First Search for Dark Sector e⁺e⁻ Explanations of the MiniBooNE Anomaly at MicroBooNE

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We present MicroBooNE's first search for dark sector e^+e^- explanations of the long-standing MiniBooNE anomaly. The MiniBooNE anomaly has garnered significant attention over the past 20 years including previous MicroBooNE investigations into both anomalous electron and photon excesses, but its origin still remains unclear. In this letter, we provide the first direct test of dark sector models in which dark neutrinos, produced through neutrino-induced scattering, decay into missing energy and visible e^+e^- pairs comprising the MiniBooNE anomaly. Many such models have recently gained traction as a viable solution to the anomaly while evading past bounds. Using an exposure of 6.87×10^{20} protons-on-target in the Booster Neutrino Beam, we implement a selection targeting forward-going, coherently produced e^+e^- events. After unblinding, we observe 95 events, which we compare with the constrained background-only prediction of 69.7 ± 17.3 . This analysis sets the world's first direct limits on these dark sector models and, at the 95% confidence level, excludes the entirety of single dark neutrino, and majority of dual dark neutrino, parameter space viable as a solution to the MiniBooNE anomaly.

The anomalous results obtained in the short-baseline115 MiniBooNE [1–3] and LSND [4] experiments have hinted 116 at the existence of possible new physics beyond the117 standard model (BSM) for two decades. The discovery₁₁₈ of one or more new particles would represent the first₁₁₉ laboratory evidence of a new paradigm in high energy₁₂₀ physics since the detection of neutrino mass through 121 oscillations with profound implications for particle,122 astrophysics, and cosmology. As such, a great deal of 123 effort has been focused on confirming, or rejecting, vari-124 ous interpretations of these anomalous excesses. These 125 efforts include the construction of the MicroBooNE₁₂₆ experiment [5], an 85 metric ton active volume liquid₁₂₇ argon time projection chamber (LArTPC) situated in₁₂₈ the Booster Neutrino Beam (BNB) at Fermilab [6]. The 129 MicroBooNE experiment has been specifically designed 130 to study the MiniBooNE anomaly as its primary goal. 131 Both detectors were located in the same neutrino beam₁₃₂ at a similar baseline.

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Possible origins of the MiniBooNE anomaly through¹³⁵ more traditional hypotheses, such as underestimated¹³⁶ backgrounds or systematic uncertainties [7, 8], have¹³⁷ been unable to explain the anomaly and its resolution¹³⁸

remains unclear. MicroBooNE has previously reported on searches for both an anomalous excess of electron-like events [9–13], a subsequent direct search for a light oscillating sterile neutrino in the BNB [14], as well as a search for neutrino-induced neutral current (NC) production of the $\Delta(1232)$ baryon resonance with subsequent Δ radiative decay [15], the dominant source of single photons predicted in both neutrino-argon and neutrino-carbon scattering below 1 GeV [16]. these placed strong bounds on their respective channels, MicroBooNE's first single-photon search was a model dependent search for NC Δ radiative decay specifically, and the strength of its result relied heavily on observing a proton in conjunction with the single-photon final state. Although this first analysis did contain a zeroproton sample, it had substantially more backgrounds. As such, any coherent-like single-photon final states without observable hadronic components were largely unconstrained. This possibility remains a viable and untested explanation of the MiniBooNE anomaly.

In recent years, there has been a rapidly growing interest in using neutrino experiments as a probe into dark sectors, hypothetical extensions of the standard model

(SM) that introduce new particles and interactions below₁₇₄ the electroweak scale. In many dark sector models, 175 new unstable particles can be abundantly produced₁₇₆ in neutrino-nucleus interactions through one or more₁₇₇ portals, renormalizable terms added to the interaction₁₇₈ Lagrangian that couple the dark sector weakly to the 179 SM, such as the vector, scalar and neutrino portals [17].180 It was highlighted that, alongside the long discussed 181 electron and photon explanations of the MiniBooNE₁₈₂ anomaly, if a dark sector particle decays to electronpositron (e^+e^-) pairs with $\mathcal{O}(100)$ MeV energies [18, 19], its signature can contribute to the excess of electron-like events observed by the MiniBooNE experiment. Although originally proposed to explain the MiniBooNE excess, the development of such models has since gained₁₈₃ traction within the community and now forms a diverse $_{184}$ class of theories that we are beginning to explore [20–30].₁₈₅

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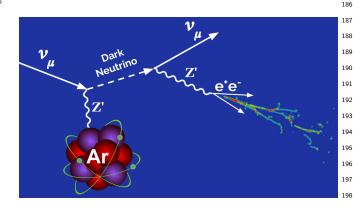


FIG. 1. An example of the dark sector models being probed. An incoming muon-neutrino from the BNB scatters off an argon target producing a dark neutrino. The subsequent decay of the dark neutrino leads to a visible e^+e^- pair, mediated via a new dark gauge boson (Z'). A simulated event is also²⁰³ shown behind the e^+e^- arrows to represent what an e^+e^- pair would actually look like in our LArTPC detector after forming electromagnetic showers.

