

# A geomagnetic polarity timescale for the Carboniferous

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## Abstract

The geomagnetic polarity pattern for the Carboniferous is incompletely known, but with the best resolved parts in the Serpukhovian and Bashkirian. Hence, data from both igneous and sedimentary units are also used in an additional polarity bias evaluation. In the Tournaisian to mid Visean interval polarity is mainly derived from palaeopole-type palaeomagnetic studies, allowing identification of polarity bias chrons. Seven polarity bias chrons exist in the Mississippian (MI1n<sub>B</sub> to MI4n<sub>B</sub>) with an additional 33 conventional magnetochrons and submagnetochrons (MI4r to MI9r). The Moscovian and Gzhelian polarity is best resolved in magnetostratigraphic studies from the Donets Basin and the southern Urals. Dispute about the reliability of these data is ill-founded, since an assessment of supporting data from palaeopole-type studies suggests that these datasets currently provide the best magnetic polarity data through the Pennsylvanian. Polarity bias assessment indicates a normal polarity bias zone in the Kasimovian. In the Pennsylvanian there are 27 conventional magnetochrons and submagnetochrons (PE1n to C11r) and one normal polarity bias chron (PE8n<sub>B</sub>). The Kiaman Superchron begins in the mid Bashkirian, with clear data indicating brief normal polarity submagnetochrons within the Superchron. The magnetochron timescale is calibrated using 31 U-Pb zircon dates and a quantitative Bayesian-based age-scaling procedure.

## Introduction

Stratigraphic changes in geomagnetic polarity (magnetostratigraphy) have been realized as having great utility in the Mesozoic and Cenozoic for global and local correlation and dating. These fossilised polarity changes are recorded as normal polarity states (geomagnetic field like today) and reverse polarity (magnetic poles opposite), with brief (100-1000's years) transitional field intervals between these two states. This bipolar state of the earth's magnetic field has been predominant for the last 2 Ga (Evans, 2006). The between-polarity transitional state has so far not been observed in Carboniferous rocks. Since the palaeomagnetic technique is largely independent of sedimentary or climatic environments; in the longer term it also has the ability for correlating chronostratigraphic and biostratigraphic boundaries into differing faunal realms. Its ability for solving dating and correlation problems in the Carboniferous has not yet shown fruit, but it has been realized in the Cenozoic (and to some extent the Mesozoic), where the technique can greatly surpass biostratigraphic methods in terms of chronostratigraphic resolution (Langereis et al. 2010; Miller & Wright 2016).

However, in much of the Palaeozoic, especially so the Carboniferous and Devonian, the magnetic polarity pattern is insufficiently known in detail. This is primarily due to two factors, firstly many Palaeozoic sediments are very weakly magnetic (low remanence intensity), so it was not until the 1990's  
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that magnetometers were available that could reliably measure most sediments during demagnetization procedures (magnetic 'cleaning' or the 'washing' of Cox & Doell 1960). Secondly, many Carboniferous successions are rather thermally mature (conodont alteration indices >2 to 3), having been buried to substantial depths, and in many cases the Fe-Ti oxides, which carry the remanence, have been subjected to substantial diagenetic modification (Johnson *et al.* 1995, 1997). Units may also be re-mineralized in various ways, both modifications potentially generating new magnetic minerals and destroying most of the originally deposited Fe-Ti oxides. This has contributed to the problem of perceived widespread remagnetisations (Van der Voo & Torsvik 2012). These factors have created slow and difficult progress, in obtaining good Carboniferous palaeomagnetic data, since many palaeomagnetists would rather focus on easier to solve problems, with a better chance of scientific return.

This work is the first comprehensive assessment of existing strands of knowledge about the geomagnetic polarity in the Carboniferous, utilizing a wide range of palaeomagnetic and magnetostratigraphic data. This work firstly examines the historical context of geomagnetic polarity investigations of the Carboniferous, explaining development of evolving palaeomagnetic paradigms, that have strongly fashioned the development and reliability of Carboniferous magnetostratigraphic and palaeomagnetic studies. Secondly it examines in detail Carboniferous palaeomagnetic palaeopoles-type datasets from extrusive and igneous rocks, which help to fill some of the data gaps in geomagnetic polarity knowledge. Thirdly, the details of existing Carboniferous magnetostratigraphic studies are examined, bringing in new stratigraphic and dating information where relevant. Lastly, these data are synthesized into a Carboniferous polarity timescale, with inbuilt uncertainty estimates. Since many of the magnetic dataset are biostratigraphically linked to regional substages, this requires relationships to international stages to be defined, which largely follows those proposed by Richards (2013), unless otherwise indicated.

## **Progress in developing a Carboniferous polarity scale**

The development of palaeomagnetism, focused on the Carboniferous, has followed several phases which relate to the development of new kinds of instrumentation along with analytical methods and magnetisation paradigms. This development can be broadly classified into three phases: a) Early reconnaissance work in the 1950's and early 1960's; b) early use and development of demagnetization techniques in the 1960's to 1980's; c) a remagnetisation 'realization' in which more complex magnetisation (and remagnetisation) models were being increasingly more widely utilized from the 1990's to the present day.

Throughout this study interval, the bulk of palaeomagnetic studies in the Carboniferous have focused on palaeopoles-type studies which use palaeomagnetic directional data to assess continental or tectonic motions. This kind of data does not directly help with knowing the detailed stratigraphic changes in magnetic polarity (since sampling is not designed to do this), but it may help understanding if certain time intervals have a preferential bias to either reverse or normal polarity (Belshé 1957; Irving & Pullai 1976; Algeo 1996). A very much smaller subset of Carboniferous palaeomagnetic studies utilize a stratigraphically well-constrained samplings strategy, to allow an estimate of the pattern of polarity change with stratigraphy. Still fewer studies are true magnetostratigraphic studies in the sense of Opdyke & Channel (1996).

### Early reconnaissance work

The 1950's saw the development of the founding principles of evaluating the fidelity of rocks as geomagnetic recorders, and means to describe the directional data (Irving 1959; Cox & Doell 1960 for reviews). These early attempts were largely focused on defining possible continental motions and the characteristics of the ancient magnetic field. The first Carboniferous palaeomagnetic studies were on volcanics and sediments by Clegg *et al.* (1954) and Belshé (1957) from Britain, Australian volcanics and sedimentary rocks (Irving 1957; Irving & Green 1958) and the Visean -Bashkirian Barnett Shale (Martinez & Howell 1956; Howell & Martinez 1957) and the late Pennsylvanian-Permian Naco Group sandstones (Runcorn 1956) from the USA. These had contributed to a substantial set of Carboniferous palaeomagnetic data by the late 1950's (Irving 1959), from which Everitt & Belshé (1960) were able to infer that the early Carboniferous was an interval of both reverse and normal magnetic polarity and the late Carboniferous was predominantly reverse polarity. This significant inference was later amplified by work on Permian rocks from Australia by Irving & Parry (1963), who extended this predominantly reverse polarity interval well into the Permian, and who coined the term for this late Carboniferous – mid Permian reverse polarity interval, as the Kiaman magnetic interval (here referred to as the Kiaman Superchron, Hounslow & Balabanov 2018). From 1959 Khramov and co-workers began their study of the Carboniferous sediments of the Moscow and Donets Basins (Khramov, 1963; Khramov *et al.* 1974).

These early studies did not use alternating field (AF) or thermal demagnetization (which all studies now use), to isolate the primary remanence since these methods were only beginning to be developed in the late 1950's, for AF 'cleaning', firstly using static AF fields (Brynjólfsson 1957; Creer 1958; As & Zijderveld 1958), and later tumbling of samples in AF (Creer 1959) with subsequent improvements (McElhinny 1966). In spite of this initial inadequacy, both polarities were present in Carboniferous rocks (Irving 1957; Belshe 1957; Clegg *et al.* 1957, Howell & Martinez 1957). However, additional magnetizations from the present day and after the time of rock formation (secondary magnetizations; Creer 1957; Irving 1959; Collinson & Runcorn 1960; Irving *et al.* 1961) contaminated these early datasets, sometimes giving smeared directions along a great circle, called by Khramov (1958) the 'circle of remagnetisation'. Some studies tried to use long term storage of samples in near zero magnetic field to attempt removal of the modern viscous components (Creer 1957 1959; Akimoto & Kushiro 1960), which contaminated the primary components. This was a process much used later by early Russian workers, and termed 'temporal demagnetisation'. By the end of the 1950's Irving (1959) had concluded that the "reconnaissance stage of palaeomagnetism" had come to an end, and that more detailed studies focusing on single stratigraphic units was needed. At this time the Carboniferous could be magnetically divided into the Kiaman and pre-Kiaman magnetic polarity intervals with the stratigraphic position of this boundary thought to be in the Westphalian (late Bashkirian) or near the base of the Stephanian (Moscovian) (Irving & Parry 1963).

### Early studies using demagnetization techniques

The first attempt in the Carboniferous to express the future vision of Irving (1959) was the study of Wilson & Everitt (1963) on the Kinghorn lavas from Fife in Scotland. This is the first true magnetostratigraphic study in the Carboniferous, since it related sampling to a detailed stratigraphic succession (see later). This was also ground breaking in several other ways, firstly that it used thermal demagnetisation to isolate the primary Carboniferous remanence using a magnetometer and furnace that the Imperial College team (and others; Stacey 1959, Irving *et al.* 1961) had been working on for some years (Wilson 1960). Thermal demagnetization had preliminary use in the 1950's mostly for rock-

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magnetic or geomagnetic palaeointensity work, with Doell (1956), Cox (1957) and Akimoto & Kushiuro (1960) applying this to directional data in Cenozoic volcanics and sediments. Secondly, the PhD of Wilson (1960) extensively used “vector heating diagrams” (a method later published in Wilson 1961) to display the directional and intensity data (what are now widely referred to as ‘Zijderveld plots’ following Symons & Stupavsky 1974; Dunlop 1979), a key method of displaying demagnetization data. Similar plots also appeared at the same time in the thesis of Everdingen (1960) (cited in Dietzel 1960) based at Utrecht. This was very significant, since Wilson’s vector heating diagrams clearly showed isolation of the primary components during demagnetization, and also the directional information of the secondary components, in the straight lines on his diagrams. Much of what Wilson discussed in using these diagrams is explained in more detail in Dunlop (1979) now widely used as standard by all palaeomagnetists. As explained by Storevedt (2016) these and other developments opened up a new era for palaeomagnetic research in which magnetisation paradigms moved from ‘simple’ to a more “realistic experimental-analytical approach”.

During these early studies on Carboniferous volcanic successions by Wilson and co-workers, Russian workers studied Carboniferous sediments from 1959 to the 1970’s from the Moscow and Donets basins, and sections adjacent to the Sea of Azov (Khramov 1967; Khramov *et al.* 1974). These largely used the methods that had existed in the west in the 1950’s, but they refined the ‘temporal’-demagnetisation methods with imposition of opposed components to correct for the modern viscous field components (Khramov *et al.* 1974; Khramov 1987). Carboniferous sediments studied were largely fined-grained clastics, with sometimes a focus on red-beds, which previous Carboniferous studies (Belshé 1957) has shown to give best results. In spite of adherence to the 1950’s paradigms these Russian dataset appear to display what looked like some sensible directional changes and datasets, with cross-section validation of magnetozones, tied to a detailed bio- and lithostratigraphy (Fig. 1). These and other datasets appeared in the Carboniferous part of the magnetostratigraphic time scale of Khramov & Rodionov (1980) and what has become known as the ‘general magnetostratigraphic scale’ (Guizikiv 2016). This had many later revisions, with the last major one being that of Khramov & Shkatova (2000), along with some later ‘time adjustments’ used by Perchersky *et al.* (2010)(Fig. 2). Unfortunately all source details and stratigraphic relationships of how the Carboniferous part of the ‘general magnetostratigraphic scale’ was constructed were not published, so it is impossible to critique and re-evaluate details, so this scale is not used in this compilation, which instead uses primary sources.

Some of the sample sets presented in these early Russian Carboniferous studies have been re-examined using modern demagnetisation techniques (Iosifidi *et al.* 2010, 2016, 2018; Iosifidi & Mikhailova 2017). For example in the Bashkirian data shown in Fig. 1, Iosifidi *et al.* (2016) re-examined a sample set covering 35 m in the C<sub>2</sub><sup>2</sup> interval (Cheremshankian, mid Bashkirian), which show very similar mean directions to the same interval studied in the 1960s (Iosifidi *et al.* 2016). However, these newer studies have shown an absence of Carboniferous normal polarity samples in the previously available sample sets (Iosifidi *et al.* 2016), except in the Aleksinian substage (mid Visean , early Warnantian). Unfortunately these new studies do not seem to have assessed the normal polarity intervals (although this is not entirely clear) or evaluated how the earlier studies of the same sample sets might have arrived at the interpreted normal polarity like shown in Fig. 1. The key issues may have been over-simple magnetisation models, Permian remagnetisations, with perhaps dual polarity Permian-Triassic remagnetisations, such as in the Priksha River sections (Iosifidi *et al.* 2018) which is in the Tullian substage (mid Visean ; Richards 2013). The extent to which these explanations might apply to the

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normal polarity intervals in the Pennsylvanian as shown in Fig. 1 is unknown. Other data is assessed here to try and evaluate the reality of these Pennsylvanian normal magnetozones.

### *The remagnetisation 'realisation' and a new magnetisation paradigm*

A widening realization of the problem of Kiaman reverse polarity remagnetisations on Carboniferous (and other) successions evolved in the 1970s and 1980's as more studies on late Palaeozoic successions were performed. This followed initial studies implicating wide-scale remagnetisations, found first in the Devonian (Chamaluan and Creer 1964; Chamaluan 1964), a phenomena that was not only restricted to sediments (Storevedt 1970). Clearer understanding of the characteristics of remagnetisations evolved in the late 1980's, when the impact of this was realized on magnetizations previously interpreted as Devonian- Carboniferous in age (McCabe & Elmore 1989; Van Der Voo & Torsvik 2012). The simple (two-component) magnetization models of the 1950's and early 1960's had evolved to an appreciation of the observational and interpretational complexity of palaeomagnetic data (Roy & Lapointe 1978, Storevedt 2016), which often come with 3-4 magnetisation components in pre-Permian datasets.

During this interval of 'realization' Tarling and Turner in Britain had a program of studies looking specifically to determine geomagnetic polarity and magnetization processes in Carboniferous sediments (Turner & Tarling 1975, Turner 1975, Perry 1979, Turner *et al.* 1979, Addison 1982, Addison *et al.* 1985, Palmer 1987). Although there were indications in some sections of probable dual polarity (Turner *et al.* 1979), the magnetizations were either largely Kiaman remagnetisations (McCabe & Channel 1994), or mixtures with what would now be considered as probably primary Carboniferous directions (Morris 1971; Turner *et al.* 1979), but were not considered so at the time. Palmer *et al.* (1985) attempted to provide a summary of existing Carboniferous polarity information (largely from paleopole-type studies), but failed to try and incorporate the distinction between Kiaman remagnetisations and potentially primary Carboniferous polarity datasets. In a sense these largely carbonate-based studies were utilizing the palaeomagnetic paradigms of the 1950's (e.g. stability indices) which was also reflected in the Carboniferous Russian datasets of the time, rather than the 'post-realization paradigms' expressed in holistic data assessments like that proposed by Roy & Lapointe (1978).

A turning point for the re-evaluation of the Carboniferous magnetic polarity stratigraphy was the assessment of Roy & Morris (1983), which excluded most remagnetized data from the North America Carboniferous datasets, and identified the base of the Kiaman Superchron around the Namurian-Westphalian boundary (mid Bashkirian), largely based on the data from the Maritime Provinces of eastern Canada (Roy & Park 1969, 1974). The later key contributions from Neil Opdyke and co-workers followed, starting with the first magnetostratigraphic studies on N. American Carboniferous sediments by DiVenere & Opdyke (1990, 1991a, 1991b) using the post-realisation paradigms of Roy & Lapointe (1978). These events set the scene for later reviews of Carboniferous magnetostratigraphy by Idnurm *et al.* (1996), Opdyke (1995) and Opdyke & Channell (1996). The magnetostratigraphic studies from the 1990's onwards form the backbone of what is described in the following sections.

### **Geomagnetic polarity bias during the Carboniferous**

The sample collection strategy of most palaeopoles-type studies (i.e. those used for palaeotectonic studies), do not usually provide sufficient information to describe the stratigraphic relationships between the individual sampling sites. Rather sites tend to be grouped into members, formations or

particular rock units, with a focus on those most likely to yield good results (often igneous units or red beds in the Carboniferous). Nevertheless, it is possible to utilize data from palaeopoles-type studies to construct an assessment of polarity bias (dominance of a particular polarity) during the Carboniferous like Irving & Pullaiah (1976) and Algeo (1996) has attempted for the Palaeozoic, and Irving & Parry (1963) and Palmer *et al.* (1985) had previously attempted for the Carboniferous. Similar approaches have been used in the Ordovician and Silurian to help define reverse and normal polarity dominance in parts of the Ordovician and Silurian respectively (Trench *et al.* 1993). Palaeopole-type sampling is in contrast to magnetostratigraphic studies, in which the stratigraphic relationship between samples is of paramount importance.

Since magnetostratigraphic data do not exist for the entire Carboniferous, the following assessment utilizes polarity bias data from Carboniferous igneous and sedimentary rocks. The first part uses data from extrusive and intrusive igneous rocks, weaving this data onto studies where stratigraphic-style sampling has been undertaken in these units. The following part examines the magnetostratigraphic data from sedimentary rocks, and integrates this with the polarity bias data from sediments. Many of the relevant palaeomagnetic and magnetostratigraphic studies in the Carboniferous have inadequate documentation of stratigraphic and biostratigraphic relationships, so these are dealt with in some detail, to allow a more holistic, integrative magnetostratigraphic evaluation, in the final sections.

## **Polarity data from extrusive and intrusive rocks**

Data collation utilized the palaeomagnetic database v4.6b (McElhinney & Lock 1976), the MagIC database (Tauxe *et al.* 2016), and all available Carboniferous (and earliest Permian) published up to early 2020. These palaeopole datasets were filtered to exclude:

- a) Data with large age uncertainties (either radioisotopic or stratigraphic). Large here is often greater than a stage age uncertainty.
- b) Data in which the original published material could not be examined, or was not sufficiently detailed to explain the data quality. This eliminated some early palaeomagnetic studies, prior to the development of competent demagnetization schemes. It also eliminated many of the Russian 'pole lists' contained in the above databases.
- c) A re-evaluation of stratigraphic and radioisotopic dates was also undertaken to revise the ages of sampled units, since their original publication. 29 of the 50 have reevaluated for ages.

The selected igneous and volcanic data are listed in Table 1, along with the percentage of reverse polarity in the studies, the age information and age confidence interval of the sampled sites (Fig. 3). Separated from the list in Table 1 (and not in Fig. 3) are some of the palaeomagnetic studies on the Tamworth Belt of the New England Orogen in eastern Australia. These complement and add to the magnetostratigraphic study of Opdyke *et al.* (2000) where sampling has been undertaken with a strong stratigraphic context built-in. These Australian studies are dealt with in a separate section below. An integration of the magnetic polarity data of Wilson & Everitt (1963) and Piper *et al.* (1991) from British late Visean volcanic units is also dealt with separately below, since these have rather better stratigraphic control than most studies, and can be stratigraphically linked together.

### *Global Polarity Bias data*

The igneous rock bias data clearly show the reverse polarity dominance of the Pennsylvanian (i.e. in the Kiaman Superchron), beginning ca. 320 Ma, a bias to rather more reverse polarity dominance during the From Lucas, Schneider *et al.* (eds) *The Carboniferous Timescale*, Spec. Publ. Geol. Soc. Lond.

Serpukhovian and Visean , and a normal polarity bias in the Tournaisian through early Visean (Fig. 3). Younger than 320 Ma in the Pennsylvanian, several studies have some normal polarity data indicating possible brief normal polarity intervals embedded in a dominance of reverse polarity.

The palaeomagnetic study of Halvorsen *et al.* (1989) from the Karkonasze granite and its metamorphic aureole, sampled the gray porphyritic granite and andalusite - cordierite hornfels of the micaceous schist country rock. The granite has an extensive range of dating methods applied, both SHRIMP, Rb-Sr and CA-ID-TIMS, suggesting a formation age at 312 Ma (Kryza *et al.* 2014), with emplacement and crystallization within less than 1 Myr duration. Both sample sets show evidence of dual polarity, although the hornfels has some contamination from a steeper (B component) magnetization. The directional data were consistent with other associated granites of mostly reverse polarity, and have a combined palaeopole  $3^\circ$  from the spline-fit pole at 310 Ma of Torsvik *et al.* (2012), all indicating the well constrained nature of this dataset. This may represent one of the normal polarity intervals in the DB4n or DB5n composite of Khramov *et al.* (1974) shown in Fig. 1.

The data of Beck *et al.* (1991) is from an extensive (~900 km long) late Paleozoic granitic plutonic belt sampled at Lago Ranco and Lago Rinihue (S. Chile), which ranged in composition from biotite granodiorite to pyroxene-bearing diorite. Their data indicate only one site (on south side of Lago Ranco) has normal magnetic polarity (8 samples) with all others from Lago Ranco reverse polarity. This site has a K/Ar date on biotite of  $309 \pm 8$  Ma (Beck *et al.* 1991). Deckart *et al.* (2014) dated a biotite-amphibole granodiorite (from Principal Cordillera) from the southern shore of Lago Ranco (FO09-38) yielding a SHRIMP U-Pb (Temora standard) age of  $305.9 \pm 2.4$  Ma, which is similar to a range of other dates from the late Carboniferous Chilean plutons with a mean ~309 Ma. Although the directional dataset is a little sparse, and the details of intrusions around the normal polarity site are not supplied, it does suggest this normal polarity event may be younger in age to that in the Karkonasze granite

The study of the Malaoba Formation volcanics by Yi *et al.* (2015) from the Tacheng Basin (NW China) found one possibly normal polarity site (a rhyolite, 9 samples) among 12 other sites (basalts, rhyolites and tuffs) of reverse polarity. An Ar/Ar date from the top of the succession yielded an age of  $304 \pm 4.7$  Ma. The authors did not use the data from this site (declination is displaced by  $\sim 80^\circ$ , but inclination is consistent), so it's not clear how reliable this single site is. If this is valid, it is possible it is the same normal polarity interval as found at Lago Ranco.

The undemagnetised samples of Clegg *et al.* (1957) and Everitt (1960) from the Shatterford sill in the English West Midlands, shows reasonably well defined dual polarity, and was originally dated as Westphalian, since its emplaced into Westphalian- C (Bolsovian, early Moscovian) clastics (Kirton 1984). This is one of number of other genetically-related intrusions in the same area which have been dated at around 295 -296 Ma using K/Ar (Fitch & Miller 1964), probably providing only a minimum age (Bolsovian is now around 315-310 Ma; Aretz *et al.* 2020). Field evidence suggests the sills were intruded into waterlogged sediment, so may be nearly contemporaneous with the Bolsovian (Kirton 1984), so the normal polarity samples may be the same age as those of the Karkonasze granite.

Edel *et al.* (2014) has reported dual polarity late Carboniferous magnetisations (normal polarity C, and D ; reverse C' and D' components) from various intrusive bodies in Corsica and Sardinia. Assuming their late Carboniferous reconstructions for Corsica and Sardinia are correct, their site polarities would be

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dominated by normal polarity magnetization (their C and D components). However, their data includes units such as the Isola Diorite and the Osani Andesite, which have well constrained crystallization ages at  $300 \pm 6.1$  Ma and  $308 \pm 3.0$  Ma respectively (Casini *et al.* 2012), yet are interpreted as exclusively normal polarity, using their tectonic reconstruction. The Edel *et al.* (2014) interpretation is that many of the late Carboniferous units acquired their magnetizations during the interval  $311 \pm 10$  to  $288 \pm 0.7$  Ma, from a remagnetisation during the emplacement of the U2 magmatic episode. The dominance of normal polarity is erroneous, with respect to their ages and data here (and in Hounslow & Balabanov 2018), and so the polarities originally inferred were not been used (Table 1; Fig. 3). However, if the C and D component directions are reverse polarity and the magnetisation ages are close to the crystallisations ages, then these would be magnetisations acquired with Sardinia and Corsica in a similar rotational position to that in the Triassic, rather than  $\sim 180^\circ$  rotated as suggested by Edel *et al.* (2014). Hence, the polarity has been inverted, giving a more sensible polarity distribution. The dual polarity Barrabisa granitoids (ca. 315.5 Ma; Table 1) fall within the Late Bashkirian- Moscovian interval (Fig. 3) possibly correspond to one of the normal polarity magnetozones in Fig. 1.

