Points      Range      Declination  Inclination  α95  Component
1 To 4      0 To 200     327.9       89.0        45.6  A
4 To 15     200 To 485   246.7       54.5        8.6   B
15 To 17    485 To 999   129.9       -35.7       7.4   C
The fitted Planes:
1 To 17     0 To 999     64.8        28.2        6.8   1
Partial AIC = 5.5625      NDF= 26  MDF= 22

```

Fig. S16. Sample data which is thermally demagnetised to 350°C then AF demagnetised to 100 mT. Linear and planar segments with a rho=4. A second parameter 'radtol' can be set which filters out (does not display) linear segments which exceed a α_{95} threshold (default is 20°, but 40° has been typically used). The above example uses radtol=50°. Assuming 3 points on a line –fit not anchored to the origin, a radtol of 40° and 50° would equate to MAD values of 6.7° and 8.3° respectively (Khoklov & Hulot, 2016).

- 5) The automatic choice of linear segments can be overridden by introducing 'break points', in case the lines 'overspill' - which can happen for strong linear segments. In this respect this manual break point is similar to conventional PCA, pinning the ends of the linear segments. Unlike conventional PCA, linear segments cannot be anchored to the origin. LINEFIND decides if a line is sufficiently co-linear with the origin. A larger rho will allow co-linear-with-origin behaviour to be more likely. Sometimes a breakpoint may be needed at the HT end of a long LT line segment, so it does not overspill into the weaker HT components. It may also be needed at the LT and MT join if the MT or LT component is long and strong.
- 6) Key strengths of LINEFIND are: A) its ability to use the measurement errors, and a noise model of the data, B) its automatic, and objective statistical selection of linear and planar segments, stemming from the uncertainties and given guidance from interpretation of the data thorough choice of rho. C) It can find small components, or components which may be unseen in the particular 2D Zijderveld projection being used- i.e. if its near-orthogonal to the page (i.e. its not dependent on chosen projection). D) Linear segments can be overridden, based on interpretation of the data with break points. E) It may give a selection of statistically significant GC planes to the data (hinting at component overlap), rooted in the measurement variance and the rho chosen.

Therefore, when combined with conventional interpretation of data using demagnetisation diagrams, the use of LINEFIND is a more objective method of analysis than conventional 'eye-balled' PCA fits. A DosBox version of Linefind (linef.exe), and its source code can be found at the following DOI:

10.13140/RG.2.2.29761.38244, or the link:

https://www.researchgate.net/publication/347327549_A_DOSBox_version_of_Linefind.

DosBox is a Dos emulator which runs on many platforms, and can be obtained from:

<https://www.dosbox.com/information.php?page=0>.

The format of files needed by Linef.exe can be created with the windows software, GM4Edit available from the DOI of Hounslow (2019), which contains additional help on these file formats.

Statistics of the demagnetisation data analysis

The statistics from the LINFIND fitting procedure applied to all the data are detailed in Table S3. Broadly these indicate the following: A) The Möjczka data has the largest set of rho values, which probably relates to the fact that this was measured on the JR6 (1 unbiased per axis; Briden & Arthur; 1981), whereas others were largely measured on the GM400, 3-axis squid system (6 unbiased per axis), B) Data from the Cheney

Longville section has the next largest rho, which may be rather noiser data, since this section contained a lot of sandier lithologies, which may not be such good field recorders as the finer-grained lithologies in the other sections. C) Average α_{95} is approximately related to the S1 to S3 classes, but this relationship varies, because of its dependency on rho and the measurement variance. D) Average α_{95} is not systematically related to the T1 to T3 classes, likely because the quality category is also dependent on the length of approach to the ChRM direction. Comparing confidence intervals between lines and planes (i.e. pole for the great circle plane) is unwise, since the planes generally contain more measurement points.

Section/Core	Class	N	rho	α_{95} (°)	equivalent MAD ¹ (°)	
Backside Beck [BB] (Ordovician)	S1					
	S2					
	S3	3	3	30.5	5.1	
	T1	4	1.5	15.3	2.6	
	“	T2	2	2.2	25.1	4.2
	T3	2	1.7	8	1.3	
Cheney Longville [CL]	S1					
	S2	5	4	14.5	2.4	
	S3	28	5	18.8	3.1	
	T1	9	4.5	15.7	2.6	
	T2	4	5	14.8	2.5	
	T3	8	2.8	23.4	3.9	
Möjca [M]	S1	2	4	6.3	1.1	
	S2	2	7	12.7	2.1	
	S3	8	4	22.9	3.8	
	T1	6	4	12	2.0	
	T2	4	1.4	21.9	3.7	
	T3	13	3	17.9	3.0	
Grabowiec-6	S1	1	7	27	4.5	
	S2					
	S3	5	1.7	15.3	2.6	
	T1	3	1.9	14.8	2.5	
	T2	5	1.9	14.8	2.5	
	T3	6	1.3	15.7	2.6	
Core-A	S1	23	1	5.6	0.9	
	S2	40	1.4	12.1	2.0	
	S3	44	1.4	13.3	2.2	
	T1	3	1.75	13.1	2.2	
	T2	2	1.4	14.1	2.4	
	T3	1	2.5	14	2.3	

Table S3. Statistics from the LINFIND fitting procedures applied to the demagnetisation classes. N= number in each category. Rho=median excess standard deviation, and α_{95} is geometric mean of the 95% confidence interval as determined by LINEFIND. For T-class data this uncertainty is on the pole to the great circle. ¹ equivalent maximum angular deviation (MAD) uses Khoklov & Hulot (2016) with 3 points anchored- i.e. their C_{MAD} (3) conversion of 6.0. For the better defined lines and planes with more data points MAD will be an underestimate.

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