This letter describes the first experimental search₂₀₈ for dark sector e^+e^- explanations of the MiniBooNE₂₀₉ anomaly and complements and improves on previous phe-210 nomenological work [21, 31, 32]. We focus on the sce-211 nario in which active neutrinos from the BNB scatter to₂₁₂ produce one or two dark neutrinos, which subsequently 213 decay into missing energy and visible e^+e^- pairs. It₂₁₄ builds on prior MicroBooNE e^+e^- searches utilizing the₂₁₅ NuMI beamline [33–36]. While any dark sector may²¹⁶ be quite rich phenomenologically, here a minimal $U(1)'_{217}$ gauge group is assumed with one or more dark neutrinos₂₁₈ being the only particles carrying the group charge. The219 combination of neutrino mixing between active and dark₂₂₀ neutrinos (the neutrino portal) as well as kinetic mix-221 ing between the dark gauge boson (Z') and the standard₂₂₂ model photon (the vector portal) gives rise to the neces-223 sary phenomenology to explain the MiniBooNE anomaly. 224

An example of how these dark sector models produce the visible e^+e^- signal is shown in Fig. 1 for a single dark neutrino. For dual dark neutrino models the outgoing muon neutrino would be replaced by a second, lighter, dark neutrino. To study these explanations, we use the phenomenological model as described in Ref. [37] where we parameterize the interaction Lagrangian that couples the Z' boson to both neutral leptons and to standard model electromagnetic (EM) charge as:

$$\mathcal{L}_{\mathrm{int}} \supset \sum_{i,j}^{3+n} g_D Z'_{\mu} V_{ij} \overline{\nu_i} \gamma^{\mu} \nu_j - e \varepsilon Z'_{\mu} J^{\mu}_{\mathrm{EM}},$$
 (1)

where n is the number of dark neutrino states, $J_{\rm EM}^{\mu}$ is the SM EM current with ε as the kinetic mixing, $V_{ij} = \sum_{k=0}^{n} U_{ki}^{*} U_{kj}$ is the interaction vertex for dark current assuming dark neutrino states have a +1 charge under the U(1)' group, and g_D is the dark sector U(1)' coupling constant which is set at 2 throughout this paper. We adopt the same simplifying assumptions as Ref. [37] in which dark neutrinos mix only with muon neutrinos. More details of the model and simplifying assumptions can be found in the Supplemental Materials. In this simplified regime $|V_{\mu 4(5)}| = |U_{\mu 4(5)}|$, the latter being a more familiar parameter from the literature in heavy neutral leptons and the primary mixing element probed in this result.

We consider models with one or two dark neutrinos. In the single dark neutrino model, the dark neutrino ν_4 is long-lived unless it decays to an on-shell Z' boson via $\nu_4 \to \nu_\mu(Z' \to e^+e^-)$. This need not be the case for ν_5 in the dual dark neutrino models since it decays via $\nu_5 \to \nu_4(Z' \to e^+e^-)$. While the decay rate for ν_4 is proportional to the small coupling between active and dark neutrinos, the decay rate for ν_5 is proportional to an $\mathcal{O}(1)$ coupling between the dark sector ν_5 and ν_4 . This allows us to explore regions of parameter space where the Z' is heavy $(m_{Z'} \gtrsim O(500) \text{ MeV})$ while still targeting dark neutrinos that have a large probability to decay inside the MicroBooNE detector after the initial neutrino interaction.

The kinematics of the decay process is also different between single and dual dark neutrino models. The latter depends on a key dimensionless parameter $\Delta = (m_5 - m_4)/m_4$, which controls the visible energy release in the $\nu_5 \to \nu_4 e^+ e^-$ decays of interest. When $\Delta \gg 1$, the decay kinematics is reduced to that of a single dark neutrino scenario, while for $\Delta < 1$, the energy release into the final state $e^+ e^-$ is suppressed. A further description of the class of models, as well as its ability to describe the MiniBooNE anomaly, can be found in the Supplemental Materials as well as references [20, 37, 38] and references therein with further

discussions of UV completions of this phenomenological₂₈₁ model in [20, 22, 39]. ²⁸²

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MicroBooNE uses a custom tune [40] of the GENIE²⁸⁴ neutrino event generator v3.0.6 [41, 42] to simulate₂₈₅ backgrounds from neutrino-argon interactions. BSM dark sector e^+e^- signal events were generated₂₈₇ with the open-source Darknews [38] generator that is₂₈₈ integrated into the first stage of MicroBooNE's LArTPC289 simulation chain. The DARKNEWS generator output₂₉₀ was also validated with a GENIE implementation of₂₉₁ the same single dark neutrino phenomenological model.292 Depending on the dark gauge boson mass, the initial₂₉₃ scattering process on the nucleus can be predominantly₂₉₄ coherent or incoherent, with the latter potentially₂₉₅ resulting in multiple visible hadronic final states, in₂₉₆ addition to the e^+e^- pair. This analysis focuses solely₂₉₇ on coherent scattering with no associated hadronic₂₉₈ component for multiple reasons: (a) this is the least con-299 strained channel by MicroBooNE's prior results, (b) the 300 Darknews generator lacks a detailed treatment of the₃₀₁ complex nuclear response of the incoherent regime, and₃₀₂ (c) when moving from carbon ($Z_{\rm C}=6$, MiniBooNE's₃₀₃ primary target) to argon ($Z_{Ar} = 18$, MicroBooNE's₃₀₄ target) we expect a substantial enhancement to coherent 305 scattering that naively scales as atomic number squared, 306 $(Z_{\rm Ar}/Z_{\rm C})^2 = 9$, helping MicroBooNE's sensitivity to₃₀₇ remain competitive by compensating for its lower active₃₀₈ mass of 85 metric tonnes compared to MiniBooNE's 818 metric tonnes.