### New England Orogen, Tamorth Belt, E. Australia

The palaeomagnetic studies in the forearc basin, that is the Tamworth Belt of the New England Orogen (Geeve *et al.* 2002; Klootwijk 2002, 2003, 2016, 2019), provides a unique set of stratigraphically well-constrained, sampling sites, which allow these to be combined with the more magnetostratigraphic-style study of Opdyke *et al.* (2000) from the northern part of the Tamworth Belt. The Mississippian and Pennsylvanian successions in most of the Tamworth Belt are divided by a late Visean to Serpukhovian uplift and hiatus, with the Pennsylvanian successions seeing glaciogenic sediments deposited with volcanogenic and fluvial units (Glen & Roberts 2012; Phillips *et al.* 2016). The palaeopoles-type studies have primarily sampled the numerous extrusive lavas and pyroclastic flows, except for that of Opdyke *et al.* (2000) which also sampled suitable clastic and volcanoclastic lithologies. These studies have covered a few sites from the Tournaisian, with rather more sites from the Visean, to the more extensive sampling by Opdyke *et al.* (2000) in the Early and Middle Pennsylvanian. The successions are divided into a number of northwest to southeast distributed fault and thrust bounded blocks, with each block having a rather different lithostratigraphy (Roberts *et al.* 1993, 1995, 2003a, 2006; Phillips *et al.* 2016). Age control on the partly marine Mississippian successions is a brachiopod zonation, age-calibrated against western European successions (Roberts *et al.* 1993), with some additional age calibration against conodonts and ammonoids (Jenkins *et al.* 1993), as well as SHRIMP dates (Roberts *et al.* 1995; Fielding *et al.* 2008). The Pennsylvanian units have a very low resolution biozonation, supported by some fossil plants (Roberts *et al.* 1993), supplemented with a large set of SHRIMP dates from volcanic units (Fielding *et al.* 2008; Roberts & James 2010).

The Mississippian in the SE of the Tamworth belt in the Rouchel, Gresford and Myall blocks has the most extensive polarity dataset (Fig. 4), based on the work of Geeve *et al.* (2002) and Klootwijk (2016, 2019). This appears to show at minimum five normal polarity magnetozones (Fig. 4), labeled here TB (for Tamworth Belt). The detailed sampling of the Gilmore Volcanic Fm (Gresford Block) by Klootwijk (2019) appears consistent with the data of Geeve *et al.* (2002) from the same unit, to define a rather more detailed pattern of polarity changes through this unit (Fig. 4).

In the NW Tamworth Belt, at the Rocky Creek syncline and Werrie syncline, the Mississippian units are the Caroda Fm and the Merlewood Fm respectively (Fig. 5). The Merlewood Fm in the Werrie Syncline

has best age control in the Kydalyne Member, which has faunas from the *Linoprotonia tenuirugosa* Subzone of the *Delepinea aspinosa* Zone (Roberts & James 2010; Fig. 5). The same brachiopod subzone is present in the limestone unit in the upper part of the Horton River section (Caroda post office bridge; Mory 1980). It seems likely that the normal polarity level Klootwijk (2002) found in the upper andesite level in the Merlewood Fm, is therefore that found in the Koorngal Dacite and unnamed ignimbrite at Horton River (Opdyke *et al.* 2000) and in the High Valley Tuff (Klootwijk 2002) at Moorabool (Fig. 5). This may be TB2n which is within the *D. aspinosa* Zone in the Isismurra Fm in the Rouchel Block (Fig. 4). The Barney Springs Andesite is present in the upper part of the Caroda Fm (Wang *et al.* 2001), and has similar trace element chemistry and polarity to the Koorngal Dacite (Roberts *et al.* 2003), suggesting likely equivalence. In addition, the similarity of the corrected (Black *et al.* 2003) SHRIMP date ( $339.5 \pm 3.7$  Ma) on the Barney Springs Andesite (Roberts *et al.* 2003a; Fig. 5) and the U-Pb date ( $342.8 \pm 2.7$  Ma) from ignimbrites in the Albano region above TB2n (Roberts *et al.* 2006; Klootwijk 2016; Rouchel Block; Fig. 4) indicates that all these levels may represent polarity magnetozone TB2n (Figs. 4, 5).

Normal polarity site 71 of Opdyke *et al.* (2000) in the lower part of the Merlewood Fm (Werrie Syncline, right of Fig. 5) may therefore be magnetozone TB1n, which in the Rouchel Block is in the early Molinacian, Isismurra Fm (Fig. 4). The magnetozone correlations in Figs. 4 and 5 in the Visean, may be rather optimistic considering the sparse sampling and uncertainties in correlation and it is equally likely that the polarity pattern is more complex than suggested in Fig. 4. Anderson *et al.* (2003) also found normal polarity in the Mamberra Andesite Member of the Gilberton Fm (approx 335 Ma ( $\pm 7$ ), ASUD 2020) which appears to be Visean (Oversby & Mackenzie 1994), and may be one of the TB3n- TB5n magnetozones, although this formation extends into the Famennian, based on the plant fossil *Leptophloeum australe*.

In the Pennsylvanian of the Rocky Creek Syncline region, from the northern part of the Tamworth Belt (Fig. 6), Opdyke *et al.* (2000) and Klootwijk (2002) define the base of the Kiaman Superchron, with the last assured normal polarity level at the Wanganui Andesite Member (Roberts *et al.* 2003a). This normal polarity magnetozone extends downwards through the underlying sandstones, but not necessarily to the base of the Clifden Fm (Fig. 6). Opdyke *et al.* (2000) also identified a normal polarity site 118, overlying the Peri Rhyolite Mbr (The Tops section, Rocky Creek; Fig. 6), but full confidence in its stratigraphic position was complicated by faulting. Klootwijk (2002) has two sites with normal polarity near the base of the Ermelo Pyroclastics (Fig. 6), also with overlying and underlying reverse polarity sites of Opdyke *et al.* (2000), indicating an additional normal magnetozone, which may be brief. These have been labeled TB6n and TB7n (Fig. 6). Overlying the Clifden Fm, the Rocky Creek Conglomerate and the Lark Hill Fm contain sites only of reverse polarity.

Anderson *et al.* (2003) have found normal polarity in the upper most site from the Routh dacite (SHRIMP  $321.9 \pm 6.6$  Ma) an age which overlaps the TB7n magnetozone in the Clifden Fm (corrected date of  $322.4 \pm 2.8$  Ma) which may represent the same normal magnetozone. In these Pennsylvanian successions additional regional correlation levels are provided by the glacial intervals of Fielding *et al.* (2008) with the C1 glacial in the Spion Kop Conglomerate inferred by them to be in the early Namurian (Serpukhovian), and the C3 glacial in the Eulowie Pyroclastics Member and upper most Rocky Creek Conglomerate (Fig. 6), inferred to be in the early Moscovian (Fielding *et al.* 2008).

In the Werrie syncline of the central Tamworth Belt, Opdyke *et al.* (2000) and Klootwijk (2003) found

only reverse polarity in the Currabubula Fm (Fig. 7). These datasets overlap in corrected SHRIMP age (~ 316 Ma) with the data from the upper part of the successions in the Rocky Creek Syncline to the north (Figs. 6, 7). The corrected SHRIMP dates from the Currabubula Fm suggest the C3 glacial is around 316 Ma (Fig. 6), and in the Eulowie Pyroclastics and upper Rocky Creek Conglomerate, somewhat similar between 314-317 Ma, confirming the reverse-only polarity in the upper part of the Rocky Creek Conglomerate.

### Volcanics from the British late Visean

The study of Piper *et al.* (1991) has provided one of the better age-constrained studies on lavas (and intrusives) interbedded in late Asbian and early Brigantian limestones in Derbyshire. This study followed earlier pioneer studies on these rocks by Belshé (1957) and Everitt & Belshé (1960), also finding both reverse and normal polarity. Piper *et al.* (1991) sampled lava's at differing stratigraphic levels over some 30 km in a limestone platform succession in the upper part of the Bee Low Limestone Fm and the Monsal Dale Limestone Fm (Fig. 8). These lavas were generated at a number of separate volcanic centres probably at separate events in time. Due to an absence of a detailed section-based biochronology, the relative age relationships of the lavas across this area are not entirely secure, being based largely on lithostratigraphy from bore-hole records (Walters & Ineson 1981). The biochronology for the limestones comes from a mixture of coral biozones, supplemented by ammonoids, foraminifera and some conodonts, mostly at the formation level (rather than section based).

The Bee Low limestone Fm contains the late Asbian coral biozone G of Mitchell (1989). Foraminifera assemblages from this formation in the Eyam Borehole (Strank 1985) indicate a mid to late Asbian age for the Bee Low Limestone, with the volcanic-bearing upper part of the formation around the Asbian-Brigantian boundary (i.e. somewhat lower than the lithostratigraphic boundary; Fig. 8). Ammonoid fauna from the Bee Low limestone Fm indicate the *G. globostriatus* subzone (B2b Subzone; Waters *et al.* 2011) of the late Asbian. The brachiopod *Davidsonina septosa* which has been considered diagnostic of the late Asbian (although also known in the earliest Brigantian, Somerville & Strank 1984) is found below the Lower Millers Dale Lava in the Cressbrooke Dale area (Fig. 8). In these units the Asbian-Brigantian boundary is usually placed at the disconformity between the Bee Low and Monsal Dale limestones (Butcher & Ford 1973; Walkden 1977; Gutteridge 1989), conforming to the Mississippian sequence 9 - 10 boundary (Herbig *et al.* 2016).

Coral-based biozones are most detailed in the Brigantian, with the Lower and Upper units of the Monsal Dale Limestone Fm corresponding to coral divisions H and I of Mitchell (1989), with the overlying coral biozone J missing or undetected. Divisions H and I correspond to the early part of the RC8 Belgium coral biozone (Poty *et al.* 2006). The overlying Eyam Limestone Fm is assigned to coral biozone K (Mitchell 1989; uppermost part of RC8 in Belgium). This agrees with the presence of the latest Visean to early Serpukhovian conodont *Lochriea mononodosa* (Smith *et al.* 2017) and an ammonoid P<sub>2</sub> subzone assemblages in the overlying Eyam Limestone Fm (George *et al.* 1976). The regional correlation of lavas within the Monsal Dale Limestone Fm is less secure above the basal part, since while correlation is possible in the central area using the Litton Tuff (Butcher & Ford 1973) and Cressbrook Dale Lava (Waters *et al.* 2009) and the Upper Dale and Hobs House coral bands as guides (Fig. 8; Butcher & Ford 1973), correlations south to the Lathkill Dale to Matlock area are less confident. Within the southern area, the Upper Matlock Lava is high in the Monsal Dale Limestone Fm, with the Conksbury Lava correlated to a level a little below the Upper Matlock Lava (Bridge & Gozzard 1981, based on dark-

coloured limestone intervals). Walters & Ineson (1981, their Fig. 8) correlated the Conksbury Lava with an ash level, below the Upper Matlock Lava, based on its relationship to the *Orionastrae placenta* band in the Wince No1 borehole.

The relative correlation of the Shacklow Wood area to the Lathkill Dale area is the least well defined, with Bridge & Gozzard (1981) correlating the Litton Tuff to the Conksbury Lava, although the relative position based on formation thickness of these units in the Monsal Dale Fm, more likely indicates a lower level closer to the Lathkill Dale Lava for the Litton Tuff correlation (Fig. 8). This suggests the Upper Matlock Lava and the Shacklow Wood Lava represent different normal polarity intervals (Fig. 8). Walters & Ineson (1981) similarly correlate the Shacklow Wood Lava to a level below the Lathkill Shell bed, and probably also below the Lathkill Dale Lava. For reasons not explained, Piper *et al.* (1991) placed both the Shacklow Wood and Lee Bottom lavas between the Lower Matlock and Winster Moor lavas, which seems erroneous. Immediately below the Lathkill Lodge Lava in the Haddon Fields borehole is the 'Brigantian' coral *Diphyphyllum lateseptatum* (Sommerville & Strank 1984) which Aitkinhead *et al.* (1985) assumed to indicate the Asbian-Brigantian boundary interval (Fig. 8), although this inference contradicts the previous proposed relationships (*D. lateseptatum* first appears in the RC7/ $\beta$  coral biozone in Belgium, in the later part of sequence 9; Poty *et al.*, 2006; 2014). These proposed correlations give three normal polarity intervals, all within the early Brigantian (Fig. 8). Alternatively, foraminiferal data from the Eyam Borehole (Strank 1985) suggests the Asbian-Brigantian boundary may be in the Chee Rock Mbr around the level of the two lower reverse polarity lava's in the Castleton-Bradwell Moor area (Fig. 8).

Everitt and Belshé (1960), Wilson (1960) and Wilson & Everitt (1963) were the first significant studies to use thermal demagnetization procedures on Carboniferous volcanic rocks, working on the Kinghorn successions in Fife (Scotland) to isolate cleaned dual polarity directions. This followed preliminary undemagnetised data published on these same sample sets by Clegg *et al.* (1957), in which they identified a magnetic polarity boundary within the succession of lavas. Following Wilson's PhD work (Wilson 1960), partly on these units, Wilson and Everitt (1963) later synthesized their data on the succession of the Kinghorn Volcanic Fm, relating the detailed sampling by Everitt to the numbered stratigraphic log information in Geike (1900) to provide a detailed polarity record through part of the late Viséan (Fig. 9). Torsvik *et al.* (1989) later measured some of the same lava's, but without the stratigraphic detail provided by the study of Wilson & Everitt (1963). However later, Wilson (1966) revised his interpretation of the fidelity of the geomagnetic field recording at Kinghorn, finding a relationship between the quantified percentage of optically observable ilmenite and polarity (see the table in Fig. 9), inferring a (mineralogically-driven) self reversal mechanism instead (see discussion in McElhinney 1973). This relationship can clearly be seen, using the basaltic petrological classification of Allan (1923) shown in the data in Fig. 9. This relationship is due to an unfortunate sampling bias to normal polarity basalts with abundant augite-olivine phenocrysts in the lower part of the Kinghorn succession, mostly above the 80 m level (Fig. 9). At this time a fierce debate about the 'reality' of geomagnetic field reversals was raging (McElhinney 1973) due to a correlation in some studies (with dubious statistics; Merrill 1985) between igneous rock petrology and polarity (Wilson & Watkins 1967). It was not until the late 1970's that evidence from baked contacts of igneous bodies was more widely accepted to prove the reality of field reversals (McElhinney & McFadden 2000). The conclusions of Wilson (1966) undoubtedly clouded the great significance of the Kinghorn data. Remarkably, the pioneering work of Wilson and Everitt (1963) still stands 6 decades later as the best stratigraphically- constrained magnetic polarity study on

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Carboniferous lavas.

The chronology of the lava succession has improved since these early palaeomagnetic studies. From Buntisland in the west to Kircaldy in the east, the lavas are easily-related to other successions in the Scottish Midland Valley, by a lithostratigraphy (Brown *et al.* 1999; Guirldham *et al.* 2003), linked to miospore zonations (Brindley & Spinner 1989; Owens *et al.* 2005), and some foraminifera data from the upper part (Lower Limestone Fm) of the Kinghorn-Kircaldy succession (Karbub 1993; Cozar *et al.* 2008). Palynological correlation suggests the base of the Lower Limestone Fm (base of Hurlet Limestone) is equivalent to the base of the 1st Abden Limestone at Kinghorn (Brindley & Spinner 1989).

The lowest part of the miospore *Bellisporites nitidus* - *Reticulatisporites carnosus* (NC) Zone can be placed above the lava succession within the Second Abden Limestone (Fig. 9; Brindley & Spinner 1989; Owens *et al.* 2005), which is the equivalent of the St Monance Little Limestone (Owens *et al.* 2005). In northern England and Scotland the base of this zone is within the late Brigantian (P2b-P2c ammonoid zones), with the underlying *Tripartites vetustus*-*Rotaspora fracta* (VF) zone extending into the late Asbian (McLean *et al.* 2018). The overlying Seafield Tower Lmst (equivalent of the Charlestown Main Limestone; Fig. 9) is of latest Brigantian age (ammonoid zone P2c; Wilson 1980; Owens *et al.* 2005).

The Asbian-Brigantian boundary is not easily located in the succession at Kinghorn, although (Karbub 1993) suggested it was located between the 1<sup>st</sup> and Second Abden limestones, but which is inconsistent with the data from Cozar *et al.* (2008) who indicated the base of the Lower Limestone Fm (Hurlet Limestone) was mid Brigantian. Unpublished foraminifera data (some from the same sections as Karbub 1993) of P. Cózar from the First and Second Abden Limestone, the St. Monans Little and St. Monans Brecciated limestones (of east Fife) indicates the Asbian-Brigantian boundary lies below the 1<sup>st</sup> Abden Limestone (Pers. Comm. Pedro Cózar, 2020).

At Kinghorn the position of the base of the underlying VF miospore zone is less clear, but may possibly extend as low as the Houston Coal, thought to be above the Dunnet Shale at Kinghorn (Allan 1923, Francis 1961; Brindly & Spinner 1989), although this is based on a correlation of this coal southwards (Francis 1961). In the section, the last VF biozone indicators are in the first Abden Limestone (Brindley & Spinner 1989), with all underlying samples from the *Raistrickia nigra*-*Triquitrites marginatus* (NM) miospore zone. The Burdiehouse Limestone and its correlatives, are a well defined marker in this region (Guirldham 1998) in the lower part of the NM miospore zone (Rex & Scott 1987), with the base of the NM biozone normally placed in the early Asbian (Owens *et al.* 2005). These data suggests the Kinghorn Volcanic Fm ranges from early Brigantian at the top, through the late Asbian, possibly extending into the mid Asbian. A K-Ar whole rock age of  $338 \pm 4$  Ma from a roadside lava sample between Burntisland and Kinghorn (Fitch *et al.* 1970) indicates an early Holkerian- late Asbian age range, using the timescale of Richards (2013). Ar-Ar dates from the lavas have failed to yield sensible ages (Monaghan & Browne, 2010).

At Kinghorn some later intrusive sills in the top part of the succession, need to be removed for magnetostratigraphic comparison, but do appear to extend the polarity data from the Derbyshire lavas down through the Asbian, with the reverse polarity in the upper 90 m of the Kinghorn Volcanic Fm probably equivalent to that seen in the late Asbian in Derbyshire (Figs. 8, 9). These two studies appear to jointly define a composite polarity from around the mid Asbian into the later parts of the early

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Brigantian (~332-335 Ma; Richards, 2013). The normal polarity in the mid to lower part of the Kinghorn Volcanic Formation may relate approximately to the level of the normal polarity Paterson Volcanics (magnetozone TB5n) in the Australian sections (Figs. 4, 9).

## **North American magnetostratigraphic data**

The most comprehensive modern Carboniferous magnetostratigraphic studies are from the Minudie-Joggins sections in Nova Scotia (DiVenere & Opdyke 1991b), Maringouin Peninsula, New Brunswick (DiVenere & Opdyke 1990), Cape Breton Island (Opdyke *et al.* 2014) and the Mauch Chunk Formation in Eastern Pennsylvania (DiVenere & Opdyke 1991a). These sections are in fluvial red-siltstones and sandstones in the Mauch Chunk Fm. In Nova Scotia and New Brunswick studies from predominantly red-sometimes grey coloured fluvial successions. Opdyke *et al.* (2014) attempted to integrate these magnetostratigraphic datasets, but for the Mauch Chunk Formation utilized a stratigraphic age analogue from Virginia, which was inappropriate to E. Pennsylvania. Hence, a revised age model for the Mauch Chunk Fm will be examined in detail, based around regional relationships, which are crucial in reconstructing a magnetostratigraphy in the late Mississippian.

### **Magnetostratigraphy of the Mauch Chunk Fm**

The magnetostratigraphy from the Mauch Chunk Formation is based around four sections (Fig. 10) from a transect across the southern Anthracite Coal Field of eastern Pennsylvania (DiVenere & Opdyke 1991a, Opdyke & DiVenere 2004). Within Pennsylvania, the Mauch Chunk Fm displays lateral facies change from much thicker fluvial red-beds in the east to much thinner red beds, interbedded with a variety of marine carbonates and clastics to the SW (Fig. 11). Within the S. Anthracite Coal Field area the Mauch Chunk Fm has upper and lower divisions, transitional in facies to the underlying Pocono Fm (Mt Carbon Member) and overlying Pottsville Fm (Tumbling Run Mbr). Filmore *et al.* (2012) formalised three divisions of the Mauch Chunk Fm in the S. Coal Anthracite Fields, with the names Lavelle, Indian Run and Hometown members, corresponding to the similarly defined, unnamed previous divisions (Wood *et al.* 1969). The Mauch Chunk Fm boundaries are based on the lowest presence of red beds at the lower boundary, and highest presence of red beds at the upper boundary (Wood *et al.* 1969). In the Anthracite Fields area, the Hometown Mbr interdigitates northwards, with the Indian Run Mbr and with the Tumbling Run Mbr (of the Pottsville Fm) southwards (Wood *et al.* 1969; Fig. 11). Consequently Mauch Chunk member and formation boundaries are probably time-transgressive within E. Pennsylvania. It is not entirely clear how the magnetostratigraphy in the south Lavelle section (Fig. 10) correlates to that in the Jim Thorpe section (Opdyke & DiVenere 2004). Opdyke *et al.* (2014) used the higher (dotted) correlation in Fig. 10, but if the Lavelle Member is ~180 m thick (Wood *et al.* 1969) and not too much is removed by faulting in the south Lavelle section, a correlation lower in the south Lavelle section is possible (solid correlation line in Fig. 10).

The only direct age dating of the Mauch Chunk Fm in the Anthracite Coal Fields area is based on the *Adiantites antique* macroflora from the upper part of the Indian Run Mbr (Fig. 10) at the Schuylkill Gap section (Jennings 1985). In Europe, *A. antique* occurs primarily in the Viséan, but is also known from the early Namurian (ca. Serpukhovian). Vertebrate tracks from the Mauch Chunk Fm provide a similar, Mississippian age assignment, but based on poorer biochronological resolution (Filmore *et al.* 2012).

In the S. Anthracite Field area the Mauch Chunk Fm appears to display continuous deposition from the underlying Pocono Fm to the overlying Pottsville Fm (Fig 11), these formations also provide bracketing

age constraints. Plant fossils described by Read (1955) from the upper part of the Pocono Fm in the S. Coal Anthracite Fields (around Pottsville and Jim Thorp) were assigned to the '*Triphylopteris*' spp. Zone of Osagean age (Read & Mamey 1964), now assigned to megafloreal Zone 2 (Wagner 1984; Eble *et al.* 2009) of the late Tournaisian-early Viséan.

The Pottsville Fm overlying the Mauch Chunk Fm has a complex regional distribution due to unconformities (Fig. 11), the most significant one of which, in the S. Anthracite Coal Field area, is within the upper part of the Pottsville Fm (Edmunds *et al.* 1999; Fig. 11). Megafloreal Zones 4 and 5 (Read & Mamey 1964) occur within the Tumbling Run Mbr. Read and Mamey's Zone 4 is within Lykens Valley No. 5 and No. 6 coals (around the mid part of the member), about 120 m above its base (Edwards *et al.* 1999) and would appear to indicate an early Langsettian (Westphalian A; Fig. 11; mid Bashkirian) or slightly older age (Eble *et al.* 2009). European substages are used here, since later this data will be related to Canadian Maritime successions, which also use these substages for division. Megafloreal Zone 5 is within Lykens Valley No. 4 coal, and is in the upper part of the member (~160 m above base of member) and is around mid Langsettian (Eble *et al.* 1999). These age constraints for the Mauch Chunk Fm in the S. Anthracite Coal Field suggest the magnetostratigraphy of DiVenere & Opdyke (1991a) and Opdyke & DiVenere (2004) may encompass, at maximum, an interval from the early Viséan to the later mid Bashkirian (i.e. to the late 'Namurian'; Fig. 11).

Additional age constraints can also be attempted using longer-range correlation to W. Pennsylvania and N. Virginia, where marine units interdigitate with the Mauch Chunk Fm, and provide potentially correlative regressive and transgressive cycles with the Mauch Chunk Fm units in W. Pennsylvania (Brezinski 1989a). Uttley (1974) examined in detail the correlation of the Mauch Chunk Fm in SW Pennsylvanian to Ohio and correlated the Loyalhanna Limestone to the Ste. Genevieve Limestone in Ohio. The two units also have similar brachiopod assemblages (Kammer & Lake 2001). Based on foraminifera and conodonts Maples and Waters (1987) placed the base of the Ste. Genevieve Limestone at the base of the Chesterian, which is equivalent to the base of 16i zone of Mamet and the base of foraminifera zone MFZ14 (Poty *et al.*, 2006). Uttley (1974) correlated the Wymps Gap limestone (Fig. 11) to the Jonathan Creek Limestone of Scatterday (1963), the lower part of which contains conodonts from the *Gnathodus bilineatus-Cavusgnathus charactus* biozone, the 2<sup>nd</sup> conodont zone of the Chesterian (Scatterday 1963; Repetski & Stamm 2009). Based on conodonts, Horowitz & Rexroad (1972) have also suggested a 'pre Glen Dean' age for the Wymps Gap Limestone. Trilobites from the Wymps Gap Limestone also place this unit in the early Chesterian (Brezinski 2009). Uttley (1974) correlated the Reynolds Limestone (Fig. 11) to the Glen Dean Fm, which is mid Chesterian in age (*Kladognathus mehli* conodont zone, Al-Tawil *et al.* 2003; Repetski & Stamm 2009). Kammer & Spinger (2008) and others have suggested the same correlation.

Two possible options (long and a short duration Mauch Chunk options) have been proposed to correlate the SW. Pennsylvania marine units to the Mauch Chunk Fm in the Anthracite Coal Fields in E. Pennsylvania. Edmunds (1996) and Ettensohn (2009) placed the Loyalhanna Limestone in the mid part of the E. Pennsylvania Mauch Chunk Fm, whereas Berg *et al.* (1983) and Brezinski (1999, 2009) place the Loyalhanna Limestone near the base (as shown in Fig. 11). In SW Pennsylvanian, N. Maryland and Ohio two major progradations of coarser clastics are represented by the units between the Loyalhanna and Wymps Gap limestones, and in the part of the Mauch Chunk Fm above the Greenbrier Fm (Uttley 1974; Brezinski 1989a, 1989b), generating two major westwards directed progradation pulses in the

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Chesterian (Fig. 11). These may relate to the sequence C5+C6 lowstand and the sequence C10+C11 lowstands of Al-Tawil *et al.* (2003) from W. Virginia. Currently there is no similar sequence stratigraphy for the Mauch Chunk Fm from E. Pennsylvania, but it is possible that the lower of these westwards-directed clastic progradations is represented by the sandstone dominated interval in the lower part of the Indian Run Mbr at Jim Thorpe (Fig. 10), and the upper progradation by the Hometown Member. Perhaps the three calcrete rich intervals (relatively wetter environment; Fig. 10) in the Indian Run Mbr may be the equivalent of the limestone units in SW Pennsylvania? This correlation option is compatible with the lower placement of the Loyalhanna Limestone, and suggests the Meramecian may be absent or condensed in the Lavelle Member or upper part of the Pocono Fm. This age option is broadly similar to that suggested by Opdyke *et al.* (2014) based on magnetostratigraphic correlation to the eastern Canadian sections.

For the long duration Mauch Chunk option, the sandstone dominated interval in the lower part of the Indian Run Mbr, may represent the upper clastic progradation in SW Pennsylvania, compatible with the mid- Mauch Chunk position of the Loyalhanna Limestone (i.e. not that shown in Fig. 11). Based on cyclostratigraphy, Kodama (2019) has proposed an accumulation rate of the Indian Run Member of 5.69 cm/Ka, which using the average Mauch Chunk Fm thickness estimates of Woods *et al.* (1969) would give a duration of ~20 Ma for the formation, compatible with the duration of the Chesterian plus Meramecian (Richards 2013). This cyclostratigraphic duration is compatible with Edmunds (1996) and Ettensohn (2009) for the mid-placement of the Loyalhanna Limestone Mbr, suggesting the magnetostratigraphy of the Mauch Chunk Fm may extend into the early Visean. Currently there is insufficient stratigraphic data internal to the N. Appalachian basin, to distinguish these two age options for the Mauch Chunk Fm.

### *Magnetostratigraphy from New Brunswick and Nova Scotia*

Studies in the Cumberland Basin in New Brunswick and Nova Scotia, and on Cape Breton Island, have defined a composite magnetostratigraphy through the Brigantian to early Langsettian (latest Visean to mid Bashkirian) (DiVenere & Opdyke 1990, 1991b; Opdyke *et al.* 2014; Fig. 12). The younger part of this interval is of strong significance, since it is around the base of the Kiaman Superchron, and key polarity marker in the Carboniferous. The age constraining biostratigraphy is largely provided by a miospore zonation, which links it to the British substages, since the two areas were in the same floral province (Utting *et al.* 2010). A major unconformity separates Mississippian from Pennsylvanian strata (Fig. 12). Whilst the palynostratigraphy has similarity to the European miospore zonations, it also has significant differences, which do not allow the British substage boundaries to be confidently located, but instead provide weaker age equivalence (Utting *et al.* 2010).

The Mississippian data provides good inter-section consistency, allowing a composite polarity to be constructed (re-labeled more simply from the composite in Opdyke *et al.* 2014; Fig. 12). Palynological data from the Middleborough Fm is lacking, and its assumed Brigantian age is based on the age for the underlying Lime-kiln Brook Fm (faunal division B of the Windsor Group; Jutras *et al.* 2016). The faunal division B has been dated as late Asbian (Giles 2008), but Mamet (1970) has assigned foraminifera Zone 15 to this division, which may be early Warnantian (approx. early to mid Asbian; Poty *et al.* 2006). Since the Lime-kiln Brook Fm is laterally transitional into the lower part of the Middleborough Fm (Jutras *et al.* 2015), it is possible the oldest part of the magnetostratigraphy (magnetozones NN1-NN2?) may extend into the Asbian rather than be entirely Brigantian. Sedimentological work suggests that the Shepody

and Claremont formations (Claremont Fm = Enrage Fm in New Brunswick) are both bounded by hiatus of unknown duration (Jutras *et al.* 2015), so the magnetostratigraphy is possibly fragmented. The magnetozone NN7n may be somewhat thicker than indicated in the polarity composite (Fig. 12), since the correlation relationship between the Spring Valley #1 core and other sections, which contain the *Reticulatisporites carnosus* Zone (such as in the Pomquet Fm) are not constrained by zonal boundaries. Opdyke *et al.* (2014) correlated the reverse magnetozone in the upper part of the Enrage Fm in the Spring Valley #1 core with the equivalent of NN5r, which seems unlikely, since NN5r magnetozone is within the older *Grandispora spinosa*–*Ibrahimisporites magnificus* Zone in the Maringouin Peninsula and Minudie-Joggins sections (Fig. 12).

The Pennsylvanian age sections in the Port Hood Fm (in the Margaree Member) from Cape Breton Island preserve normal magnetozone NN8n overlain by a reverse polarity interval NN8r–NN9r (Fig. 12). Opdyke *et al.* (2014) suggested the base of the Boss Point Fm (at the Joggins section) related to this reverse polarity interval seen in the upper part of the Broad Cove Chapel section (similar to correlation shown in Fig. 12). However, the correlation of NN9n is not well constrained by miospores since all these units are within the *Reticulatisporites saetosa* Zone, and its possible NN9n (within the Chignecto Bay Mbr of the Boss Point Fm; Fig. 12) may fall above the section at Broad Cove Chapel. It is likely magnetozone NN8n is of Yeadonian age and so is probably the equivalent to magnetozone DB1n in the Lower Bashkirian in the Donets Basin (Fig. 1).

In the New Glasgow Fm from the eastern part of the Cumberland Basin, Buchan & Chandler (1999) found nine out of 13 sites were normal polarity, with all located within the lower part of the distal member of this formation. The New Glasgow Fm overlies the Boss Point Fm, which is a mappable unit from west to east across the Cumberland Basin. The New Glasgow Formation has similar sedimentology to the Polly Brook Fm, the alluvial-fan, lateral equivalent of the better dated Langsettian age units from the western part of the Cumberland Basin (Allen *et al.* 2011). Palynology of the New Glasgow Fm indicates probable Langsettian age, but is presently poorly defined (Buchan and Chandler 1999; Allen *et al.* 2011.). The New Glasgow Fm also has a macroflora dominated by cordaitaleans with minor pteridosperms and ferns, similar to those from the Little River and Joggins formations (with reverse polarity; Fig. 12) from the SW Cumberland Basin (Falcon-Lang 2006). These relationships suggest an additional substantive normal magnetozones may exist above the reverse polarity Little River Fm, perhaps equivalent to magnetozone DB2n in the Donets Basin (Fig. 1), since a substantial thickness of Langsettian age strata overly the Little River Fm in the SW Cumberland Basin (Allen *et al.* 2011).

## Tengiz, Kazakhstan

A magnetostratigraphy was determined by Ratcliffe *et al.* (2013) across the Mississippian–Pennsylvanian boundary using core from the Tengiz carbonate platform in Kazakhstan (Fig. 13). Their study utilised a range of sequence stratigraphic, trace element geochemistry and isotope geochemistry to constrain the correlations between a number of wells. The magnetostratigraphic datasets are from platform-top well T-220 and platform flank well T-5056 (Fig. 13). Existing correlation relationships on this platform have largely used down-hole log based data to construct a sequence stratigraphic model adapted over the whole platform (Collins *et al.* 2006; Kenter *et al.* 2006), tied to a large number of log-correlation tie points. In the data from the younger part of the carbonate platform studied by Ratcliffe *et al.* (2013) these surfaces are the Lvis13, Serp\_SSB and Bash\_SSB boundaries, which cover an interval from the late Viséan to the late Bashkirian (Kenter *et al.* 2006; Fig. 13). Chemostratigraphic (trace element based) From Lucas, Schneider *et al.* (eds) The Carboniferous Timescale, Spec. Publ. Geol. Soc. Lond.

divisions (Packages 1 to 6, based on core and cuttings) allowed additional constraining relationships between wells to support these inter-well correlations shown in Fig. 13. Foraminifera data (Brenckle & Milkina 2003) indicate a broad assignment of some parts of the cored intervals to the Russian substages (Figs. 13 and 14b). The sequence stratigraphic divisions also have a foraminiferal biostratigraphy (Brenckle & Milkina 2003) linked to the traditional Russian substage/horizon divisions. This can in turn be linked to the sequence stratigraphic model (Weber *et al.* 2008) to allow a reasonably well-defined chronostratigraphic model for the boundaries (Kenter *et al.* 2006).

The 'Serp\_SSB' sequence boundary links the two wells together, with normal polarity magnetozones TZ5n in both wells spanning this boundary, and underlying normal magnetozones TZ4n, and overlying normal magnetozones TZ6n also detected in both wells (Fig. 13). In addition, whole rock carbon isotope data linked to the chemostratigraphic packages, sequence stratigraphy and magnetostratigraphy in well T-220 allow refinements of the chronostratigraphy particularly in the Bashkirian (Fig. 14a). This suggests that the polarity succession in well T-220 probably starts around the base Serpukhovian (based on correlation to the Antler Basin record; Fig. 14a) and ends around the mid Bashkirian (later part of Akavasian substage), based on isotope correlations to the Askyn River section (Fig. 14a). The sequence boundary Serp\_SSB is close to the Serpukhovian-Bashkirian boundary, clearly shown by the large negative  $\delta^{13}\text{C}_{\text{carb}}$  excursion in the late Serpukhovian in the correlated sections (Fig. 14a). These correlations indicate the uppermost magnetozones TZ6n at Tengiz is mid Akavasian age (Fig. 14a).

## Synthesis of a polarity timescale

### Mississippian

To infer geomagnetic polarity in the Tournaisian- Visean (prior to the oldest parts of the magnetostratigraphic data from the Mauch Chunk Fm), there are data available from: 1) the rather widely spaced sampling from the Tamworth Belt in Australia (Fig. 4); 2) polarity bias data from palaeopole-type studies (Fig. 15). 3) The study of Liu *et al.* (1991) covering the Devonian-Carboniferous boundary, and 4) earliest Carboniferous data synthesized by Kolesov (2007) from NE Asian sections.

To add to the polarity bias data for igneous rocks (Table 1; Fig. 3) polarity bias data has been collated from sedimentary rocks (Table 2). This dataset started with the MagIC and palaeomagnetic databases, but filtered the data set to exclude; a) poorly dated studies, or studies with a wide chronostratigraphic range of sampling (ideally <12 Ma range), b) studies which had limited sampling, c) studies whose details could not be scrutinized, and d) suspected remagnetisations. The Serpukhovian- mid Bashkirian interval has adequate magnetostratigraphic data, so was excluded in this analysis (Fig. 15). A magnetozones that is based solely on polarity bias data, with widely spaced stratigraphic sampling, is given the subscript B (i.e. bias magnetozones M1r<sub>B</sub>, Fig. 15). This is the case for much of the Tournaisian to mid Visean interval. A bias magnetozones implies that it probably contains (as yet unknown), opposite polarity sub-magnetozones and so this designation also expresses the type of data that contribute to the final polarity timescale.

The study by Liu *et al.* (1991) is currently the best biostratigraphically constrained magnetostratigraphic study in the Carboniferous, although it is on a condensed succession. Reverse polarity occurs across the Famennian-Tournaisian boundary (base *Siphonodella sulcata* Zone) with an overlying normal polarity magnetozones (here called M1n<sub>B</sub>; Fig. 15) beginning in the last part of the *S. duplicata* Zone in the

Daposhang section (Fig. 15). A similar polarity pattern (Kolesov 2007) seems to occur in the Kamenka section in NE Russia (Gagiev *et al.* 1991), and the Kozhim section in the northern Urals, although in the former MI1n<sub>B</sub> begins in the *S. sulcata* Zone, perhaps indicating an issue with the conodont zonations (Kolesov 2005). This early Tournaisian bias magnetozones MI1n<sub>B</sub> seem to be supported by polarity bias data from igneous rocks (Fig. 15). The magnetozones TB1n from the Tamworth Belt (four sampling levels) is younger than MI1n<sub>B</sub>, since the Lower *crenulata* Zone occurs in the underlying Brushy Hill Limestone in the *sol* brachiopod zone (Mory & Crane 1982; Fig. 4). A sketch outline (without details) of a Tournaisian magnetostratigraphy has also been described by Kolesov (1984), from southern Belgium but other studies have indicated a pervasive early Permian remagnetisation of most Devonian and Carboniferous sediments in Belgium (Thominski *et al.* 1993; Zegers *et al.* 2003), so its reliability is unclear.

Overlying magnetozones MI1r<sub>B</sub> to MI4n<sub>B</sub> are zones of polarity bias, since there appear to be no true magnetostratigraphic studies in this interval. Magnetozones MI2n<sub>B</sub> and MI4n<sub>B</sub> seem to be borne out by the equivalent TB1n and TB2n magnetozones from the New England Orogen Tamworth Belt (Figs. 4,5, 15), although the polarity pattern is likely incomplete. MI1r<sub>B</sub> is perhaps the least well constrained of these bias zones, although Kolesov (2005, 2007) shows reverse polarity from the Uttykeli and Khurendzha suites, from the Omolon Massif in NE Russia, of mid to late Tournaisian age, similar to that inferred here. The sparse data over the ~30 myr interval of the early Tournaisian to the mid Visean has clearly proved a problematic interval for palaeomagnetic studies. The complexity of missing magnetozones in the MI1<sub>B</sub> to MI4<sub>B</sub> bias zones, can only be guessed at, but comparing the polarity in the late Visean from the Australian dataset, with that based on the composite magnetostratigraphy of sections in Fig. 16, likely gives a rough guide (see upper part of Fig. 15).

The key datasets for age constraining a late Visean to Serpukhovian polarity timescale come from the Maritimes Basin and Tengiz, since these have the best age control, and are relatively long records (Fig. 16). For this reason, the following evaluation starts with the most-securely dated younger units, moving to the older, less securely dated, Mauch Chunk Fm.

Major normal magnetozones NN5n and NN6n-NN7n (in Pendleian and Arnsbergian strata), from the Cumberland- W. Cape Breton basins are the likely equivalent of Serpukhovian magnetozones TZ3n and TZ5n from Tengiz (Fig. 16). It is probable that there is part of magnetozones NN7r (Maritimes Basin) and TZ5n (at Tengiz ) missing, close to the top of the Serpukhovian due to hiatus. Underlying magnetozones NN5n and TZ3n is a dominantly reverse polarity interval (containing thinner magnetozones NN3n, NN4n and TZ2n), which supports the validity of this correlation. The Brigantian- Pendleian boundary is within the early Serpukhovian (Savastopulo & Barham 2014), like these magnetostratigraphic correlations also suggest. Within the Serpukhovian, the relative duration and pattern of polarity changes is rather different between Tengiz and the Maritimes Basin, which is most likely due to the hiatus detected in the E. Canadian sections, which bound the Shepody and Claremont formations (Jutras *et al.* 2015; Fig. 12).

The polarity data from volcanic units at Kinghorn and Derbyshire cover a short, probably mid Asbian to early Brigantian interval, whose ages indicate some overlap (Fig. 16). The reverse polarity in the youngest part of the succession at Kinghorn (intrusive sill-data removed from Fig. 9) is probably the reverse polarity seen in the older lava's in Derbyshire (Fig. 16). Although there are substantial sampling gaps in the Kinghorn succession (Fig. 9), the normal polarity dominance is clear, which appears to be a useful marker (referred to as the MI5n 'marker') for this mid to late Asbian interval. Hence, it is likely

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the Kinghorn normal polarity interval is the equivalent of the NN1n to NN2n magnetozone interval in the Maritimes Basin (Fig. 16). Maritimes Basin magnetozones NN1- NN2 are potentially late Asbian, if the base of the Middleborough Fm is the lateral equivalent of the Brick-Kiln Limestone in the Maringouin Peninsula- Minudie area as suggested by Jutras *et al.* (2016). If this Kinghorn to Maritimes Basin correlation is correct, there must be at least three normal magnetozones above in the Brigantian as seen in the Derbyshire data (Fig. 16), so one or more must be missing between the Brigantian and Pendleian age units in the Canadian sections. These relationships provide a basis for correlations in the Brigantian, using the Mauch Chunk data, which occupies this age interval.

The magnetostratigraphy of the Mauch Chunk Fm, whilst probably the longest record of magnetic polarity in the Mississippian, and has two possible age options. The Hometown Mbr may be either within the Morrowan (early Bashkirian) or within the late Chesterian (late Serpukhovian), depending upon the severity of lateral interdigitation with the overlying Pottsville Fm (Fig. 11). Mauch Chunk correlation Option 1 places the Hometown Mbr in the early Bashkirian and Option-2 places this member in the late Serpukhovian (Fig. 16). Option-1 attempts to use the thickness estimates of the Mauch Chunk Fm as a guide to how much polarity data may be missing in the unsampled interval between MC9n and MC9r. The merit in Option-1 is the reverse polarity dominance in the TZ1r-TZ2r interval (at Tengiz), is like that seen between MC6r to MC8r (Fig. 16). Also in the youngest parts, a tentative reverse polarity magnetozone (probable equivalent of MC11r ) is seen in TZ6n at Tengiz (Figs. 13, 16).

In the lower part of the Indian Run Mbr, magnetozone MC4n is likely the equivalent of the NN1n-NN2n and Kinghorn- normal polarity interval (the chron MI5n 'marker'; Fig. 16). This would be compatible with the lower placement of the Loyalhanna Mbr in the Mauch Chunk Fm (like shown in Fig. 11), as a proxy for the base of the Chesterian (base Chesterian equates to base of foraminifera zone MFZ14; Poty *et al.* 2006 and base of the late Asbian, Cf6 $\delta$  Biozone; Cozar & Somerville, 2004). The correlation alternative of placing the Loyalhanna Mbr in the mid part of the Mauch Chunk Fm (and so the base Chesterian) seems unlikely, since the Asbian and Brigantian strata from other studies would need to fit within the mid and missing part of the Indian Run Mbr (Fig. 16).

Mauch Chunk correlation Option-2 is similar to that used by Opdyke *et al.* (2014) to interrelate the Maritimes Basin and Mauch Chunk datasets, and conveniently places the base of the Pottsville Fm close to the base of the Bashkirian. This is the preferred option here, and the polarity composite is constructed using this in Fig. 16. Between the MI5n 'marker' and MC10n three normal magnetozones are seen in the Derbyshire data and five in the Mauch Chunk Fm (although conceivably more may be present in the unsampled mid part of the Indian Run Mbr). Mauch Chunk correlation Option 2 minimizes the amount of missing thickness and magnetozones in both the Maritimes Basin and Mauch Chunk data. This correlation option indicates the relatively short duration of NN7r and equivalents (e.g. MC11r, MI9r). However, with this option there is no constraint on the relative thickness of magnetochron PE1n, which may be relatively longer than shown on the polarity composite. However, the Tengiz carbon isotope data compared to the Askyn River section, suggests little is missing at this boundary (Fig. 14).

Opdyke *et al.* (2000) was rather ambiguous about how the normal polarity intervals in the lower part of the Clifden Fm from the Rocky Creek Syncline (Fig. 6) may relate to other data and age, since much depends on the SHRIMP dates from these successions, and the Carboniferous timescale. The original  
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SHRIMP date for the Wanganui Andesite was  $319.2 \pm 2.8$  Ma (Fig. 6), but using the corrections of Black *et al.* (2003a) gives  $322.4 \pm 2.9$  Ma for the andesite. Using the array of SHRIMP dates from  $\sim 325$  to  $\sim 321$  Ma from the Clifden Fm, and Arnsbergian U-Pb dates of  $\sim 324$ - $325$  Ma (Pointon *et al.* 2012; Jirasek *et al.*, 2018) would push the base of this formation to around the Serpukhovian- Bashkirian boundary, shown as Option A in Fig. 16. The alternative, based only on the magnetostratigraphy would suggest Option B in Fig. 17. The succession of corrected SHRIMP dates from the Rocky Creek Syncline sections fits rather better with Option A in Fig. 16, which would place the Clifden Fm in the Late Serpukhovian- Early Bashkirian. The anomalously young corrected SHRIMP date of  $317.8 \pm 1.9$  Ma in the upper part of the Clifden Fm may relate to problems in correlating the Lochiel Downs- Darthula section (Fig. 6) and uncertainties in the constraining SHRIMP-based age control (Roberts *et al.* 2003a).

### Pennsylvanian

The Tengiz carbon isotope correlations suggest that magnetozone TZ6n occupies the early to mid part of the Akavasian (Fig. 14), with the magnetostratigraphy not extending to the base Askynbashian at Tengiz. In the Donets Basin data the equivalent Sverokeltmenian-Prikmian boundary (Figs. 1, 14) occurs within the magnetozone DB1n, with an underlying reverse polarity interval in the Sverokeltmenian (i.e. Akavasian at Tengiz; Fig. 14b). This indicates that the reverse interval underlying DB1n is magnetochron PE2r (Fig. 17). The relative thickness of PE2r is unconstrained in any section. Iosifidi & Khramov (2013) also seemed to have detected PE2n and the underlying reverse magnetozone PE1r from Spitsbergen sections, although sampling and stratigraphic details are insufficient to be sure. Reverse and normal polarity intervals of Morrowan age (early Bashkirian) were also confirmed by Nick *et al.* (1991) on reddened palaeosols from the Black Prince Limestone from SW Arizona, but stratigraphic details are insufficient to link to the data in Fig. 17.

Equivalent age magnetozones to the Donets Basin magnetozone DB1n are known from the Maritimes Basin (Fig. 17). Opdyke *et al.* (2000) proposed that the Wanganui 'reversal' (Figs. 16, 17) represents the start of the Kiaman Superchron, although based on the corrected SHRIMP dates, this is too old, since the Wanganui reversal probably represents the base of PE1r (Fig. 16). Russian workers have long considered the Kiaman Superchron to start at the top of magnetozone DB5n in the Donets Basin (Fig. 1; Khramov & Rodionov 1980). This placement had probably evolved from Irving & Parry's (1963) placement of the base of the Kiaman in the "Westphalian or between the Westphalian and Stephanian", that was based on data from the better-dated European data at the time. A placement of the base Kiaman around the Namurian-Westphalian boundary (mid Bashkirian) was also proposed by Irving & Pullaiah (1976), consistent with that proposed here below.