This coherent scattering tends to produce very₃₁₁ forward-going dark neutrinos due to the low momentum312 exchange with the argon nucleus. When considering the313 subsequent dark neutrino decay, the model parameter space is broadly split into two regimes: for very prompt₃₁₄ decays, with proper lifetime $c\tau \ll 10$ cm, the scattering³¹⁵ on argon ($\approx 1.38 \text{ g/cm}^3$) and subsequent decay both₃₁₆ occur predominately inside the detector, and when the₃₁₇ proper lifetime grows, $c\tau >> 10$ cm, the portion of dark₃₁₈ neutrino decays originating from neutrino scattering on argon inside the detector is dwarfed by the flux of $_{319}$ long-lived dark neutrinos from neutrino scattering on the 320 compacted soil and clay ($\approx 2.25 \text{g/cm}^3 \text{ see } [43-45] \text{ and}_{321}$ the Supplemental Materials) upstream of the detector322 and subsequently entering the MicroBooNE TPC. Both₃₂₃ regimes are crucial for probing this class of models and 324 are fully simulated in this analysis.

A high statistics sample of dark sector signal events were generated and passed through MicroBooNE's₃₂₇ simulation and reconstruction suite. This sample was₃₂₈ distributed to cover the heavy $(m_{Z'} > m_{4(5)})$ and light₃₂₉ $(m_{Z'} << m_{4(5)})$ kinematic regions of interest of, as well₃₃₀ as single and dual dark neutrino scenarios. Due to the₃₃₁ high dimensionality of the dual dark neutrino parameter₃₃₂

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space, these simulated events were then reweighted to any point in our target model parameter space based on underlying true kinematic distributions.

The selection of e^+e^- events and rejection of cosmic and neutrino backgrounds is built upon the framework developed for MicroBooNE's first generation NC Δ radiative search [15]. Utilizing the Pandora pattern recognition reconstruction framework [46], ionization charge hits are first clustered and matched across three 2D projected views of the MicroBooNE active TPC volume into 3D reconstructed objects. These are then classified as tracks or showers based on a multivariate classifier score and aggregated into candidate neutrino interactions. While sufficiently overlapping e^+e^- pairs are primarily reconstructed as single showers and are practically identical to true photons in our detector, e^+e^- pairs with wider opening angles can be reconstructed as two-shower or one-shower plus one-track events. This analysis utilizes all three of these expected event topologies and is the first search for anomalous events consistent with MiniBooNE that targets events with two connected showers.

Once reconstructed, a series of four boosted decision trees (BDTs) are trained using the XGBoost [47] algorithm to reject various background categories while selecting dark sector e^+e^- events:

- Cosmic BDT: As a surface detector, there is a need to reject cosmogenic activity that enters in time with the TPC readout. It is trained on cosmic ray data events collected when no neutrino beam was present.
- Charged Current ν_{μ} BDT: Although the most common interaction in the 99.5% ν_{μ} BNB, CC ν_{μ} events are not a major background due the BDT being able to easily identify and veto long muon tracks [48].
- Charged Current ν_e BDT: As the primary source of non-photon EM showers, it is vital to reject CC ν_e interactions to build confidence that selected events are truly e^+e^- . This BDT relies heavily on the calorimetric capabilities of a LArTPC [49], separating electrons from photons and e^+e^- pairs by their energy deposition (dE/dx) at the start of the showering process.
- Neutral Current π^0 BDT: Neutral pions are by far the largest background to any photon or e^+e^- search. If one daughter photon from a $\pi^0 \to \gamma\gamma$ decay is not reconstructed or leaves the detector, the remaining shower can be indistinguishable from our signal.

Events with BDT scores above a chosen threshold are kept, optimizing signal efficiency × purity over a large sample of dark sector parameter space, covering both single and dual dark neutrino scenarios, as well as light and heavy dark gauge boson masses. For a detailed discussion into the most discriminating variables of each BDT, the training process as well as the final BDT classifier distributions and critical threshold values, please see the dedicated discussion included in supplemental materials.

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As this analysis is focused on coherent BSM signals, we implement recently developed tools to search for low energy deposits of charge (down to 10 MeV) in the backwards direction of the reconstructed object to help veto any events with evidence of proton activity [50]. The total efficiency for selecting dark sector e^+e^- events varies across the model parameter space but is typically in the range 20-40%. Figure 2 shows a representative efficiency as a function of true e^+e^- energy and overall angle with respect to the neutrino beam.

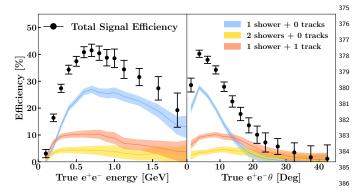


FIG. 2. Relative signal efficiency as a function of (a) true e^+e^- energy and (b) true $e^+e^ \theta$ (angle with respect to neutrino beam) for a sample of dark sector signal events generated uniformly inside the TPC and covering a wide range of model parameter space. Shown also is the how the efficiency varies when one splits the total efficiency into the three re-392 constructible topologies this analysis targets. Total efficiency is highly dependent on exact model parameters chosen which highly dependent on exact model parameters chosen which influence these variables as well as the e^+e^- energy asymmetry. For more information see the Supplemental Materials.

Prior to unblinding the signal region, the analysis³⁹⁸ underwent a series of validations in background rich³⁹⁹ sideband channels designed to ensure our modeling of⁴⁰⁰ NC π^0 backgrounds was sufficient. In all sidebands⁴⁰¹ studied, we observe good agreement between data and⁴⁰² our simulation prediction within assigned uncertainties.⁴⁰³ In addition to these dedicated background studies, a high⁴⁰⁴ statistics two-shower sideband in which both daughter⁴⁰⁵ photons of π^0 decays were successfully reconstructed⁴⁰⁶ is also employed to constrain the large cross-section⁴⁰⁷ uncertainties associated with NC π^0 production. For⁴⁰⁸ more details of this sideband and the constraint see the⁴⁰⁹

Supplemental Materials and [15, 51].

Background Events		Uncertainty	Breakdown
$NC 1\pi^0$	42.3	Cross-Section	24.2%
Out of TPC	11.8	Detector	19.9%
$CC 1\pi^0$	6.9	Flux	8.0%
Cosmic	5.2	MC Stat.	5.5%
Other	10.2	Geant4	1.2%
TT			

Unconstrained Background: $76.4 \pm 25.0 \ (32.8\%)$ Constrained Background: $69.7 \pm 17.3 \ (24.9\%)$

TABLE I. Breakdown of expected background events and the total associated systematic uncertainties for the final e^+e^- selection both before and after applying the high statistics NC π^0 constraint. Percentages refer to the uncertainty on the predicted number of events.

The effect of five categories of systematic uncertainties on our predictions are included: neutrino flux, neutrino interaction cross-sections, secondary interactions of hadrons outside of the target nucleus during Geant4 simulation [52], the detector response model, and finite Monte-Carlo simulation statistics. The effect of the detector response is estimated in the same manner [53] as our prior NC Δ radiative analysis [15]. The exact breakdown of expected backgrounds after selection as well as the effect of the NC π^0 constraint on the predicted background and its uncertainty can be found in Table. I. As with prior MicroBooNE photon searches, the vast majority of remaining backgrounds are NC π^0 events in which one photon either left the detector $(\approx 21\%)$ or failed to be reconstructed ($\approx 79\%$), leaving a single shower to mimic the e^+e^- signal.

After unblinding data corresponding to an exposure of 6.87×10^{20} protons-on-target (POT) we observe 95 e^+e^- candidate events in the signal region, consistent with a constrained background-only prediction of 69.7 \pm 17.3 at the 1.5 σ level. For the three signal event topologies studied (single-shower, one-shower plus onetrack, two-shower) we observe 52, 26, and 17 data events with constrained background predictions of 47.7 ± 11.7 , 15.5 ± 7.9 , and 9.5 ± 4.6 respectively. Distributions showing observed spectra in total reconstructed visible energy (defined as the sum of energies across all reconstructed showers and tracks), reconstructed primary shower angle, and NC π^0 background rejection BDT can be seen in Fig. 3. In all cases we see agreement between observed data and our background-only prediction within assigned uncertainties.

Studies of the sensitivity to these dark sector e^+e^- signals are performed using the total event rate, total reconstructed visible energy, reconstructed angle, and NC π^0 rejection BDT score. The most sensitive variable tested, the NC π^0 rejection BDT score, is presented here in more

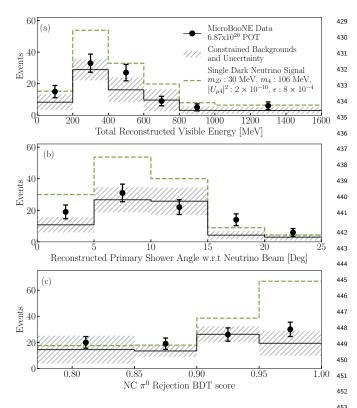


FIG. 3. Final selected data and constrained background pre- $_{454}$ diction for the e^+e^- signal region. Shown is the reconstructed $_{455}$ total energy (a), the angle the primary reconstructed shower makes with respect to the incoming neutrino beam (b) and the NC π^0 BDT rejection score (c). Shown also in green is 457 a single dark neutrino example signal stacked on top of the 458 background prediction, chosen to lie just within our predicted 459 95% C.L. sensitivity, with $m_4=106$ MeV, $m_{Z'}=30$ MeV, $_{460}$ $|U_{\mu 4}|^2=2.0\times 10^{-10}$, and $\epsilon=8\times 10^{-4}$. Equivalent spectra for pre-constrained backgrounds can be found in the Supplemental Materials.