### **Late Bashkirian, Moscovian and younger**

We have seen that several lines of evidence from other studies have suggested the potential reliability of most parts of the magnetostratigraphy from the Donets Basin (Fig. 1) of Khramov *et al.* (1974). These include, similar palaeomagnetic mean directions, using modern demagnetization techniques (Iosifidi *et al.* 2010, 2016) and the multi-section nature of the original data (Fig. 1). Evidence has been previously examined that suggests that normal polarity intervals from igneous rocks may represent the equivalent of magnetochrons in the late Bashkirian- Moscovian. In addition, a probable Langsettian age normal magnetozone in the New Scotland Fm in Nova Scotia (Buchan & Chandler 1999) is the probable equivalent of magnetochron PE4n and DB2n (Fig. 17). This range of support suggests that the Bashkirian to Moscovian magnetic polarity data of Khramov *et al.* (1974) probably provides the best record (at this

time) of polarity changes through the early and mid Pennsylvanian, in spite of its short-comings with respect to out-dated demagnetization techniques.

In addition, Kim *et al.* (1992), Doh & Piper (1994) and Lee *et al.* (1996) have found both polarities in palaeopoles-type datasets (directions passing a fold test) from South Korean Moscovian age red beds of the Manhang and Yobong formations and Hongjom Series of the Pyongan Supergroup (Lee 2010; Table 2; Fig. 17b). Their sample sets are dominated by reverse polarity but with around 5%-15% of the samples showing normal polarity. Those data from the Manhang Fm are from the lower parts of the formation. These Korean Carboniferous magnetizations are challenging to isolate from later remagnetisations, but the careful rock-magnetic and demagnetisation procedures employed by Lee *et al.* (1996) clearly allowed their isolation. These formations are primarily early Moscovian (primarily based on fusulinids; Lee 2010), but the lower parts of the Yobong Fm may extend into the late Bashkirian, so the normal polarity intervals could be the equivalent to DB3n or DB4n (chrons PE5n, PE6n) in the Donets Basin (Figs. 1, 17). The Shatterford sill and Karkonosze granite dual polarity datasets may also be the equivalent of DB4n and PE6n (Table 1; Fig. 17b).

The Kasimovian was widely thought to be entirely reversed polarity, even in the Russian compilations (Fig. 2), yet palaeopoles type studies from Argentina, China and E. Canada (Tables, 1,2; Fig. 17) show a grouping with less than 100% reverse polarity at about 305 Ma. In the Kasimovian to Gzhelian age Dzhingilsaj section (Ferghana, Uzbekistan) the oldest two levels were also interpreted as normal polarity (Davydov & Khramov 1991). These two levels span the *Montiparus montiparus* and *Protriticites psudomontiparus- Obsoletes obsoletes* foraminifera zones in the mid Kasimovian, adding further support to a probable normal polarity interval in the mid Kasimovian. This is referred to as polarity bias zone PE8n<sub>B</sub>, since most support is from polarity bias data (Fig. 17b).

The late Gzhelian contains a brief normal magnetozone, named the “Kartamyshian” by Davydov & Khramov, (1991), for which there is evidence from other palaeopole-type studies and several magnetostratigraphic studies (reviewed by Hounslow & Balabanov 2018). Hounslow & Balabanov (2018) referred to this as CI1n, the base of their magnetostratigraphic numbering scheme for the Cisuralian.

However, other studies (e.g. Steiner 1988) have failed to find normal polarity intervals in the Late Bashkirian to Gzhelian, such as Opdyke *et al.* (2000) from the New England Tamworth Belt (Fig. 17) and Opdyke *et al.* (2014) from the Moscovian to Gzhelian Spring valley #1 Core (Nova Scotia). Magnus & Opdyke (1991) collected 197 samples from 1280 m of the Minturn Fm of late Atokan-Desmoinesian age (mostly Moscovian, Itano *et al.* 2003) from the Arkansas River Valley in Colorado and found only reversed polarity. In addition, Diehl and Shive (1981) detected no normal polarity in 549 samples covering 86% of the Casper Fm in Wyoming. Based on fusulinids the Casper Fm ranges in age from the late Desmoinesian through Missourian and Virgilian (most of Kasimovian-Gzhelian) and into the earliest Permian (Burns & Nestell 2009), although possible hiatuses may be present (Diel & Shive 1981).

This conflicting situation is much like the early Permian, where there is evidence for some brief normal polarity intervals, but some detailed-sampling in red-bed based magnetostratigraphic studies have failed to detect any normal polarity (Hounslow & Balabanov 2018). The reasons for the between-study disparities are un-resolved. In some cases it may be due to stratigraphic complexity due to hiatus, or the logistical difficulty of sampling and so finding, brief normal polarity intervals, in a dominance of reverse

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polarity. However, it most likely relates to the mechanisms of haematite generation in red-beds, which can be both short-term (at or close to time of deposition; Besly & Turner 1983; Turner *et al.* 1985), as well as long term lasting millions of years (Larson *et al.* 1982; Johnson *et al.* 1997). Resolving this issue for the Kiama Superchron goes to the crux of the 'superchron paradigm', in which exclusively single-polarity intervals are thought to punctuate earth history (Driscoll & Evans, 2016).

## A geomagnetic polarity timescale

A magnetostratigraphic age model for the interval between magnetochrons MI4r and PE7r is constructed using the methods proposed by Hounslow (2016), and used by Hounslow & Balabanov (2018) and Hounslow *et al.* (2018). This firstly generates a statistical composite of the magnetostratigraphic data in an arbitrary section-height scale (Hounslow 2016). Data used is from the Mauch Chunk (all 4 sub-sections, Fig. 10), Tengiz, Donets Basin (3 sub-sections, Fig. 1) and Maritimes Basin. In the Maritimes Basin data (Fig. 16) the Middleborough, Claremont Fm and the Port Hood Fm data are only used, since in the compositing process, other section data do not constrain the relative thickness of any magnetostratigraphic unit. The Rocky Creek Syncline data from the Clifden Fm and Ermelo Pyroclastics is not used because of the age and correlation uncertainty, and the irregular sampling. There are three magnetostratigraphic units PE1n, PE2n and PE2r without a relative thickness constraint. To constrain PE1n, the base of PE1r is placed at the top of the Mauch Chunk Fm Lavelle section (Beury Road; Fig. 10). To constrain PE2n and PE2r the base of PE2r is placed at the base of the Davidovka section in the Donets Basin (Fig. 1). The conversion of the statistical composite to time should partly correct for any distortions introduced by these unconstrained magnetostratigraphic units. A linear sediment accumulation model is used for all sections (Hounslow 2016), and the statistical optimization procedure shows that the Port Hood Fm and the Middleborough have the poorest overall match ( $D_j$  of 0.28 and 0.13 respectively; Fig. 18; Hounslow 2016). The overall optimization mis-match across all sections is around 8% ( $D_s = 0.078$ ; Figs. 18a, b). Additional magnetostratigraphic constraints at formation top were used for the Claremont and Middleborough formations (MI9r, MI6r.1n respectively), and at formation bases for the Port Hood and Middleborough formations (PE3n.1n, MI5n.1n respectively), and section base at Tengiz (MI7n.1r). The magnetostratigraphic units at Tengiz were sectioned into two sets at the Serp\_SSB disconformity (Figs. 16, 18).

U-Pb radioisotopic dates are attached to the polarity scale as detailed in Table 3, as shown on the magnetostratigraphic scale in Figs. 16 and 17, and displayed with their uncertainty in Fig. 18c. To derive the magnetostratigraphic ages a Bayesian method (Haslett & Parnell 2008; Parnell *et al.* 2008) is used as implemented in the Bchron functions in R. This constructs an age model from piecemeal linear segments, simulating sedimentation by small increments, random in duration and sedimentation rate. The method handles radioisotopic date uncertainties (as normally distributed values,  $\sigma_R$ ; Table 3) and uses procedures to deal with radioisotopic date outliers, flagging dates with a probability of being an outlier ( $P_{out}$  in Table 3). The measure of stratigraphic uncertainty in placing the radioisotopic date onto the optimised composite scale is also included (i.e.  $e_s$ ; Hounslow 2016; Table 3), as the 'uniform range' ( $d_{max}-d_{min}$  of Parnell *et al.* 2008). The uncertainty on each of the magnetostratigraphic units derived from the initial statistical compositing procedure is scaled to Ka as  $\pm\sigma_T$  in Fig. 18a, using the appropriate magnetostratigraphic unit duration from the age model. These methods therefore incorporate all the major uncertainties into the final GPTS (Fig. 18c, Table 4).  $\sigma_T$  can be thought of as a kind of inter-section magnetostratigraphic unit placement uncertainty, which is impacted by between-section changes in sedimentation rate and sampling density.

Confidence intervals ( $C_{95}$ , Fig. 19d) on the magnetostratigraphic ages are obtained from Monte-Carlo simulations, using the 95% highest posterior density region (HPD) from the age model (Haslett & Parnell 2008; Fig. 18c; Table 4). The HPD represents the shortest interval of time that encompasses 95% of the simulations (Parnell et al., 2011). The HPD confidence intervals ( $C_{95}$ ) may be overly pessimistic in intervals without age control points (Blaauw & Christen 2011), which is to some extent expressed by the generally lower  $2\sigma_T$  compared to  $C_{95}$  (Fig. 19c,d). Those magnetostratigraphic units outside the MI4r and PE7r interval have their base ages estimated as outlined in Table 4, and shown in Fig. 19a. The scaling to age rather compresses the mid Bashkirian magnetostratigraphic units and expands the late Serpukhovian magnetostratigraphic units (Fig. 18c). The differences between the estimated age of stage bases here (Fig. 19a) and those of Aretz et al. (2020) relate to the different scaling methods and where the radioisotopic dates are attached to that scale.

The percent reverse polarity using a four-chron wide window (with magnetostratigraphic durations), clearly shows the polarity dominance switch-over into the base of the Kiaman Superchron at 319–317 Ma. The midpoint of the polarity bias change (using mid way between the 20 and 80% shoulders of %reverse polarity; Fig. 18d) is within PE3r, so the proposed base of the Kiaman Superchron is at the base of PE3r.1r at  $318.6 \pm 0.33$  Ma (Fig. 19a, Table 3). This mid-point approach is similar to that used by Haneda et al. (2020) to place the Matuyama–Brunhes boundary within its transition interval. The end of the Kiaman Superchron is in the mid Permian at  $266.66 \pm 0.76$  Ma (Hounslow & Balabanov, 2018). Although normal polarity intervals have long been suspected in the Kiaman Superchron since the early days of palaeomagnetism (Irving & Pullaiah 1976), many researchers have persisted with the concept of a 100% reverse polarity superchron (e.g. Opdyke 1995; Pavlov & Gallet, 2005), like that originally envisaged by Irving & Parry (1963). Evidence presented here and in Hounslow & Balabanov (2018) indicate that there is strong evidence for normal magnetostratigraphic units in the Kiaman Superchron, which hopefully will spur further studies to fully validate their existence, and utilize them for detailed stratigraphic studies.

## Conclusions

The development of a Carboniferous GPTS has shown hesitant progress over the last 60 years, partly dictated by modifications in the sediment magnetisation paradigms that apply to Carboniferous sediments, concomitant with recognition of remagnetisations. Improvements in instrument sensitivity and data analysis and the key contributions of Neil Opdyke have also played a significant part in this progress. Geomagnetic polarity changes in the Carboniferous are currently one of the least well understood parts of the Palaeozoic. However, there are insufficient magnetostratigraphic and palaeomagnetic data to attempt a full, detailed polarity zonation of the entire Carboniferous (Fig. 19a). However, a preliminary set of 37 magnetostratigraphic units can be identified in the Mississippian and 28 in the Pennsylvanian. This first comprehensive review, takes a bottom-up approach by integrating all biostratigraphic, chemostratigraphic and radioisotopic data with the connected geomagnetic polarity information.

For most of the Tournaisian and into the mid Visean, geomagnetic polarity is only known in outline, largely based on studies from the Tamworth Belt of the New England Orogen in E. Australia and polarity bias data from some other studies. This age interval can best be characterized by polarity bias magnetostratigraphic units (MI1n<sub>B</sub> to MI4n<sub>B</sub>), which have an implicit assumption that the polarity pattern in these is almost certainly incomplete (grayed intervals in Fig. 19a). If the reversal rate in the Tournaisian to mid

Visean is like the younger parts of the Carboniferous, there may be a minimum ca. 40 other magnetozones to discover in the initial 23 Ma of the Mississippian.

The interval from the late Visean to the late Moscovian has a fuller dataset of polarity changes, which allows a reasonable estimate of geomagnetic polarity to be made, and related to a set of 31 radioisotopic dates to generate a GPTS. Two short intervals at the Serpukhovian-Bashkirian boundary and in the late Akavasian substage (Marsdenian substage), have an incompletely known pattern of polarity changes. The late Brigantian may also have an incomplete polarity pattern. The various parts of the procedures leading to the final GPTS, allow uncertainties from the intersection composite to be specified as  $\sigma_T$  (Fig. 19c), as well as confidence interval on the chron ages to be specified as  $C_{95}$  (Fig. 19d).

The base of the Kiaman Superchron at magnetochron PE3r is placed in the mid Bashkirian within the Yeadonian and Prikamian substages, rather than using the traditional Russian position in the late Moscovian. Conventional palaeomagnetic studies supported by the Donets Basin magnetostratigraphy clearly show that the Kiaman Superchron is not exclusively reverse polarity, but has at least eight normal polarity chrons within it. An additional bias magnetochron PE8n<sub>B</sub>, is present in the mid Kasimovian.

Sediment-based magnetostratigraphic data from red-beds has been a key, but not exclusive, source of data in this review. Future improvements need to move to sediment successions which have a better biostratigraphy, or thick lava successions with good geochronological control. Probably the previous strong focus on red-beds has created some of the misconceptions about the 100% reverse polarity state of the Kiaman Superchron, which may in part be due to the undeniably prolonged haematite diagenesis-magnetisation processes in some types of red-beds. The resulting polarity changes throughout the Carboniferous, provides a rich set of chronostratigraphic markers to help solve many (e.g. glacially-related) environmental problems that require high stratigraphic resolution at the global scale.

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## Figure Captions

Fig. 1. Magnetostratigraphic data from the Donets Basin of Khramov *et al.* (1974), showing the magnetic polarity and the sample declination in the four sections. Each of the sections is drawn with the same thickness scale. F<sub>1</sub> to N<sub>1</sub> are the standard limestone codes, and C<sub>1</sub><sup>5</sup> to C<sub>2</sub><sup>1</sup> are the standard formation codes as shown in Davydov *et al.* (2010), along with the regional substage names. The correlated European regional substages and million year ages (Ma) are derived from Fig. 2 of Davydov *et al.* (2010). The limestones are often considered chronostratigraphic horizons, although this may not always be the case (Davydov *et al.* 2010). The composite polarity column is constructed using the formations and limestone units, except for around the L<sub>1</sub> limestone, where L<sub>1</sub> may be diachronous with respect to the two apparent normal magnetozones at this level.

Fig. 2. The Russian 'general magnetostratigraphic scale' in the summaries of Khramov & Rodionov (1980), its major update formalised into the Russian stratigraphic code (Khramov & Shkatova 2000). Stage relationships of these two scales indicated. Rightmost is the modification of the 2000 scale by Molostovsky *et al.* (2007) and its 'time-matching' to the ICS 2008 timescale as used and listed in Table 1 of Perchersky *et al.* (2010), with the apparent magnetostratigraphic relationships to the Khramov & Shkatova (2000) scale indicated.

Fig. 3. Polarity bias data for the Carboniferous, using data from intrusive and extrusive igneous rocks (Table 1). The age boundaries of the ICS 2012 stage boundaries (Richards 2013) are placed on the age axis. The bias of some 0% or 100% data points has been adjusted to provide non-overlapping age confidence intervals (as in Table 1).

Fig. 4. Polarity (N=normal, R=reverse) of palaeomagnetic sampling sites from the Tamworth Belt of the New England Orogen from Geeve *et al.* (2002) and Klootwijk (2016, 2019), placed onto the adapted Fig. 3. of Klootwijk (2016). Those palaeomagnetic sites without polarity label, did not work. The stages are placed according to the correlations in Roberts (1975), Roberts *et al.* (1993) and the Belgium substages according to Poty *et al.* (2014). SHRIMP U-Pb dates are those outlined in Klootwijk (2016), with the adjusted age in [...], according to the factors in Black *et al.* (2003a), to convert the AS3 and SL13 standards to be consistent with the Temora-1 standard. \$=SL13 standard, \*=AS3 standard. Klootwijk (2016) data through the mid Visean Gilmore volcanic are sufficiently close for a mini-magnetostratigraphy. Klootwijk (2016) sites 2-6 in the Foybrook Andesite (Waverly Fm), where moved to near the base of the Waverly Fm in Klootwijk (2019). The normal polarity magnetozones through the Visean have been labelled (TB for Tamworth Belt).

Fig. 5. Polarity of palaeomagnetic sampling sites in the Caroda Formation, from the northern and central Tamworth Belt of the New England Orogen. The Caroda Fm has brachiopod fauna of the *tenuirugosus* Subzone of the *Delepinea aspinosa* Zone (Roberts 1975). CA1, CA2, CA3 and CA5 are volcanic member codes of Roberts *et al.* (2003a). Age of the Caroda Fm is mid to possibly late Visean (Mory 1980, Roberts *et al.* 2003a), based largely on conodonts (Jenkins *et al.* 1993). Barney's Spring Andesite SHRIMP age from supplementary information in Roberts *et al.* (2003a). The Trevallyn section is assigned to the Caroda Fm based on Opdyke *et al.* (2000) and Roberts *et al.* (2003a), not the alternative in Klootwijk (2002). Biozones and ages on Merlewood Fm based on Roberts *et al.* (1993) and Roberts & James (2010).

Fig. 6. Polarity (N=normal, R=reverse) of palaeomagnetic sampling sites from the Tamworth Belt of the northern New England Orogen. Site numbers in black from Opdyke *et al.* (2000) and those in blue from Klootwijk (2002). The original site positions have been simplified onto a smaller number of sections by lateral correlation, using the correlations in the original publications. Data from The Pine Cliffs section (shown on the Arizona-Barneys springs section) is the correlation based on Roberts *et al.* (2003a), not the alternative in Klootwijk (2002). For the SHRIMP dates, red dots are dates in shown section, blue correlated from other section. The position of the C1 and C3 glaciations are those described by Fielding *et al.* (2008) and Birgenheier *et al.* (2009). Ages on the composite polarity column are adjusted SHRIMP dates according to conversion factors in Black *et al.* (2003a), converting SHRIMP AS3 and SL13 standards to be consistent with the Temora-1 standard. Likely disconformities based on Roberts *et al.* (2003a). Ticks on composite polarity column are correlated site locations from all sections.

Fig. 7. Polarity (N=normal, R=reverse) of palaeomagnetic sampling sites from the Werrie Block of the Tamworth Belt (New England Orogen, E. Australia). Site numbers of Opdyke *et al.* (2000) in black, and Klootwijk (2003) in blue on the composite section of Opdyke *et al.* (2000). Volcanic level labels after McPhie (1984) and Opdyke *et al.* (2000). Note, the interval with glacial sediments (C3 glacial of Fielding *et al.* 2008) is placed in the Currabubula Fm below the Canna Creek Tuff (CC) according to McPhie (1984). SHRIMP dates as in caption to Fig. 4. ). Ticks on composite polarity column are correlated site locations.

Fig. 8. Polarity data of Derbyshire lavas from Piper *et al.* (1991) redrawn onto logs generalised from Walters & Ineson (1981). [...] site numbers of Piper *et al.* (1991). Bonsall Moor-Matlock log from Walters and Ineson (1981) Fig. 2 (3 left-hand logs), Lathkill Dale log from their Fig. 8; Shacklow Wood log, from their Fig. 10; Cressbrook Dale-Great Rocks Dale log from their fig. 14 (rightmost 5 logs); Castleton-Bradwell Moor log from their Fig. 16 (left-most 3 logs). The lithostratigraphy is from Waters *et al.* (2009) with other stratigraphic details from Butcher & Ford (1973), Walters & Ineson (1981), Gatliff (1982) and Gutteridge (1989). Correlations in the southern locations based on Bridge & Gozzard (1981), and for the Litton Tuff from Butcher & Ford (1973). Thicknesses of formation units in the Lathkill Dale-Matlock area from Cox & Harrison (1980) and Bridge & Gozzard (1981). Waters *et al.* (2011) place the Conksbury and Lathkill lavas (and limestone between and below) within the Fallgate Volcanic Fm (not followed here for sake of simplification).

Fig. 9. Magnetic polarity data from the Kinghorn Volcanic Formation, from Wilson & Everitt (1963) and Torsvik *et al.* (1989) placed onto the log of Allan (1923). Bed numbers are those of Allan (1923) with [...] indicating equivalent bed numbers used by Wilson & Everitt (1963), if different. Part of the log in the Lower Limestone Fm is based on Karbub (1993), and the stratigraphic distances to units below the Kinghorn Volcanic Fm is based on Francis (1961). Positions of sampling sites of Torsvik *et al.* (1989) are based on transposing their map locations onto Fig. 1 of Brindley & Spinner (1989), giving an approximate position (their site numbers shown), based on the bed outcrop details in Allan (1923). On the log bed numbers the b and s postfix are the percent ilmenite values of Wilson (1966) divided into a big (>9% ilmenite) and small (<9% ilmenite) classes. The tabular data is the number of flow or intrusive units with petrological classifications of Allan (1923) in the two % ilmenite classes (from Wilson 1966), and the number of units with reverse (R) or normal (N) polarity.

Fig. 10. Magnetic polarity data for the Mauch Chunk Fm, E. Pennsylvania (DiVenere & Opdyke 1991, Opdyke & DiVenere 2004). The sedimentary log for the Jim Thorpe section is from Epstein *et al.* (1974); their section 19, place alongside the thickness scale of Opdyke & DiVenere (2004).

Lithostratigraphy is that of Filmore *et al.* (2012), using the lower, middle, upper divisions of the Mauch Chunk Fm of Wood *et al.* (1969) and Jennings (1985). The position of *Adiantitites antiquus* in the Schuylkill Gap section is based on correlation of the logs of this section (shown) in Jennings (1985) against that in DiVenere & Opdyke (1991). Mauch Chunk member thickness and missing strata estimates based on Wood *et al.* (1969) and Epstein *et al.* (1974). Calcrete levels in the Schuylkill Gap section from Levine & Slingerland (1987), which can be related to the re-drafted log of DiVenere & Opdyke (1991). There is no calcrete data from the lower part of the Schuylkill section. The Hometown Mbr is largely calcrete-free (Levine & Slingerland 1987). Two possible options of how to relate the south Lavelle and Jim Thorp sections are shown.

Fig. 11. Lithostratigraphy and age control for the Mauch Chunk Fm and adjacent units in a SW to ENE transect across Pennsylvania. Lithostratigraphic relationships based on Berg *et al.* (1983), Brezinski (1989a, 1999), Edmunds (1996), Edmunds *et al.* (1999). Magafloral zones are those of Read and Mamay (1964) slightly modified as outlined by Eble *et al.* (2009), with relationship to Pennsylvanian strata as in Edmunds *et al.* (1999). LM=Loyalhanna Mbr and equivalents, DVL=Deer Valley Limestone, RL=Reynolds Limestone, WGL=Wymps Gap Limestone, PS=Patton (red) Shale, CLS=Campbells Ledge Shale. The greyed 3a megafloral zone is not known from Pennsylvania, but from the equivalent interval in Virginia (Eble *et al.* 2009).

Fig. 12. Magnetic polarity data from the Cumberland Basin and Cape Breton Island sections, New Brunswick and Nova Scotia, derived from DiVenere & Opdyke (1990, 1991b) and Opdyke *et al.* (2014). Lithostratigraphy and hiatus in the Cumberland Basin sections from Jutras *et al.* (2015, 2016). Magnetostratigraphic zones re-labelled from Opdyke *et al.* (2014) to NN after New Brunswick- Nova Scotia. It is not clear how the hiatus on the basal Enragé Fm in the Maringouin Peninsula of Jutras *et al.* (2015), relates to the Maringouin Peninsula column data. Sloping boundary on polarity composite= uncertainty in placing the biozone boundary. N.D=no palynology data, a-t Z.= *S. acadensis*- *K. triradiatus* Zone.