detail. The selected e^+e^- candidates were fit alongside₄₆₁ the two-shower NC π^0 sidebands in order to constrain₄₆₂ backgrounds using the combined Neyman-Person (CNP)₄₆₃ χ^2 test statistic [54]. This test statistic includes all sys-464 tematic uncertainties discussed above through a covari-465 ance matrix. Excellent agreement with the constrained 466 background-only prediction was observed with a $\chi^2_{\rm CNP}$ of 467 3.08 for 4 degrees-of-freedom (d.o.f.). Given this agree-468 ment we calculate limits from a fit to the NC π^0 rejection₄₆₉ BDT score on both single and dual dark neutrino models $_{470}$ of dark sector production in Fig. 4. We present the lim-471 its for six representative values of the mass of the dark₄₇₂ gauge boson, Z', and the dual dark neutrino result for 473 fixed relative mass gap $\Delta = 1$. This is both for simplic-474 ity of presentation and to allow direct comparison to the 475 published MiniBooNE preferred regions of model param-476 eter space that can explain the MiniBooNE low-energy 477 anomaly [37]. Limits are obtained for each Z' mass₄₇₈ individually by placing cuts on $\Delta\chi^2_{\rm CNP}$ < 5.99, corre-479

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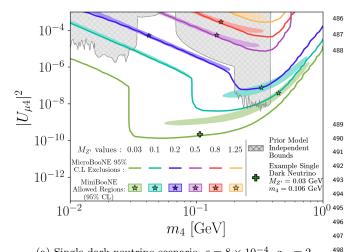
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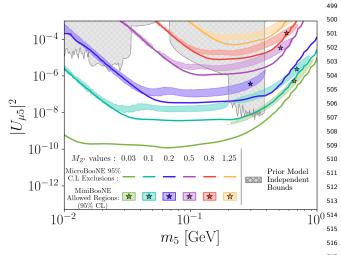
sponding to a 95% C.L assuming Wilk's theorem for 2 d.o.f. MicroBooNE places world leading exclusion limits on this class of models, with the MicroBooNE 95% confidence level exclusion limits ruling out this interpretation of the MiniBooNE anomaly for the single dark neutrino scenario and extremely constrains the broader dual dark neutrino scenario. We note that the results of this paper are robust to the simplifying assumptions made in the parameter space. While variations in the values of g_D , ϵ , and no longer enforcing $V_{\mu 4,5} \equiv U_{\mu 4,5}$ would alter the MiniBooNE-allowed regions, they would also shift the MicroBooNE limits in the same direction and magnitude, thereby preserving the overall conclusions.

This letter has presented the first dedicated analysis exploring e^+e^- explanations for the short-baseline Mini-BooNE anomaly, as suggested by non-minimal dark sector models. By expanding beyond past single-photon searches at MicroBooNE to a broader e^+e^- signal topology containing also two-shower and one-shower, onetrack events, a significant increase in efficiency is achieved for very forward-going, coherently-produced e^+e^- pairs. For these models predicting such a signature, we find no evidence of a signal consistent with MiniBooNE's observation. This result holds for both scenarios involving light Z' mediators and those with heavier Z' states. Our analysis significantly constrains the majority of the model phase space motivated by the MiniBooNE anomaly. Nevertheless, alternative dark sector models with different mediators, such as scalar mediators or those dominated by incoherent scattering leading to distinct energy and angular distributions, are not constrained by this result and remain promising avenues for future exploration.

This document was prepared by the MicroBooNE collaboration using the resources of the Fermi National Accelerator Laboratory (Fermilab), a U.S. Department of Energy, Office of Science, Office of High Energy Physics HEP User Facility. Fermilab is managed by Fermi Forward Discovery Group, LLC, acting under Contract No. 89243024CSC000002. MicroBooNE is supported by the following: the U.S. Department of Energy, Office of Science, Offices of High Energy Physics and Nuclear Physics; the U.S. National Science Foundation; the Swiss National Science Foundation; the Science and Technology Facilities Council (STFC), part of the United Kingdom Research and Innovation; the Royal Society (United Kingdom); the UK Research and Innovation (UKRI) Future Leaders Fellowship; and the NSF AI Institute for Artificial Intelligence and Fundamental Interactions. Additional support for the laser calibration system and cosmic ray tagger was provided by the Albert Einstein Center for Fundamental Physics, Bern, Switzerland. We also



(a) Single dark neutrino scenario, $\varepsilon = 8 \times 10^{-4}$, $q_D = 2$



(b) Dual dark neutrino scenario, $\Delta = 1$, $\varepsilon = 8 \times 10^{-4}$, $g_D = 2$.

FIG. 4. The resulting MicroBooNE 95% C.L. exclusion con- 519 tours for a single dark neutrino model (a) and representative 520 $\Delta = 1$ dual dark neutrino scenario (b). Each solid line rep-521 resents MicroBooNE limits for one of six values of the $\rm dark^{522}$ gauge boson mass (m_{Z^\prime}) varying from 30 MeV to 1.25 GeV, 523 with everything above the solid line being excluded at above 524 the 95% C.L. The corresponding MiniBooNE allowed regions 525 for each dark gauge boson mass indicate the 95% C.L. pre- 526 ferred region of parameter space that offers a viable explana- 527 tion to the MiniBooNE anomaly, taken from [37], with each 528 star showing the MiniBooNE best fit for that single value⁵²⁹ of M_{Z^\prime} . The green plus in (a) indicates the example model⁵³⁰ point used in Fig. 3. The gray solid regions indicate model- 531 independent limits on dark neutrinos [55-60] prior to this result. The 30 MeV Z^{\prime} contour in (a) corresponds to the dark⁵³³ neutrino model introduced in Ref. [18]. 535

acknowledge the contributions of technical and scientific staff to the design, construction, and operation of the Mi-540 croBooNE detector as well as the contributions of past⁵⁴¹ collaborators to the development of MicroBooNE analy-542 ses, without whom this work would not have been pos-543 sible. For the purpose of open access, the authors have 544

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- [1] A. A. Aguilar-Arevalo et al. (MiniBooNE), Unexplained Excess of Electron-Like Events From a 1-GeV Neutrino Beam, Phys. Rev. Lett. 102, 101802 (2009), arXiv:0812.2243 [hep-ex].
- [2] A. A. Aguilar-Arevalo et al. (MiniBooNE), Significant Excess of Electronlike Events in the MiniBoone Short-Baseline Neutrino Experiment, Phys. Rev. Lett. 121, 221801 (2018).
- A. A. Aguilar-Arevalo et al. (MiniBooNE), Updated MiniBooNE neutrino oscillation results with increased data and new background studies, Phys. Rev. D 103, 052002 (2021).
- [4] C. Athanassopoulos et al. (LSND), Evidence for neutrino oscillations from muon decay at rest, Phys. Rev. C 54, 2685 (1996), arXiv:nucl-ex/9605001.
- R. Acciarri et al. (MicroBooNE), Design and Construction of the MicroBooNE Detector, J. Instrum. 12 (02), P02017, arXiv:1612.05824 [physics.ins-det].
- A. A. Aguilar-Arevalo et al. (MiniBooNE), Neutrino flux prediction at MiniBooNE, Phys. Rev. D 79, 072002 (2009).
- V. Brdar and J. Kopp, Can standard model and experimental uncertainties resolve the MiniBooNE anomaly?, Phys. Rev. D 105, 115024 (2022), arXiv:2109.08157 [hep-
- K. J. Kelly and J. Kopp, More ingredients for an Altarelli cocktail at MiniBooNE, JHEP 05, 113, arXiv:2210.08021 [hep-ph].
- P. Abratenko et al. (MicroBooNE), Search for an Excess of Electron Neutrino Interactions in MicroBooNE Using Multiple Final-State Topologies, Phys. Rev. Lett. 128, 241801 (2022), arXiv:2110.14054 [hep-ex].
- [10] P. Abratenko et al. (MicroBooNE), Search for an anomalous excess of charged-current quasielastic ν e interactions with the MicroBooNE experiment using Deep-Learningbased reconstruction, Phys. Rev. D 105, 112003 (2022), arXiv:2110.14080 [hep-ex].
- P. Abratenko et al. (MicroBooNE), Search for an anomalous excess of inclusive charged-current ν_e interactions in the MicroBooNE experiment using Wire-Cell reconstruction, Phys. Rev. D 105, 112005 (2022), arXiv:2110.13978
- [12] P. Abratenko et al. (MicroBooNE), Search for an anomalous excess of charged-current νe interactions without pions in the final state with the MicroBooNE experiment, Phys. Rev. D **105**, 112004 (2022), arXiv:2110.14065 [hep-
- [13] P. Abratenko et al. (MicroBooNE), Search for an Anomalous Production of Charged-Current ν_e Interactions Without Visible Pions Across Multiple Kinematic Observables in MicroBooNE, (2024), arXiv:2412.14407 [hep-ex].
- [14] P. Abratenko et al. (MicroBooNE), First Constraints on Light Sterile Neutrino Oscillations from Combined Appearance and Disappearance Searches with the Micro-BooNE Detector, Phys. Rev. Lett. 130, 011801 (2023),

arXiv:2210.10216 [hep-ex].