Fig. 13. Magnetostratigraphic data for wells from the Tengiz carbonate platform, Kazakhstan (adapted from Ratcliffe *et al.* 2013). Codes for sequence stratigraphic boundaries (in grey box) and the chemostratigraphic divisions, based on trace element geochemistry, are shown with greyed numbers. See Kenter *et al.* (2006) for sequence stratigraphic codes.

Fig. 14. A) Carbon isotope datasets for the two wells (T-220, T-5853) from the Tengiz carbonate platform (from Ratcliffe *et al.* 2013) compared to similar aged units from the Antler Basin, Idaho, USA (Batt *et al.* 2007), the Askyn River section, S. Urals (Kuleshov *et al.* 2018) and the Pennsylvanian/Bashkirian in Arrow Canyon, Nevada, USA (Saltzman, 2003). Codes for sequence stratigraphic boundaries in grey box, and their relationship to the Russian horizons based on Weber *et al.* (2008). Isotope correlations (shown in grey) and other correlations in black. The Arrow Canyon and Askyn River sections shows the abbreviated conodont zones: nod= *Declinognathodus noduliferous*; sin= *Idiognathoides sinuatus*; b-s= *Neognathodus symmetricus*- *N. bassieri* Zone; k-s= *Idiognathodus klapperi* - *I. sinuosus* Zone; con= *Id. convexus*; Prof= *Profusulinella*; Fus= *Fusulinella*. Corr= *Id. corrugates*; ask= *N. askynensis*; marg= *D. marginodosus*; don= *D. donetzianus*. Biostratigraphic correlation (dotted lines) between the Arrow Canyon section and the Bashkirian in the Askyn River based on Nemirovskaya & Alekseev (1994), Groves *et al.* (1999), Kulagina *et al.* (2009) and Hu *et al.* (2017). B) The table shows the stratigraphic relationships of the Russian horizons and the British substages (based on Richards 2013 and Kuleshov *et al.* 2018).

Fig. 15. Summary magnetostratigraphic and polarity bias data for the early - mid Mississippian. Igneous and sedimentary rock polarity bias data from Tables 1 and 2. Magnetostratigraphic zones and polarity bias zones from Lucas, Schneider *et al.* (eds) *The Carboniferous Timescale*, Spec. Publ. Geol. Soc. Lond. 46

in the composite polarity labelled MI (for Mississippian). The detailed late Visean and Serpukhonian magnetostratigraphy is from Fig. 16, expressing the incompleteness of the Tamworth Belt polarity data. Daposhang magnetostratigraphy, conodont zones and radioisotopic date from Liu *et al.* (1991, 2012). At ages older than MI4r, magnetozones are intervals of reverse or normal polarity bias. The Lower *crenulata* conodont zone for the Brush Hill Limestone (in *sol* brachiopod Zone) from Mory & Crane (1982).

- Fig. 16. Summary magnetostratigraphic data for the late Mississippian and earliest Pennsylvanian. Magnetozones in the composite polarity, labelled MI (for Mississippian) and PE (for Pennsylvanian), beginning the PE magnetozones in the first zone which overlaps the base of the Bashkirian. Two options of how to correlate the Mauch Chunk Fm magnetostratigraphy are shown, with Option-2 the preferred one, which is in part used to scale the composite polarity column. A/F= ammonoid or foraminifera zone (estimated position). Radioisotopic dates on composite polarity from Table 3.
- Fig.17. a) Summary magnetostratigraphic data for the Bashkirian and Moscovian, based around the composite from Fig. 1. The earliest part of the Bashkirian is shown in Fig. 16. B) Polarity bias data for the Moscovian to Gzhelian from Tables 1 and 2. Radioisotopic dates on composite polarity column from Table 3.
- Fig. 18. Optimised composite (a, b) and age model for the Carboniferous. c) Optimised composites based on methodology in Hounslow (2016). A) The standard deviation ( $\sigma_T$ ) for the levels used in the optimised scaling procedure (scaled to Ma, using the final age model). This is a measure of the correlated magnetozone misfit (i.e. the small divergences between the original relative position of sections and the final optimized composite position in b)). No  $\sigma_T$  values for a corresponding level shown in b) indicate the level was not used to constrain the optimised model, but simply scaled with the section. B) The original section data shown on the y-axis (in a relative height scale), along with the final composite position of the levels on the x-axis. Minor scatter in the y-axis relates to the degree of between-section misfit expressed in the overlying panel as  $\sigma_T$ . Numbers in brackets (..) next to section names are the  $D_j$  values of Hounslow (2016), which express the general mis-fit of the section data to the optimised composite. i.e. the Schuylkill Gap data has a mean residual of 14% per average 'chron width' in the optimised model.  $D_s$  is the average of the  $D_j$  values across all sections. C) The Bchron age model for the Carboniferous, showing the scaling of optimised position to Ma, using the radioisotopic dates. Magnetochrons in scale of optimised composite on x-axis and in Bchron-scaled Ma on y-axis. Radioisotopic dates used to scale the optimised composite have analytical uncertainties on the y-axis ( $\sigma_R$ ) and stratigraphic uncertainties ( $e_s$ ) on the x-axis (values from Table 3). D) Percent reverse polarity using a 4-chron width window. The mid point between the 20% and 80% crossing points is used to fix the base of the Kiaman Superchron.
- Fig. 19. a) Summary geomagnetic polarity timescale (GPTS) data for the Carboniferous, with the CONOP-based ages for the stage bases (blue arrows) from Aretz *et al.* (2020). The stage names shifted to the right express stage boundary positions entirely based on Aretz *et al.* (2020), whereas those on the left, and regional substages are based on this magnetozone-stage position from this work (Gzhelian to base Asselian from Hounslow & Balabanov, 2018). b) type of data which has contributed to the timescale, as an expression of 'data-confidence'. c) Uncertainty from the statistical compositing procedure expressed as the age scaled value of two times the  $\sigma_T$  value, which is a kind of 'between-section magnetozone position' uncertainty. d) The confidence interval for the age of magnetochron bases, which is the shortest interval of time that encompasses 95% of the simulated ages at the position of the magnetochron. Abbreviated substages in oldest to youngest order: Pendleian, Arnsbergian, Syuranian, Akavasian, Prikamian, Cheremshankian, From Lucas, Schneider *et al.* (eds) *The Carboniferous Timescale*, Spec. Publ. Geol. Soc. Lond.

Melekessian, Vereian (base Moscovian), Kashirian, Podolskian, Myachokovian, Rusavkinian (base Gzhelian), Pavlovoposadian, Noginian, Melekhovian, base Asselian and Permian.

Unit/Formation/Component	Stratigraphic age	Date (Ma)	Date type	Mid-age (Ma)	CI (Ma)	%R	Palaeomag. ref	Age reference
Gazar Fm, various volcanics	-	283(3)	U-Pb.IZ	283	3	100	Zhao <i>et al.</i> 2020	
Shuangbaotang Fm, Dunhuang Block volcanics & sediments	-	280.6-291.4	U-Pb.IZ	286	5.4	98	Xu <i>et al.</i> 2019	
Panjal Traps	-	289(3)	U-Pb.IZ	289	3	100	Stojanovic <i>et al.</i> 2016	
Featherbed volcanic field, Com p C2, Drummond basin	-	300-280	Rb/Sr.w	290	10	100	Klootwijk 1993; Giddings 1993	Korsch <i>et al.</i> 2009
Tichka Plutonic Complex, diorites	-	291(5)	Rb/Sr	291	5	100	Martin <i>et al.</i> 1978	Gasquet <i>et al.</i> 1992
Holy island dyke	-	294(2)	Ar/Ar	294	2	100	Liss <i>et al.</i> 2004	
Cracow Volcanics, comp A	-	294.2(2.1)	U-Pb.SZ	294.2	2.1	100	Nawrocki <i>et al.</i> 2008	
Whin sill	-	295(6)	K-Ar	295	6	100	Liss <i>et al.</i> 2004	
Middleton dyke	-	297(0.5)	U-Pb.b	297	0.5	100	Liss <i>et al.</i> 2004	
Shan-Thai-Malay block, Woniusi Fm	Gzhelian- Asselian			299*	5	100	Huang & Opdyke 1991	Xiaochi 2002
Isola Rossa Diorite, C comp	-	300.1(6.1)	U-Pb.Z	300.1	6.1	100	Edel <i>et al.</i> 2014	
Molaoba Formation, Tacheng	U. Carboniferous (Ar-Ar near top)	304 (4.7)	Ar/Ar.G	305	4.7	94	Yi <i>et al.</i> 2015	
Strzegom Granite	-	302.9-308.4	U-Pb.SZ	305.65	3	100	Halvorsen <i>et al.</i> 1989	Mikulski <i>et al.</i> 2013
Lago Ranco Granites, Principal Cordillera.	-	305.9(2.4)	U-Pb.SZ	305.9	2.4	86	Beck <i>et al.</i> 1991	Deckart <i>et al.</i> 2014
Queensferry sill, Midland Valley Sill	Stephanian	308(5)	U-Pb.TZ	308	5	100	Torsvik <i>et al.</i> 1989	Monaghan <i>et al.</i> 2014
Osani Andesite, C comp.	-	308 (2.9)	U-Pb.Z	308	2.4	100	Edel <i>et al.</i> 2014	
Yetholme/Durandel adamellite	-	310 (6.8)	K/Ar, Rb/Sr	310	6.8	100	Facer 1976	Facer 1978
Shatterford sill	Late Bolsovian			312	2.5	59	Everitt 1960	Kirton 1984
Karkonosze Granite	-	312.4 (0.3)	Rb-Sr,Sz,TZ	312.4	0.5	50	Halvorsen <i>et al.</i> 1989	Kryza <i>et al.</i> 2014
Dutchmans Creek Gabbro	-	314 (2)	U-Pb.IZ	314	2	100	Dooley 1983	Mobley <i>et al.</i> 2014
Westphalian volcanics,	I. Westph.B to Westph. C or D	-		312.5*	2.5	100	Westphal <i>et al.</i> 1987	
Clouds Creek Diorite	-	313(4)	Rb/Sr.bW	313	4	100	Dooley 1983	Fullagar <i>et al.</i> 1997)
Barrabisa granitoids, U2,C comp.	-	320-311	U-Pb.IZ	315.5	4.5	60	Edel <i>et al.</i> 2014	
Boomi Creek, N. Tamworth terrane	none	318(2.2)	U-Pb.IZ	318	2.2	100	Pisarevsky <i>et al.</i> 2016	
Flamanville Granite	-	318.1(1.5)	U-Pb.IZ	318.1	1.5	100	Van der Voo & Klootwijk 1972; Cogné 1988	Erwan <i>et al.</i> 2018
Kullatine Formation, Hasting Terrane	e. Namurian	-		320.5	2.5	48	Schmidt <i>et al.</i> 1994)	Birgenheier <i>et al.</i> 2009
Spessart basement & intrusives(A'-A comp)	-	323(5)	Ar/Ar, U-Pb	323	5	86	Edel & Wickart 1991	Will <i>et al.</i> 2017
Bergstrasser Odenwald intrusives,A-A' components <sup>1</sup> -		330-320	Ar/Ar and others	325	5	71	Edel & Wickert 1991; Zwing & Bachtadse 2000	Stein 2001

Argyllshire, camptonite & monchiquite dykes	-		K/Ar.w	326	5	100	Esang & Piper 1984	Baxter & Mitchell 1984
Central Black Forest rhyolites, FN5,6,7	-	336-320	U-Pb.TZ	328	5	100	Edel 1987; Edel & Schneider 1995	Schaltegger 2000
Hardwood Ridge Volcanic Mbr ("Boiestown, Royal Road")	I. Asbian- e. Brigantian	328	U-PB.TZ	328*	1	85	Seguin <i>et al.</i> 1985	Jutras <i>et al.</i> 2018; Park & Hinds 2019)
Kudowa Granitoid	-	329(4.5)	U-Pb.SZ (SL13)	332.29 <sup>\$</sup>	4.5	83	Halvorsen <i>et al.</i> 1989	Mikulski <i>et al.</i> 2013
Champ du Feu Massif diorites <sup>2</sup> , Comp.C1,c2,c3	-	331(5)	K.Ar, Ar/Ar.m; U-Pb	331	5	61	Edel <i>et al.</i> 1986	Tabaud <i>et al.</i> 2014
Kinghorn-Burntisland	I. Asbian- m. Brigantian			332.5*	1.5	71	Torsvik <i>et al.</i> 1989	Monaghan <i>et al.</i> 2014
Derbyshire lavas and intrusions	I. Asbian – e. Brigantian			332.5*	1.5	66	Piper <i>et al.</i> 1991	Harwood 2005
Upper Visean tuffs etc, Roanne	upper Visean			334.5*	2.5	67	Edel <i>et al.</i> 1981	
Kap Kolthoff & Kap Graah series	-	335.6 (3.1)	Ar/Ar.P	335	3	57	Hartz <i>et al.</i> 1997	
Star of Hope Fm, Comp C3	50-70% from base Visean			336*	2	100	Klootwijk <i>et al.</i> 1993	Sobczak <i>et al.</i> 2019
Magerøy dykes	-	337(0.4)	Ar/Ar.P	337	0.4	0	Roberts <i>et al.</i> 2003	
C and C' comp high Mg-K Granodiorites	-	338 (2)	U-Pb.TZ	338	2	50	Edel <i>et al.</i> 2014	Paquett <i>et al.</i> 2003
Punta del Agua Formation, basalts andesites	m. Visean	336(1.3)	U-Pb.TZ	336	1.3	100	Geuna & Escosteguy 2004	Cesari <i>et al.</i> 2011
Garleton Hills Volcanic Fm, B Comp	Chadian- e. Holkerian	343(1)	U-Pb	343	1	87	Rother & Storetvedt 1991	Monaghan <i>et al.</i> 2014
Silver Hills Volcanics, Comp C3	Tournasian	344(3)	U-Pb,K/Ar	344	3	0	Klootwijk <i>et al.</i> 1993	Henderson <i>et al.</i> 1998
Merlewood Fm, Werrie syncline	Molinacian- e. Livinian			344*	2.5	0	Schmidt 1988	Roberts & James 2010
Tin Serririne basin intrusive, Comp. B	-	347.6(8.1)	K/Ar.w	347.6	8	0	Derder <i>et al.</i> 2006	
Cockermouth Lavas, units 3,7	CM Miospore zone, Mid-late Ivorian			348	1.7	0	Oppenheim <i>et al.</i> 1994	Butterworth & Butcher 1983; Clayton & Turnau 1990
Admiralty Intrusives	-	364-346	K/Ar, Rb/Sr	355	9	0	Fiorett& Lanza 2000	
Admiralty extrusives, Gallipoli Volcanics	-	355.8(2.9)	U-Pb.SZ	355.8	2.9	0	Rolf & Henjes-Kunst 2003	
Frankenstein intrusive complex, Comp C	-	363(4)	Ar/Ar, Pb-Pb	363	4	75	Zwing & Bachtadse 2000	
Achala Batholith	-	368(2)	U-Pb.TZ	368	2	43	Geuna <i>et al.</i> 2008	

Table 1. Selected polarity bias data for intrusive and extrusive igneous rocks covering the Carboniferous and earliest Permian. <sup>1</sup>Excluding sites od14,13,12,25,10. <sup>2</sup> sites 8,13,19,20,38,3,4,10,11,12,18,36. Stratigraphic age abbreviations= l.=Late, m.=Mid, e.=Early, U.=Upper. Date type abbreviations: TZ= ID TIMS on zircon, SZ=SHRIMP on zircon, IZ=laser ICPMS on zircon, w=whole rock, P=on plagioclase, m=on mica, b=on Baddeleyite, G=on groundmass, bw= biotite whole rock, <sup>\$</sup>= SHRIMP age adjusted for standard using values in Black *et al.* (2003a). CI=confidence interval, which is quoted uncertainty on radioisotopic date, or half likely age range for a stratigraphic age. %R= percentage of sites or samples with reverse polarity (samples used, if quoted). Palaeomag.ref= original palaeomagnetic source, Age reference= newer reference for age determination (if none shown, using original source). \* stratigraphic age shown (from Richards 2013 timescale).

Location/ Age	Lithology, Lithostratigraphy	N <sub>n</sub> /R <sub>n</sub>	D <sub>m</sub> /FT/S	Ma [CI]	References (source and updated age)
Donbass, Kartamysh Fm / l. Gzhelian- e. Asselian	Red and grey clastics	0/31+	1/F+/PP	298.75 [1.7]	Iosifidi <i>et al.</i> 2010; Davydov <i>et al.</i> 2012
Donbass, Suhoj-Jaz, mid Kartamysh Fm / l. Gzhelian- e. Asselian	Red beds	13/32*	0/F+/PP	299.05 [0.55]	Khramov 1963, Khramov & Davydov 1984, Davydov & Leven 2003
Wescogame Fm, PS1,PS2, Arizona/ Virgillian	Red beds	0/126 <sup>+</sup>	1/0/MS	300.45 [1.6]	Steiner 1988; Blakey & Middleton 2012
Ukraine, Avilov & Araucarite Fms/ m. Gzhelian	Red and grey clastics	0/12 <sup>+</sup>	1/F+/PP	300.95 [2.25]	Iosifidi <i>et al.</i> 2010
Moscow, Shchelkovo Series, Gzhel Quarry/ 10% of e. Rusavkinian;	Red and brown clays	0/14 <sup>+</sup>	1/ 0/PP	302.65 [0.15]	Iosifidi & Mikhailova 2017; Briand <i>et al.</i> 1998
Noginsk,Russia, 25% of Dorogomilovskii substage(Rusavkinian)/ e. Gzhelian	Red clays	0/5 <sup>+</sup>	1/ 0/PP	302.75 [0.25]	Iosifidi <i>et al.</i> 2018
Wyoming, Lower Casper Formation/ l. Desmoinesian- Missourian-Virgilian	Red sands/siltstone, Limestones	0/549	1/0/MS	302.85 [4.15]	Diehl & Shive 1981; Burns & Nestell 2009
New Brunswick, Hurley Creek Formation/ L. Bolsovian	Red beds	0/20 <sup>x</sup>	1/0/PP	303.5 [0.5]	Roy <i>et al.</i> 1968; Gibling <i>et al.</i> 2019
N. Italy, Auernig Group/Stephanian	Limestones	0/119*	2/0/PP	304.75 [0.25]	Manzoni <i>et al.</i> 1989
Moskva River, Moscow, 25% of Khamovnicheskii/ m. Kasimovian	Clastics?	0/9 <sup>+</sup>	1/0/PP	304.75 [0.25]	Iosifidi <i>et al.</i> 2018; Briand <i>et al.</i> 1998
NW Bulgarian, Zverino red quartzites/ Stephanian	Red beds	0/71*	2/0/PP	304.75 [2.75]	Nozharov <i>et al.</i> 1980
Nova Scotia, Miminegash & Egmont Bay Fms/ Stephanian	Red beds	2/96*	1/F+/PP	304.75 [2.75]	Pan & Symons 1993; Ziegler <i>et al.</i> 2002; Tanner <i>et al.</i> 2005; Allen <i>et al.</i> 2011
Pennsylvanian, Glenshaw Fm/ Missourian	Grey clastics	0/23*	1/0/PP	304.85 [1.15]	Kodama 2009
SW Ningxia, Taiyuan Fm, B Comp./ Stephanian	Limestones	0/16 <sup>+</sup>	1/F+/PP	305.5 [1.5]	Huang <i>et al.</i> 2001
Nova Scotia, Morien (pictou) Group/ Bolsovian	Clastics, concretions	0/21*	1/F+/PP	312.25 [2.75]	Scotese <i>et al.</i> 1984; Gibling <i>et al.</i> 2019
Colorado, Minturn Formation, Comp.Mh/ Atokan to Desmoinesian	Red beds	0/111*	1/0/MS	311 [4]	Magnus & Opdyke 1991; Itano <i>et al.</i> , 2003.
Algeria, Illizi Basin, Edjeleh Fold, Comp.C/Moscovian	Red beds	0/32*	1/F+/PP	311.1 [4.1]	Derder <i>et al.</i> 2001
Algeria, Lower El Adeb Larache Fm/Moscovian	Marl, Limestone	0/61*	1/0/PP	311.1 [4.1]	Henry <i>et al.</i> 1992
Russia, Tver Oblast, River Volga/ 25% of e. Kashirian	Red clays	0/20 <sup>+</sup>	1/0/PP	312.65 [0.15]	Iosifidi <i>et al.</i> 2018; Iosifidi & Mikhailova 2017
Russia, Pushchino, Nara River/ e. Vereian	Brown-red clays	0/14 <sup>+</sup>	1/0/PP	313.35 [0.35]	Iosifidi <i>et al.</i> 2018; Iosifidi & Mikhailova, 2017
S. Korea, Manhang Fm, High-T/ e. Moscovian	Red beds	12/81*	1/F+/PP	313.5 [1.5]	Doh & Piper 1994; Lee 2010
S. Korea, Hongjom Fm/ e. Moscovian	Red beds	3/16*	1/0/PP	313.5 [1.5]	Kim <i>et al.</i> 1992; Lee 2010
Moscow, Rzhev layers/ 5% of Kashirian	Red clays	0/20 <sup>+</sup>	1/0/PP	312.1 [0.1]	Iosifidi & Mikhailova 2017
Algeria, Upper "Dembaba" fm/ e. Moscovian	Limestone	0/113*	1/F+/PP	313.5 [1.5]	Amenna <i>et al.</i> 2014

Pushchino, Nara river/ 5% of Vereian	Red clays	0/14 <sup>+</sup>	1/0/PP	313.4 [0.1]	Iosifidi & Mikhailova 2017
S. Korea, Manhang, Yobong Fms/ e. Moscovian	Red beds	1/22 <sup>*</sup>	1/F+/PP	313.75 [1.75]	Lee <i>et al.</i> 1996; Lee 2010
Algeria, "Hassi Bachir" Fm, Comp. C/ l. Bashkirian	Red beds	0/69 <sup>*</sup>	1/F+/PP	314.5 [2.5]	Derder <i>et al.</i> 2009; Wendt <i>et al.</i> 2009
Kyrgyzstan, West Kyrgyz Range/ Bashkirian	Red beds	48/1 <sup>*</sup>	1/F+/PP	319.2 [4]	Bazhenov <i>et al.</i> 2003
<b>Serpukhovian- early Bashkirian not shown</b>					
Kyrgyzstan, Dungurma Fm, North Tien Shan/ l. Visean- Serpukhovian	Red beds	0/86 <sup>*</sup>	1/F+/PP	329 [6]	Bazhenov <i>et al.</i> 2003
Borovichi, Egol'skaya group, Novgorod region/ E. Venevian	Limestones, Red clays	0/23 <sup>+</sup>	1/0/PP	330.5 [0.5]	Iosifidi <i>et al.</i> 2018
Newfoundland, Jeffrey's Village Mbr, Comp A/ Brigantian	Red beds	29/3 <sup>+</sup>	1/F+/PP	331 [2]	Murthy 1985; Utting & Giles 2008
Nova Scoria, Green oaks Fm, Windsor Grp/ Brigantian	Limestone, red beds	0/41 <sup>*</sup>	1/F+/PP	331.25 [2.25]	Scotese <i>et al.</i> 1984; Giles 2009
SW Ningxia, Zhongning, Chouniugou Fm, B Comp/ l. Visean	Limestones	1/32 <sup>+</sup>	1/0/PP	333.95 [3.05]	Huang <i>et al.</i> 2001
Novgorod, Lubytno, /Aleksinian	Clastics, limestones	6/10 <sup>+</sup>	1/0/PP	334.5 [0.5]	Iosifidi <i>et al.</i> 2016
Boksitogorsk, mine no. 13/ 10% of Tulian	Brown and red clays	0/3 <sup>+</sup>	1/0/PP	338.5 [2.5]	Iosifidi & Mikhailova 2017
Donbass, Volnovakha, Mokrovolnovakhsкая Series/ Styl'sky Stage (~e. Tulian)	Limestones	0/11 <sup>+</sup>	1/0/PP	339.5 [1.5]	Iosifidi <i>et al.</i> 2016; Davydov <i>et al.</i> 2012
Newfoundland, Spout Fall Fm, Comp A/ Arundian	Red beds	37/0 <sup>+</sup>	1/F+/PP	343.5 [0.5]	Murthy 1985; Gibling <i>et al.</i> 2019
Newfoundland, Terrenceville Fm, Comp A/ Tournaisian	Red beds	12/28	1/F0/PP	353 [6]	Kent 1982
Nova Scotia, Horton Grp & Cheverie Fm/ l. Famennian- l. Tournaisian	Red beds	0/12 <sup>*</sup>	1/0/PP	354 [7]	Scotese <i>et al.</i> 1984; Jutras <i>et al.</i> 2006

Table 2. Studies showing polarity bias data in the early Mississippian and Pennsylvanian. N<sub>n</sub>, R<sub>n</sub>=Number of specimens or sites with normal (N<sub>n</sub>) or Reverse polarity (R<sub>n</sub>). + Stratigraphic levels, \* specimens (in most sources it is not clear how many sites or samples/specimens represent different stratigraphic levels). D<sub>m</sub>/FT/S= demagnetisation method/fold test/study type. D<sub>m</sub>=1, if full demagnetisation applied to all samples, with principle component or great circle extraction, D<sub>m</sub>=2, pilot demagnetisations of simple magnetisation behaviour, with stable point averaging, or single step. D<sub>m</sub>=0, none or temporal demagnetisation. F+= fold test positive (or demonstrate pre-folding magnetisation), F-= fold test negative, 0= no fold test. S=PP or MS for palaeopole- type sampling or magnetostratigraphic style sampling respectively. Ma= Million years for mid of study interval with confidence [CI] interval or like range of study (using timescale in Richards (2013) for most, but using timescale in Davydov *et al.* (2010), for Russian stage or Fm units as shown in their Fig. 2). l.=late, e.= early.