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- [15] P. Abratenko et al. (MicroBooNE), Search for Neutrino-610 Induced Neutral-Current Δ Radiative Decay in Micro-611 BooNE and a First Test of the MiniBooNE Low Energy612 Excess under a Single-Photon Hypothesis, Phys. Rev.613 Lett. 128, 111801 (2022), arXiv:2110.00409 [hep-ex]. 614
- [16] E. Wang, L. Alvarez-Ruso, and J. Nieves, Photon emis-615 sion in neutral-current interactions at intermediate ener-616 gies, Phys. Rev. C 89, 015503 (2014).
- [17] G. Lanfranchi, M. Pospelov, and P. Schuster, The Search⁶¹⁸ for Feebly Interacting Particles, Ann. Rev. Nucl. Part.⁶¹⁹ Sci. 71, 279 (2021), arXiv:2011.02157 [hep-ph]. ⁶²⁰
- [18] E. Bertuzzo et al., Dark Neutrino Portal to Explain Mini-621 BooNE excess, Phys. Rev. Lett. 121, 241801 (2018). 622
- [19] P. Ballett et al., U(1)' mediated decays of heavy ster-623
 ile neutrinos in MiniBooNE, Phys. Rev. D 99, 071701624
 (2019).
- [20] P. Ballett, M. Hostert, and S. Pascoli, Dark Neutrinos₆₂₆ and a Three Portal Connection to the Standard Model,₆₂₇ Phys. Rev. D **101**, 115025 (2020), arXiv:1903.07589 [hep-₆₂₈ ph].
- [21] C. A. Argüelles, M. Hostert, and Y.-D. Tsai, Testing⁶³⁰ New Physics Explanations of the MiniBooNE Anomaly⁶³¹ at Neutrino Scattering Experiments, Phys. Rev. Lett.⁶³² 123, 261801 (2019), arXiv:1812.08768 [hep-ph].
- [22] A. Abdullahi, M. Hostert, and S. Pascoli, A dark seesaw⁶³⁴ solution to low energy anomalies: MiniBooNE, the muon⁶³⁵ (g-2), and BaBar, Phys. Lett. B 820, 136531 (2021),⁶³⁶ arXiv:2007.11813 [hep-ph].
- [23] W. Abdallah, R. Gandhi, and S. Roy, Understanding the $_{538}$ MiniBooNE and the muon and electron g-2 anomalies $_{539}$ with a light Z' and a second Higgs doublet, JHEP $\mathbf{12}$, $_{640}$ $_{188}$, arXiv:2006.01948 [hep-ph].
- W. Abdallah, R. Gandhi, T. Ghosh, N. Khan, S. Roy, 642
 and S. Roy, A 17 MeV pseudoscalar and the LSND, 643
 MiniBooNE and ATOMKI anomalies, JHEP 10, 086, 644
 arXiv:2406.07643 [hep-ph].
- [25] A. Hammad, A. Rashed, and S. Moretti, The dark Z' and 646 sterile neutrinos behind current anomalies, Phys. Lett. B₆₄₇
 827, 136945 (2022), arXiv:2110.08651 [hep-ph].
- [26] A. Datta, S. Kamali, and D. Marfatia, Dark sector origin₆₄₉ of the KOTO and MiniBooNE anomalies, Phys. Lett. B₆₅₀ **807**, 135579 (2020), arXiv:2005.08920 [hep-ph].
- [27] B. Dutta, S. Ghosh, and T. Li, Explaining $(g-2)_{\mu,e,652}$ the KOTO anomaly and the MiniBooNE excess in an₆₅₃ extended Higgs model with sterile neutrinos, Phys. Rev.₆₅₄ D **102**, 055017 (2020), arXiv:2006.01319 [hep-ph].
- [28] W. Abdallah, R. Gandhi, and S. Roy, Two-Higgs dou-656 blet solution to the LSND, MiniBooNE and muon657 g-2 anomalies, Phys. Rev. D $\bf 104$, 055028 (2021),658 arXiv:2010.06159 [hep-ph].
- [29] C. Herwig, J. Isaacson, B. Jayatilaka, P. A. N. Machado,660 A. Reinsvold Hall, and M. Safdari, Shedding light on the661 MiniBooNE excess with searches at the LHC, Phys. Rev.662 D 109, 075049 (2024), arXiv:2310.13042 [hep-ph]. 663
- [30] M. A. Acero *et al.*, White paper on light sterile neu-664 trino searches and related phenomenology, J. Phys. G665 **51**, 120501 (2024), arXiv:2203.07323 [hep-ex].
- [31] V. Brdar, O. Fischer, and A. Y. Smirnov, Model-667 independent bounds on the nonoscillatory explanations668 of the MiniBooNE excess, Phys. Rev. D 103, 075008669 (2021), arXiv:2007.14411 [hep-ph].
- [32] C. A. Argüelles, N. Foppiani, and M. Hostert, Effi-671 ciently exploring multidimensional parameter spaces be-672