1.Code, age (Myr ago)	2. $\pm 2\sigma_R$ (Myr)	3. $\pm e_s$	4. Location	5.Biostratigraphy, stratigraphy {magnetostratigraphic relative position}	6. $P_{out}$	7. Ref.
n1 coal 307.26	0.37	10% of PE7r	Butovskaya shaft, Donets Basin	Base Peskovian, Base C <sub>3</sub> a and <i>Pr. ovoides</i> - <i>Praeobs. Burkemensis</i> Zone {top of DB5r in Fig. 1}	0.00	Davydov <i>et al.</i> 2010
01DES-362 307.66	0.37	10% PE7r	Dal'ny Tulkas Quarry, S. Urals	Upper Myachkovian, 70% thro Ng. roundyi CZ, {76% into PE7r }	0.00	Schmitz & Davydov 2012
06US0-2.0 308.0	0.37	15% PE7r	Usolka section, S. Urals	mid Myachkovian { 59% into PE7r}	0.01	Schmitz & Davydov 2012
01DES-481 308.36	0.38	15% PE7r	Usolka section, S. Urals	mid Myachkovian { 58% into PE7r }	0.01	Schmitz & Davydov 2012
01DES-351 308.5	0.36	10% PE7r	Dal'ny Tulkas Quarry, S. Urals	Lower Myachkovian { 20% into PE7r}, base of Ng. roundyi CZ	0.01	Schmitz & Davydov 2012
m3 coal 310.55	0.37	10% of PE6r	Zasyadko Shaft, Donets Basin	Mid C <sub>2</sub> <sup>m</sup> c, 20% into Podolskian {79% into PE6r}	0.01	Davydov <i>et al.</i> 2010
13(a) coal 312.01	0.37	75% of PE6n.2n	Krasnolimanskaya Shaft, Donets Basin	Mid C <sub>2</sub> <sup>m</sup> b, 68% into Kashirian {14% into PE6r}	0.01	Davydov <i>et al.</i> 2010
13(b) coal 312.18	0.37	75% of PE6n.2n	Zdanovskaya Shaft, Donets Basin	Mid C <sub>2</sub> <sup>m</sup> b, 65% into Kashirian {12% into PE6r}	0.01	Davydov <i>et al.</i> 2010
11 coal 312.23	0.37	20% of PE6n.1r	Kirov Shaft, Donets Basin	Base C <sub>2</sub> <sup>m</sup> b, little above L1 in Fig. 1, 10% into Mar'evsky {50% into PE6n.1r}	0.01	Davydov <i>et al.</i> 2010
k7 coal 313.16	0.37	20% PE5r	Pereval'skaya Shaft, Donets Basin	Upper C <sub>2</sub> <sup>m</sup> a, 75% into Kamensky {96% into PE5r}	0.01	Davydov <i>et al.</i> 2010
k3 coal 314.40	0.37	20% PE5r	Pereval'skaya Shaft, Donets Basin	Base C <sub>2</sub> <sup>m</sup> a, base Kamensky {56% into PE5r}	0.13	Davydov <i>et al.</i> 2010
Z1 Tonstein 313.78	0.38	15% of PE5r	Furst Leopold Coal Mine, Ruhr Basin, Germany	10–12 m beneath the Ägir marine band. So in topmost Duckmantian {30% in PE5r}	0.12	Pointon <i>et al.</i> 2012
Bed 32 317.54	0.38	20% PE4n	Kijuch section, S. Urals	Cheremshankian, 75% into Asatauian- Tashastian interval {10% into PE4r}	0.01	Schmitz & Davydov 2012
T75 317.63	0.39	15% PE4n	Zwartberg Coal Mine, Belgium,	Tonstein 75, mid RA miospore zone, 73% into Langsettian. {74% into PE4n}	0.01	Pointon <i>et al.</i> 2012, Paproth <i>et al.</i> 1983
Bed 9 318.63	0.40	10% PE3n	Kijuch section, S. Urals	Prikamian, 60% into Askynbashian, {70% into PE3n.2n};	0.01	Schmitz & Davydov 2012

Bed 2 319.09	0.38	25% PE3n	Kijuch section, S. Urals	Prikamian, 10% into Askynbashian, {45% into PE3n}	0.01	Schmitz & Davydov 2012
B9 324.54	0.46	25% of MI9n	Oakenclough Brook, Pennines Basin, England	Namurian E2b2 ammonoid subzone, 50% into Arnsbergian {50% into MI9n}	0.01	Pointon <i>et al.</i> 2012
GP-1 325.64	0.40	100% MI8r.1n	Julius Fučík Mine, Upper Silesia Basin, Czech Rep	35% into E2a Zone, Jaklovec Mbr, coal 365, Gabriela, 14% into Arnsbergian {80% into MI8r}	0.01	Jirasek <i>et al.</i> 2018
GP-2 325.58	0.45	100% MI8r.1n	Julius Fučík Mine, Upper Silesia Basin, Czech Rep	30% into E2a Zone, Jaklovec Mbr, coal 335, Eleonora, 14% into Arnsbergian {82% into MI8r}	0.01	Jirasek <i>et al.</i> 2018
GP-4 327.00	0.49	100% MI8r.1n	Zárubek Mine, Upper Silesia Basin, Czech Rep	90% into E1c Zone, Upper Hrosov Mbr, coal 252, Flora, 97% into Pendleian {base of MI8r.2r}	0.00	Jirasek <i>et al.</i> 2018
Karel Coal 327.58	0.39	15% of MI8r.1r	Staric 2 core, Upper Silesia Basin, Czech Rep.	15% into E1c zone, 80% into Pendleian, Coal 106, close to base Hrusov Mbr of Ostrava Fm {16% into MI8r}	0.01	Gastaldo <i>et al.</i> 2009, Jirasek <i>et al.</i> 2018
MOW 327.35	0.39	15% of MI8r.1r	Main Ostrava Whetstone tuffite, Upper Silesia Basin, Czech Rep.	12% into E1c Zone, 78% into Pendleian, at base of Hrusov Mbr of Ostrava Fm {13% into MI8r}	0.01	Gastaldo <i>et al.</i> 2009, Jirasek <i>et al.</i> 2018
C11 coal 328.14	0.40	15% of MI8n	Yuzhno-Donbasskaya, Ugledar, Donets Basin, Ukraine	Middle C <sub>1</sub> <sup>v</sup> <sub>g2</sub> zone- Foraminifera, <i>Betpakodiscus</i> <i>cornuspiroides</i> , lower Steshevian {base MI8n.1r}	0.01	Davydov <i>et al.</i> 2010, Somerville 2008
Ludmilla coal 328.48	0.41	15% of MI8n	Staric 2 core, Upper Silesia Basin, Czech Rep.	50% into E1B Zone, 50% into Pendleian, Coal 043, Petrkovice Mbr of Ostrava Fm {base MI8n}	0.01	Gastaldo <i>et al.</i> 2009, Jirasek <i>et al.</i> 2018
W13 332.5	0.40	50% of MI5r	Watrissse Quarry, Anhée Sud, Belgium	Lower part of Anhee Fm, in foram. zone MFZ14 ( <i>H.</i> <i>bradyana</i> interval zone), late Asbian. {base MI5r}	0.01	Pointon <i>et al.</i> 2014; Somerville 2008
02VD-0 333.87	0.39	15% of MI5n	Base Bed 21-8, Verkhnyaya Kardailovka, S. Urals	Guisiken Fm, base Mikhailovian, 40% into <i>L.</i> <i>mononodosa</i> Zone, near the base of the <i>E.</i> <i>assymetrica</i> foram. Zone (foram. zone MFZ14, mid Asbian {base MI5n.1r}	0.01	Schmitz & Davydov 2012; Sevastopulo & Barham 2014
W18 335.59	0.44	100% of MI4r.2r	Watrissse Quarry, Anhée Sud, Belgium	15m above base Bonne River Fm, early Asbian, 20% into MFZ13 {base of MI4r.1n}	0.00	Pointon <i>et al.</i> 2014;
SHRIMP dates used shown below						
320.7	3.0	20% PE1r	Rocky Creek Syncline, E. Australia	Upper Clifden Fm, Peri and Easton Arms Rhyolites combined (CL2) {60% into PE1r.1r}	0.01	Fig. 6; Opdyke <i>et al.</i> 2000
322.4	2.8	20% PE1n	Rocky Creek Syncline, Pinnacles section, E. Australia	mid Clifden Fm, Wanganui Andesite (CL10) {50% into PE1n}	0.00	Fig. 6; Opdyke <i>et al.</i> 2000
325.5	3.2	20% of MI9r	Rocky Creek Syncline, Pinnacles section, E. Australia	mid Clifden Fm, unnamed dacite {base PE1n}	0.01	Fig. 6; Opdyke <i>et al.</i> 2000
324.5	3.2	20% MI9r	Rocky Creek Syncline, E. Australia	Upper part of Ermelo Pyroclastics {50% into MI9r}	0.01	Fig. 6; Opdyke <i>et al.</i> 2000

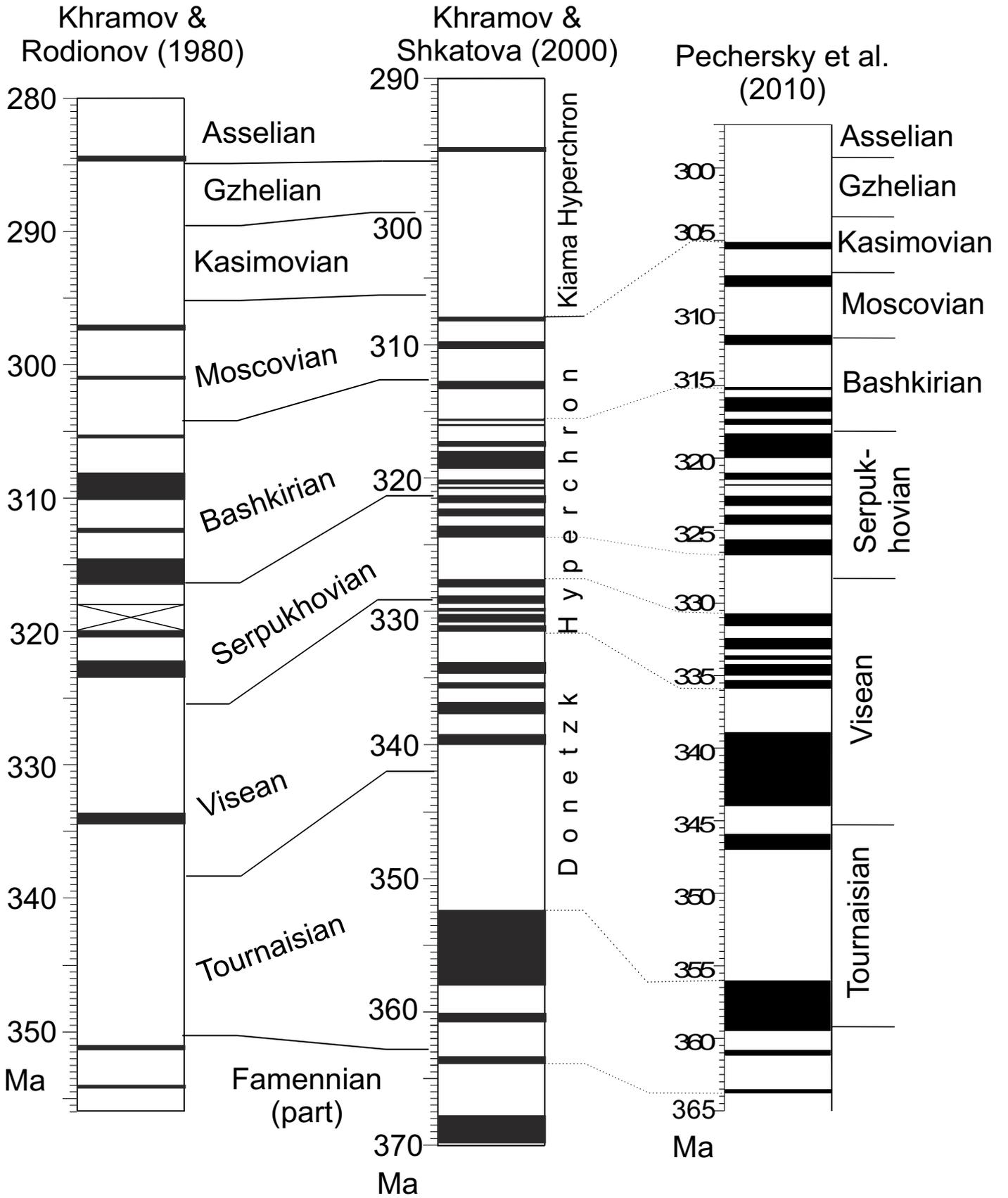
Table 3. Carboniferous radioisotopic dates used to constrain the GPTS. **Column 1:** Analysis code and date (in Ma). **Column 2:**  $\pm 2\sigma_R$  = two-sigma error on age (includes tracer calibration, and  $^{238}\text{U}$  decay constant, except for SHRIMP dates). **Column 3:**  $\pm e_s$  = estimated stratigraphic error in placing the date onto the magnetostratigraphy in percent of magnetozones widths. **Column 4:** section name, location. **Column 5:** Stratigraphic age or location, {..}= correlated chron position of date (see Figs. 16 and 17). **Column 6:**  $P_{\text{out}}$ , probability (0 to 1.0 range) the date is an outlier (from Bchron). The low values here all indicate none are likely outliers'. **Column 7:** source reference for the radioisotopic and age information. PE7r length here is to top of Bol'shaya Kalitva section in Fig. 1.

Chron	Age (Ma)	Chron duration (Ma)	C <sub>95</sub> (Ma)	σ <sub>T</sub> (Ka)	Chron	Age (Ma)	Chron duration (Ma)	C <sub>95</sub> (Ma)	σ <sub>T</sub> (Ka)
C1r.1r	298.69	0.73 <sup>H</sup>	0.37	123	MI8r.2r	326.84	1.55	0.39	-
C1n	298.77	0.081 <sup>H</sup>	0.37	140	MI8r.1n	327.05	0.17	0.35	-
PE8r	305.1	~6.3 <sup>a</sup>	-	-	MI8r.1r	327.68	0.66	0.32	144
PE8nB	~305.23	~0.13 <sup>a</sup>	-	-	MI8n.2n	327.93	0.22	0.35	-
PE7r	309.41	~4.2 <sup>a</sup>	0.74	-	MI8n.1r	328.12	0.18	0.33	-
PE7n	309.85	0.23	0.73	-	MI8n.1n	328.53	0.46	0.36	80
PE6r	312.17	2.00	0.27	-	MI7r	329.18	1.32	0.84	-
PE6n.2n	312.24	0.067	0.26	-	MI7n.2n	329.29	0.12	0.87	15
PE6n.1r	312.63	0.54	0.37	123	MI7n.1r	329.60	0.40	0.99	-
PE6n.1n	312.70	0.063	0.38	74	MI7n.1n	329.75	0.15	1.00	-
PE5r	315.13	3.30	1.07	8	MI6r.2r	330.56	0.75	1.07	-
PE5n	315.34	0.16	1.10	29	MI6r.1n	330.80	0.23	1.07	-
PE4r	317.46	1.17	0.29	41	MI6r.1r	331.44	0.38	0.97	25
PE4n.2n	317.58	0.13	0.29	27	MI6n.2n	331.67	0.16	0.92	3
PE4n.1r	317.69	0.16	0.32	0	MI6n.1r	331.89	0.13	0.84	19
PE4n.1n	317.94	0.33	0.38	-	MI6n.1n	332.06	0.10	0.76	5
PE3r.2r	318.37	0.40	0.39	50	MI5r	332.601	0.38	0.39	18
PE3r.1n	318.39	0.02	0.39	37	MI5n.2n	333.38	0.88	0.50	286
PE3r.1r	318.60	0.15	0.33	35	MI5n.1r	333.79	0.39	0.44	218
PE3n.2n	318.80	0.21	0.33	76	MI5n.1n	334.21	0.73	0.61	239
PE3n.1r	318.85	0.045	0.33	86	MI4r.4r	334.48	0.23	0.66	-
PE3n.1n	319.72	1.58	0.80	-	MI4r.3n	334.70	0.18	0.66	-
PE2r	319.78	0.10	0.84	41	MI4r.3r	335.079	0.25	0.62	26
PE2n	320.49	0.86	1.12	-	MI4r.2n	335.19	0.082	0.61	5
PE1r.2r	321.10	0.59	1.16	-	MI4r.2r	335.42	0.14	0.47	9
PE1r.1n	321.16	0.06	1.16	-	MI4r.1n	335.57	0.14	0.41	16
PE1r.1r	322.32	0.91	1.03	16	MI4r.1r	~336 <sup>F</sup>	~0.43	-	-
PE1n	323.59	0.98	0.76	28	MI4n <sub>B</sub>	~338.2 <sup>F</sup>	~2.2	-	-
MI9r	324.07	0.35	0.65	196	MI3r <sub>B</sub>	~340 <sup>F</sup>	~1.8	-	-
MI9n.3n	324.76	0.62	0.43	-	MI3n <sub>B</sub>	~342 <sup>F</sup>	~2	-	-
MI9n.2r	324.80	0.051	0.44	-	MI2r <sub>B</sub>	~343 <sup>F</sup>	~1	-	-
MI9n.2n	324.94	0.17	0.47	-	MI2n <sub>B</sub>	~349 <sup>F</sup>	~6	-	-
MI9n.1r	325.04	0.11	0.48	-	MI1r <sub>B</sub>	~354.7 <sup>F</sup>	~5.7	-	-
MI9n.1n	325.19	0.12	0.48	131	MI1n <sub>B</sub>	~357.5 <sup>S</sup>	~2.8	-	-

Table 4. Ages of Carboniferous magnetochron bases and durations. C<sub>95</sub>: 95% Highest posterior density intervals on the chron age, estimated using Bchron (shown in Figs. 18, 19). σ<sub>T</sub>: standard deviation of the chron position in the sections for the chron (from the optimisation method), scaled by the duration of the optimised chron. σ<sub>T</sub> is a measure of the uncertainty in defining the chron position in the optimised composite. <sup>a</sup> Based on base of *subcrassulus* Zone from Davydov *et al.* (2012), and average duration of PE7n, PE6n.2n and PE6n.1n. <sup>F</sup> estimated from Fig. 15. <sup>H</sup> from Hounslow & Balabanov (2018). <sup>S</sup> based on the base of the *sandbergi* Zone from Davydov *et al.* (2012).



Fig. 2



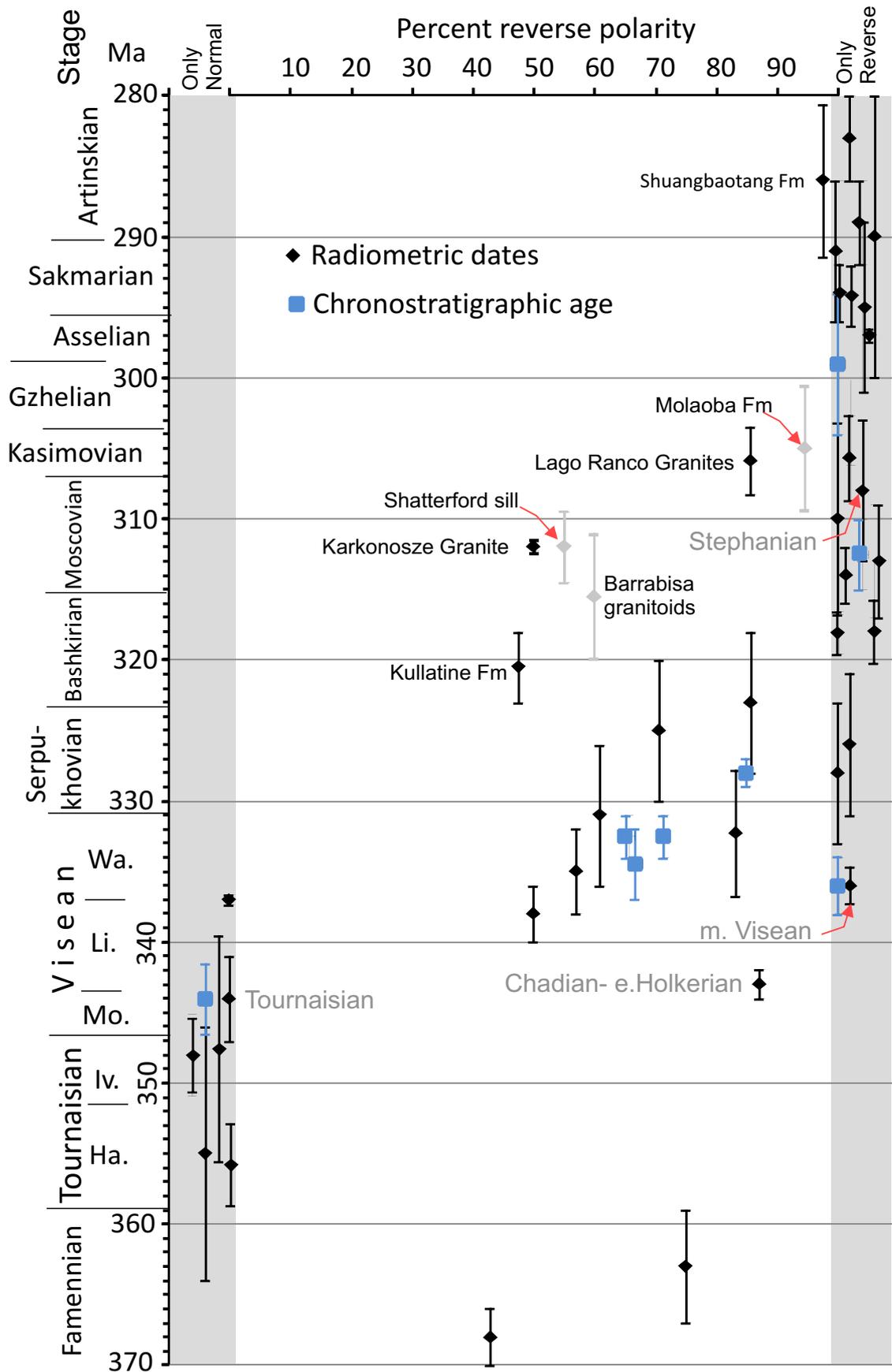


Fig. 3

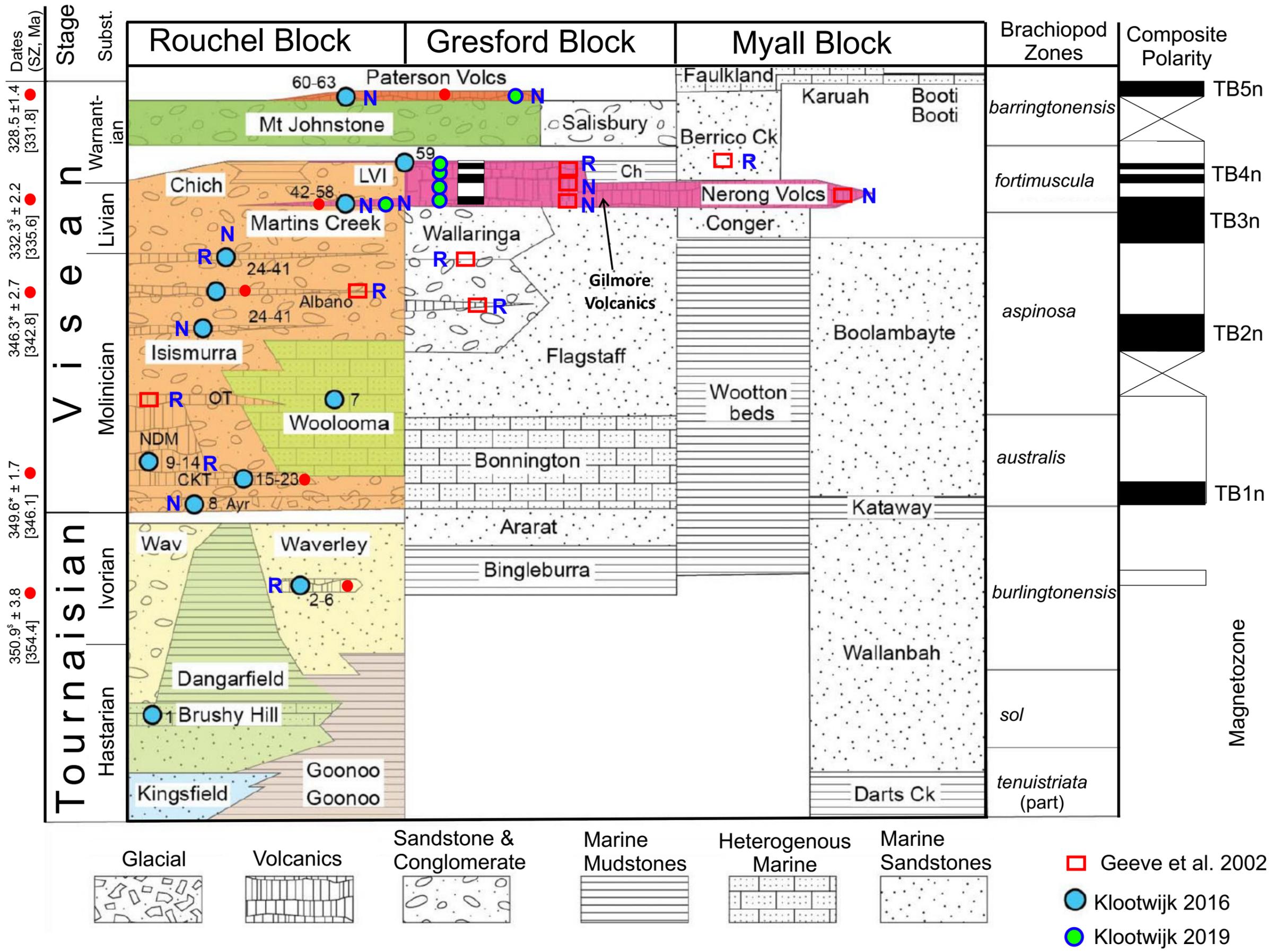
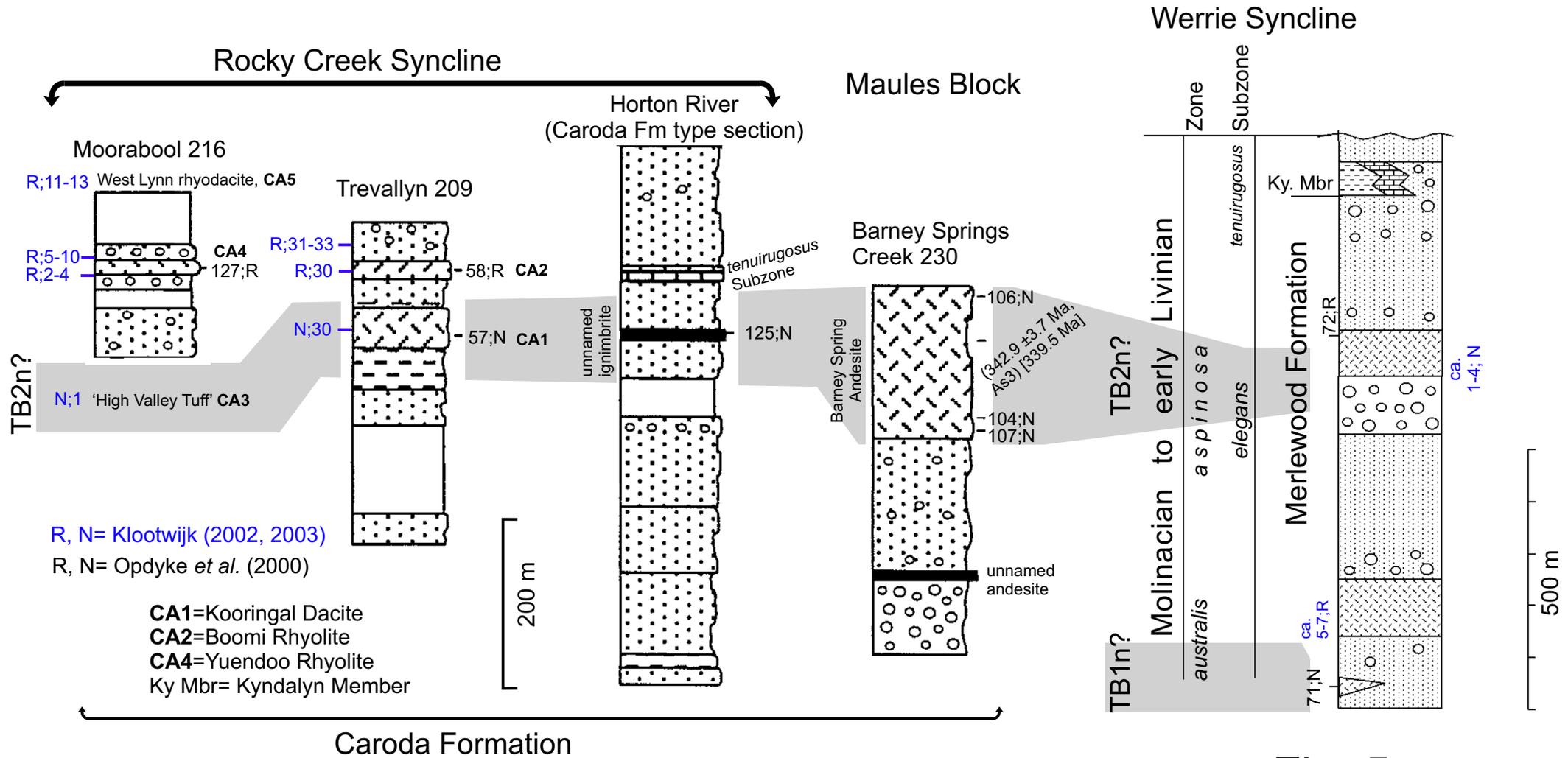
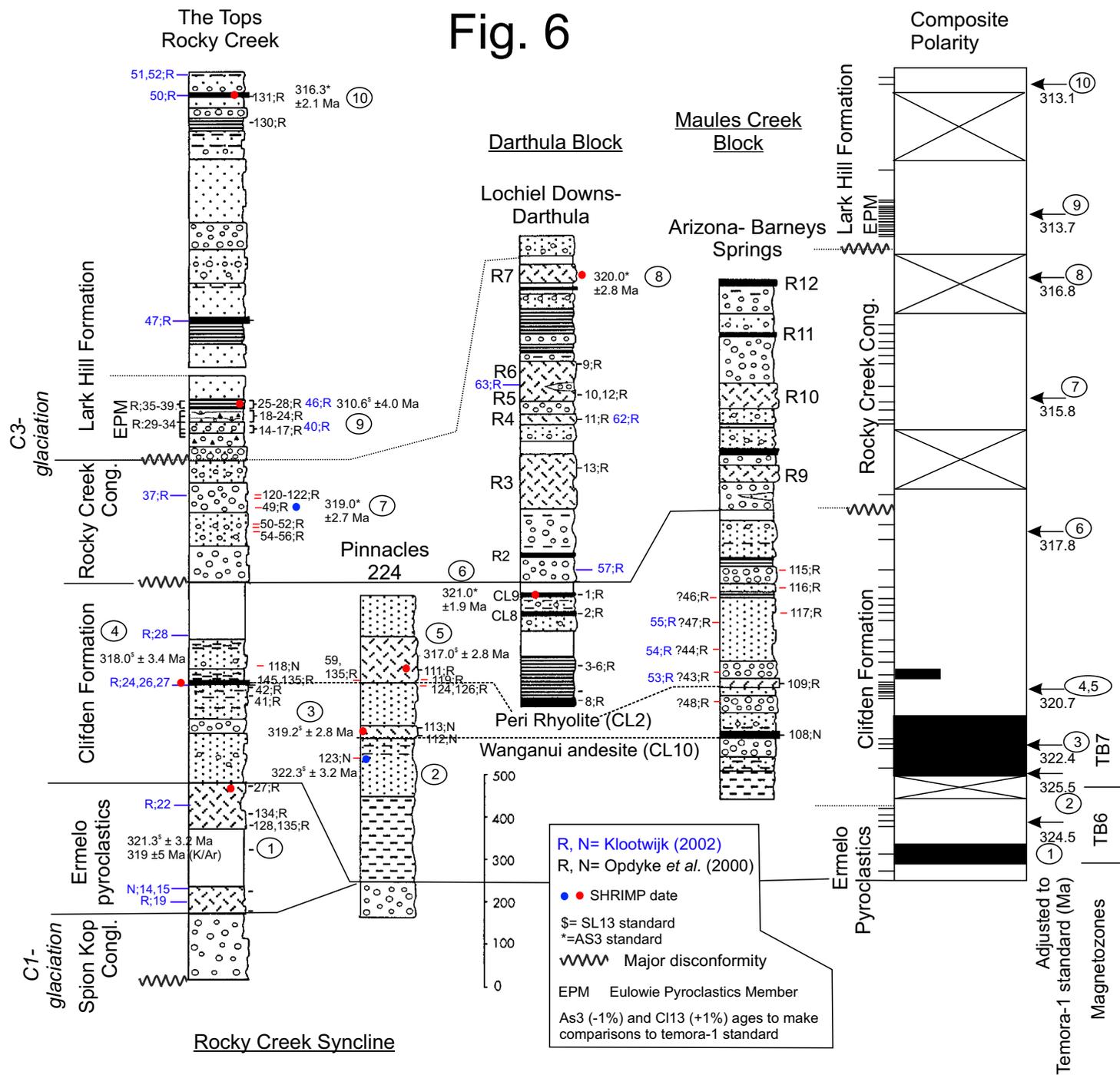


Fig. 4



**Fig. 5**

# Fig. 6



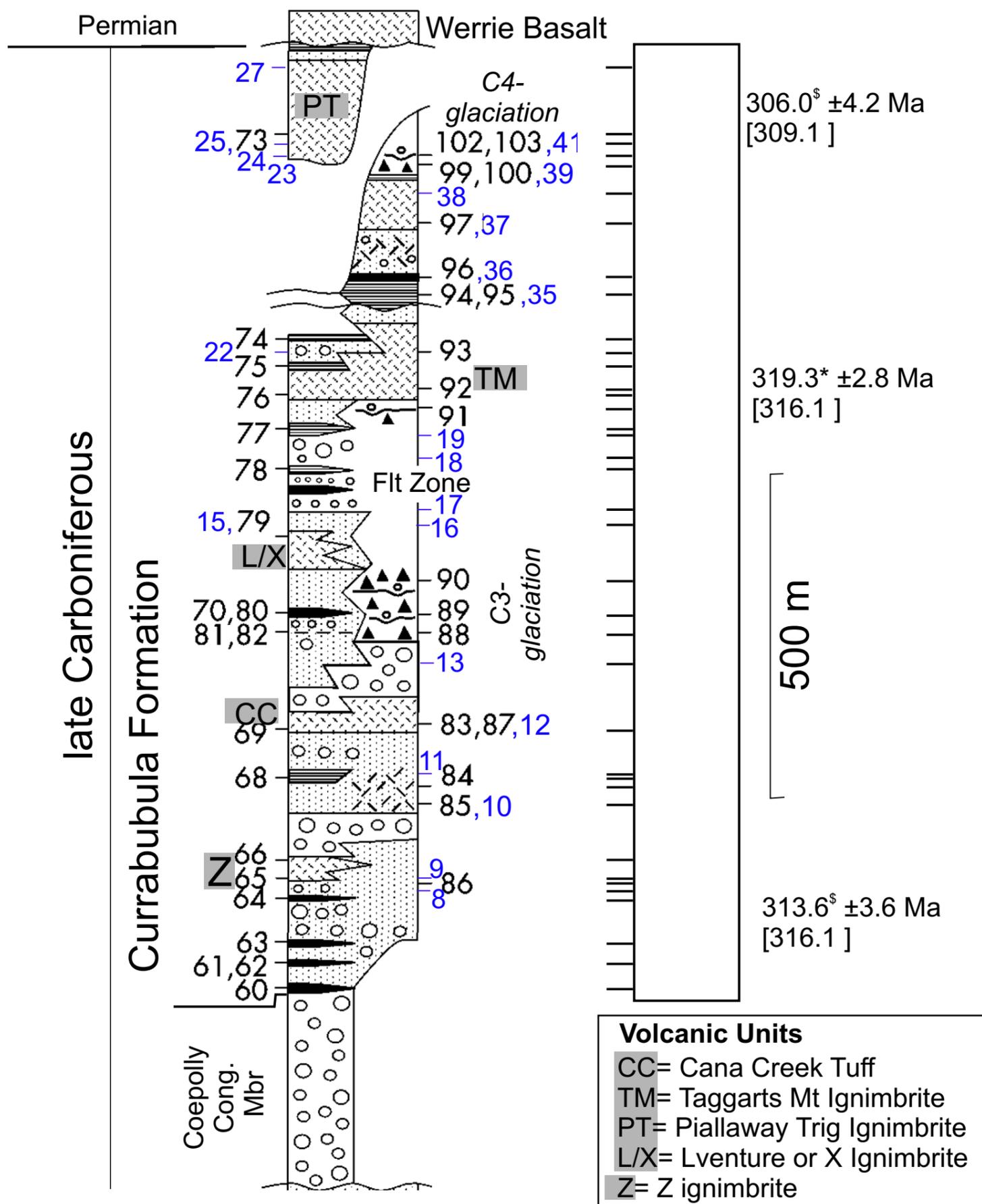
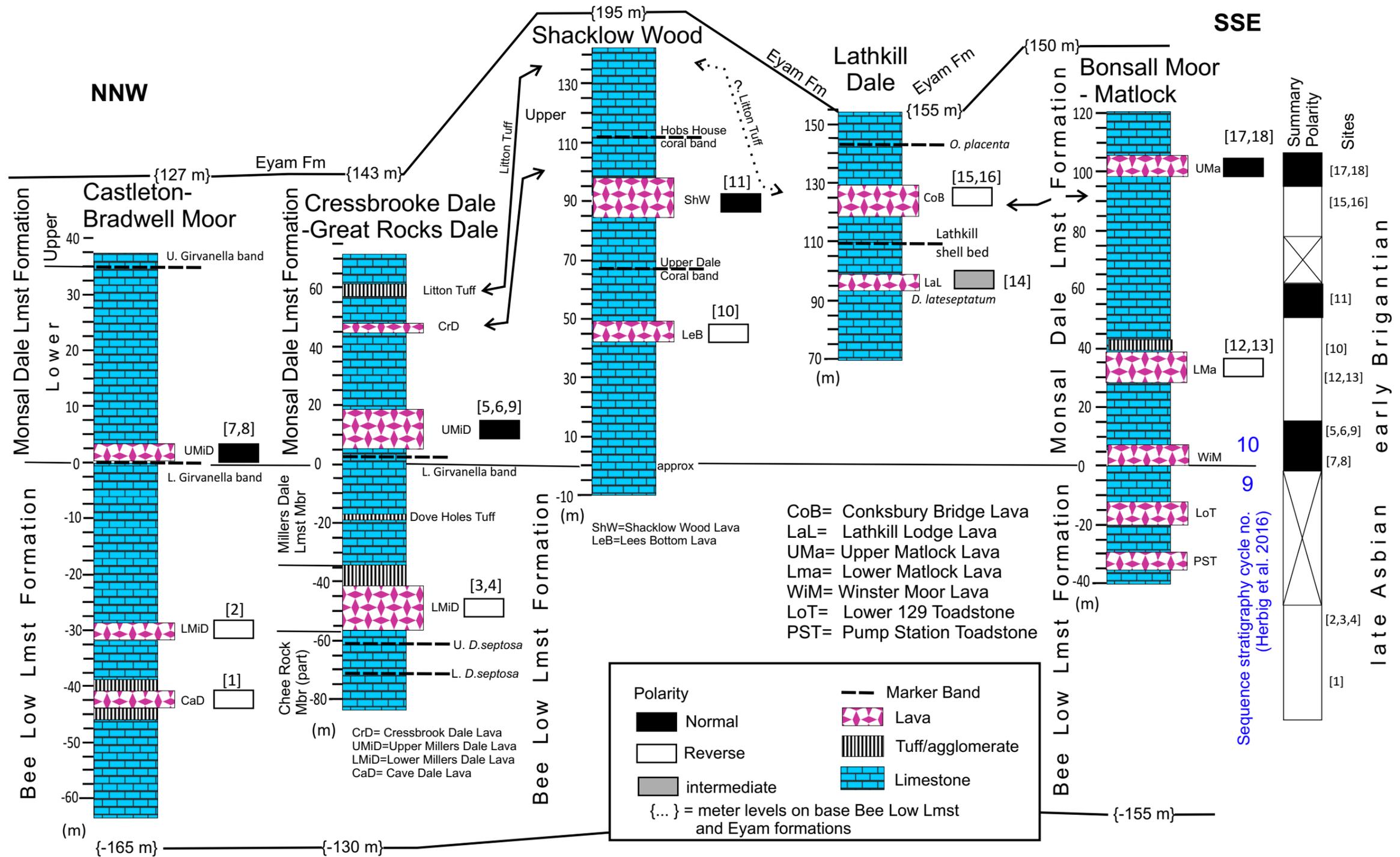