- yond the Standard Model, Phys. Rev. D **107**, 035027 (2023), arXiv:2205.12273 [hep-ph].
- [33] P. Abratenko et al. (MicroBooNE), Search for Heavy Neutral Leptons in Electron-Positron and Neutral-Pion Final States with the MicroBooNE Detector, Phys. Rev. Lett. 132, 041801 (2024), arXiv:2310.07660 [hep-ex].
- [34] P. Abratenko et al. (MicroBooNE), Search for a Higgs Portal Scalar Decaying to Electron-Positron Pairs in the MicroBooNE Detector, Phys. Rev. Lett. 127, 151803 (2021), arXiv:2106.00568 [hep-ex].
- [35] P. Abratenko et al. (MicroBooNE), First Search for Dark-Trident Processes Using the MicroBooNE Detector, Phys. Rev. Lett. 132, 241801 (2024), arXiv:2312.13945 [hep-ex].
- [36] P. Abratenko et al. (MicroBooNE), Search for the production of Higgs-portal scalar bosons in the NuMI beam using the MicroBooNE detector, (2025), arXiv:2501.08052 [hep-ex].
- [37] A. M. Abdullahi, J. Hoefken Zink, M. Hostert, D. Massaro, and S. Pascoli, A panorama of new-physics explanations to the MiniBooNE excess, arXiv (2023), arXiv:2308.02543 [hep-ph].
- [38] A. M. Abdullahi, J. Hoefken Zink, M. Hostert, D. Massaro, and S. Pascoli, DarkNews: A Python-based event generator for heavy neutral lepton production in neutrino-nucleus scattering, Comput. Phys. Commun. 297, 109075 (2024), arXiv:2207.04137 [hep-ph].
- [39] E. Bertuzzo, S. Jana, P. A. N. Machado, and R. Zukanovich Funchal, Neutrino Masses and Mixings Dynamically Generated by a Light Dark Sector, Phys. Lett. B 791, 210 (2019), arXiv:1808.02500 [hep-ph].
- [40] P. Abratenko *et al.* (MicroBooNE), New CC0 π genie model tune for microboone, Phys. Rev. D **105**, 072001 (2022).
- [41] C. Andreopoulos et al., The GENIE Neutrino Monte Carlo Generator, Nucl. Instrum. Meth. A 614, 87 (2010), arXiv:0905.2517 [hep-ph].
- [42] J. Tena-Vidal et al. (GÉNIE), Neutrino-nucleon crosssection model tuning in GENIE v3, Phys. Rev. D 104, 072009 (2021), arXiv:2104.09179 [hep-ph].
- [43] S. Baker, A. Elwyn, J. Lach, and A. L. Read, Tunneling Beyond the Fermilab Site, FERMILAB-TM-1218 (1983).
- [44] J. Cossairt, A. Elwyn, W. Freeman, and S. Butalla, A study of the transport of high energy muons through a soil shield at the Tevatron, FERMILAB-PUB-88/146 (1988).
- [45] J. D. Cossairt, Assessment of neutron skyshine near unmodified Accumulator Debuncher storage rings under Mu2e operational conditions, FERMILAB-TM-2481-ESH 10.2172/1001998 (2010).
- [46] R. Acciarri et al. (MicroBooNE), The Pandora Multi-Algorithm Approach to Automated Pattern Recognition of Cosmic Ray Muon and Neutrino Events in the Micro-BooNE Detector, Eur. Phys. J. C 78, 82 (2018).
- [47] T. Chen and C. Guestrin, Xgboost: A scalable tree boosting system, CoRR abs/1603.02754 (2016), 1603.02754, https://github.com/dmlc/xgboost.
- [48] P. Abratenko et al. (MicroBooNE), Calorimetric classification of track-like signatures in liquid argon TPCs using MicroBooNE data, JHEP 12, 153, arXiv:2109.02460 [physics.ins-det].
- [49] C. Adams *et al.* (MicroBooNE), Reconstruction and Measurement of $\mathcal{O}(100)$ MeV Energy Electromagnetic Activity from $\pi^0 \to \gamma \gamma$ Decays in the MicroBooNE

LArTPC, J. Instrum. **15** (02), P02007, arXiv:1910.02166693 [hep-ex].

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674

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678

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680

681

682

683

684

685

686

687

- [50] P. Abratenko et al. (MicroBooNE), First search for neu-695 tral current coherent single-photon production in micro-696 boone (2025), arXiv:2502.06091 [hep-ex]. 697
- [51] P. Abratenko et al. (MicroBooNE), Measurement of neu-698 tral current single π0 production on argon with the Mi-699 croBooNE detector, Phys. Rev. D 107, 012004 (2023),700 arXiv:2205.07943 [hep-ex].
- [52] S. Agostinelli et al. (GEANT4), GEANT4 A Simulation₇₀₂ Toolkit, Nucl. Instrum. Meth. A 506, 250 (2003). 703
- [53] P. Abratenko et al. (MicroBooNE), Novel approach for 704 evaluating detector-related uncertainties in a LArTPC 705 using MicroBooNE data, Eur. Phys. J. C 82, 454 (2022), 706 arXiv:2111.03556 [hep-ex].
- [54] X. Ji, W. Gu, X. Qian, H. Wei, and C. Zhang, Com-708
 bined Neyman-Pearson chi-square: An improved approx-709
 imation to the Poisson-likelihood chi-square, Nucl. In-710
 strum. Meth. A 961, 163677 (2020), arXiv:1903.07185711
 [physics.data-an].

- [55] R. S. Hayano et al., HEAVY NEUTRINO SEARCH US-ING K(mu2) DECAY, Phys. Rev. Lett. 49, 1305 (1982).
- [56] A. Aguilar-Arevalo *et al.* (PIENU), Search for heavy neutrinos in $\pi \to \mu\nu$ decay, Phys. Lett. B **798**, 134980 (2019), arXiv:1904.03269 [hep-ex].
- [57] M. Daum, B. Jost, R. M. Marshall, R. C. Minehart, W. A. Stephens, and K. O. H. Ziock, Search for Admixtures of Massive Neutrinos in the Decay π⁺ → μ⁺ Neutrino, Phys. Rev. D 36, 2624 (1987).
- [58] E. Cortina Gil *et al.* (NA62), Search for K^+ decays to a muon and invisible particles, Phys. Lett. B **816**, 136259 (2021), arXiv:2101.12304 [hep-ex].
- [59] A. V. Artamonov *et al.* (BNL-E949), Study of the decay $K^+ \to \pi^+ \nu \bar{\nu}$ in the momentum region $140 < P_\pi < 199$ MeV/c, Phys. Rev. D **79**, 092004 (2009), arXiv:0903.0030 [hep-ex].
- [60] E. Fernández-Martínez, M. González-López, J. Hernández-García, M. Hostert, and J. López-Pavón, Effective portals to heavy neutral leptons, JHEP 09, 001, arXiv:2304.06772 [hep-ph].