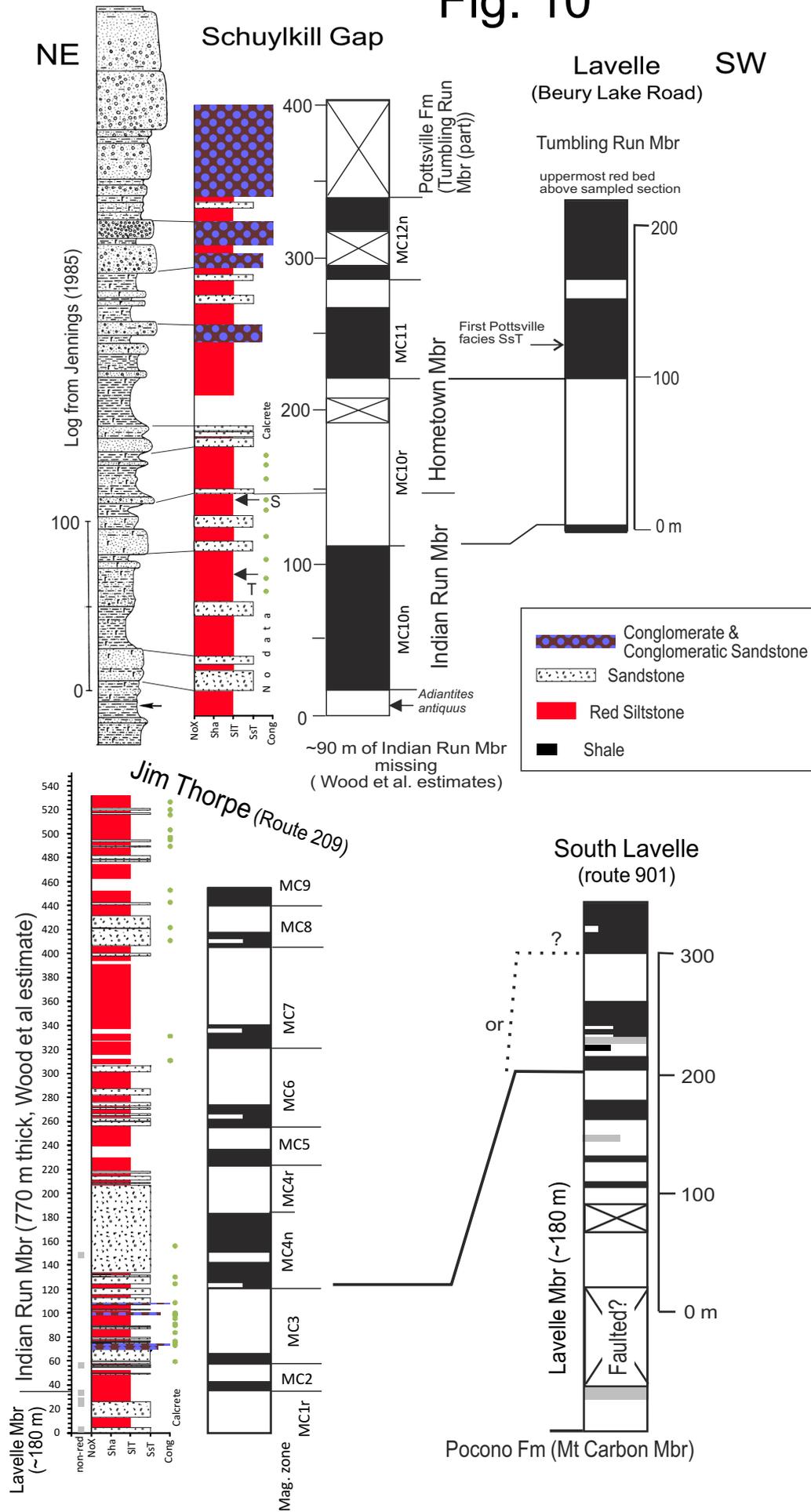
Fig. 7

Fig. 8





# Fig. 10



# Fig. 11

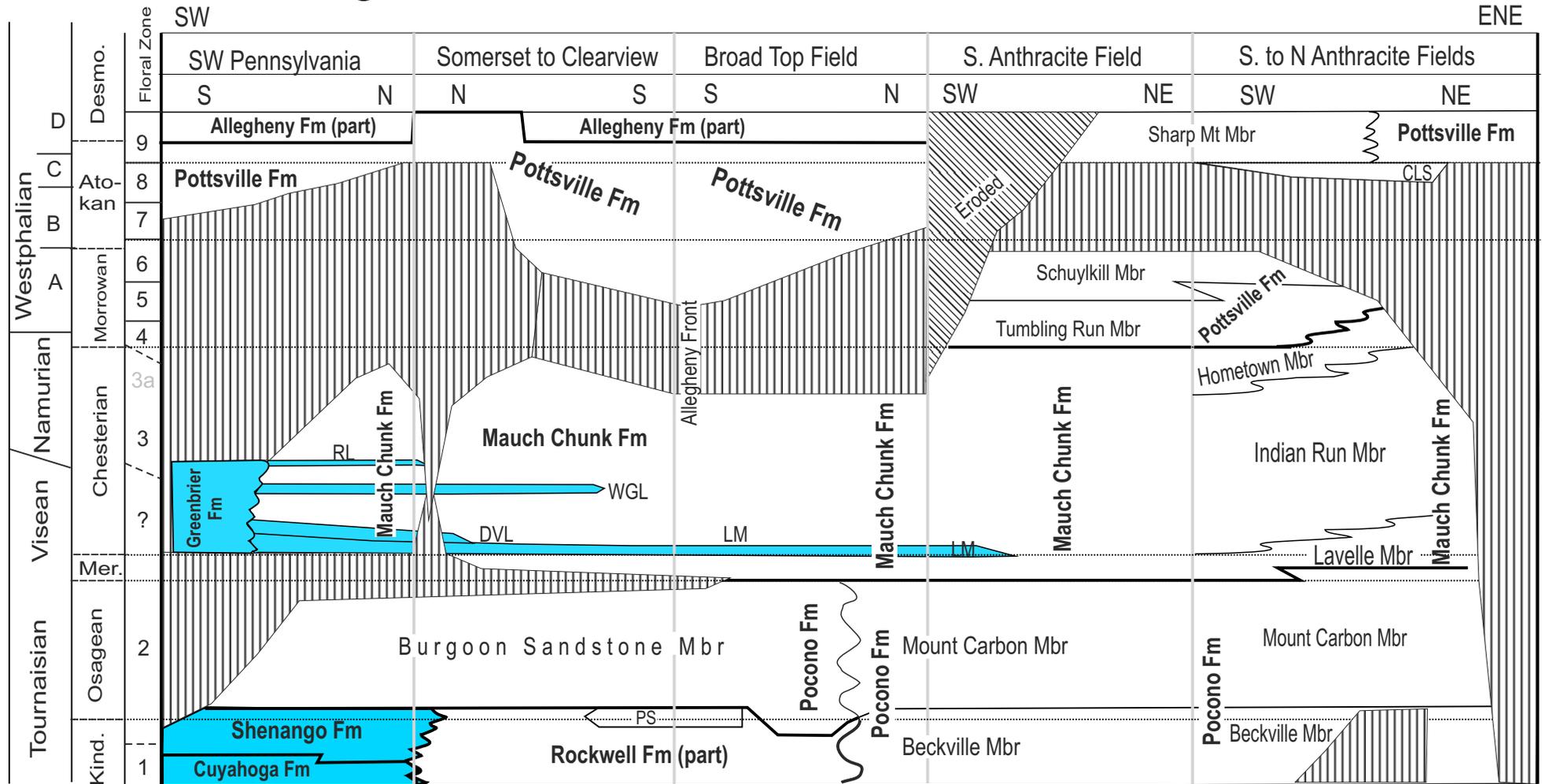




Fig. 13

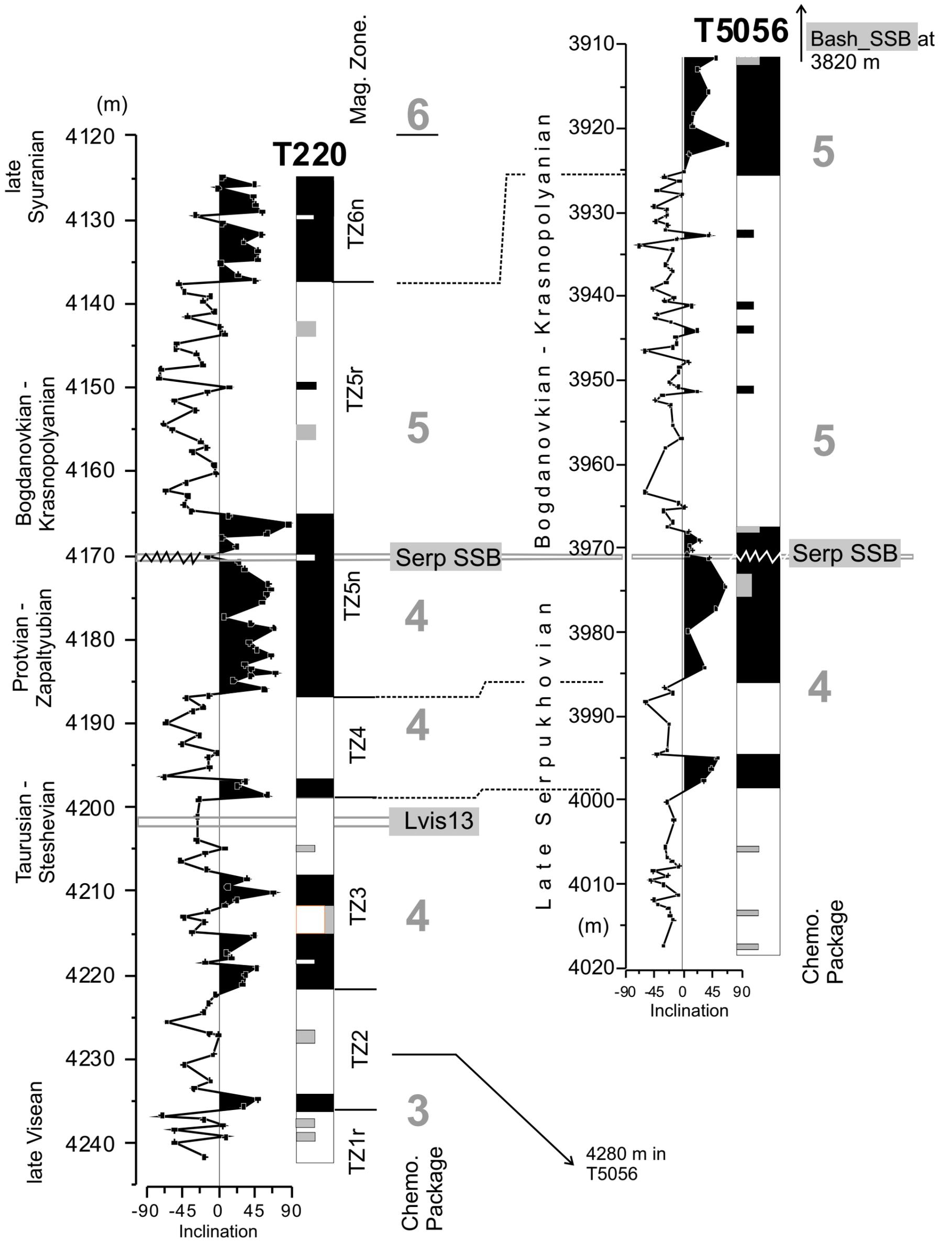


Fig. 14

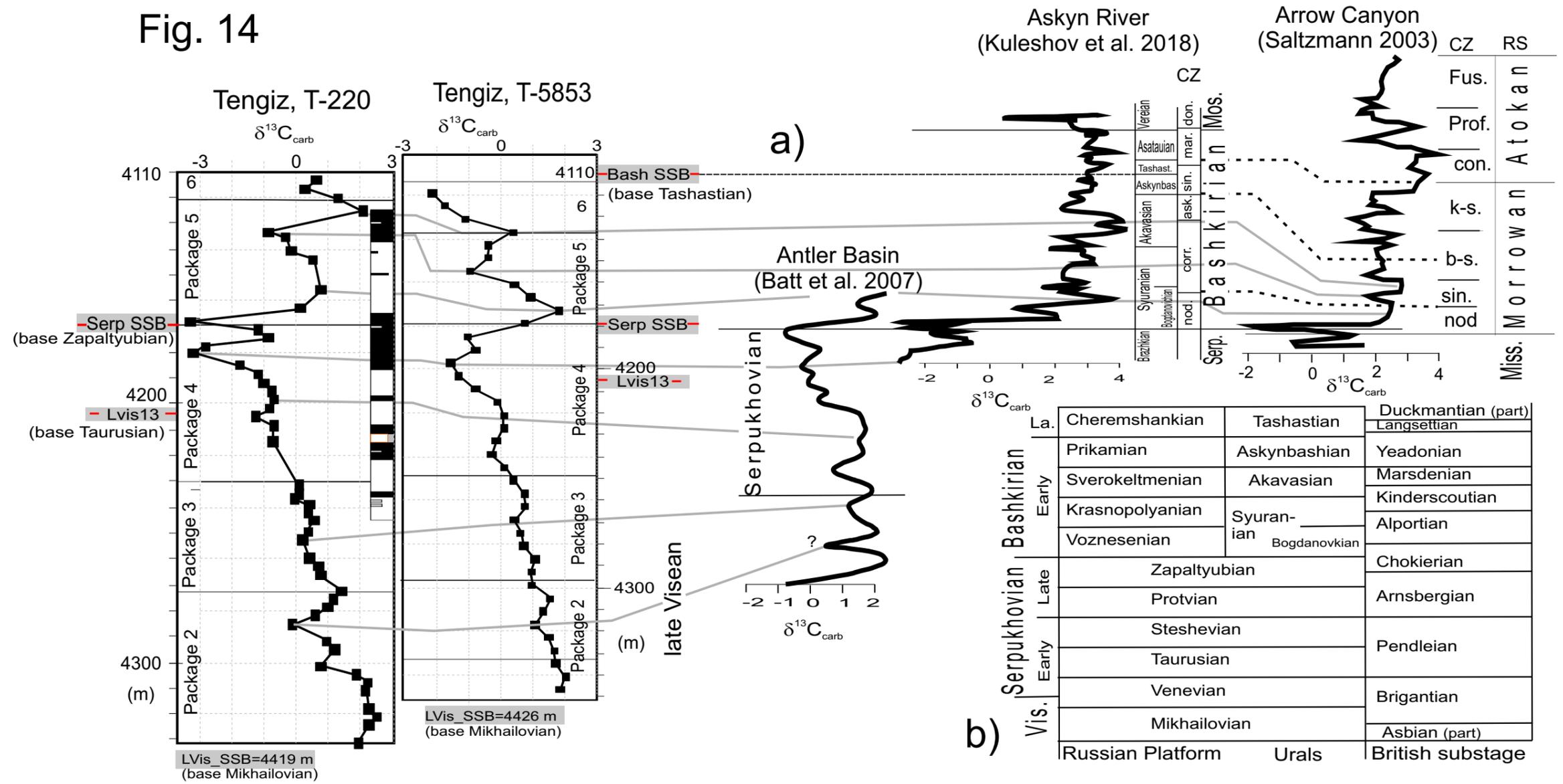
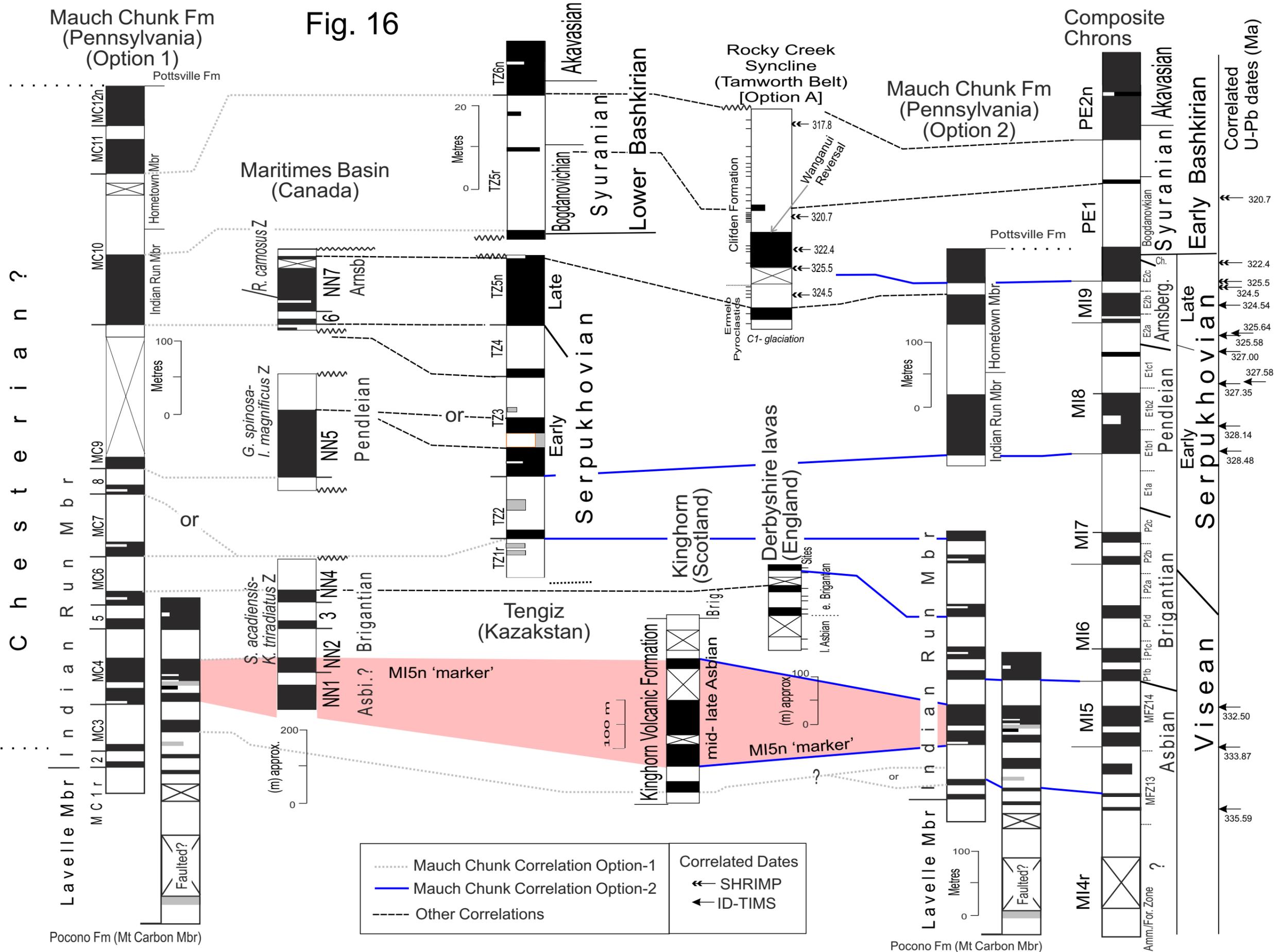




Fig. 16



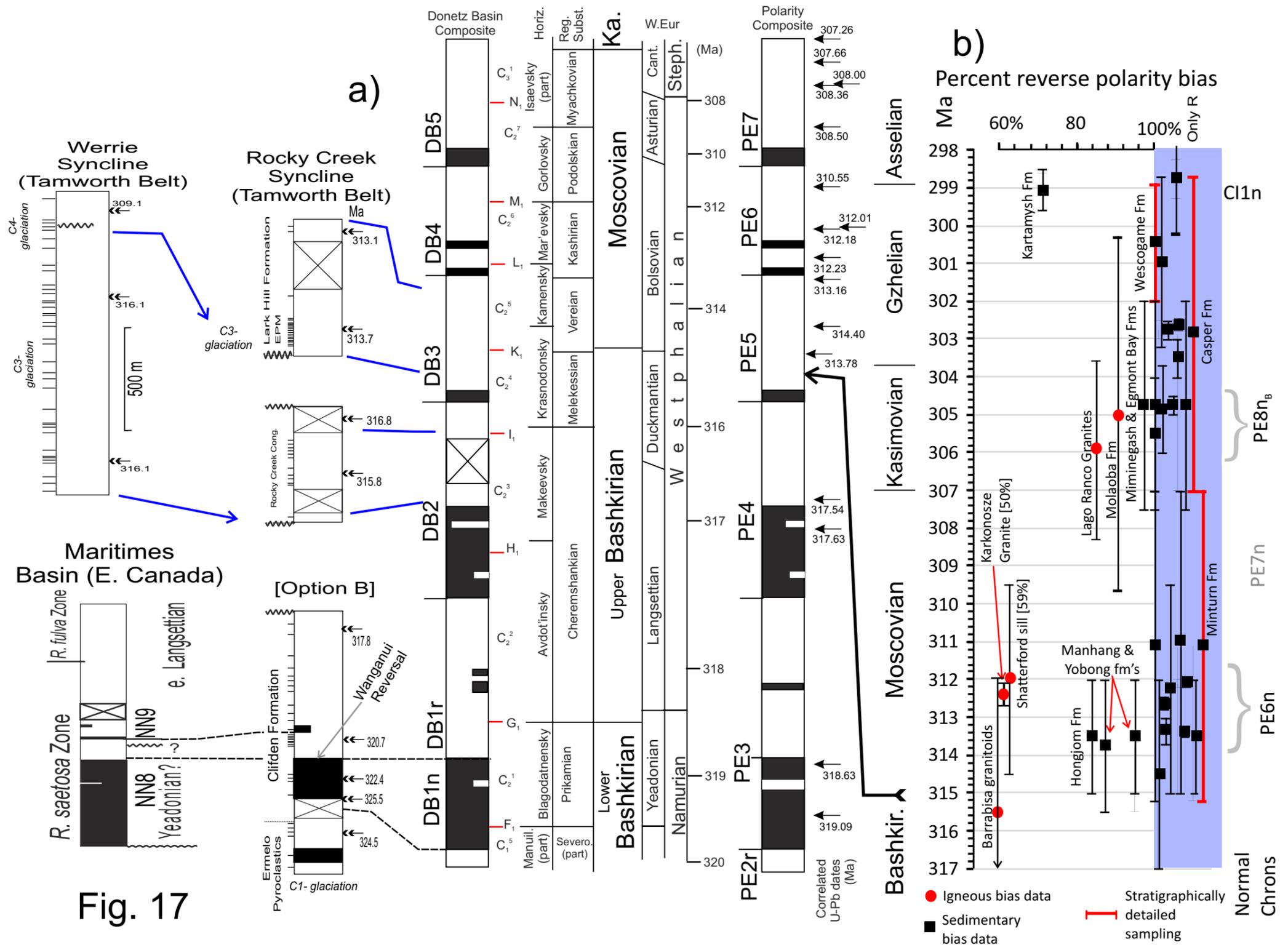
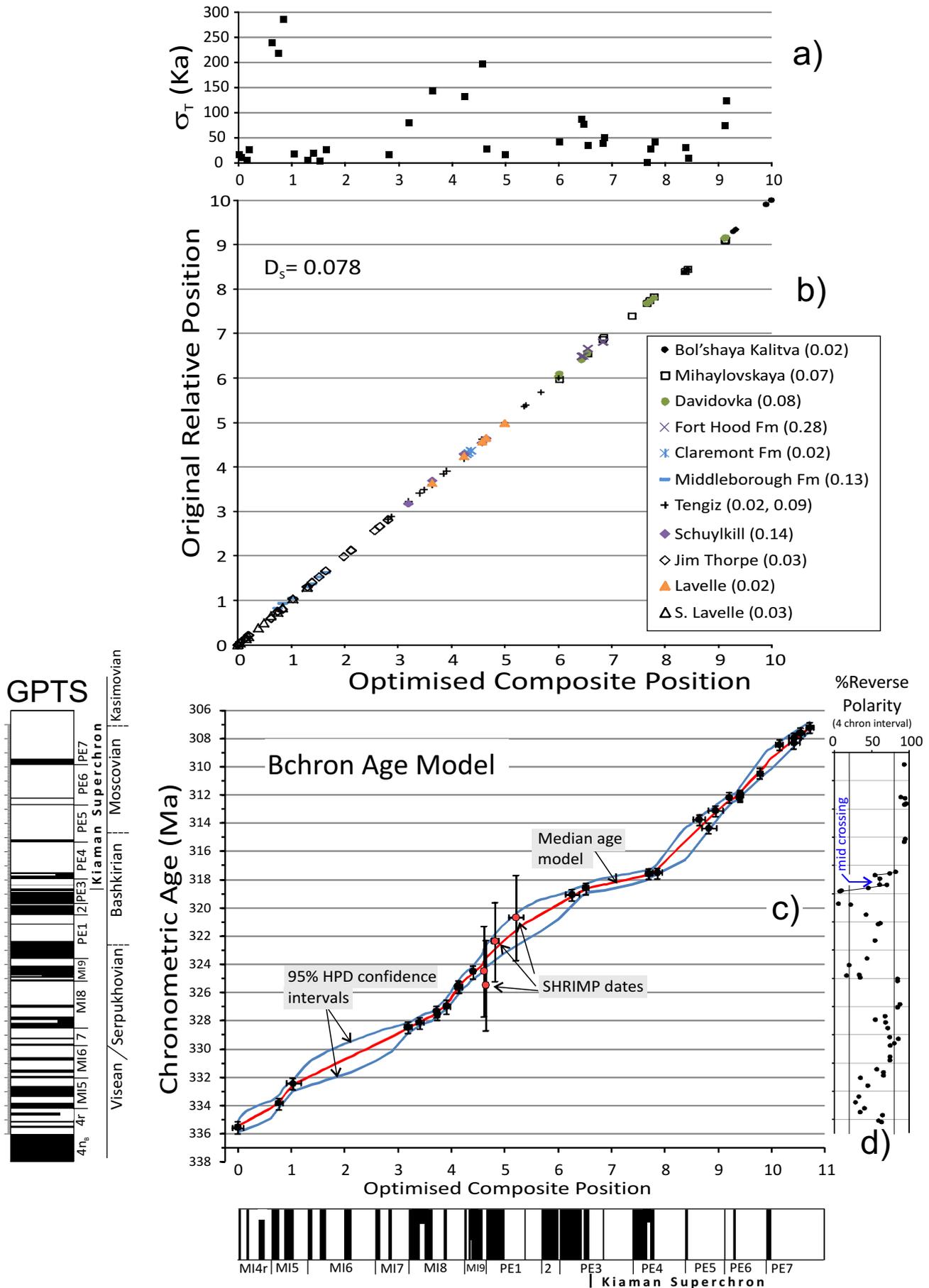


Fig. 17

Fig. 18



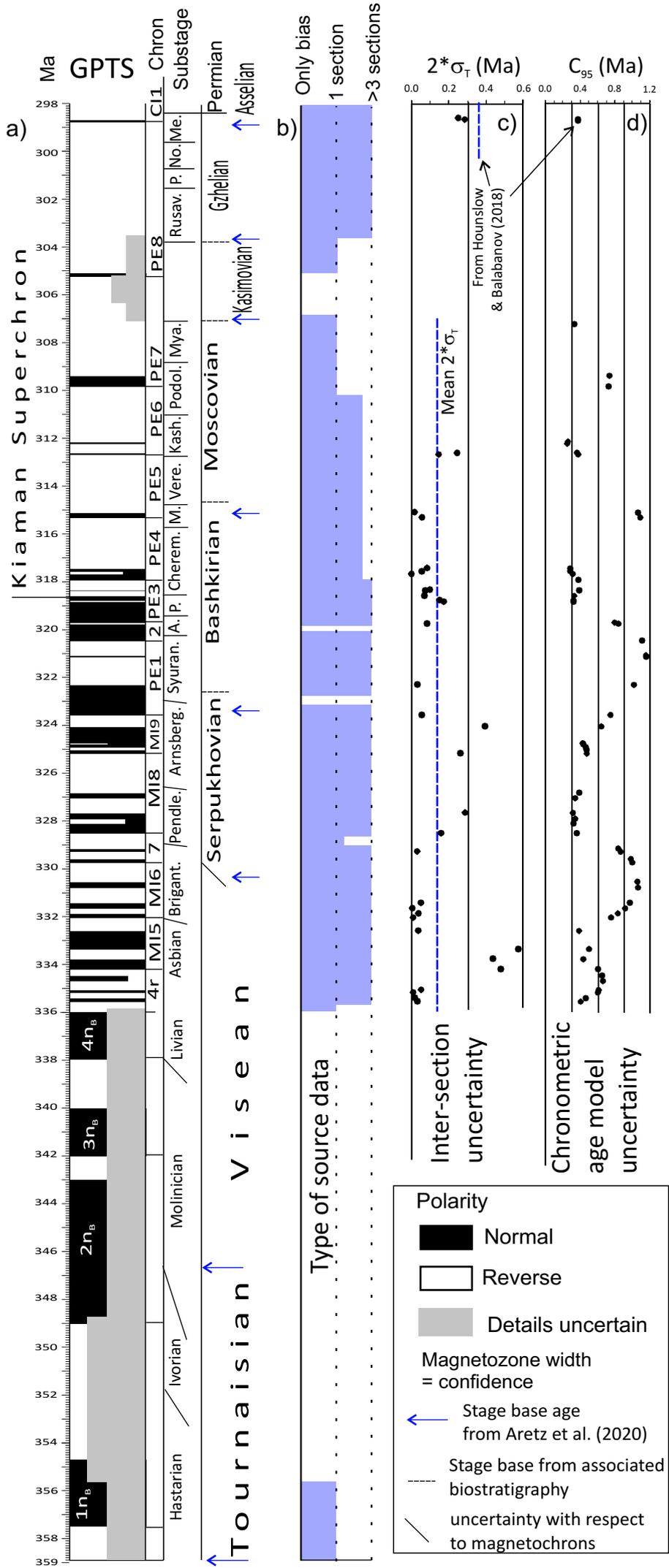


Fig. 19